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DI TORINO



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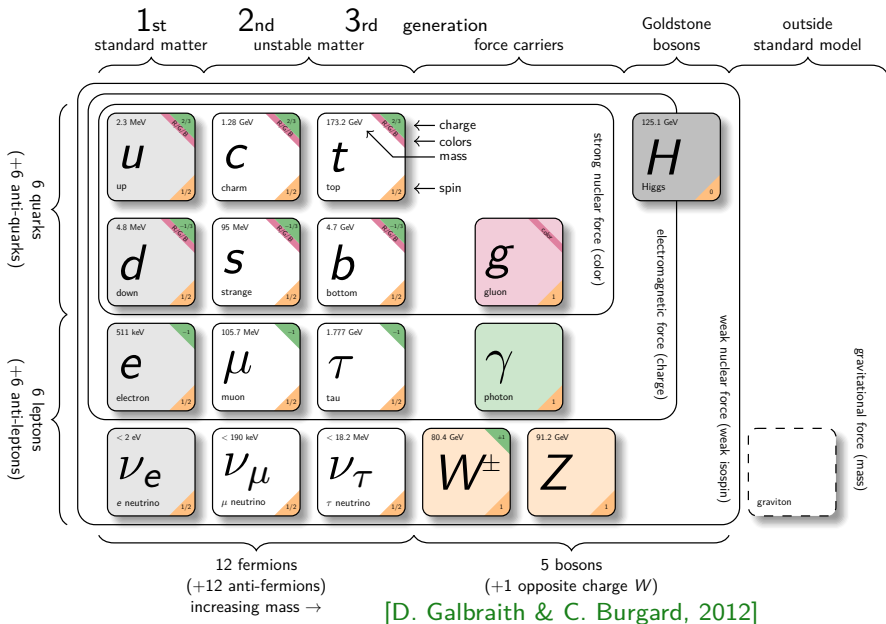
gariazzo@ific.uv.es



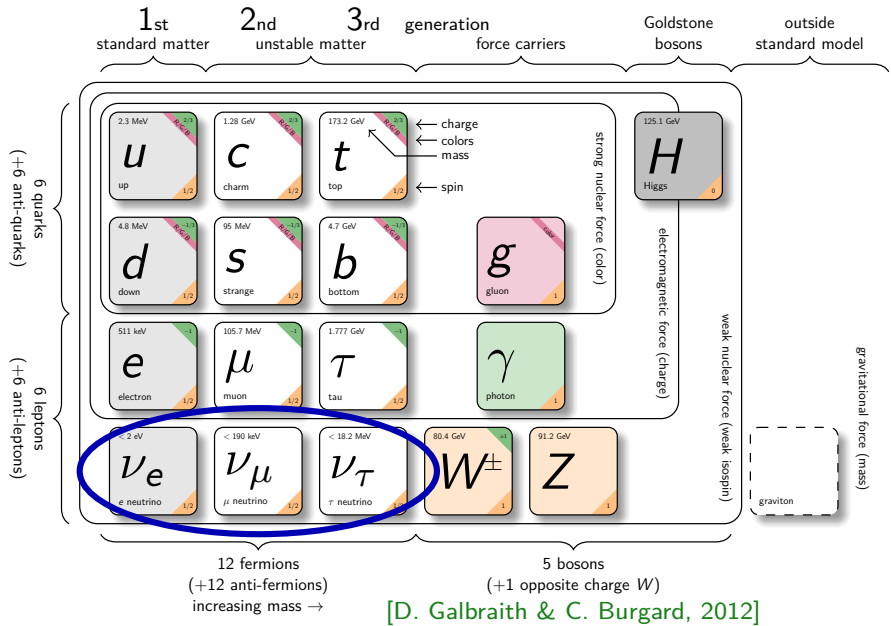
Neutrinos and cosmology

IFIC Newcomers Fest 2025, 12/12/2025

The Standard Model of Particle Physics

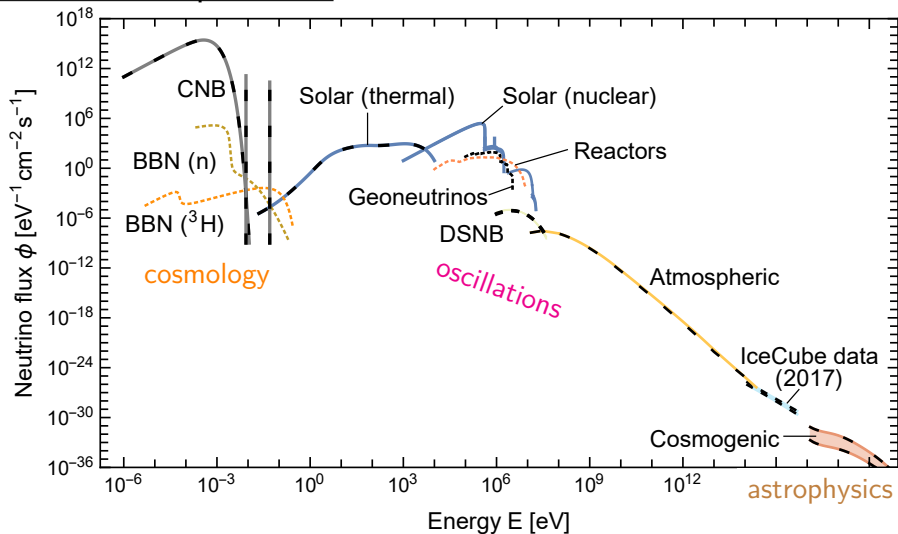


The Standard Model of Particle Physics



[D. Galbraith & C. Burgard, 2012]

Neutrino spectrum



neutrinos at all energies provide valuable information!

Three Neutrino Oscillations

$$\nu_\alpha = \sum_{k=1}^3 U_{\alpha k} \nu_k \quad (\alpha = e, \mu, \tau)$$

$U_{\alpha k}$ described by 3 mixing angles θ_{12} , θ_{13} , θ_{23} and one CP phase δ

Current knowledge of the 3 active ν mixing: [JHEP 02 (2021) update]

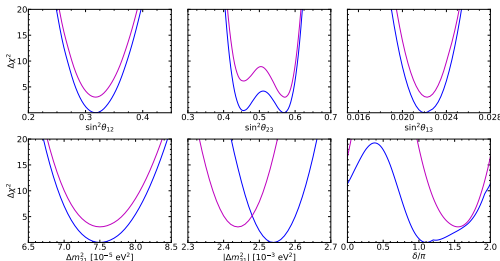
NO/NH: Normal Ordering/Hierarchy, $m_1 < m_2 < m_3$

IO/IH: Inverted O/H, $m_3 < m_1 < m_2$

$$\begin{aligned} \Delta m_{21}^2 &= (7.50^{+0.22}_{-0.20}) \cdot 10^{-5} \text{ eV}^2 \\ |\Delta m_{31}^2| &= (2.54 \pm 0.03) \cdot 10^{-3} \text{ eV}^2 \text{ (NO)} \\ &= (2.44 \pm 0.03) \cdot 10^{-3} \text{ eV}^2 \text{ (IO)} \end{aligned}$$

$$\begin{aligned} 10 \sin^2(\theta_{12}) &= 3.18 \pm 0.16 \\ 10^2 \sin^2(\theta_{13}) &= 2.200^{+0.069}_{-0.062} \text{ (NO)} \\ &= 2.225^{+0.064}_{-0.070} \text{ (IO)} \\ 10 \sin^2(\theta_{23}) &= 4.55 \pm 0.13 \text{ (NO)} \\ &= 5.71^{+0.14}_{-0.17} \text{ (IO)} \end{aligned}$$

$$\begin{aligned} \delta/\pi &= 1.10^{+0.27}_{-0.12} \text{ (NO)} \\ &= 1.54 \pm 0.14 \text{ (IO)} \end{aligned}$$



mass ordering
still unknown

δ still unknown

see also: <http://globalfit.astroparticles.es>

Neutrinos and their masses

Normal ordering (NO)

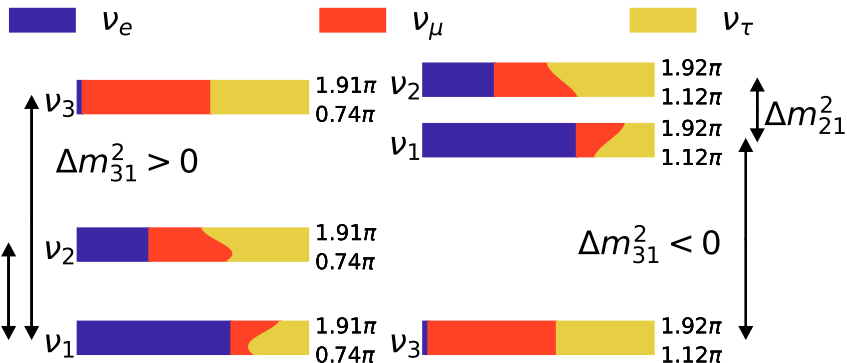
$$m_1 < m_2 < m_3$$

$$\sum m_k \gtrsim 0.06 \text{ eV}$$

Inverted ordering (IO)

$$m_3 < m_1 < m_2$$

$$\sum m_k \gtrsim 0.1 \text{ eV}$$



Absolute scale unknown!

Can we constrain the mass ordering using bounds on $\sum m_\nu$?

■ Phenomenology of neutrino physics

What can we learn about neutrino mixing and masses?

■ **Global fit of mixing parameters:** [JHEP 02 (2021) and update, in prep.]
with M. Tórtola, J. Valle, O. Mena

■ **Mass ordering:**

- review [Frontiers 5 (2018)],
- Bayesian methods [JCAP 03 (2018), JCAP 10 (2022)],
- DUNE phenomenology [PRD 100 (2019)]

with M. Tórtola, O. Mena

■ **Absolute mass scale:**

- cosmological constraints [PRD 106 (2022), PRD 108 (2023)],
- tension terrestrial-cosmological observations [PDU 40 (2023)]

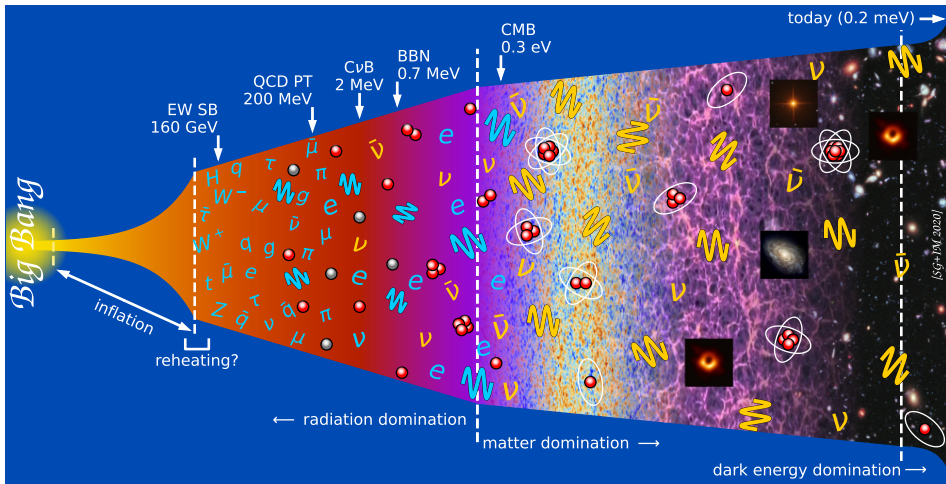
with O. Mena

■ **Additional states:**

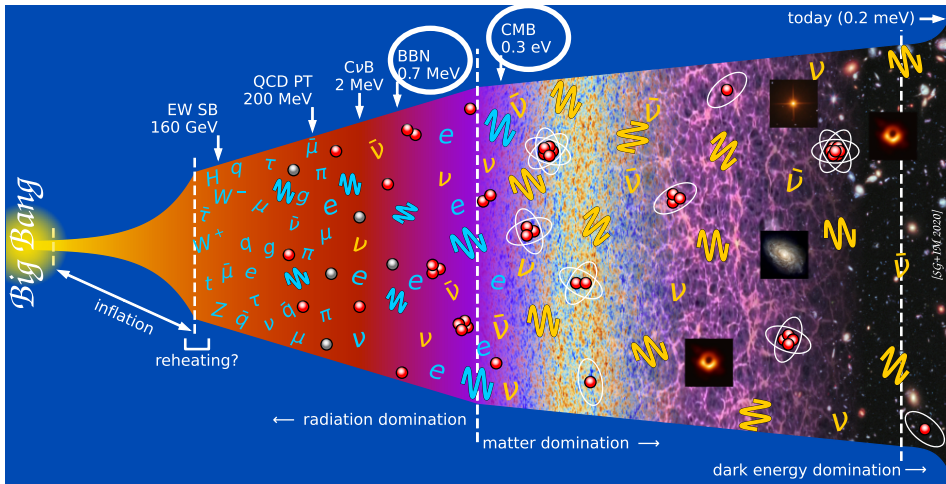
light sterile neutrinos at short baselines

[JPG 43 (2016), JHEP 06 (2017), PLB 782 (2018)]

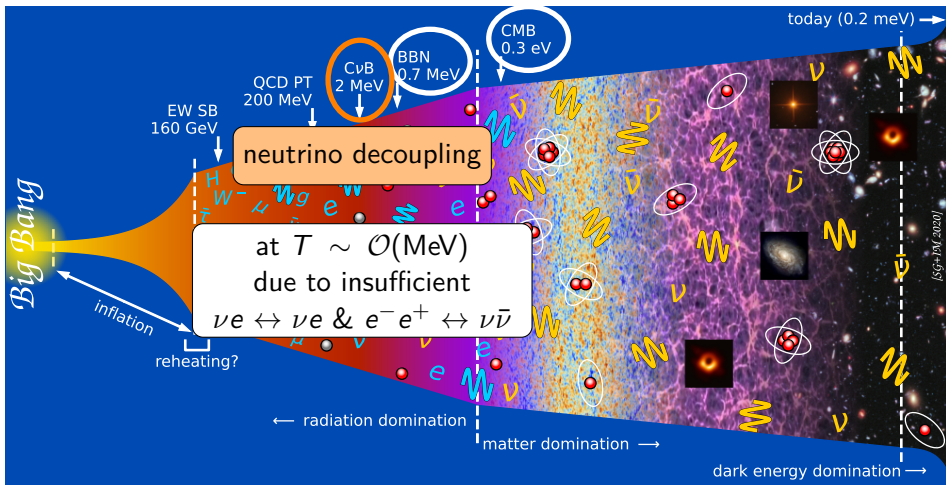
History of the universe



History of the universe



History of the universe



Phenomenology of relic neutrinos

Phenomenology related to relic neutrinos is very rich:

■ Decoupling processes:

- standard decoupling [JCAP 04 (2021)], asymmetries [in preparation]
- modified interactions [PLB 820 (2021), JCAP 08 (2025)]
or neutrino number [JCAP 07 (2019), PRD 104 (2021)],
- modified cosmology: low-reheating [PRL 135 (2025), in preparation],
additional particles [JCAP 12 (2023)]

with S. Pastor, G. Barenboim, N. Barbieri, P. Muñoz, A. Sánchez-Vargas

- Neutrino propagation: clustering in the MW [JCAP 01 (2020)],
also including modified interactions [in preparation]

with S. Pastor, P. Ghedini

- Direct detection of relic neutrinos: PTOLEMY [JCAP 07 (2019)]

- Indirect constraints on ν properties after decoupling:

N_{eff} [PRD 108 (2023)], chemical potential [in preparation], ...

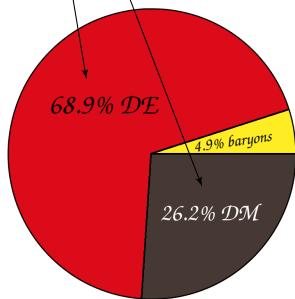
with O. Mena, P. Ghedini, D. Wang, ...

Cosmological parameters

General Relativity + Homogeneity isotropy

Cosmological evolution

Λ CDM model



[Planck Collaboration, 2018]

Λ CDM model described by 6 base parameters:

$\omega_b = \Omega_b h^2$ baryon density today;

$\omega_c = \Omega_c h^2$ CDM density today;

τ optical depth to reionization;

θ angular scale of acoustic peaks;

n_s tilt and

A_s amplitude of the power spectrum of initial curvature perturbations.

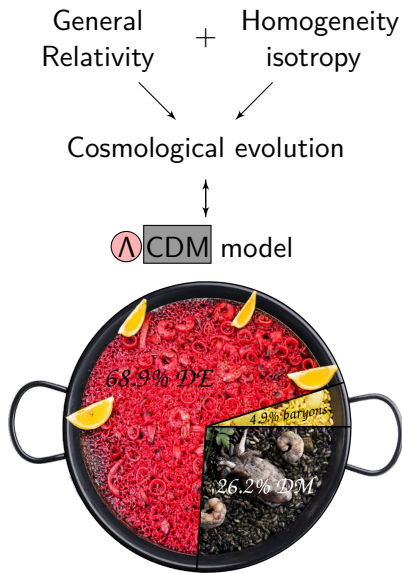
Other quantities can be studied:

H_0 Hubble parameter today;

σ_8 mean matter fluctuations at small scales;

...

Cosmological parameters



[Planck Collaboration, 2018]

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Cosmology

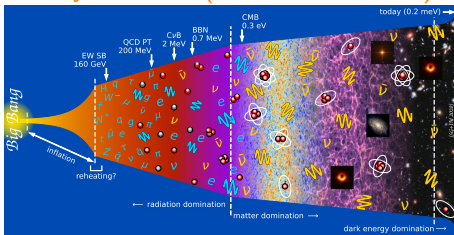
Universe content and properties through cosmological observations?

with O. Mena, D. Wang

- **Neutrino mass constraints:**
 - updated bounds [PRD 110 (2024), PRD 108 (2023), ...],
 - bayesian methods [EPJC 80 (2020), PRD 99 (2019)],
 - non-standard scenarios [PRD 112 (2025)]
 - [book contribution (2025)]
- **Dark energy/dark matter:** [PRD 105 (2022), PRD 106 (2022)]
- **Curvature:** [PDU 33 (2021)]
- **Dark radiation:**
 - review [Universe 8 (2022)]
 - axions [PRD 107 (2023)]
 - interacting neutrinos [JCAP 12 (2020), JCAP 06 (2022)]
- **Hubble tension:** [book contribution (2023)]

Summary

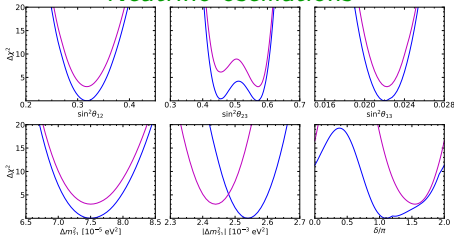
Early universe (ν dec., BBN, ...)



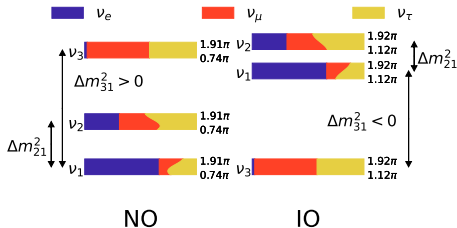
Universe content and history



Neutrino oscillations



Neutrino masses and mixings



Thanks for your attention!