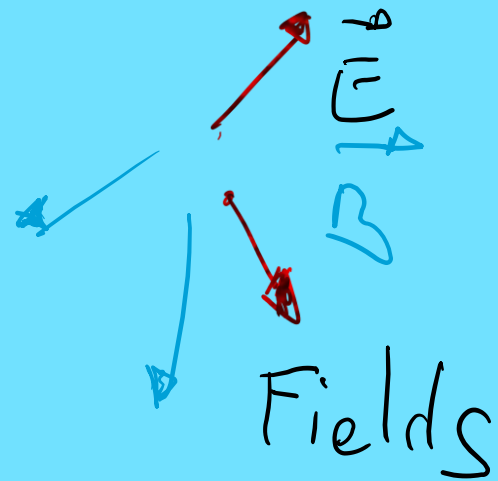


Maxwell's equations are not causal

Sources



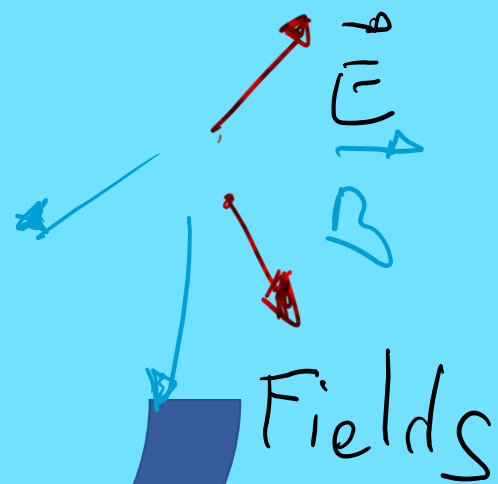
Maxwell equations

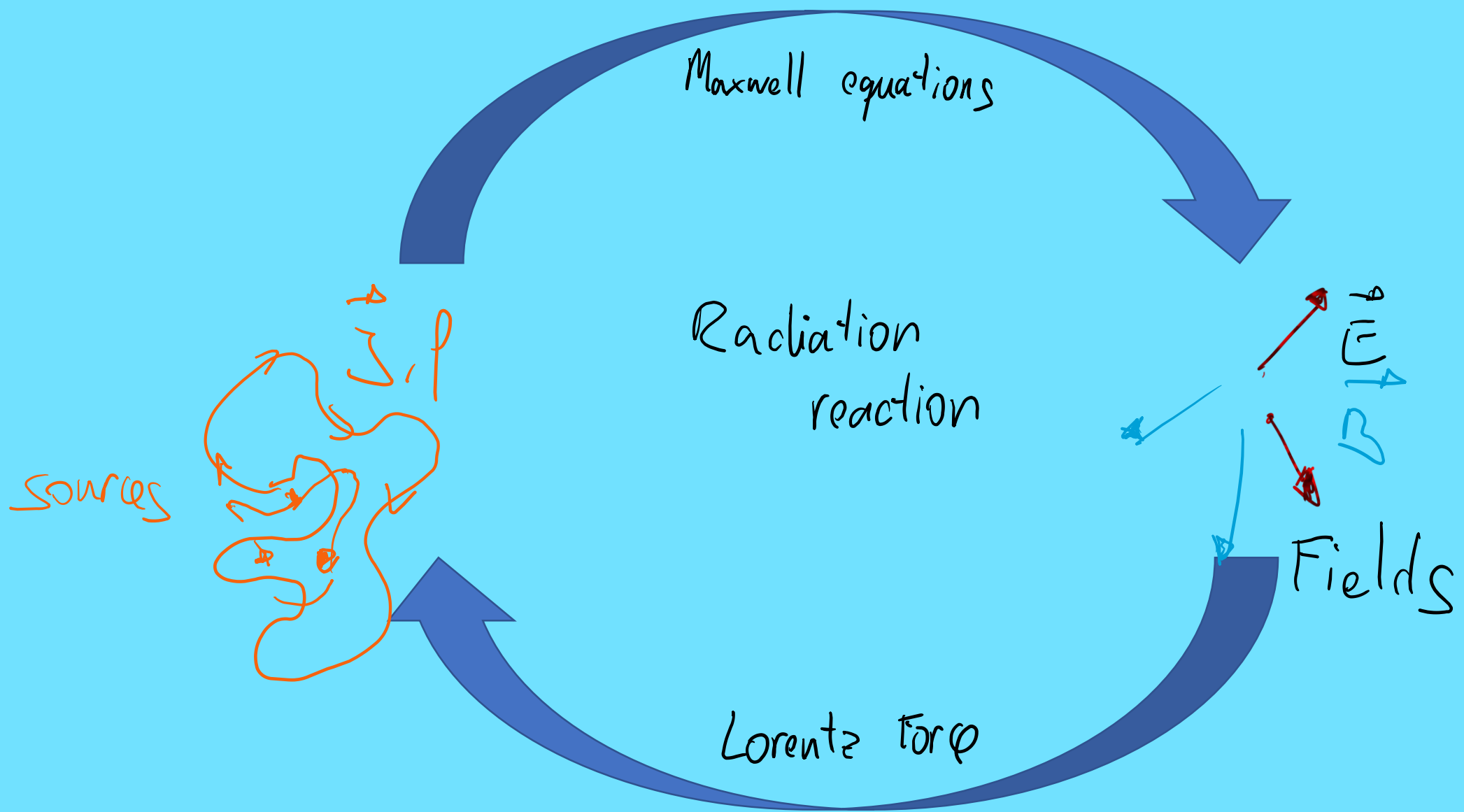


Sources



Lorentz Force





Maxwell equations are not causal

- This happens at very small scales: the classical radius of the electron
 - $d \sim 10^{-15} \text{ m}$
 - $\tau \sim 10^{-22} \text{ s}$
- It leads to “runaway” solutions and also to problems with causality (electron moves before any wave arrives)
- Quantum Electrodynamics does not seem to solve the problem
- Long standing (and sometimes forgotten) debate still (mostly unsolved)
- Possible solution: set a time scale in Maxwell’s equations

Maxwell equations are not causal

Corrections to formulas:

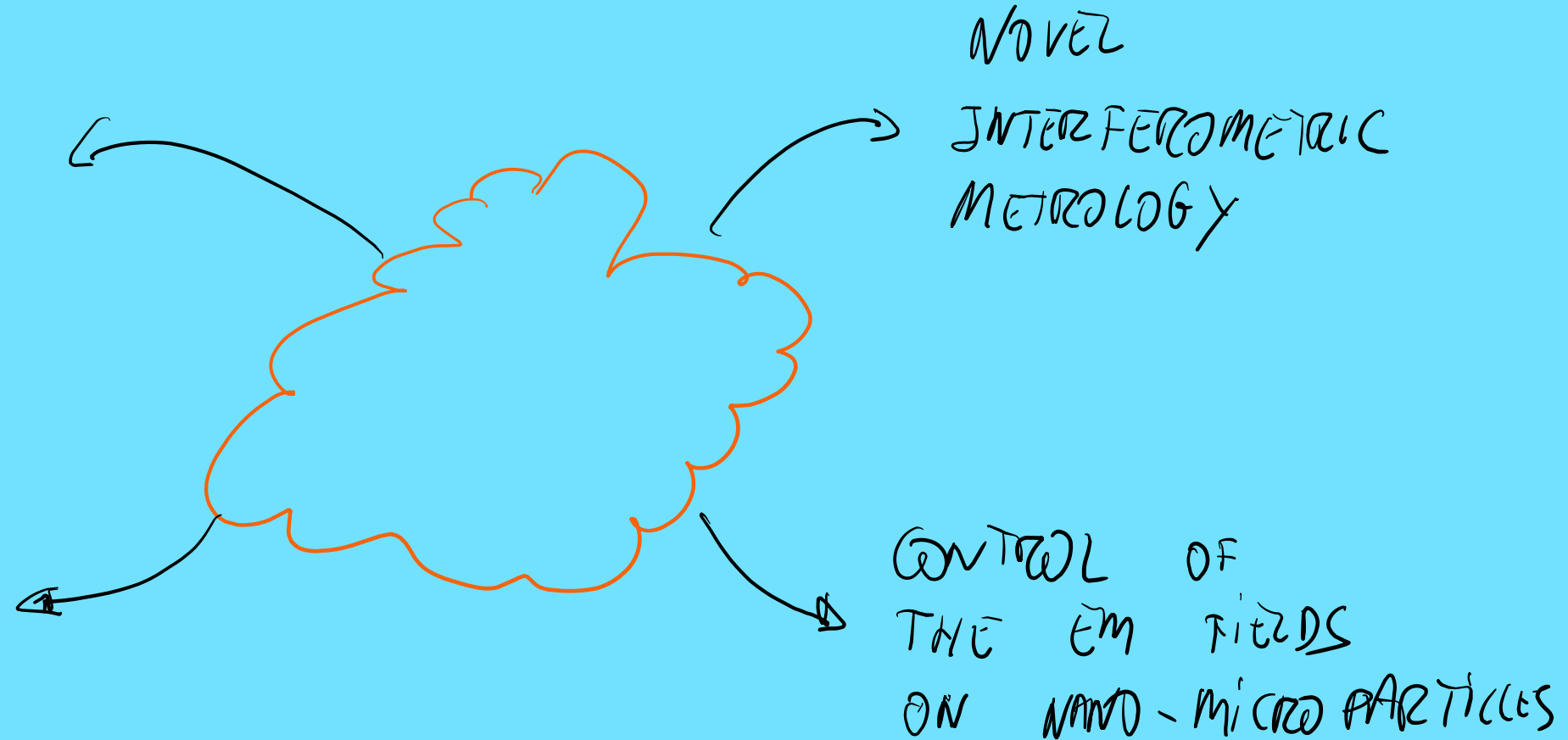
$\gamma \times \omega \sim 10^{-8}$
An arrow points from γ to 10^{-23} .
An arrow points from ω to 'freq' and 10^{15} .

- For example, one should measure the decrease in the Larmor frequency in the 8th significant digit
- One would need extremely sensitive equipment to measure these effects (no experiment to date has been able to do so)

Wishlist

ROTATING
PARTICLES

QUANTUM
EMITTERS

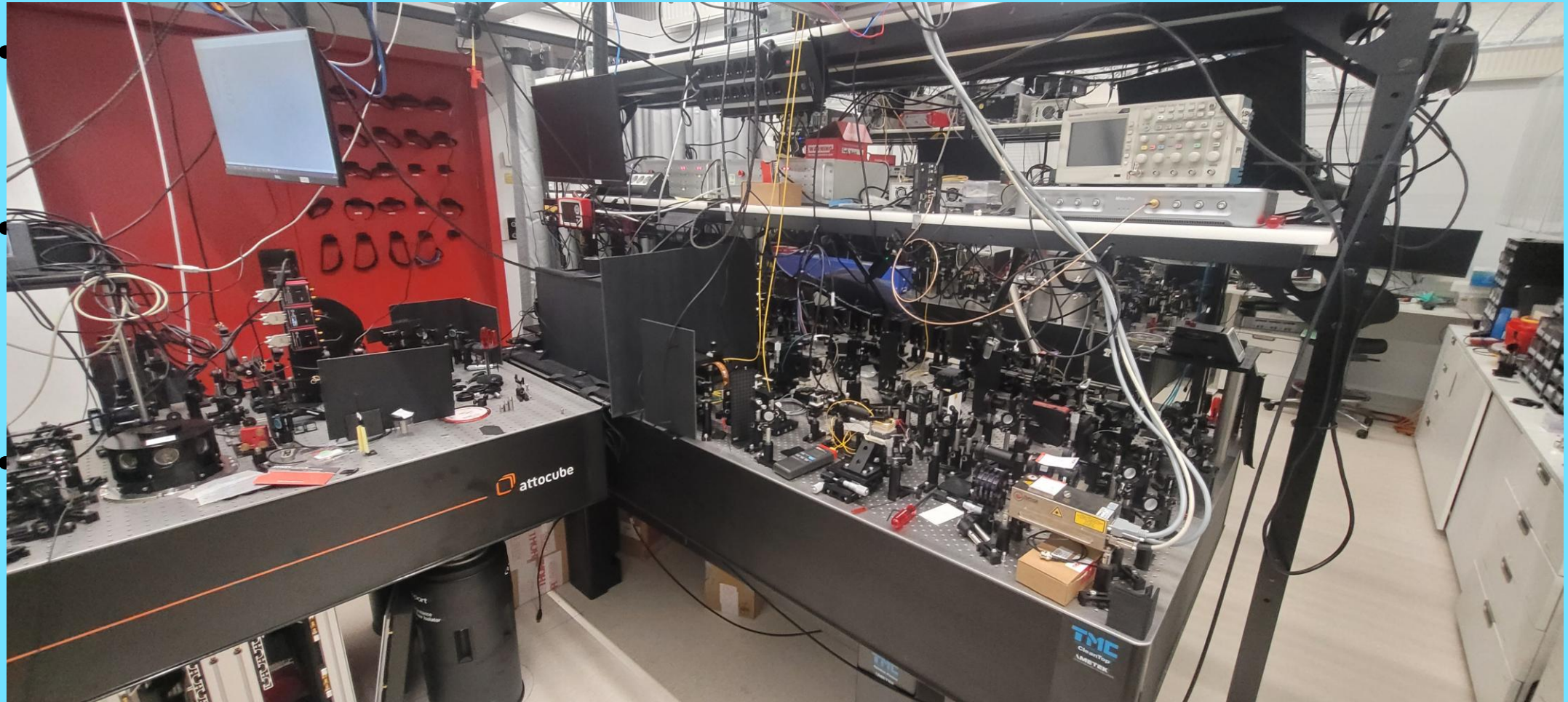


Quantum Nanophotonics Lab

(and how we contribute to Particle Physics)

Gabriel Molina Terriza
Ikerbasque Research Professor
Materials Physics Center

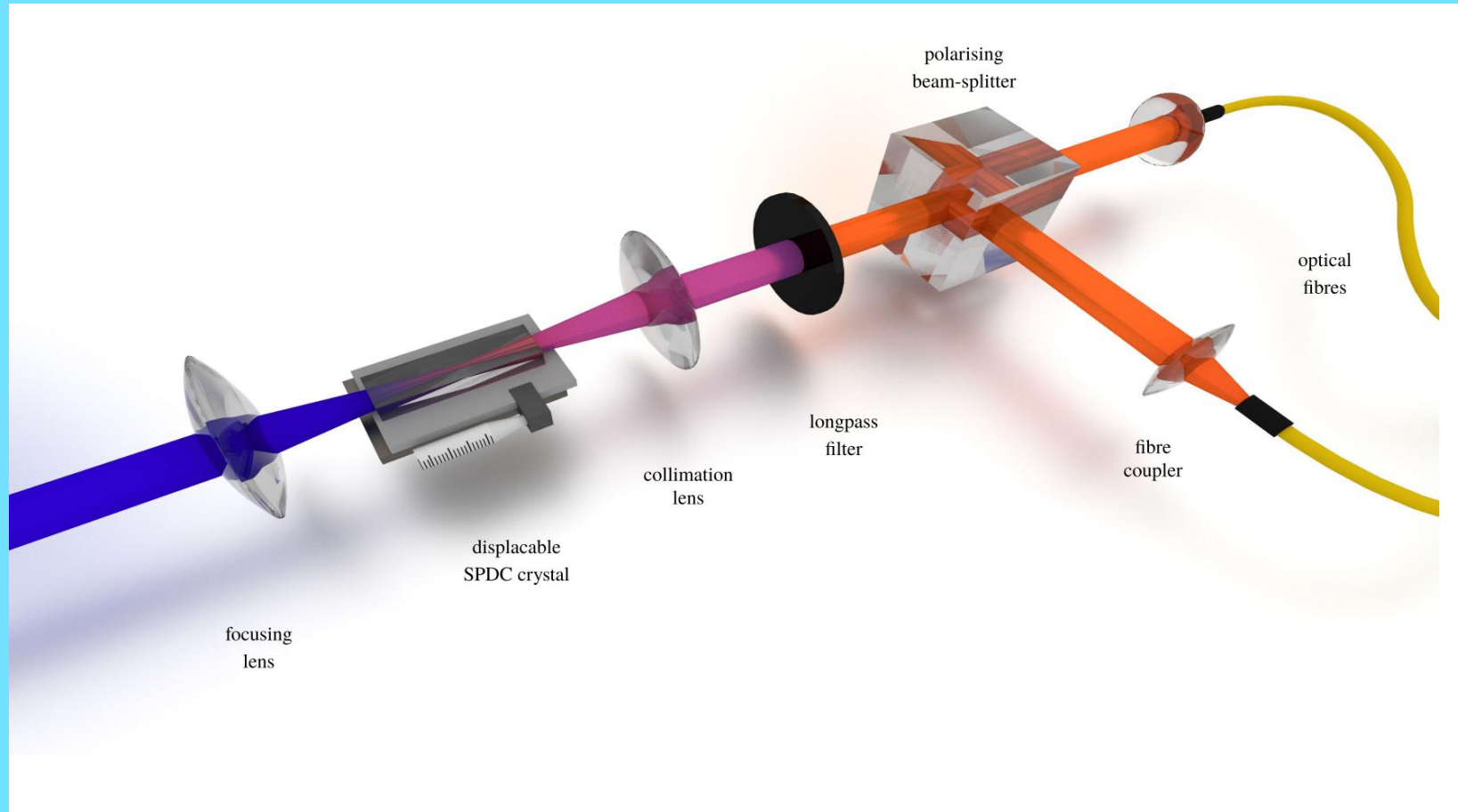
Laboratory



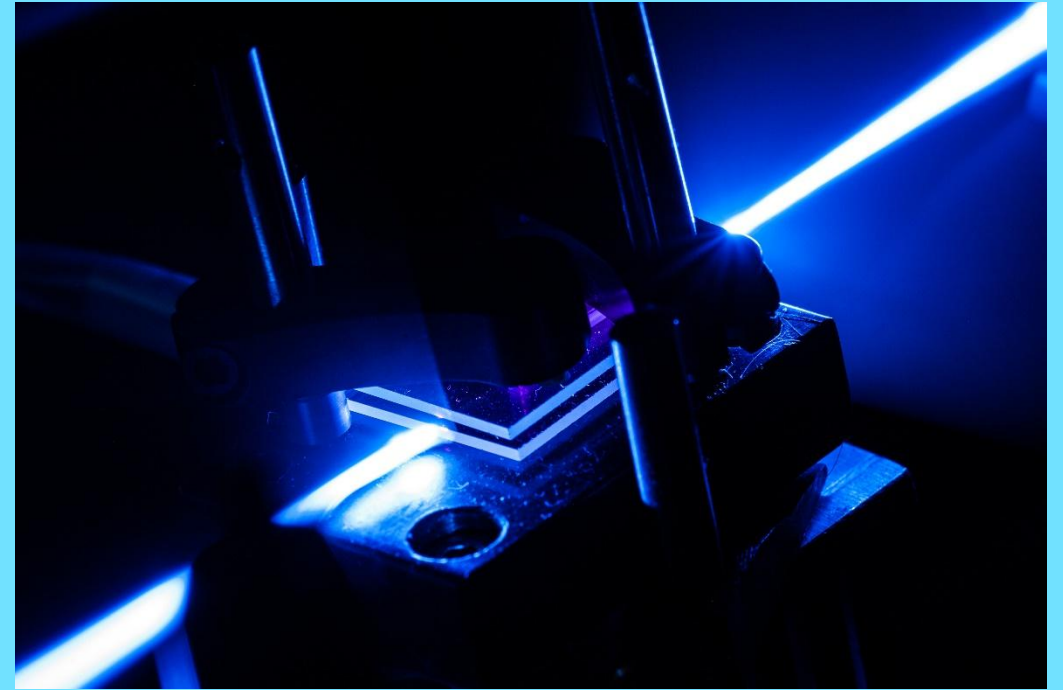
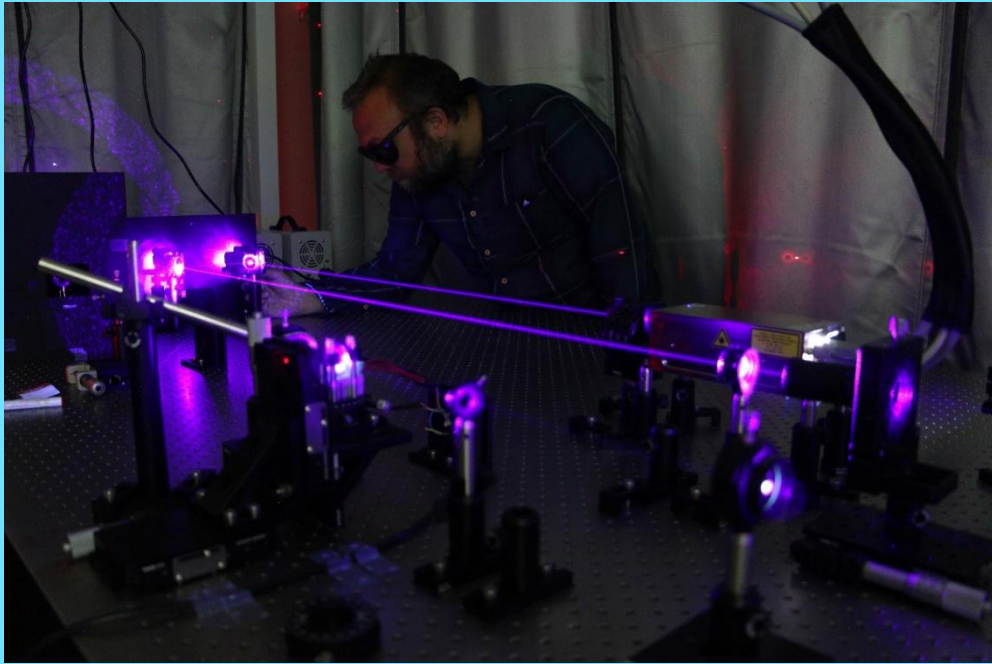
Topics we cover at the Quantum Nanophotonics Laboratory

1. Quantum and classical photonics
2. Optical forces: trapping and levitation
3. Control of the electronic and nuclear spin of Nitrogen Vacancy centers in diamond

Entangled photons



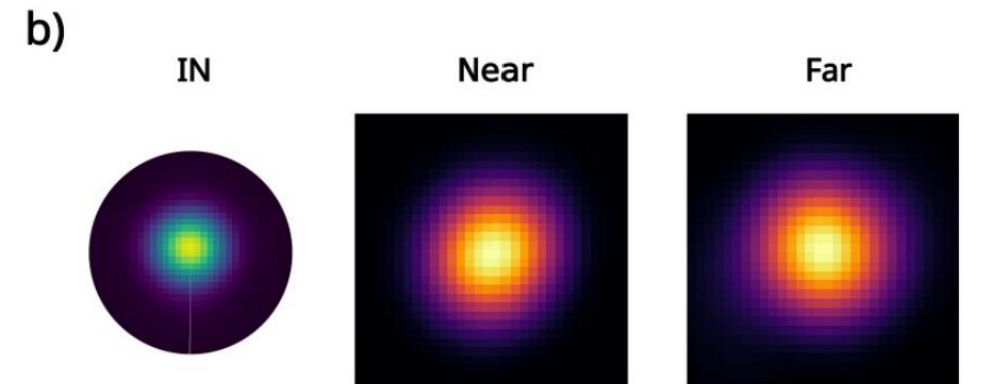
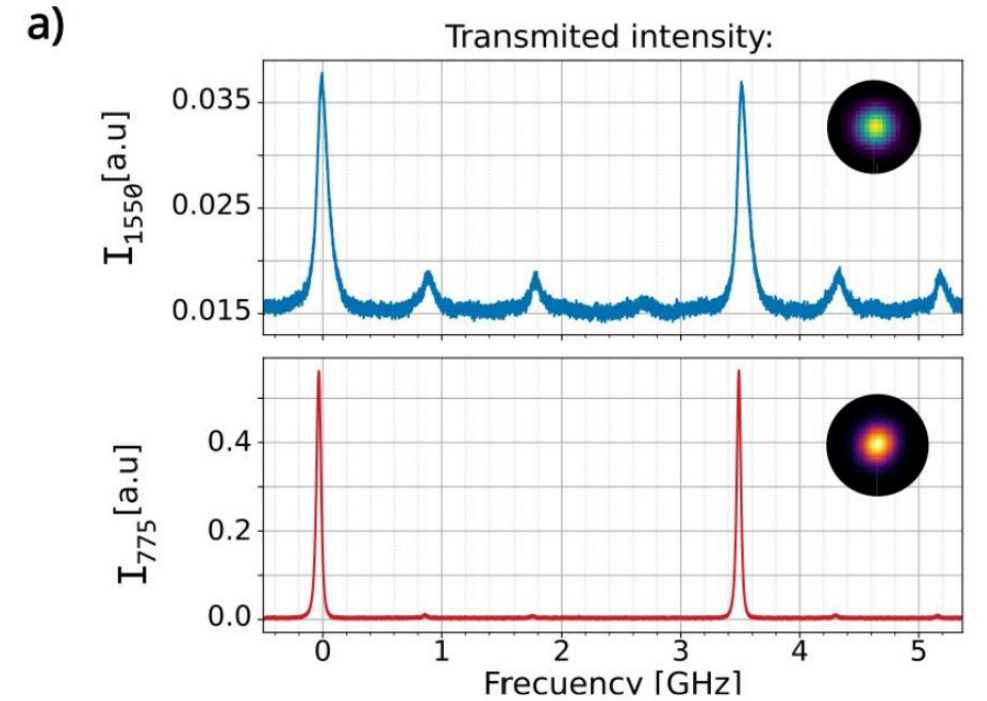
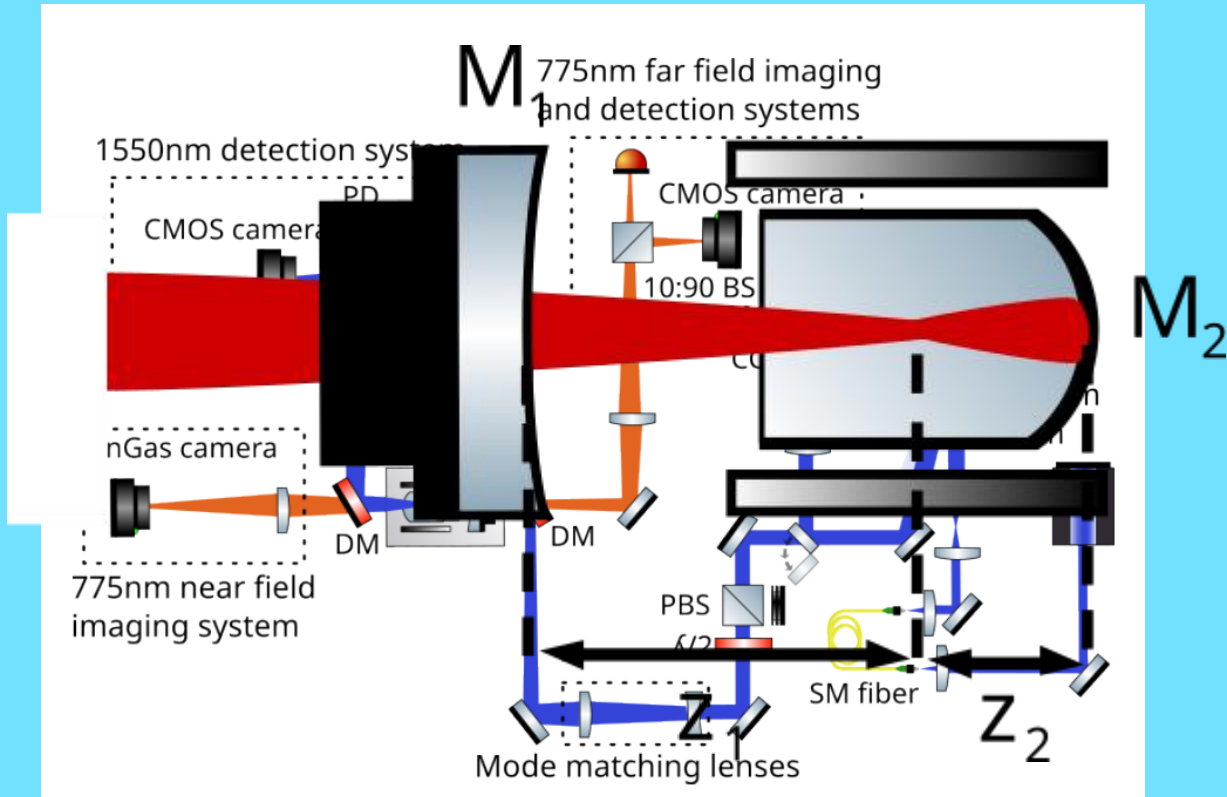
Entangled photons at QNanoLab: Quantum communications and sensing



Entangled photons

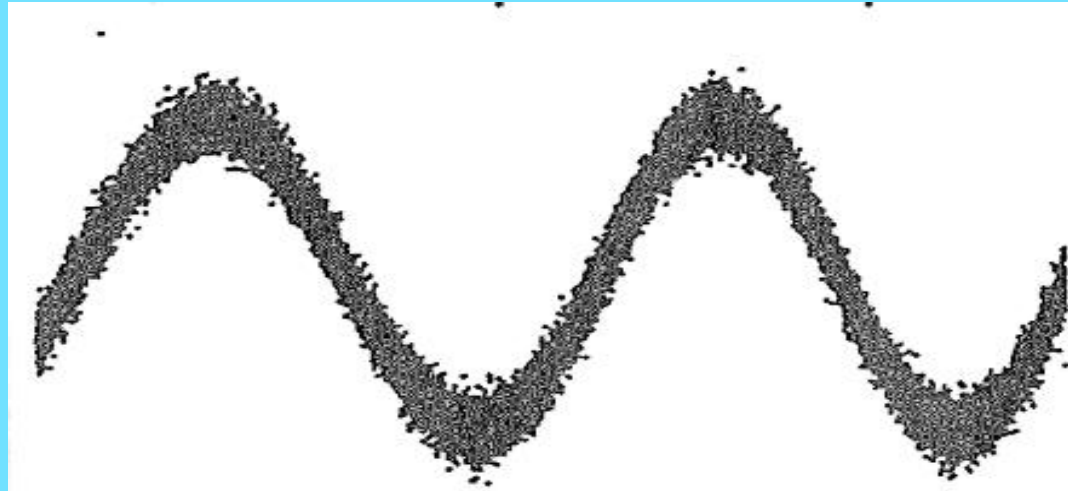
Nonlinear Down Conversion

Quantum squeezing

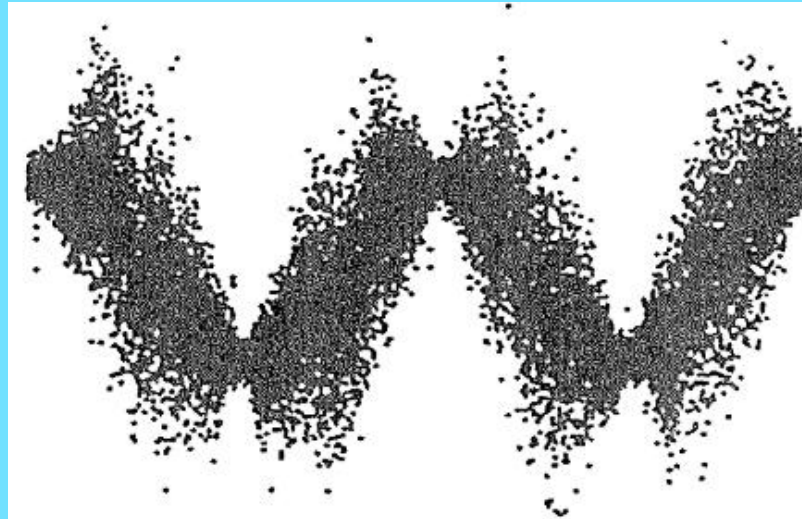


Shot noise limits some measurements

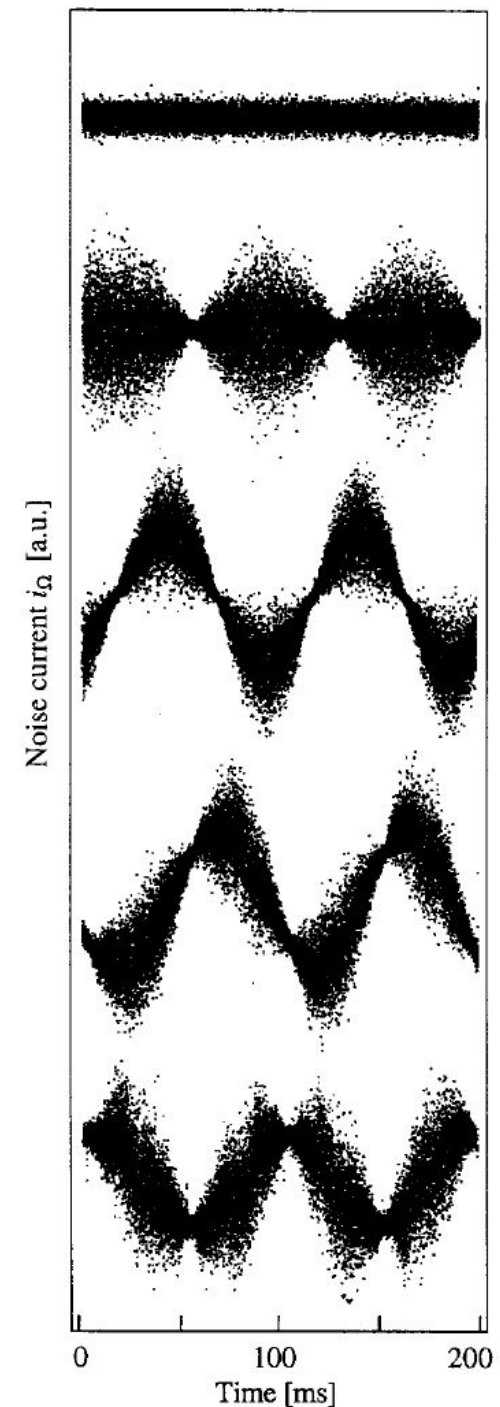
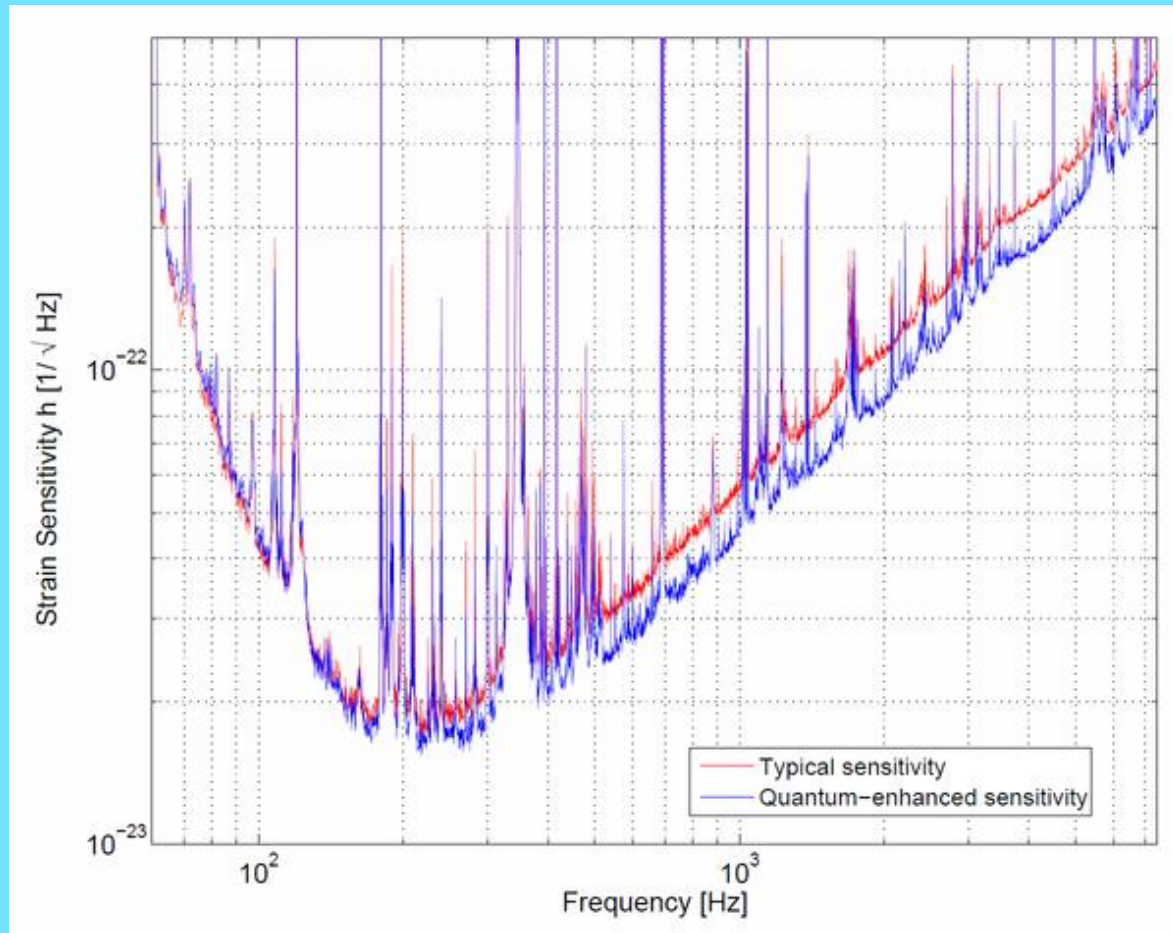
- In very sensitive applications the shot noise of the light that it is used for measuring is the main source of noise



Quantum squeezing provides a (partial) solution



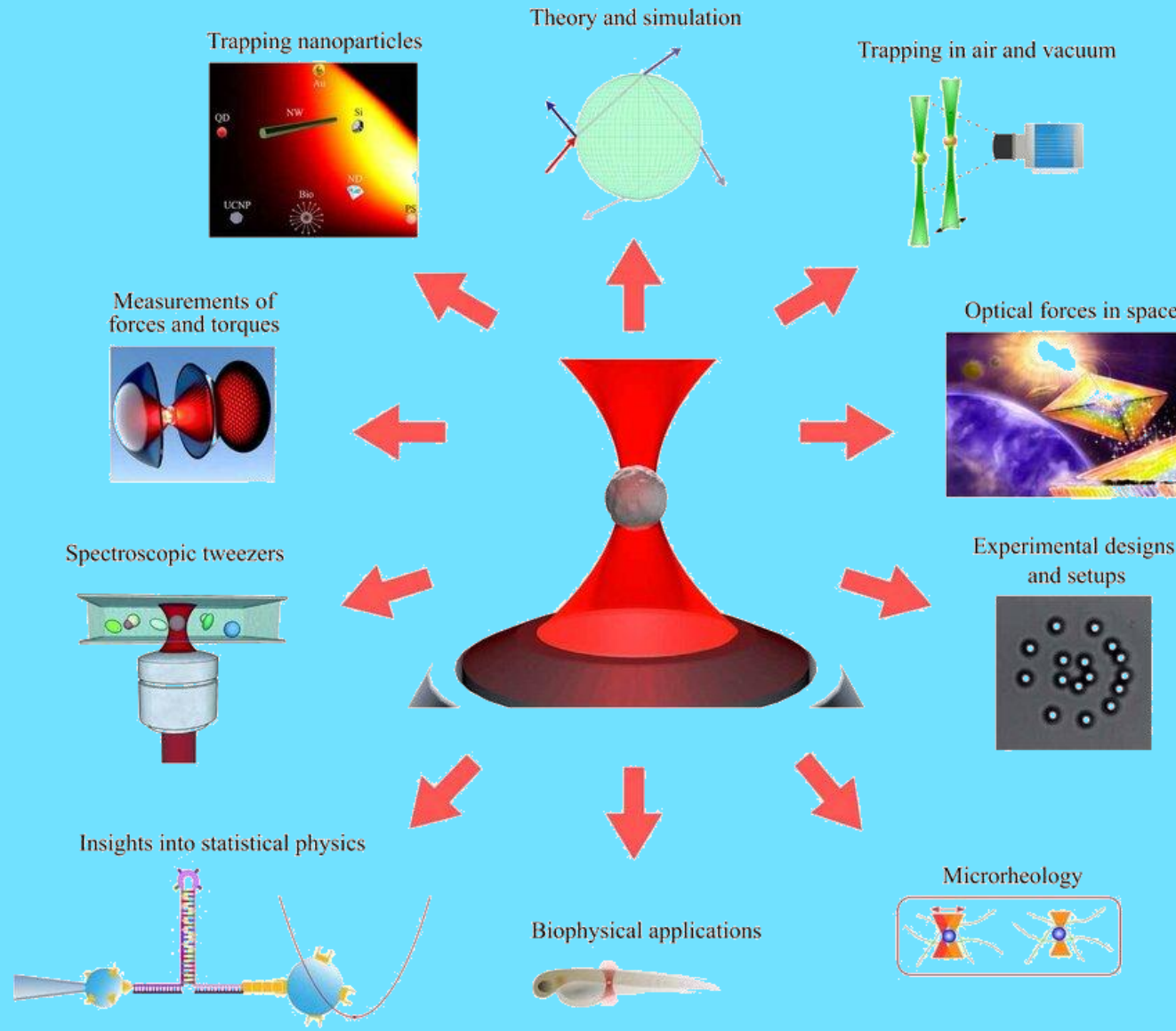
Quantum squeezing is
already helping to detect
gravitational waves

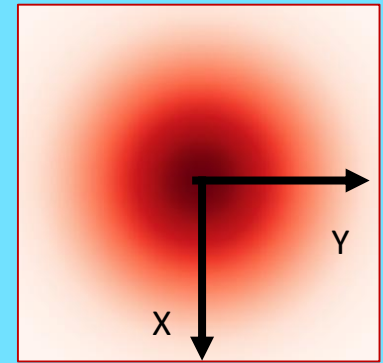
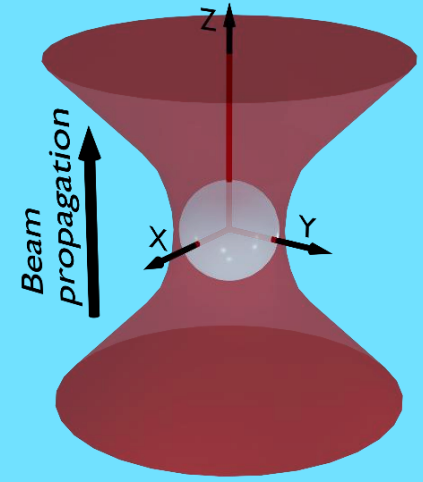


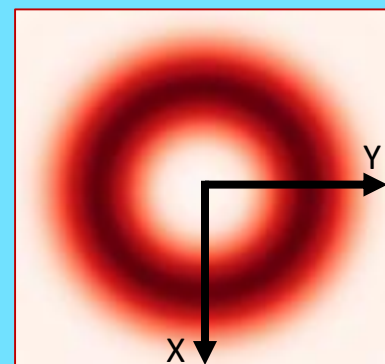
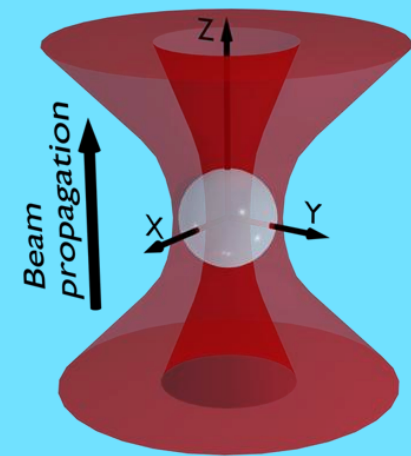
Optical levitation



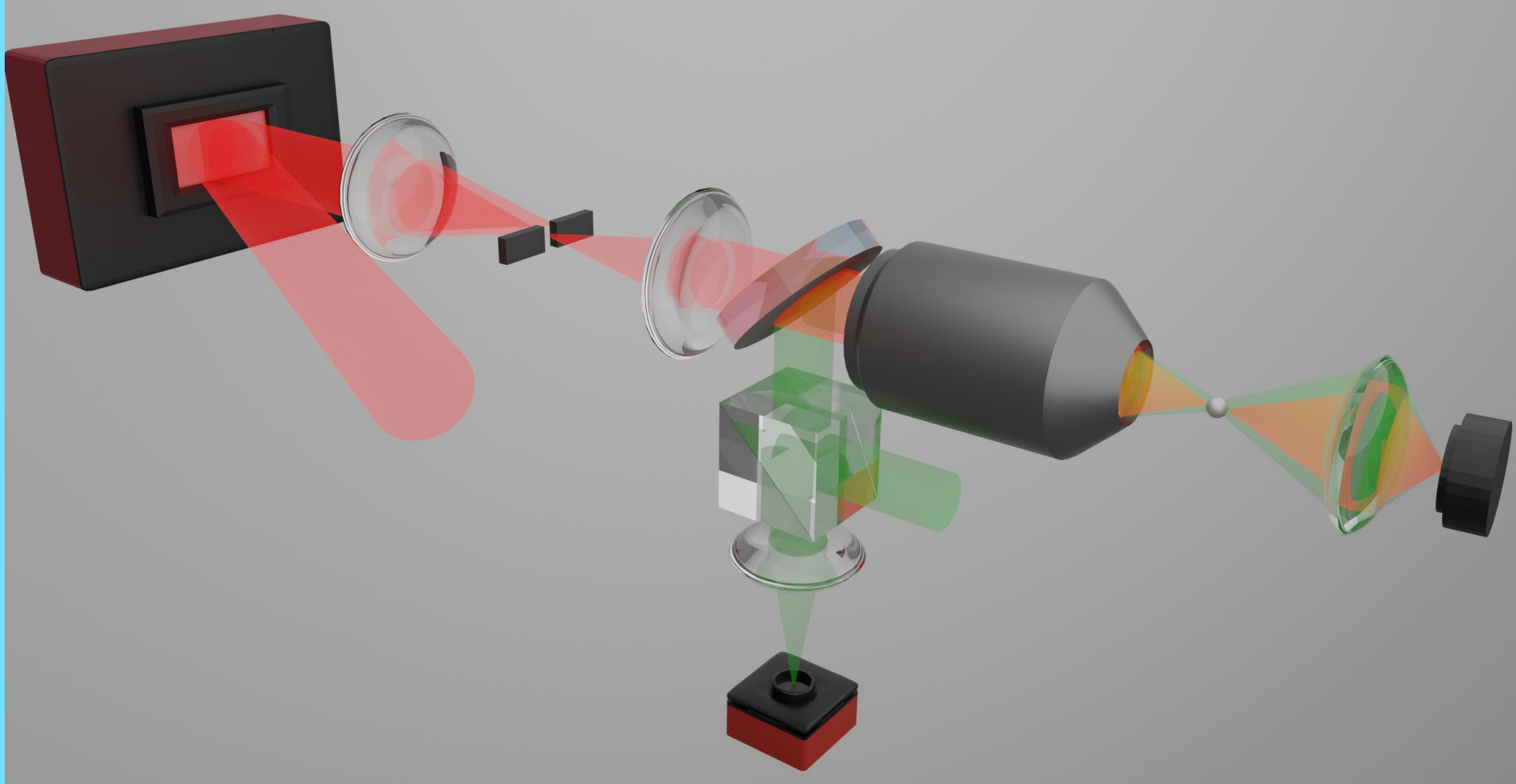
You can trap and manipulate particles with **optical forces**







Levitation Experiments



Tektronix

TDS 2024C

FOUR CHANNEL
DIGITAL STORAGE OSCILLOSCOPE

200 MHz
2 GS/s

Tek



DISPLAY

Type

Vectors

Persist

Off

Format

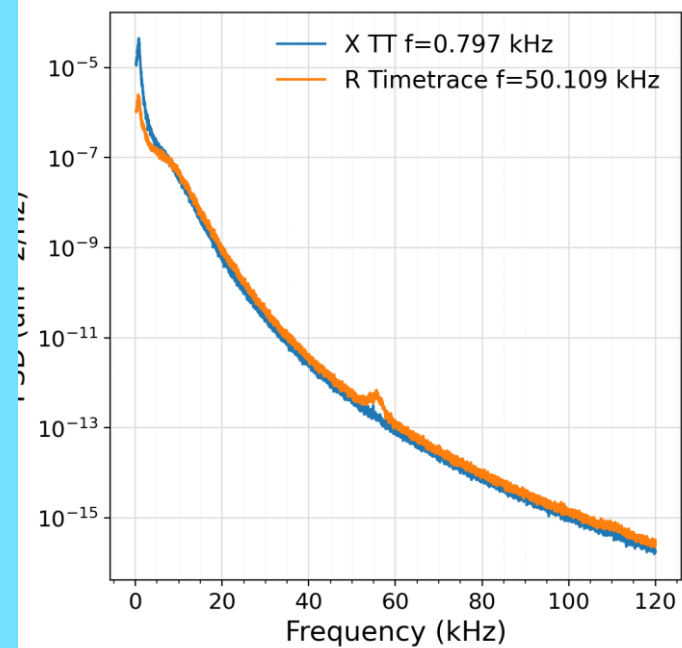
XY

CH1 5.00mV

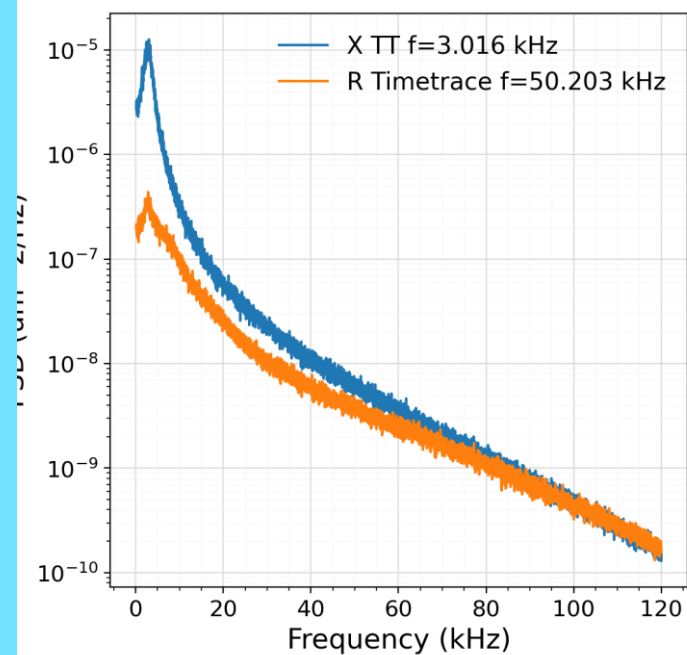
CH2 5.00mV/μs

XY Mode

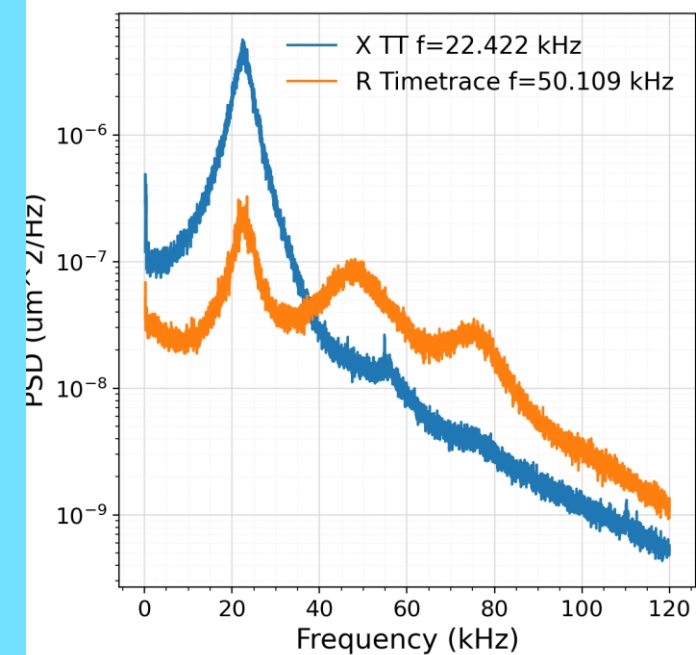
LG1 - R- 1013hPa



LG1 - R- 140hPa



LG1 - R- 12hPa



Moore and Garasci

arxiv.org/pdf/2008.13197

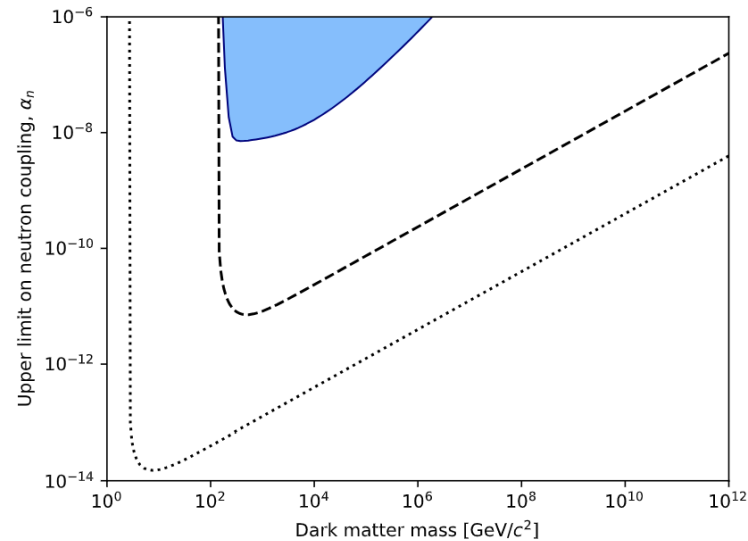


Figure 8. Background free sensitivity projections to a dark matter candidate interacting with neutrons through a long-range force, assuming a massless mediator. Existing constraints from an initial demonstration of such a search (with a few ng-days exposure) are shown by the blue filled region [82]. Substantial improvement over the initial search is possible with longer exposures and larger arrays of spheres. Projections for the sensitivity of a 100 element array of spheres with 1 yr exposure at the current sensitivity (dashed) and at the SQL (dotted) are also shown.

Moore and Garasci

arxiv.org/pdf/2008.13197

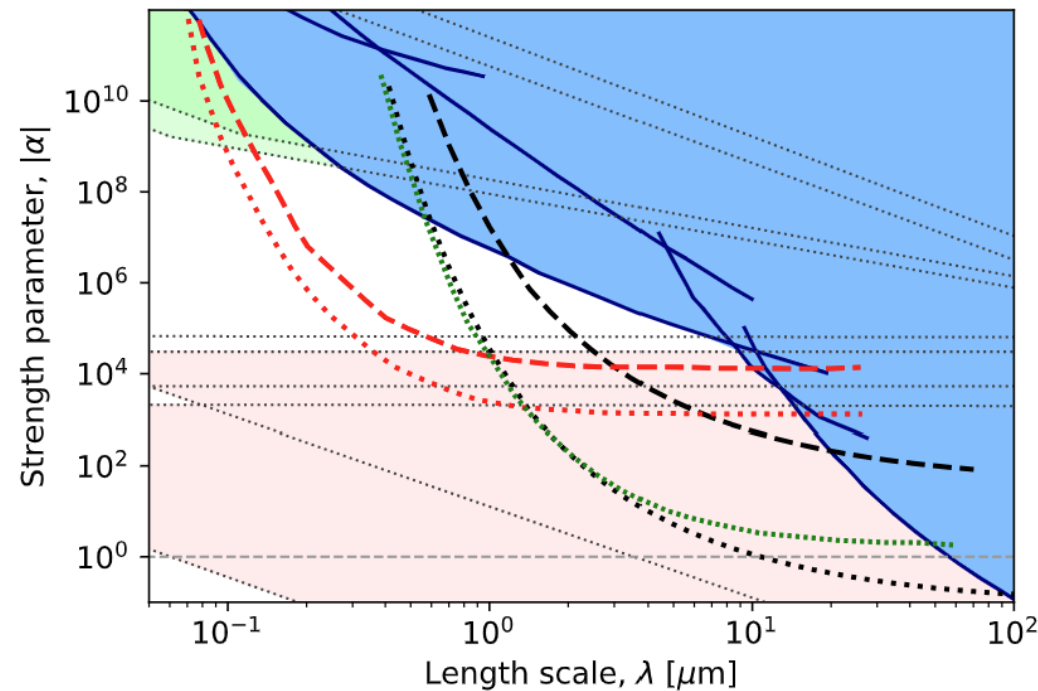
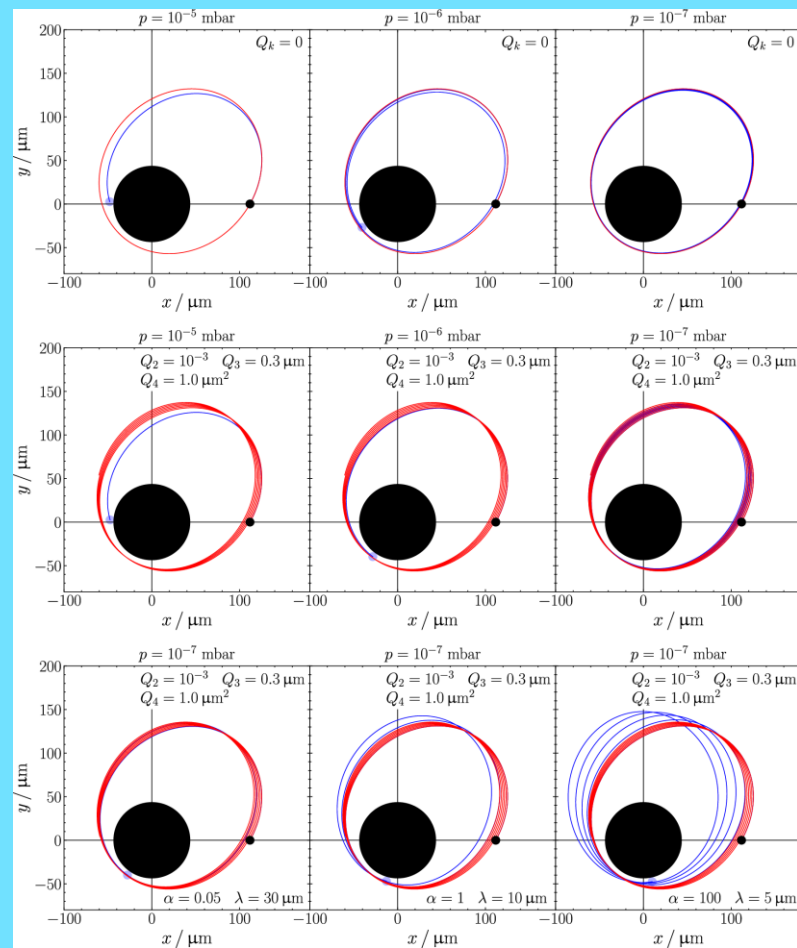


Figure 4. Background free sensitivity projections to deviations from Newton's law for example optically levitated masses. Existing limits are denoted by the blue region [10, 49, 86], with examples of specific theoretical predictions for such deviations



Towards a realistic setup for a dynamical measurement of deviations from Newton's $1/r^2$ law: the impact of air viscosity

J. Baeza-Ballesteros^{1,2,a} , A. Donini^{1,2}, G. Molina-Terriza^{3,4,5}, F. Monrabal^{3,4}, A. Simón^{3,6}



Moore and Garasci

arxiv.org/pdf/2008.13197

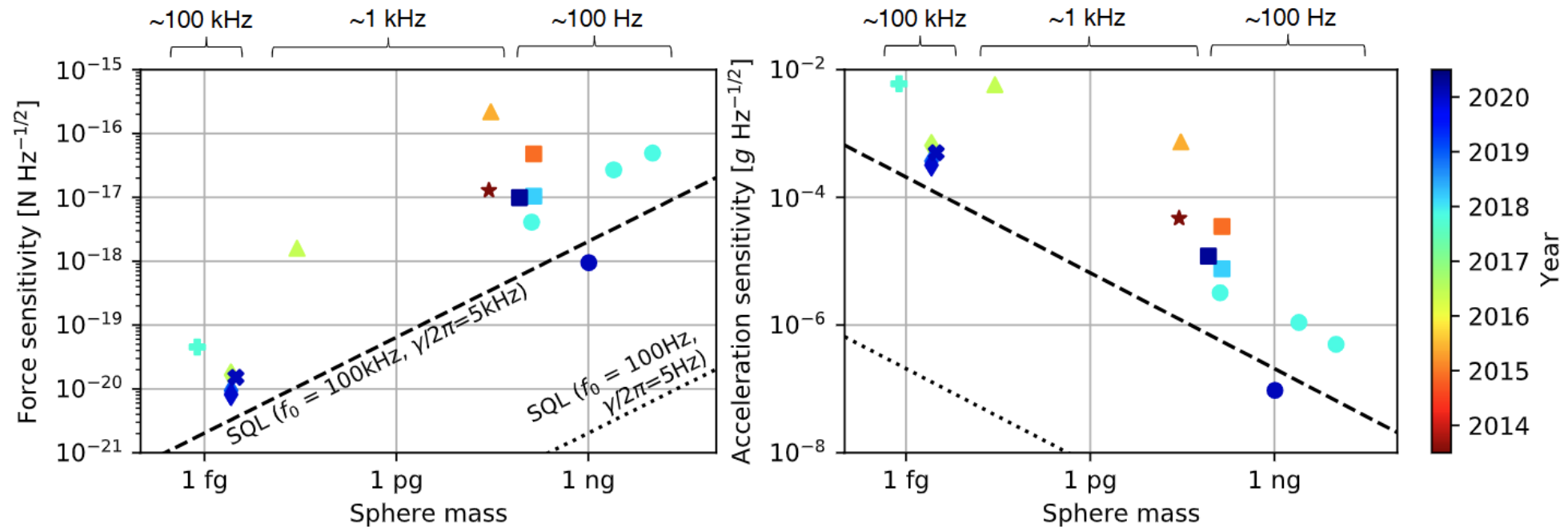
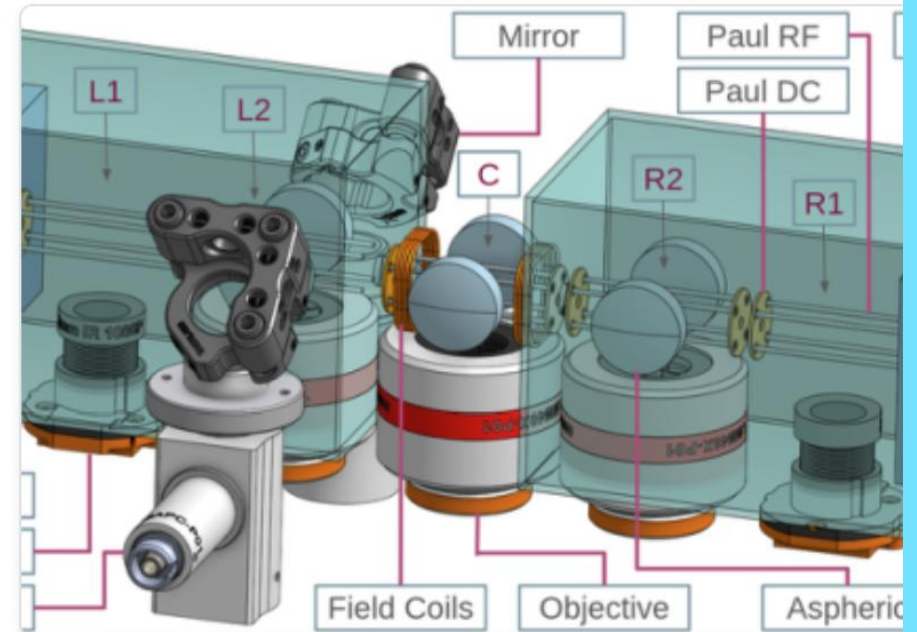


Figure 2. Compilation of measured force and acceleration sensitivities for optically levitated masses range between ~ 1 fg to ~ 10 ng from a variety of groups: ETH Zurich

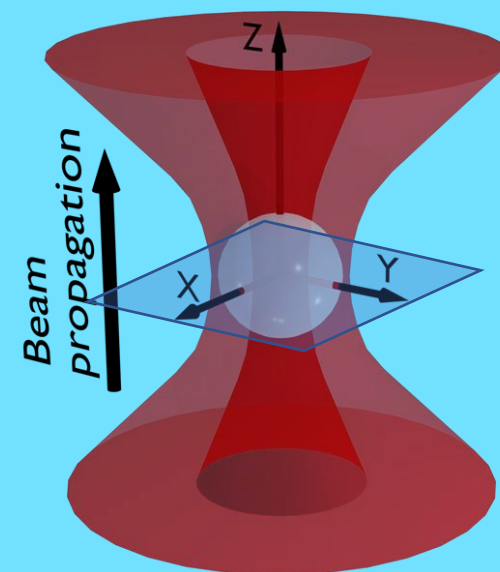
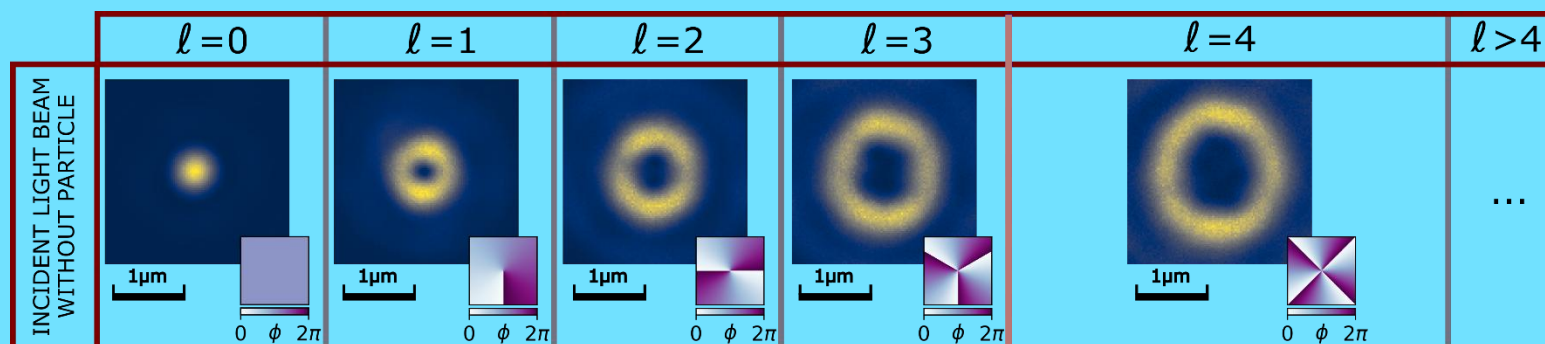
LEVITAS - Levitodynamics Experiments in Vacuum Interaction and Thermal Atmospheric Sensing

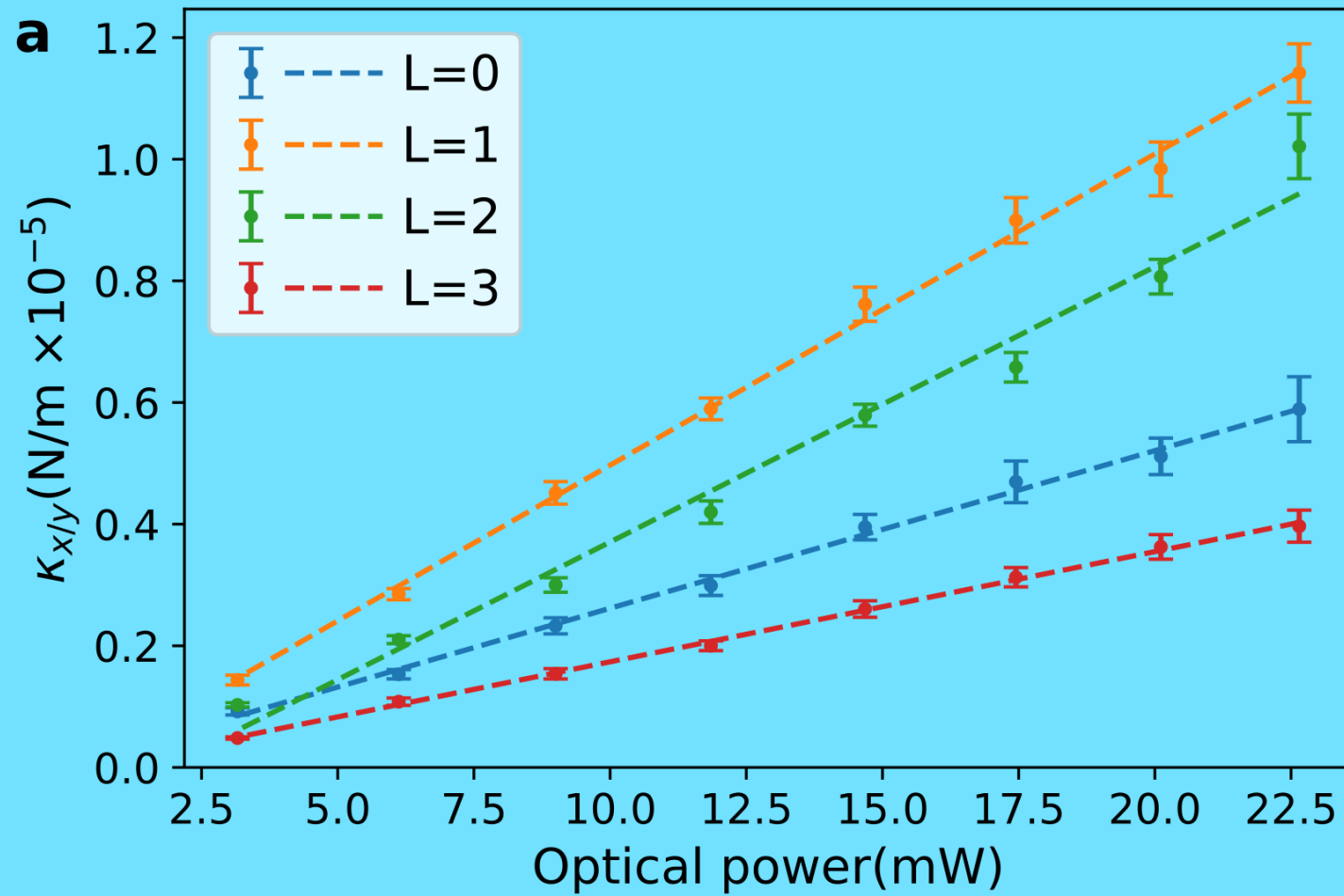
Programme Reference	Status	Country	Start Date	End Date
IPD-0545	Closed	United Kingdom	2024	2025

Programme: FutureEO Prime Contractor: UNIV WARWICK

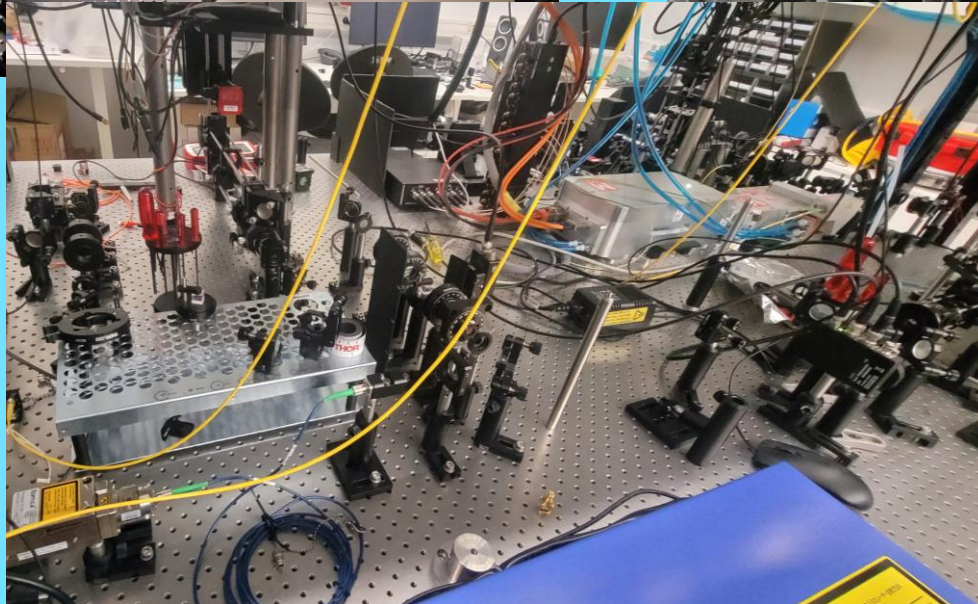
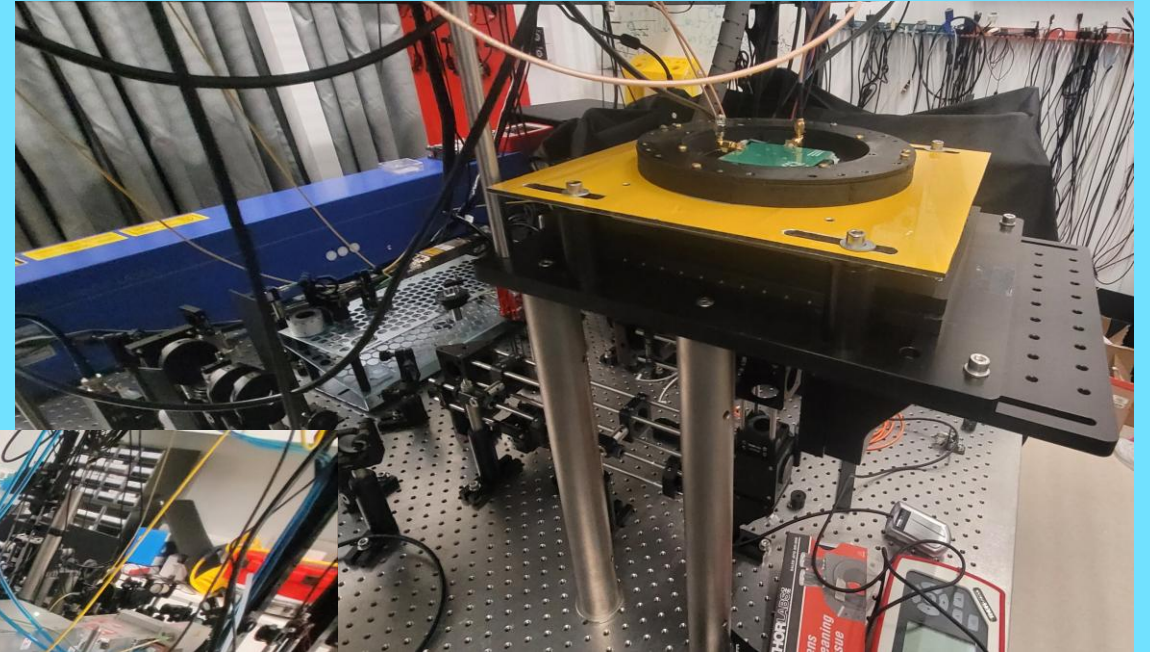
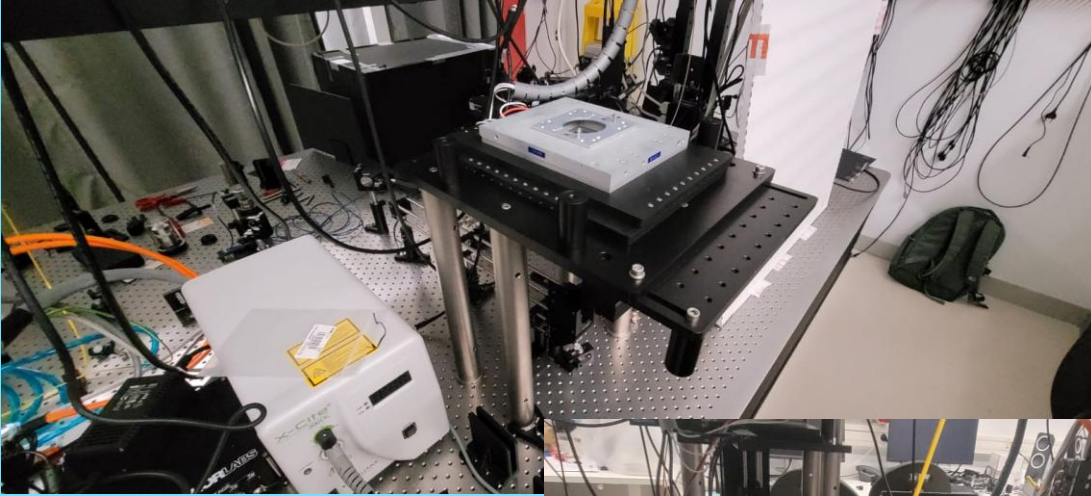


 Related Content





Quantum magnetic sensing with diamond



NV-Center Working Principles

- 532 nm
- 12 ns, 680 nm

Singlet state

Excited states

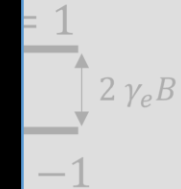
QUANTUM SENSOR RECIPE

LASER

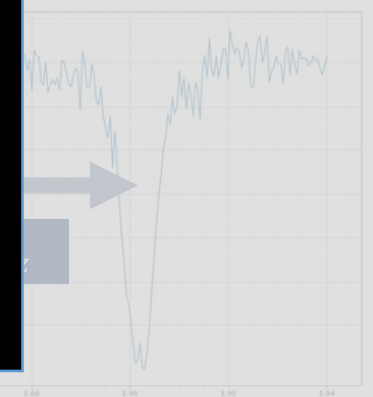
Collection system

Diamond

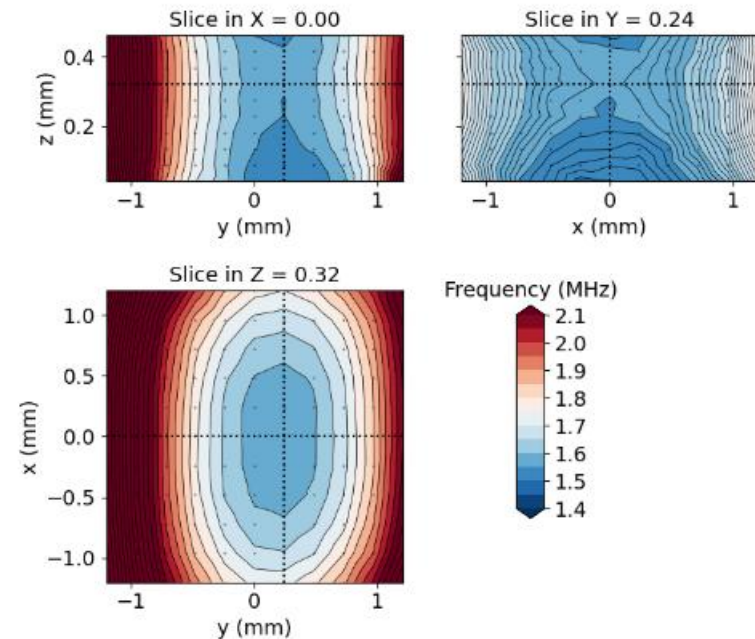
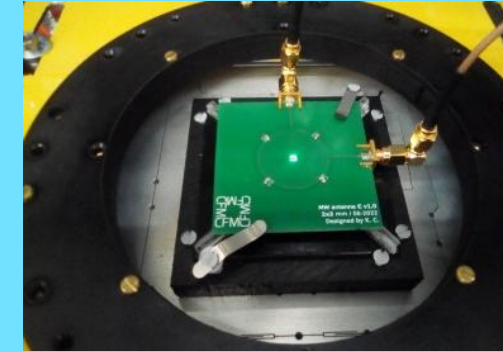
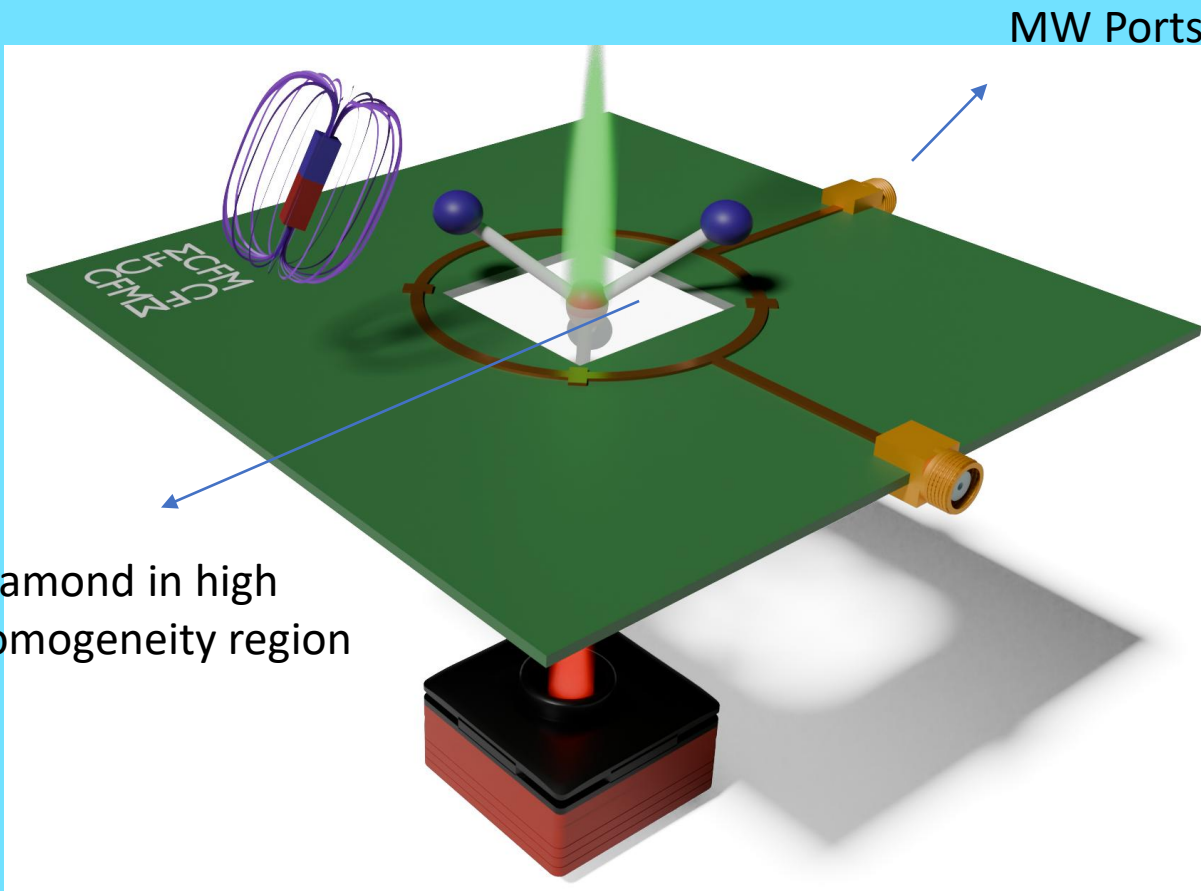
Microwave delivery



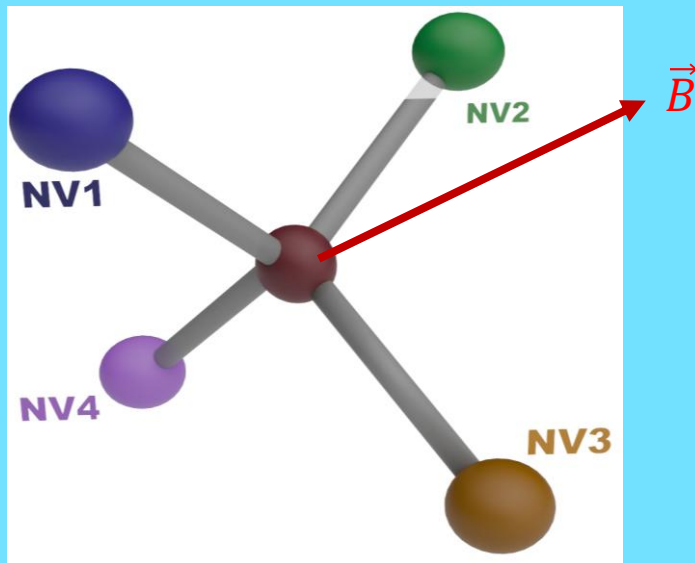
$m_s = 0$



Antenna for Bias Free Magnetometry

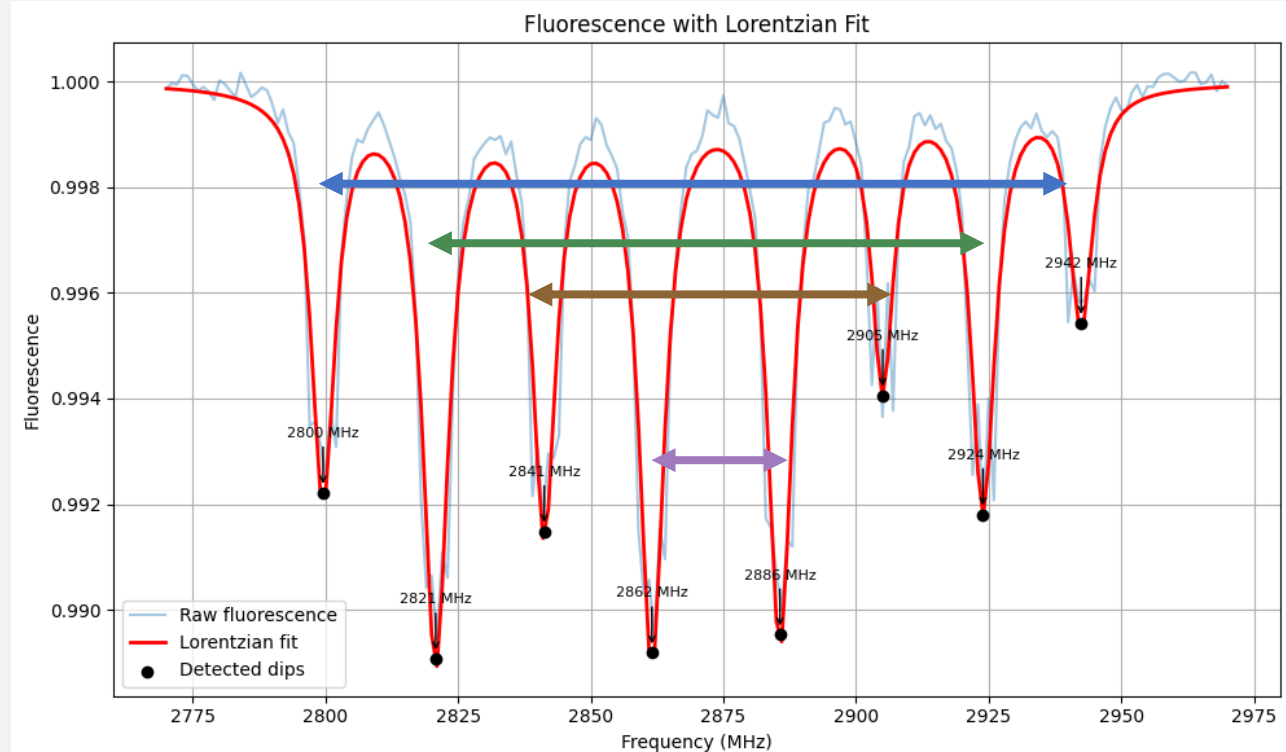


Common Magnetometry Limitations



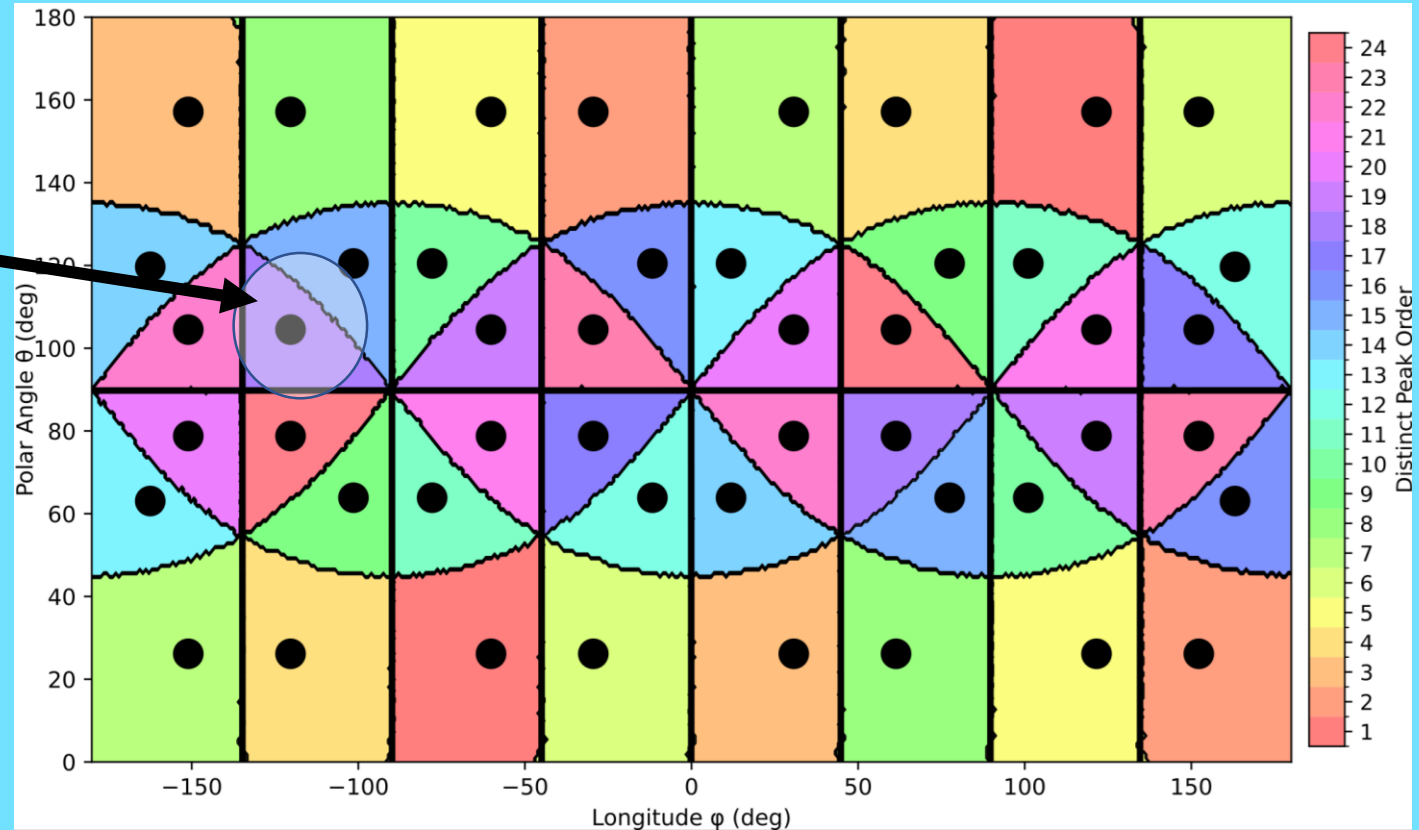
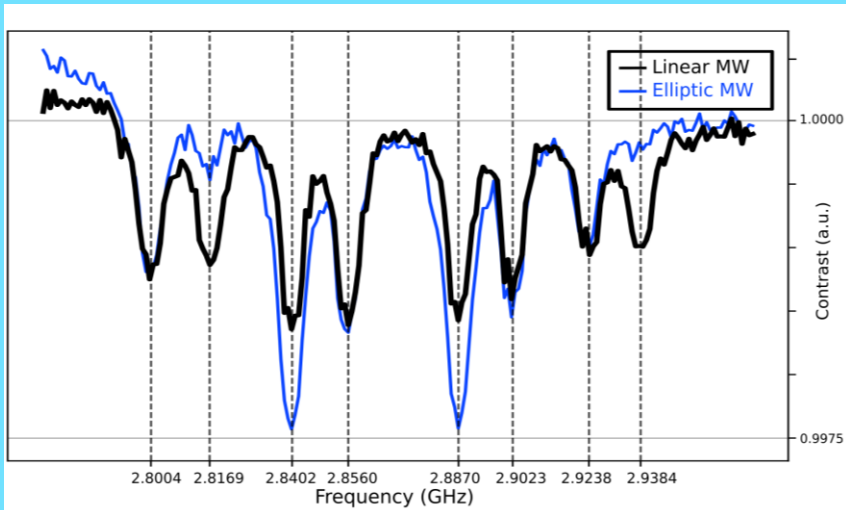
$$\begin{pmatrix} B_z^1 \\ B_z^2 \\ B_z^3 \\ B_z^4 \end{pmatrix} = \begin{pmatrix} -\frac{\sqrt{2}}{\sqrt{3}} & 0 & \frac{1}{\sqrt{3}} \\ \frac{\sqrt{2}}{\sqrt{3}} & 0 & \frac{1}{\sqrt{3}} \\ 0 & \frac{\sqrt{2}}{\sqrt{3}} & \frac{1}{\sqrt{3}} \\ 0 & -\frac{\sqrt{2}}{\sqrt{3}} & \frac{1}{\sqrt{3}} \end{pmatrix} \begin{pmatrix} B_x \\ B_y \\ B_z \end{pmatrix}$$

Linearly polarized MW



Limitations

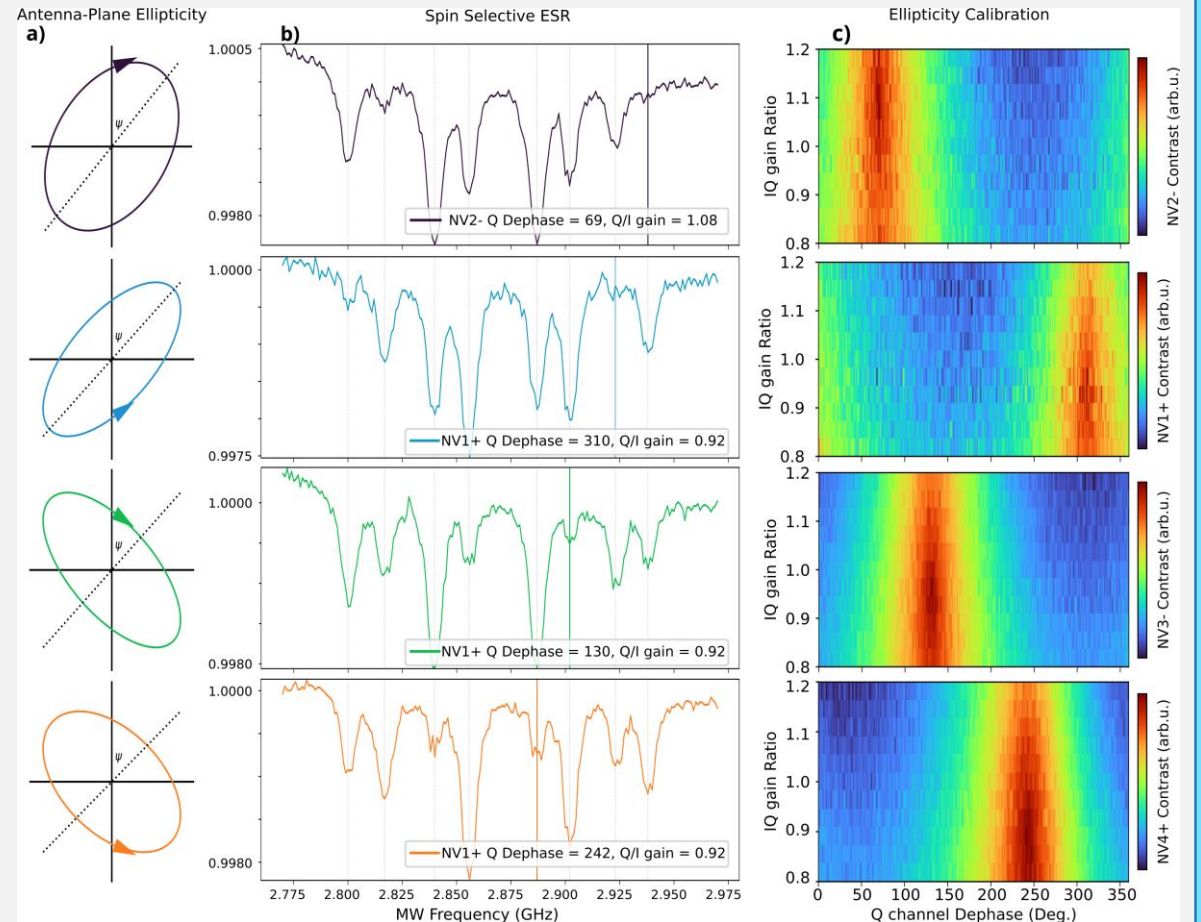
- **Single ESR:**
 - No information on axis producing peaks.
 - 48 possible solutions.
 - Typical magnetometry is limited to a single coloured area.



Solution: Elliptically polarized MW

Proof of concept for vector magnetometry protocol

- Individual MW configurations optimal for each crystal axis found in agreement with theoretical predictions.
- The spectrum response to each polarized microwave pulse reveals the axis orientation of each peak.
- Independent of B .



With our protocol you can measure:

1. The full vector magnetic field with a bulk diamond sample
2. The orientation of a single NV center with respect to a given external magnetic field