

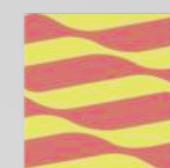
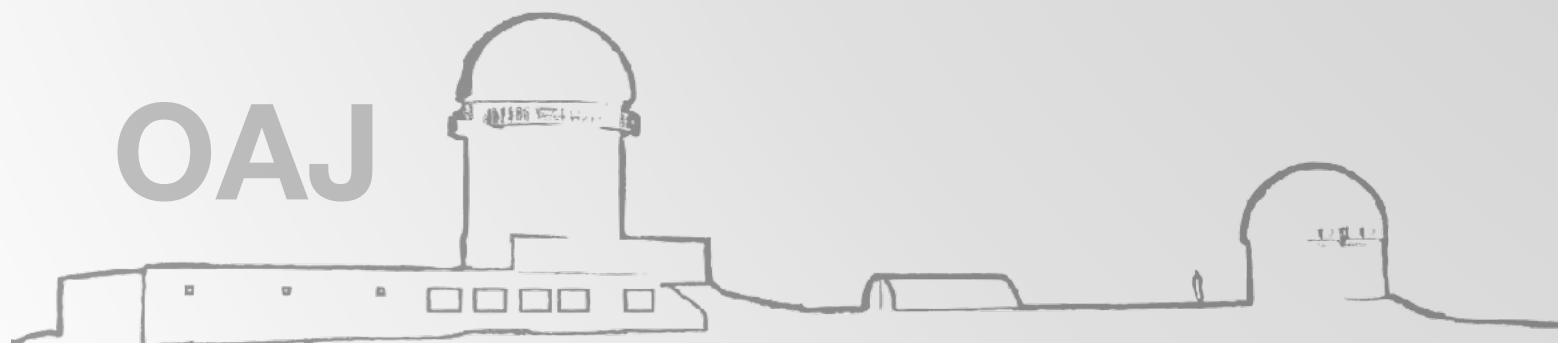
JPCam

The 1.2 Gpix camera for J-PAS

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**GOBIERNO
DE ARAGON**



MINISTERIO
DE CIENCIA, INNOVACIÓN
Y UNIVERSIDADES



Plan de Recuperación,
Transformación
y Resiliencia



Financiado por
la Unión Europea
NextGenerationEU

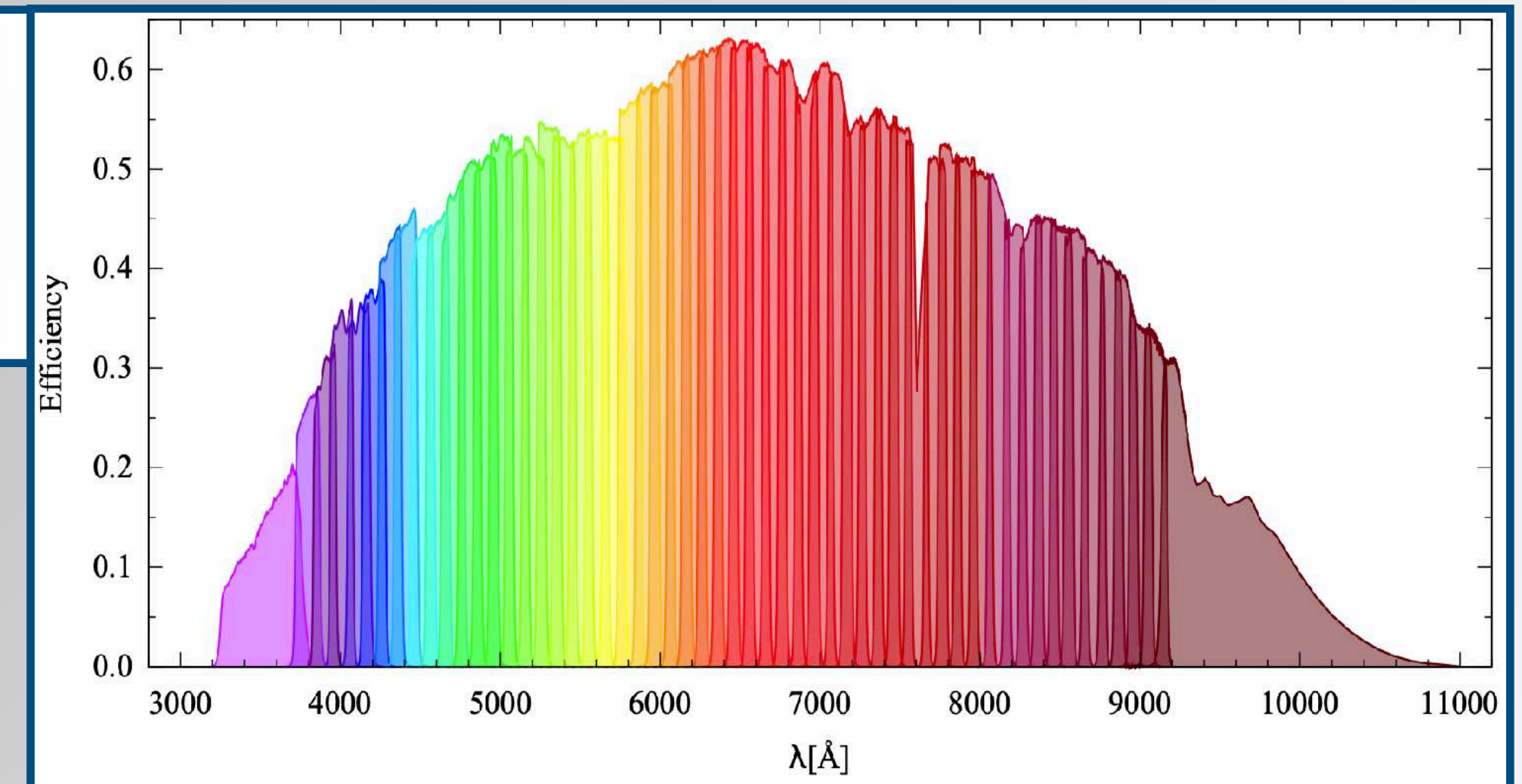
J-PAS

- J-PAS (<http://www.j-pas.org>) is a narrow-band, very wide field survey designed to produce photometric redshifts with a precision of 0.3% for over 14 million luminous red galaxies in the redshift interval $0.1 < z < 1.2$.
- It will image thousands of deg^2 of Northern Sky with 54 contiguous, narrow-band filters plus 3 medium- and broad-band filters.
- $m_{AB} (5\sigma \text{ in } 3'') = 21.3\text{-}22.4$ (with the narrow-bands).

Legacy data set:

J-PAS will provide low resolution spectroscopy ($R \sim 50$) for every observed pixel over thousands of deg^2 of the sky.

- | | | |
|--------------------------|---|------------------------------------|
| - BAOs | - High redshift galaxies. | - Galaxies in clusters and groups. |
| - 2D stellar populations | - Morphological studies. | - PN2 and MW halo stars. |
| - 2D star formation rate | - Mergers and interacting galaxies. | - Minor bodies as moving objects. |
| - Environment | - AGN and QSO at low and high redshift. | - ... |

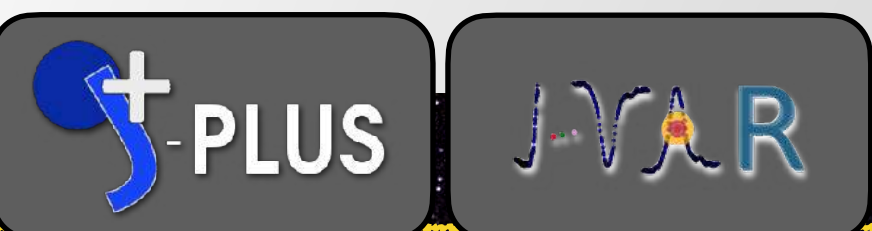


To carry out the J-PAS survey, it is required:

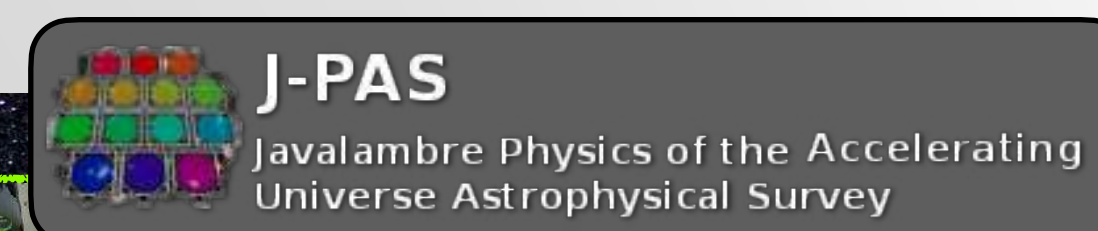
- Large *etendue* (2.5m diameter telescope + large field of view)
- Good image quality (good site and seeing limited telescope and instrument)
- Large number of optical filters (the J-PAS filter system)

The OAJ is conceived to carry out large astronomical surveys, starting with the **J-PAS** and **J-PLUS**. To this purpose, the OAJ has two unique telescopes and instrumentation of unusually large FoV and a data center.

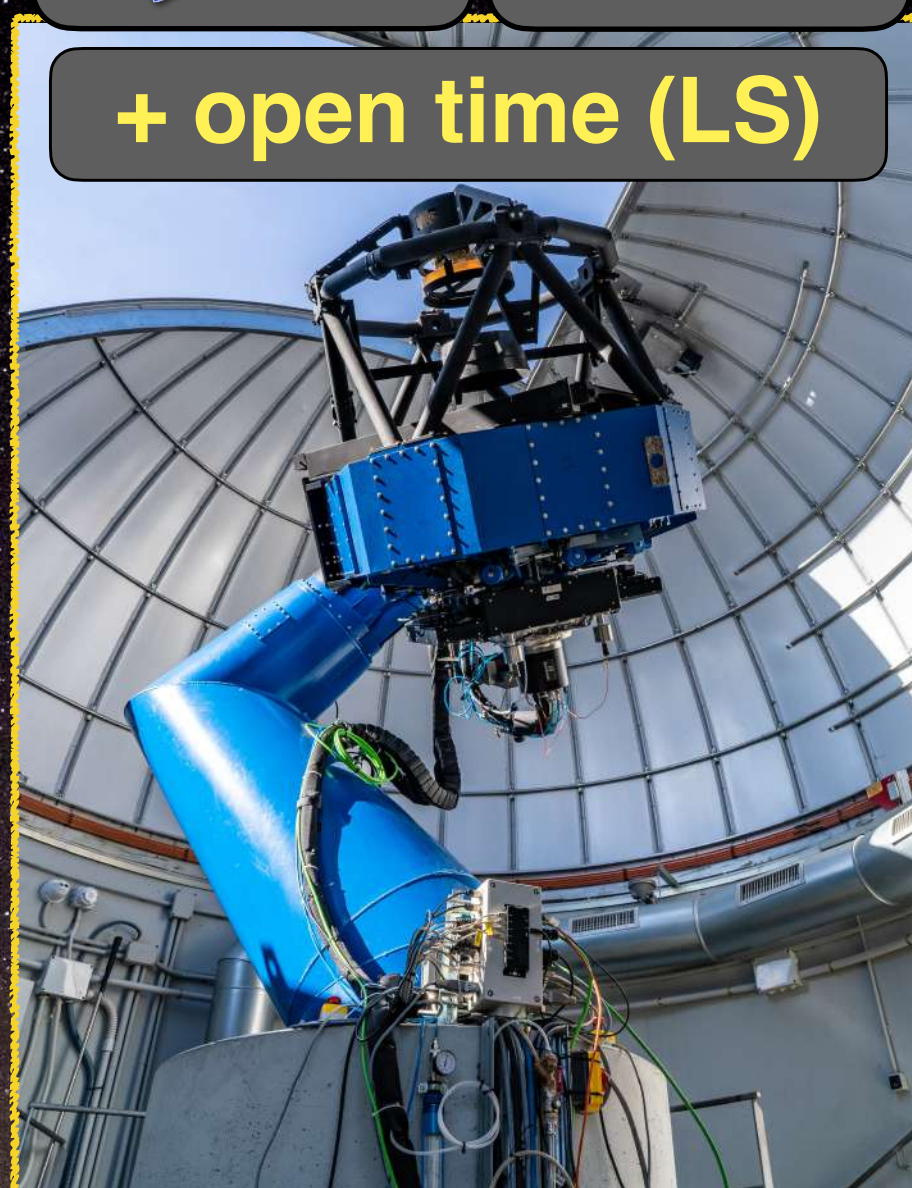
Located in Javalambre mountain (Teruel, Spain). 1957 m above sea level.



+ open time (LS)



+ open time (2025)



JAST80

Ritchey-Chrétien plus Field
Corrector
Diameter = 83 cm
FoV = 2 deg \emptyset
F/4,5
M2 hexapod



JST250

Ritchey-Chrétien plus Field Corrector
Diameter = 2,55 m
FoV = 3 deg \emptyset
F/3,5
M2 hexapod



UPAD

- Main **storage**: 1.1 PB of disk storage + 4 PB in a robotic tape library
- Data **processing**: 21 nodes with more that 450 cores
- External **Data Access** System: Redundant web-servers (> 30 TB of storage)





JST250

Outstanding Facility (OAJ-OF-1)



J-PAS
Javalambre Physics of the Accelerating
Universe Astrophysical Survey

+ MOUs with:



eROSITA-DE



JWST-PEARLS



WEAVE-QSO



EUCLID

+ open time

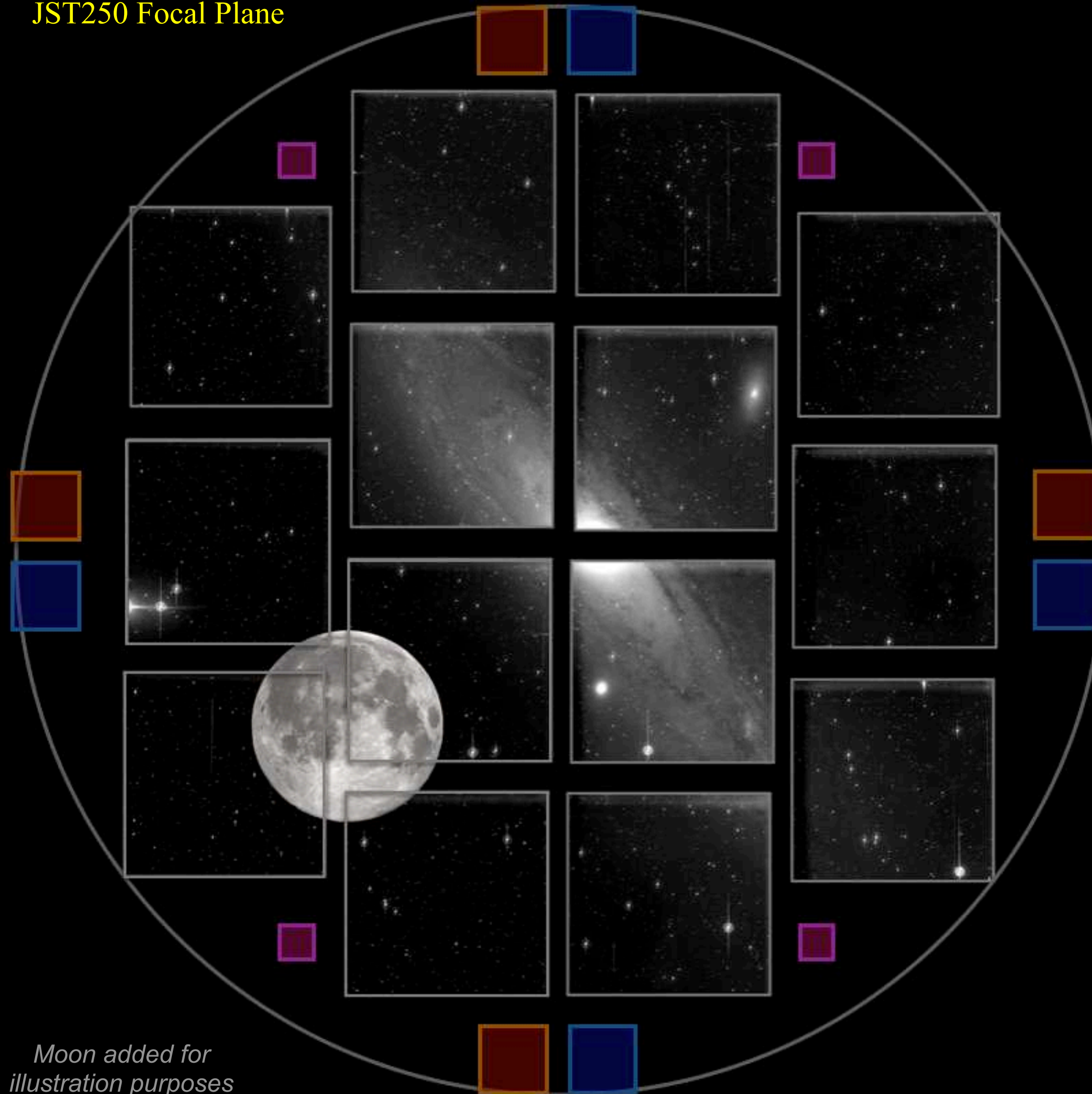
2025: Shared-Risk Programs + DDTs

2026: Legacy Surveys

JST250

Optical configuratio	Ritchey Chrétien like, equipped with a field corrector
M1 diameter	2.55 m
M2 diameter	1.18 m, with an hexapod system
Field Corrector	3 aspherical lenses
FoV diameter	3 deg (476 mm physical size)
Effective collecting	3.75 m ²
Etendue	26.5 m ² deg ²
EE50 (radius)	<4.75 microns (polychromatic) over the whole FoV
Focal length	9098 mm
Plate scale	22.67 arcsec/mm
Mount	Altazimuthal
Focus	Cassegrain

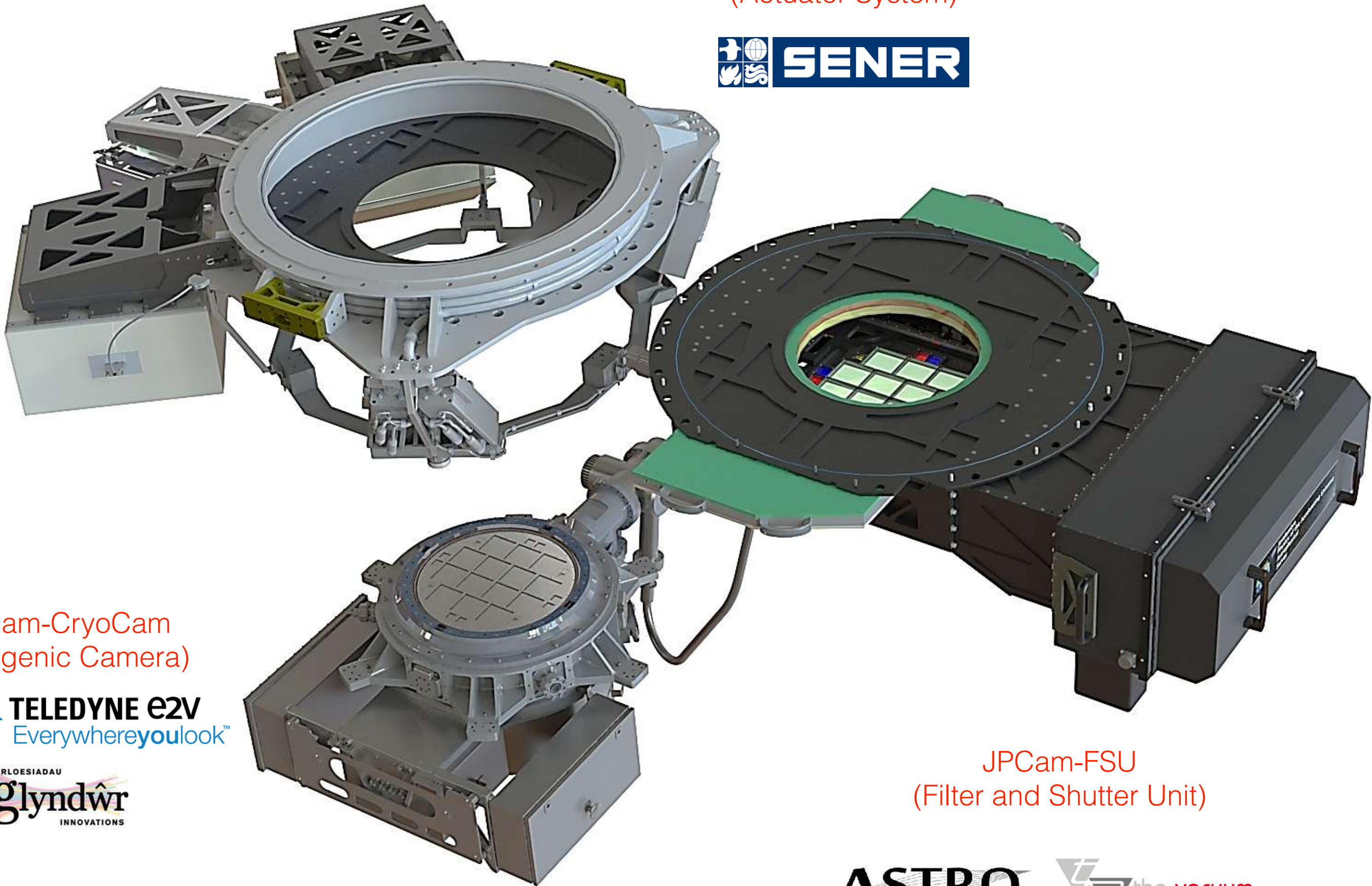
JST250 Focal Plane



*Moon added for
illustration purposes*



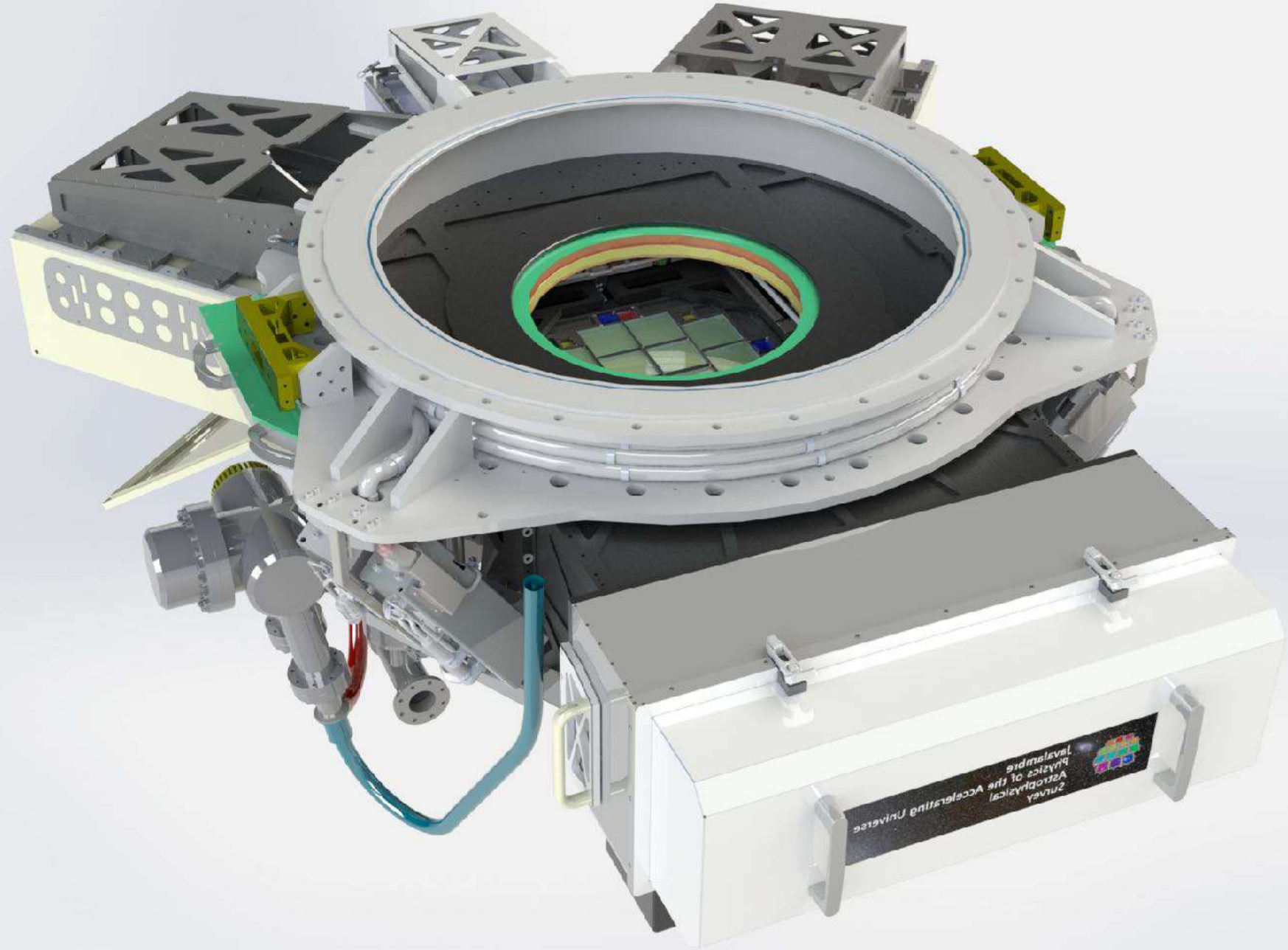
JPCam-AS
(Actuator System)



JPCam-CryoCam
(Cryogenic Camera)

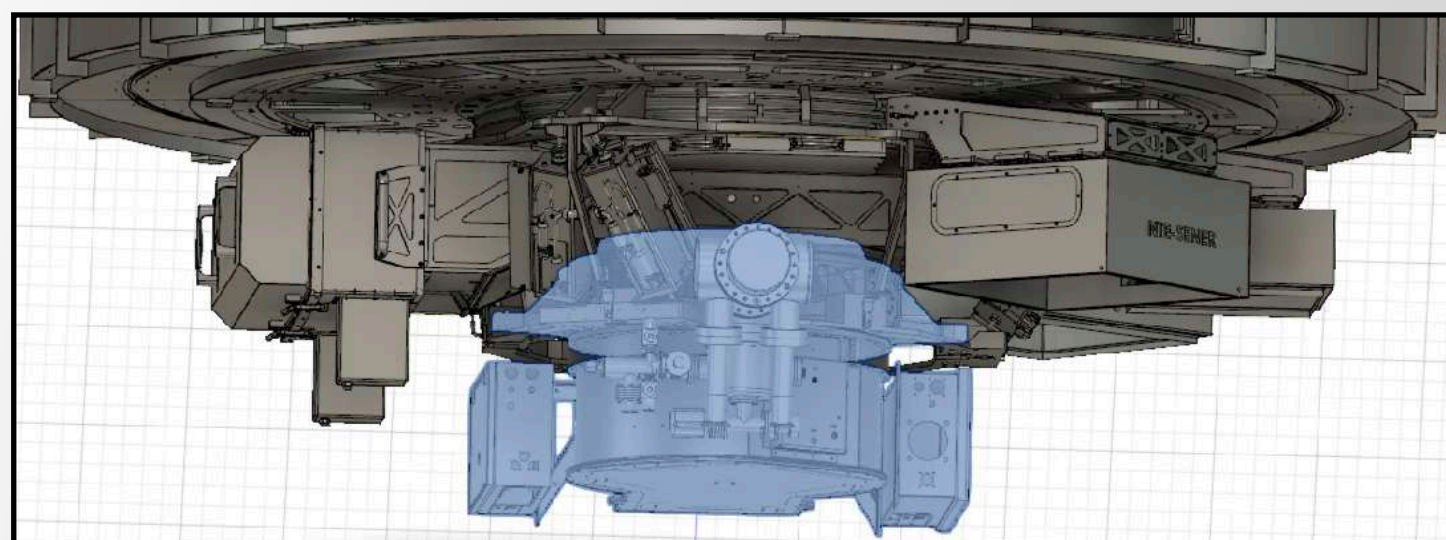


JPCam-FSU
(Filter and Shutter Unit)



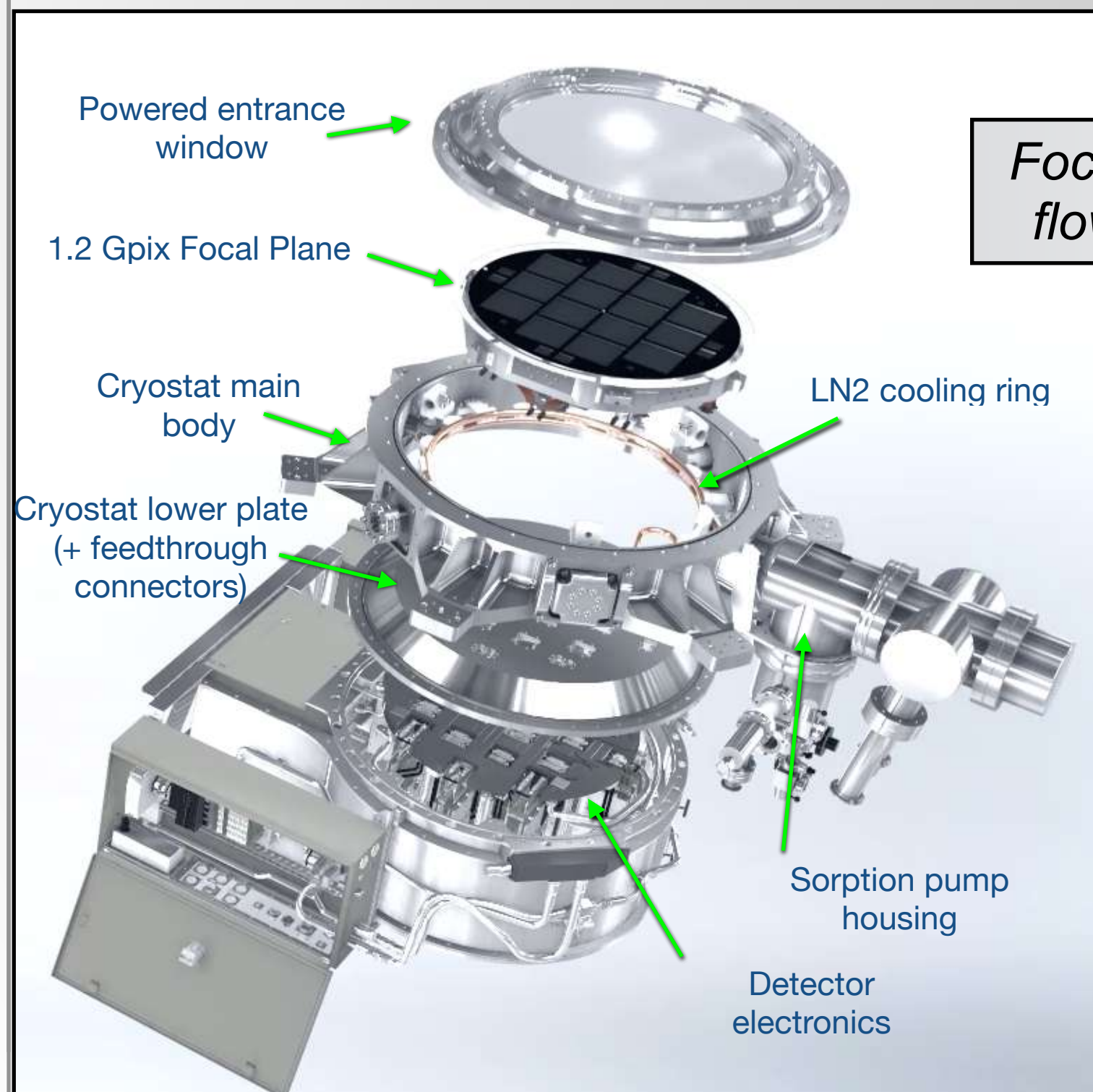
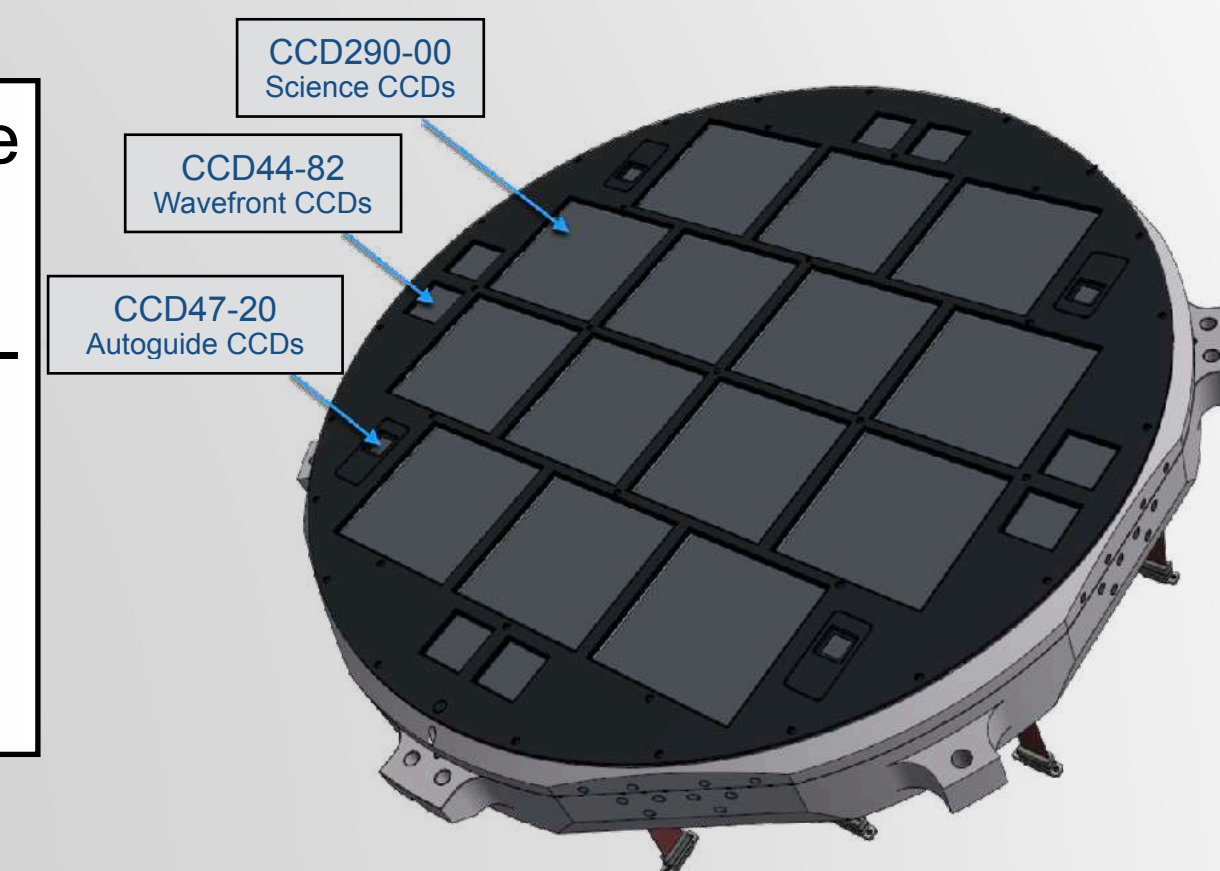
CCD format	14 × 9216 × 9232 pix, 10 μm pix ⁻¹ 1.2 Gpix camera
Pixel scale	0.2265" pix ⁻¹
Unvignetted FoV	3.4deg ² – (14×) 0.48deg × 0.51deg
Read out time (633kHz)	10.9 s (full frame) – 6.1 s (2x2 binning)
Read out noise (633kHz)	5.5 e ⁻ (RMS)
Read out time (400kHz)	16.4 s (full frame) – 8.9 s (2x2 binning)
Read out noise (400kHz)	4.3 e ⁻ (RMS)
Gain	2.274 e ⁻ ADU ⁻¹
Minimum exposure time	0.1 s
Exposure homogeneity	1 ms
Full well	> 125 000 e ⁻
Dark current	0.001 e ⁻ pix ⁻¹ s ⁻¹

JPCam Cryogenic Camera subsystem



1.2 Gpix focal plane mosaic composed by a total of 26 CCDs of three different types:

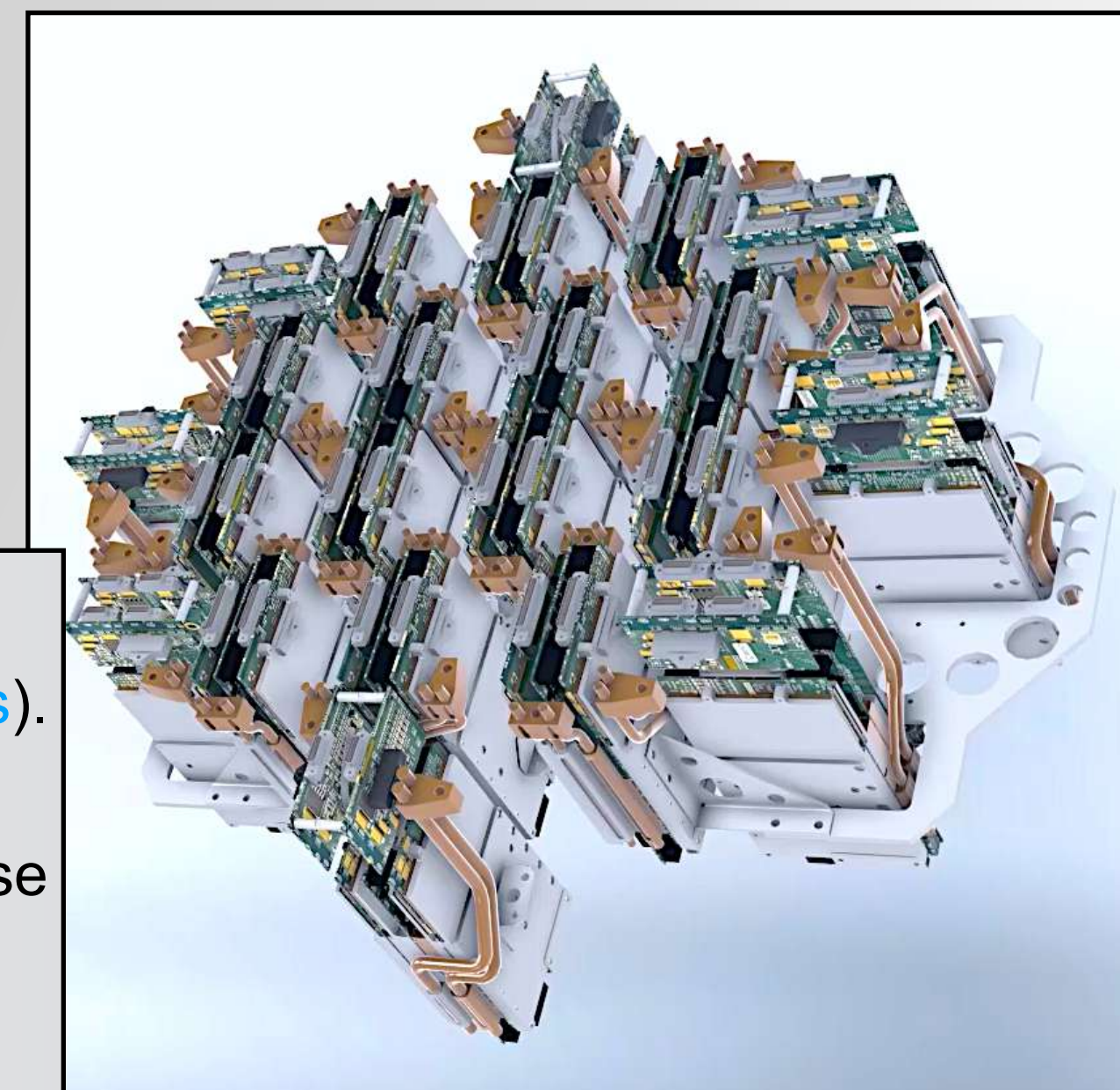
- **14 × CCD290–99** (**Science CCDs**). 9.216 x 9.232, 10μm pixel, non-inverted, full frame, deep depletion, astro multi-2 CCDs.
- **8 × CCD44–82** (**Wavefront Sensors**). 2048 X 2048 Frame-transfer.
- **4 × CCD47–20** (**Autoguide CCDs**). 1024 X 1024 Frame-transfer.



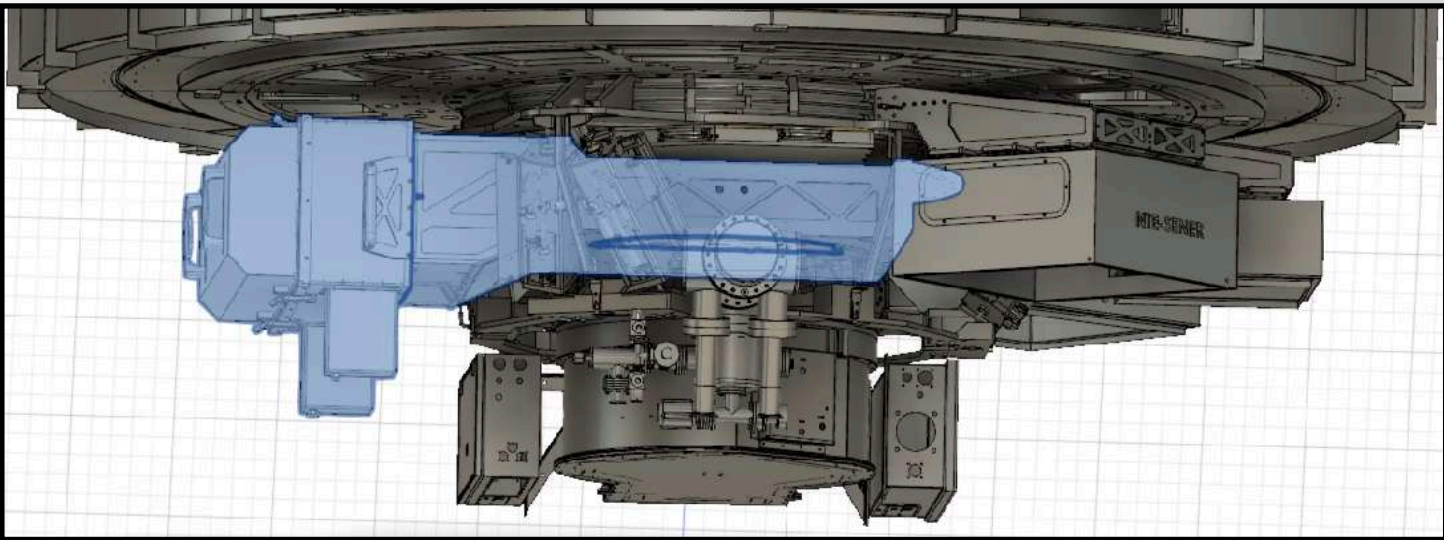
Focal Plane Array inside a continuous flow cryostat. LN2 cooled at -110°C.

Includes **detector electronics** to control all CCDs:

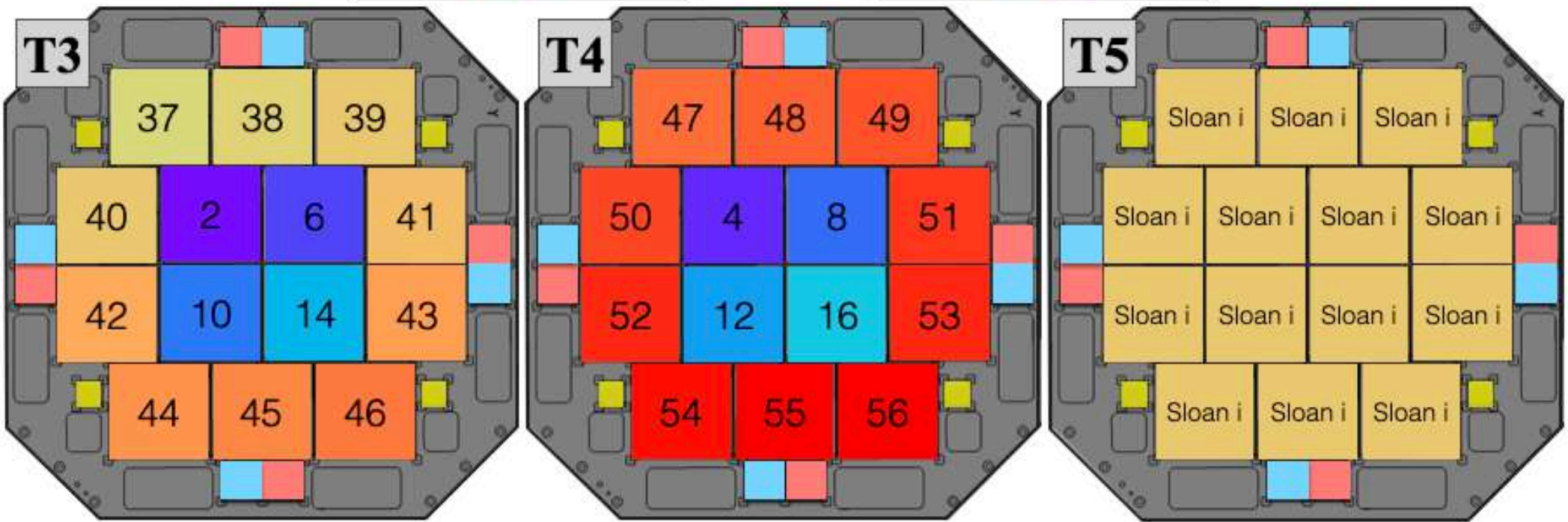
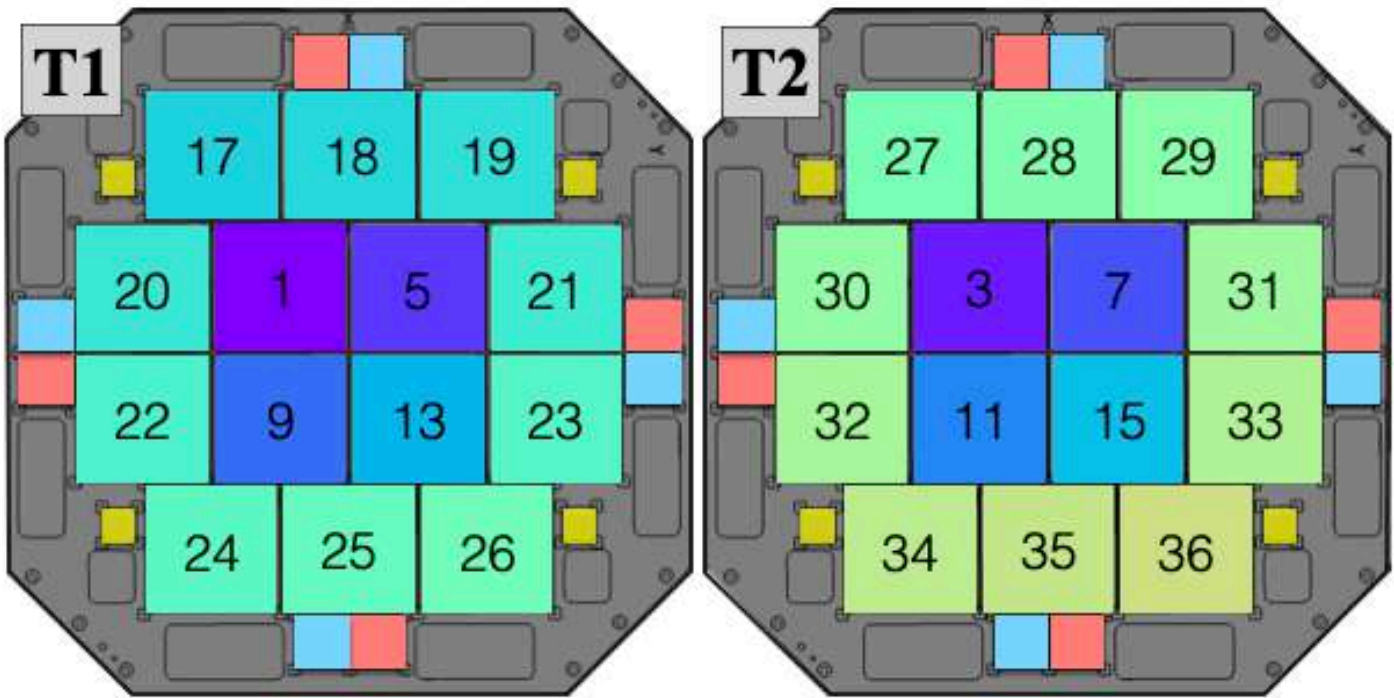
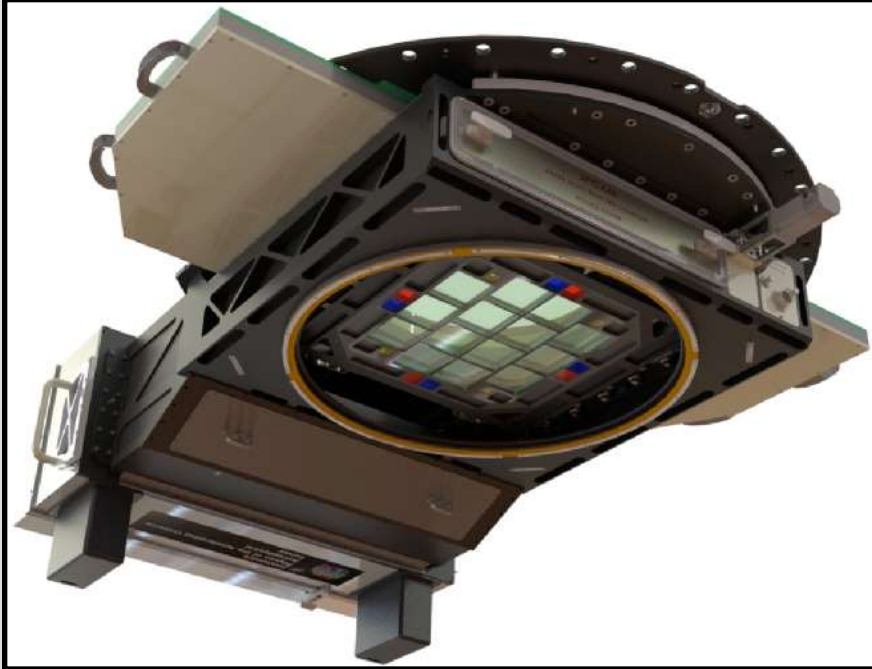
- 22 CCD drive modules + interface module (**more than 70 PCBs**). Glycol water cooling system.
- Power and data handling electronics (**224-channel** low noise electronics)
- Over **50 FPGAs** to handle the 2.4 GBytes of data per frame
- Digital CDS (Correlated Double Sampling) readout



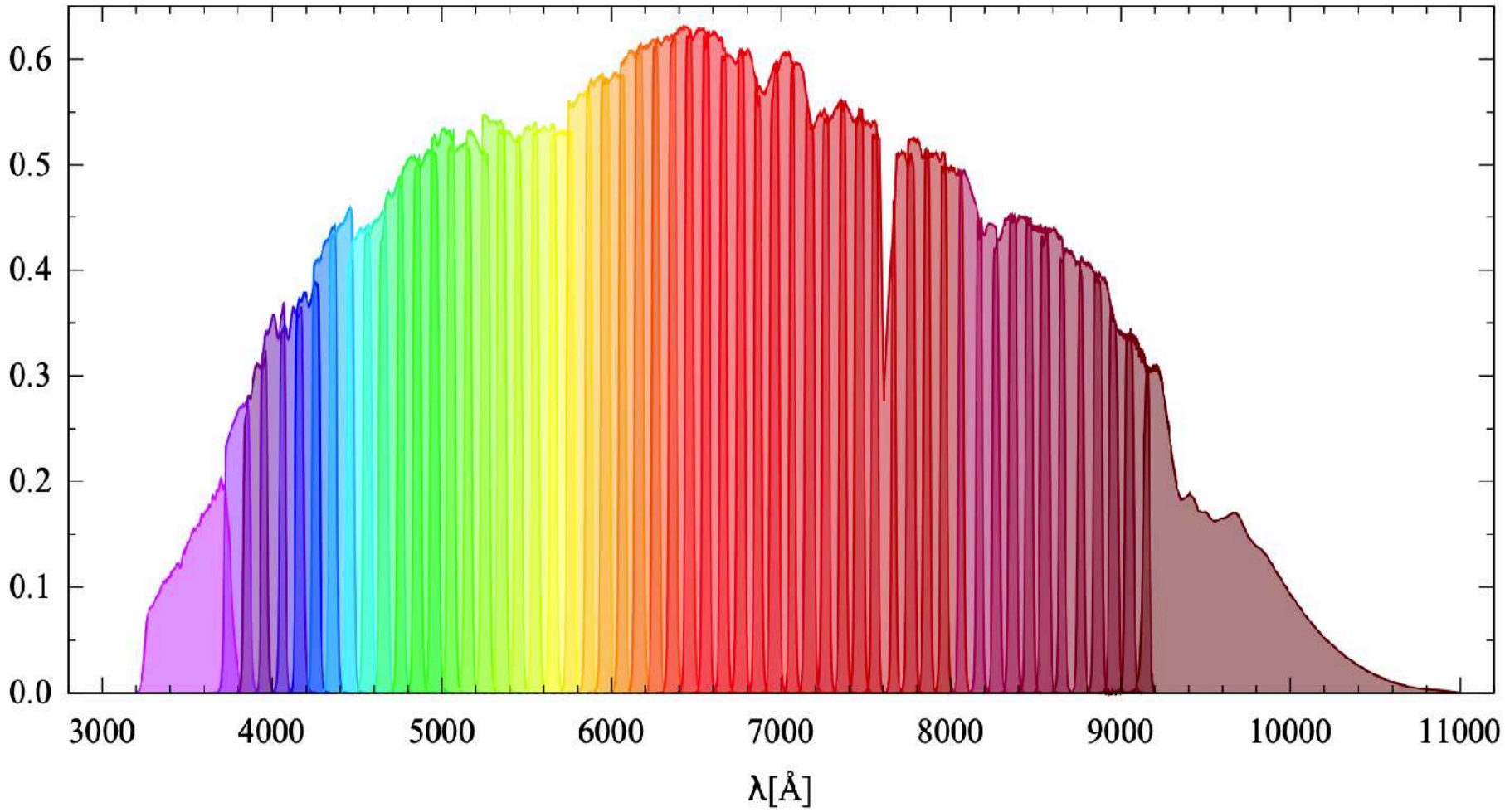
Filter and Shutter Unit



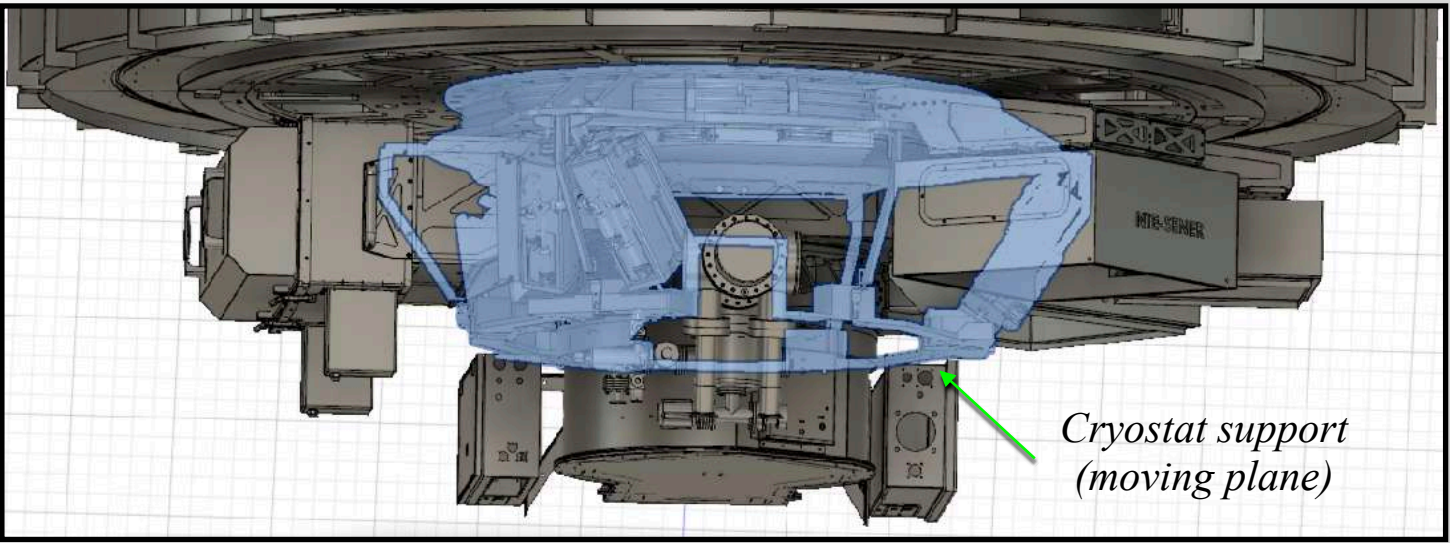
- Sealed to maintain a GN2 atm.
- “Two-curtain” 515mm shutter
- Admit 5 filter tray assemblies
- Filter tray change in <40s.



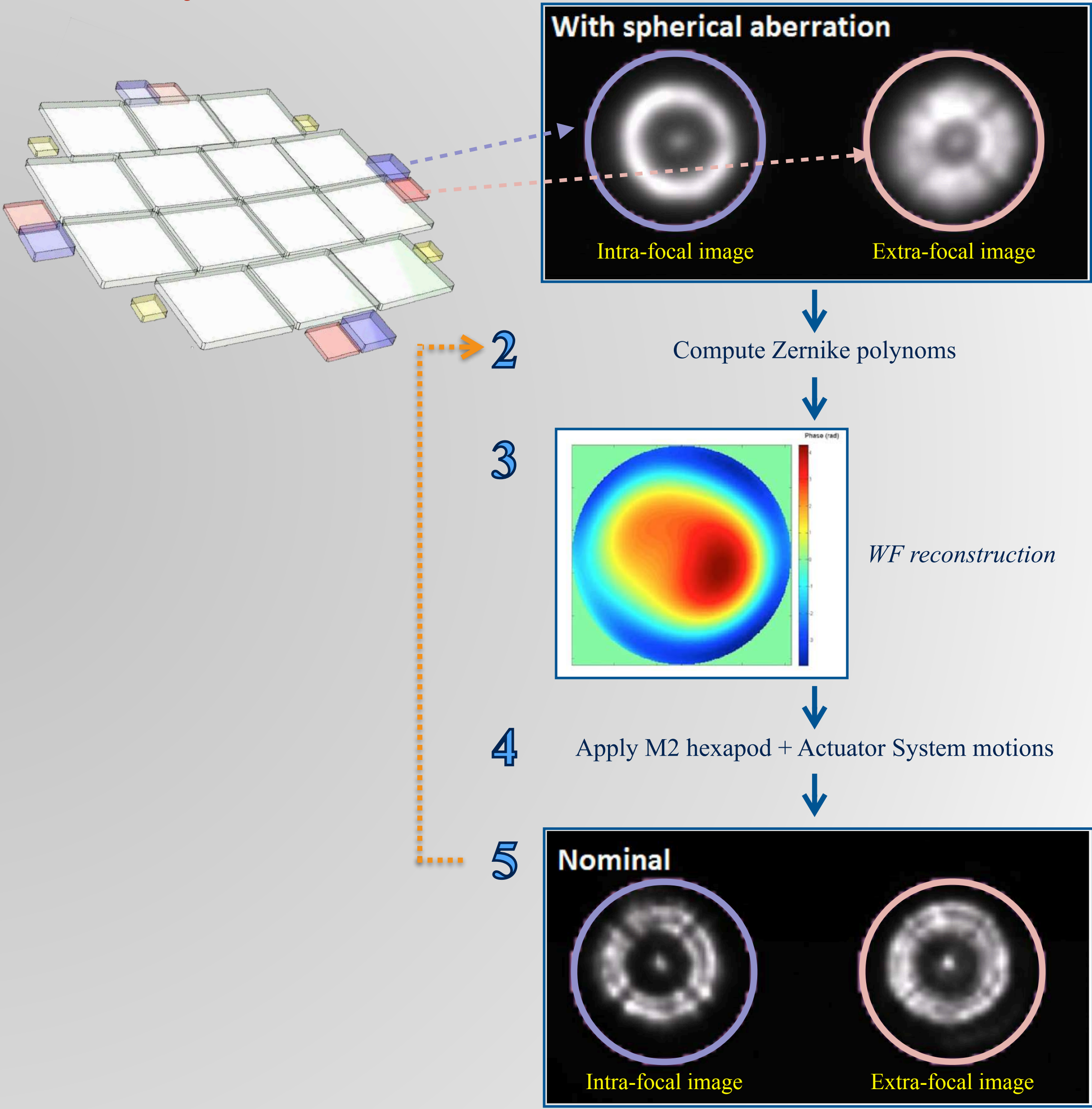
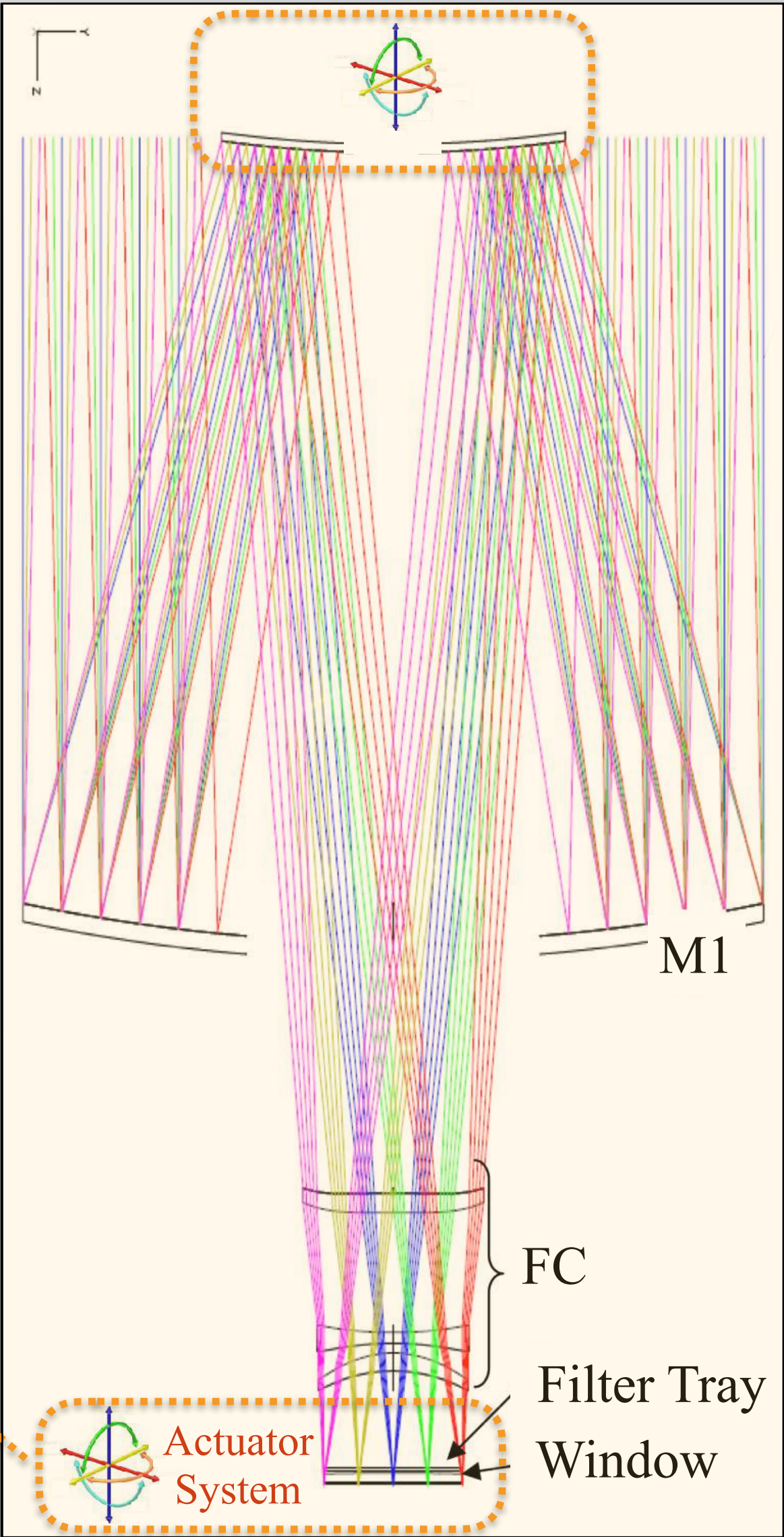
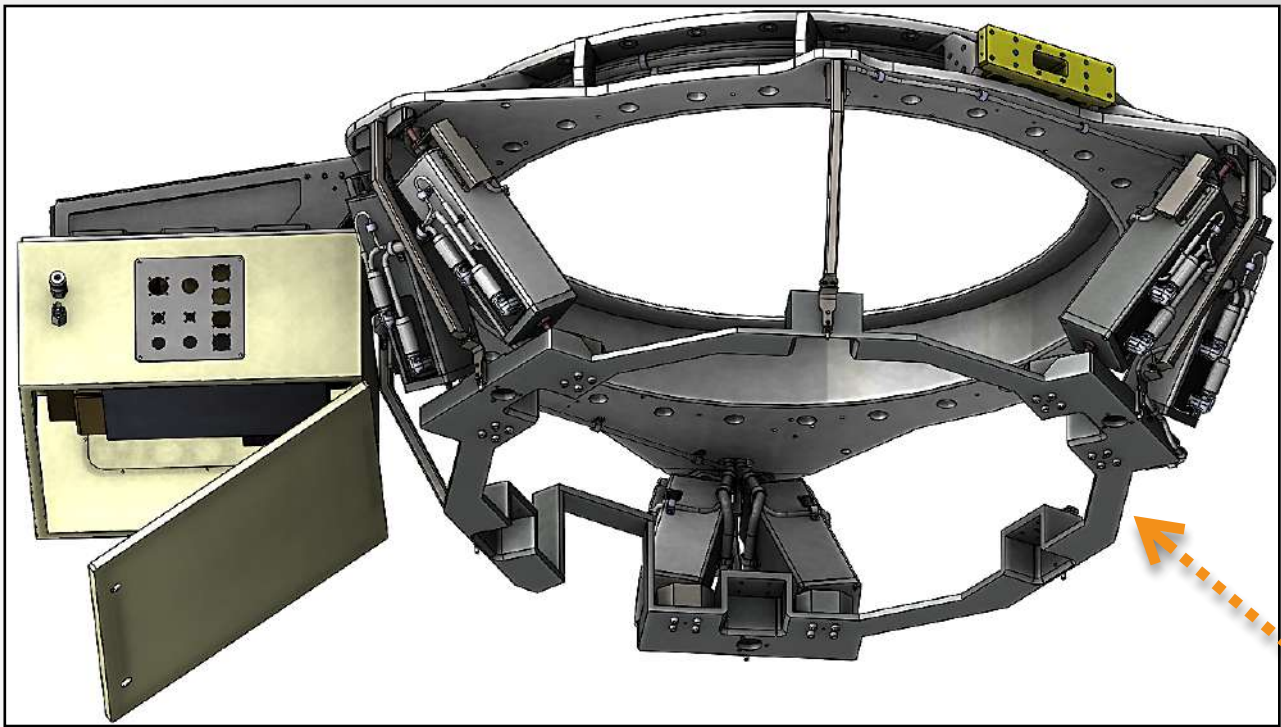
# Filter	CW (nm)	FWHM (nm)
1	348,5	49.5
2	378,5	15.5
3	390,0	14.5
4	400,0	14.5
5	410,0	14.5
...
54	900,0	14.5
55	910,0	14.5
56	1007,5	188.8



Actuator system and IQ control system



CryoCam weight	580 kg
Relative focus accuracy	$\pm 4 \mu\text{m}$
Relative tilt accuracy	$\pm 1 \text{ arcsec}$
Response time	$< 3 \text{ s}$



GOAL LA2.A1: The aim is to boost the completion of the JPCam, the 1.2Gpix panoramic camera integrated on the JST250 telescope at the ICTS Observatorio Astrofísico de Javalambre (OAJ) and start scientific operation → [The J-PAS Survey](#).

H2.1.1 – Completion of JPCam commissioning → [Milestone achieved in Nov. 2023](#)



H2.1.2 – Specialized technological assistance contract



To optimize JPCam performances: 2 year [specialized technological assistance](#) with the cryogenic camera manufacturer.

Contract's signature: 01/04/2023 — End of service: ~~31/3/2025~~ 31/9/2025

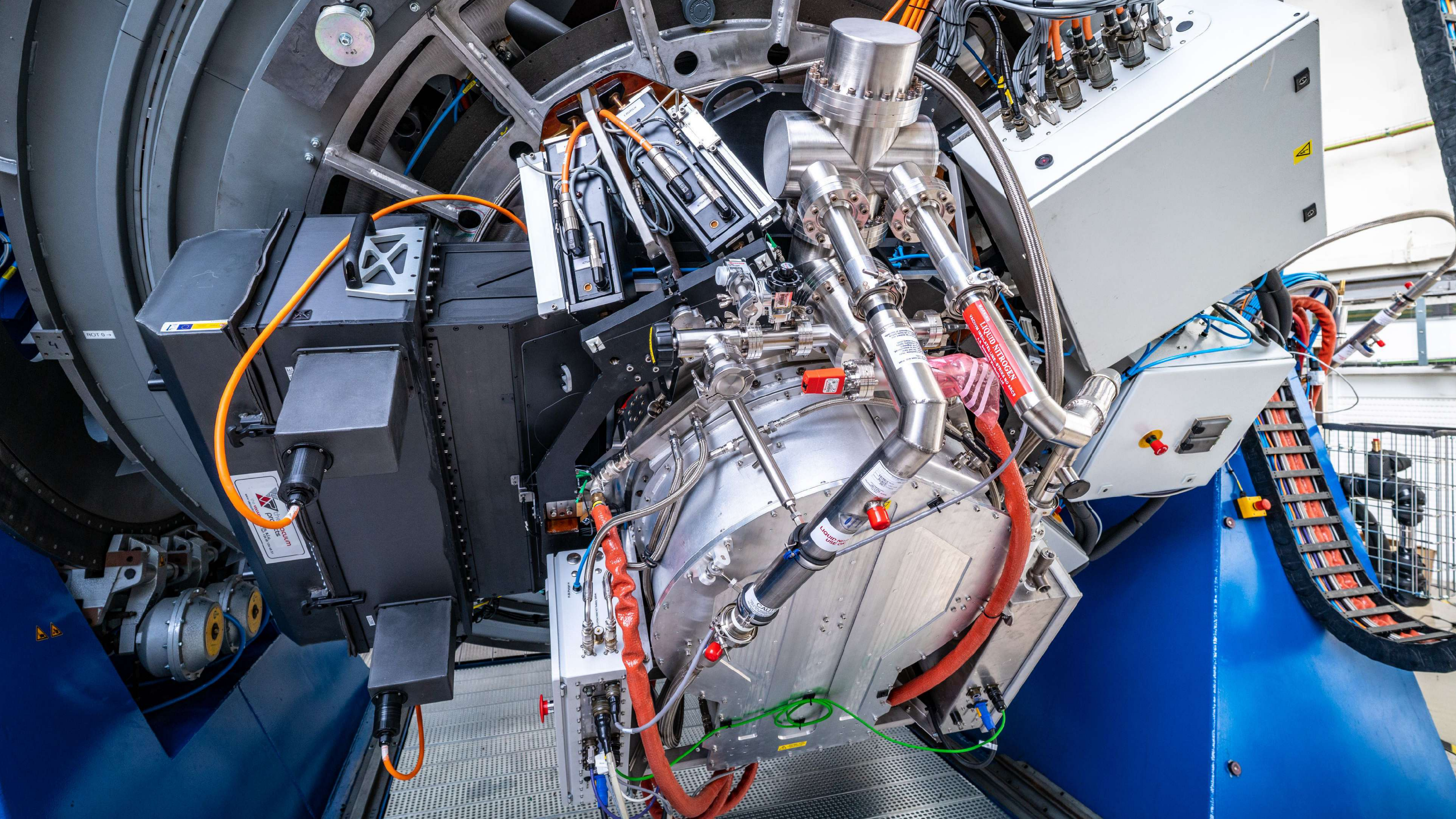
H2.1.3 – Acquisition of a set of auxiliary systems and equipment for JPCam and JST250



To maximize efficiency and availability during scientific operation: [acquisition of a set of auxiliary systems and equipment](#) for the operation of JPCam

Contract's signature: 28/12/2022 — End of supply: 30/8/2024

+ electronics engineer



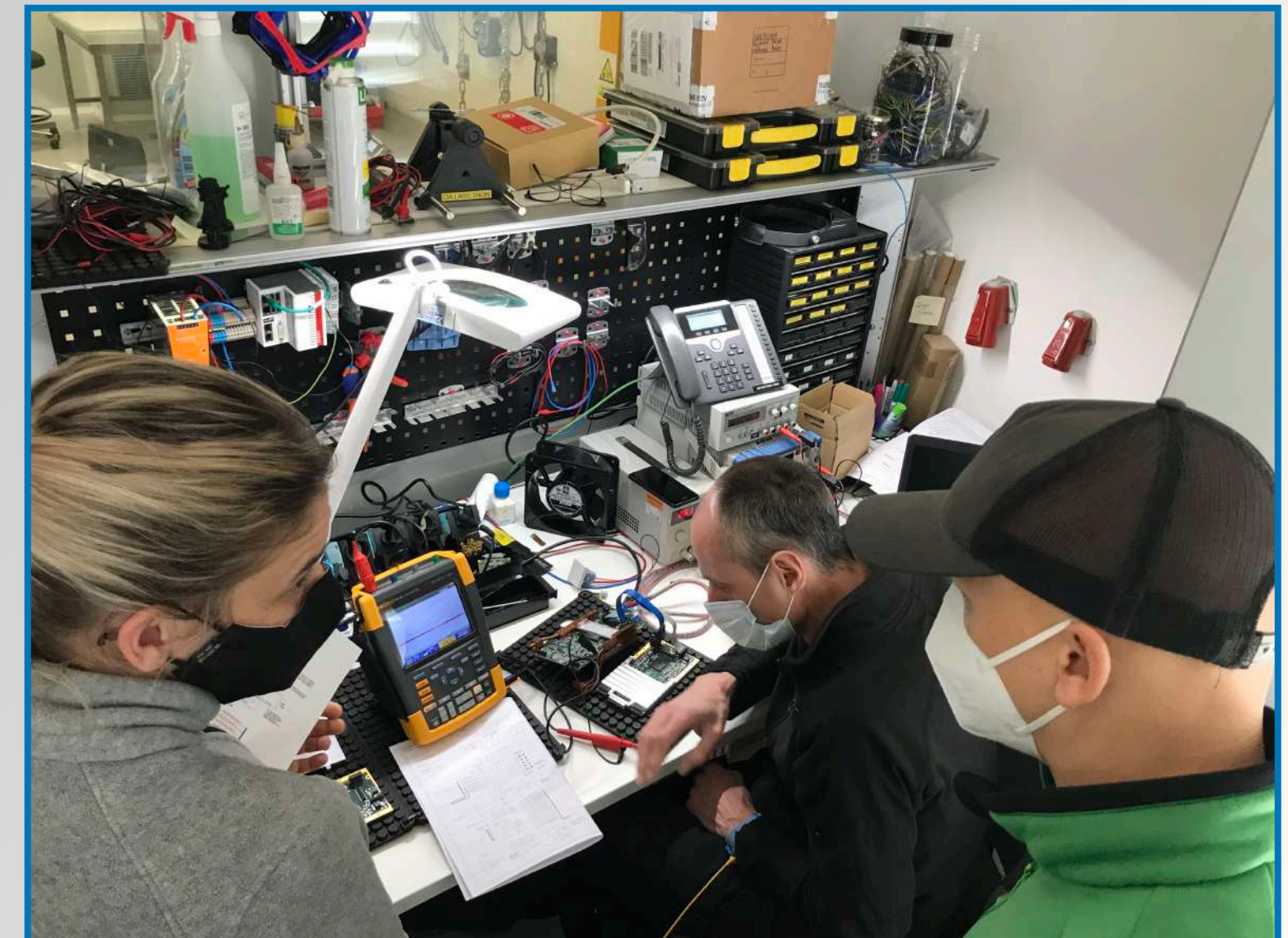
Complex JPCam electronics has represented the **most important commissioning challenge**.

Four Science Drive Modules and the Interface Module have developed faults during first operation.

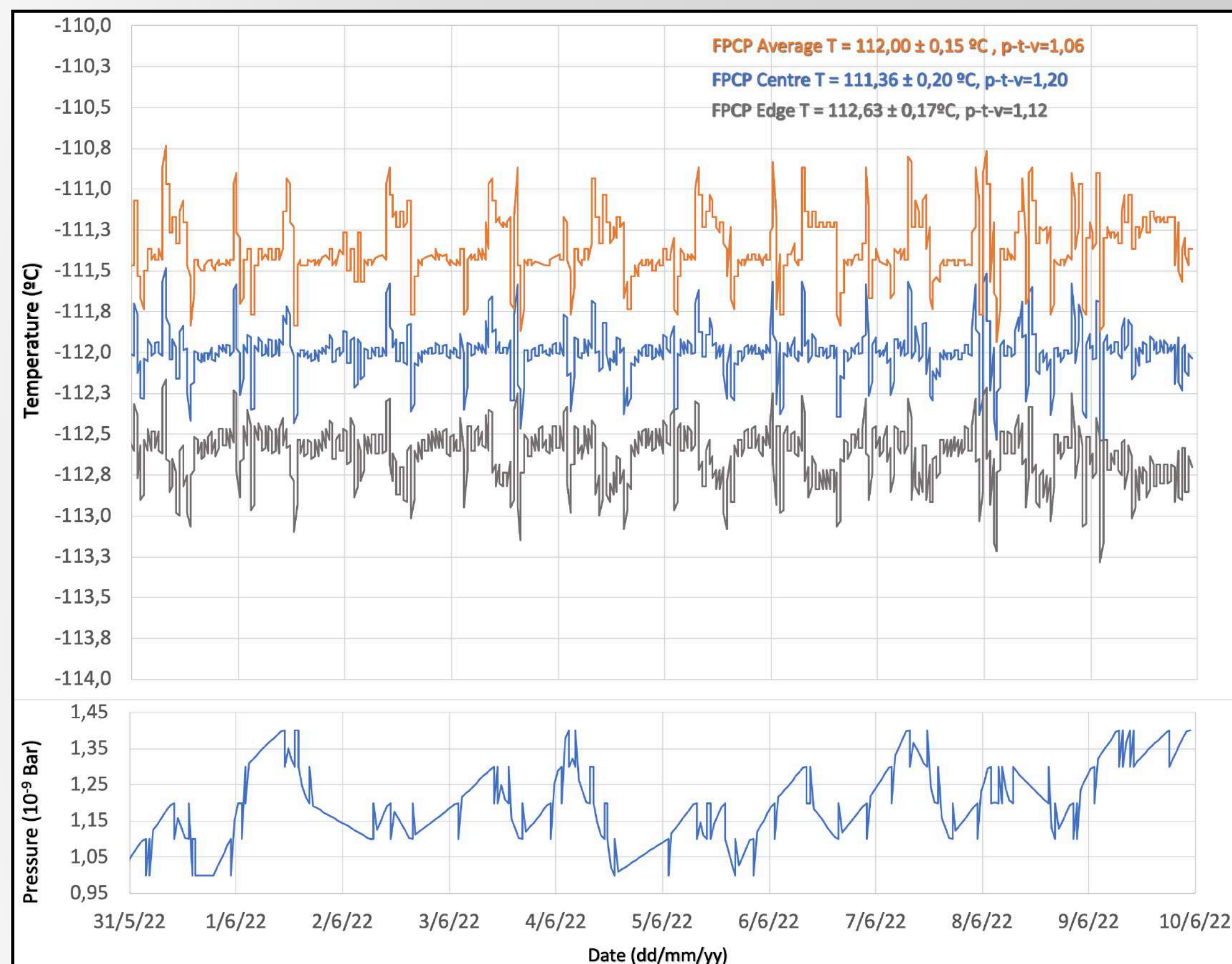
The root cause of these failures has been identified and a modification to the electronics has been implemented to prevent further failures.



Part of the commissioning performed with a reduced number of CCDs

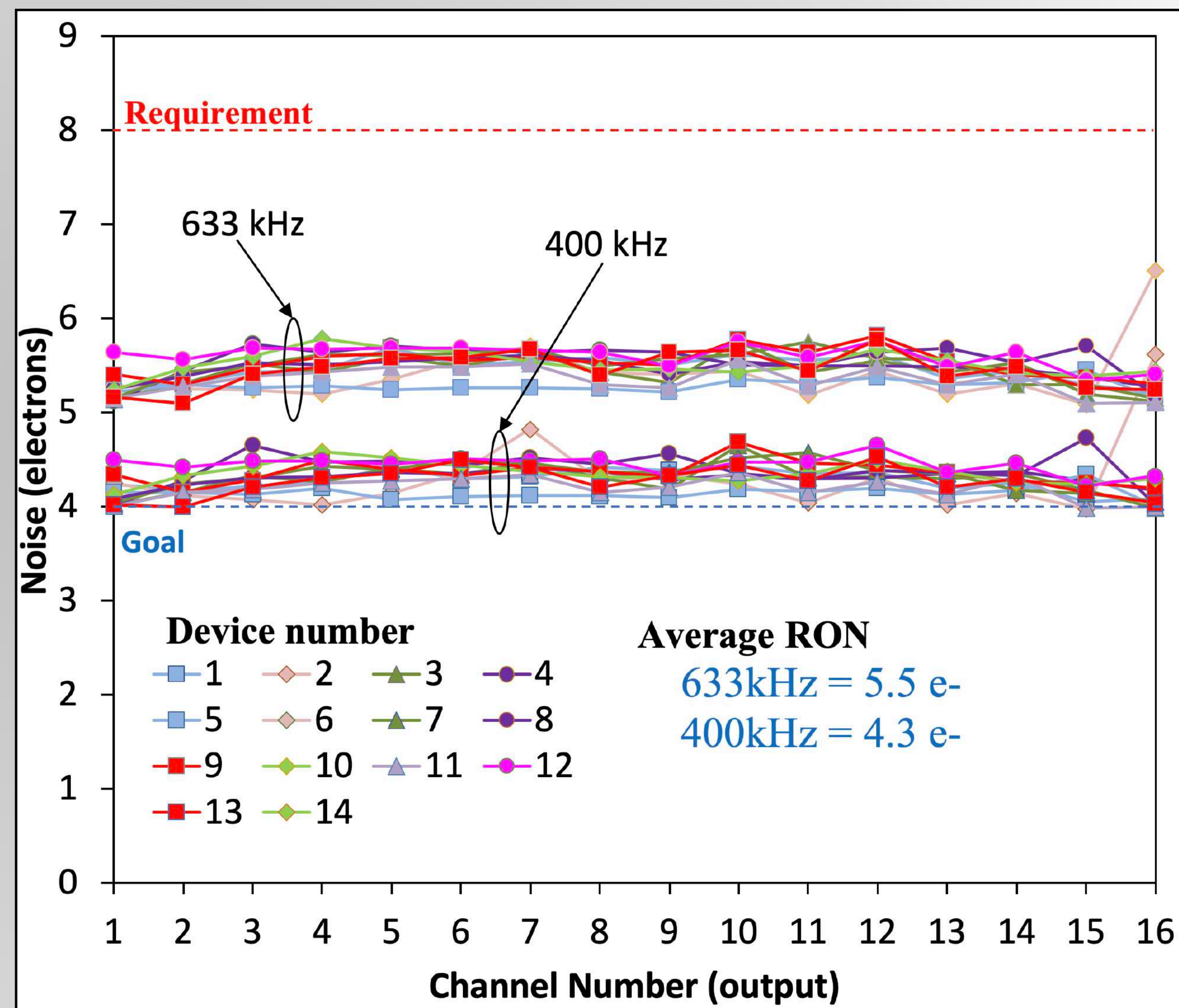


Cryogenic stability



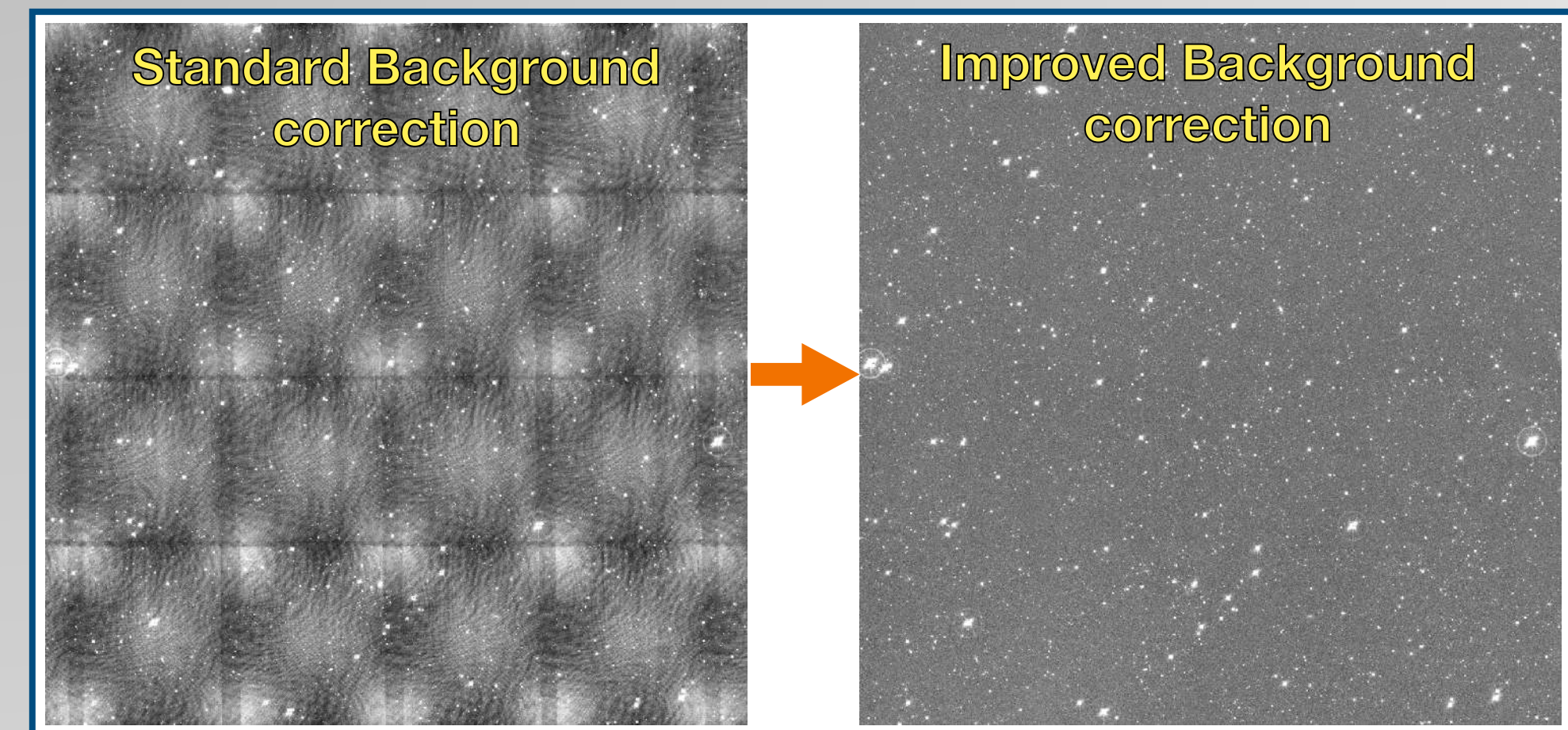
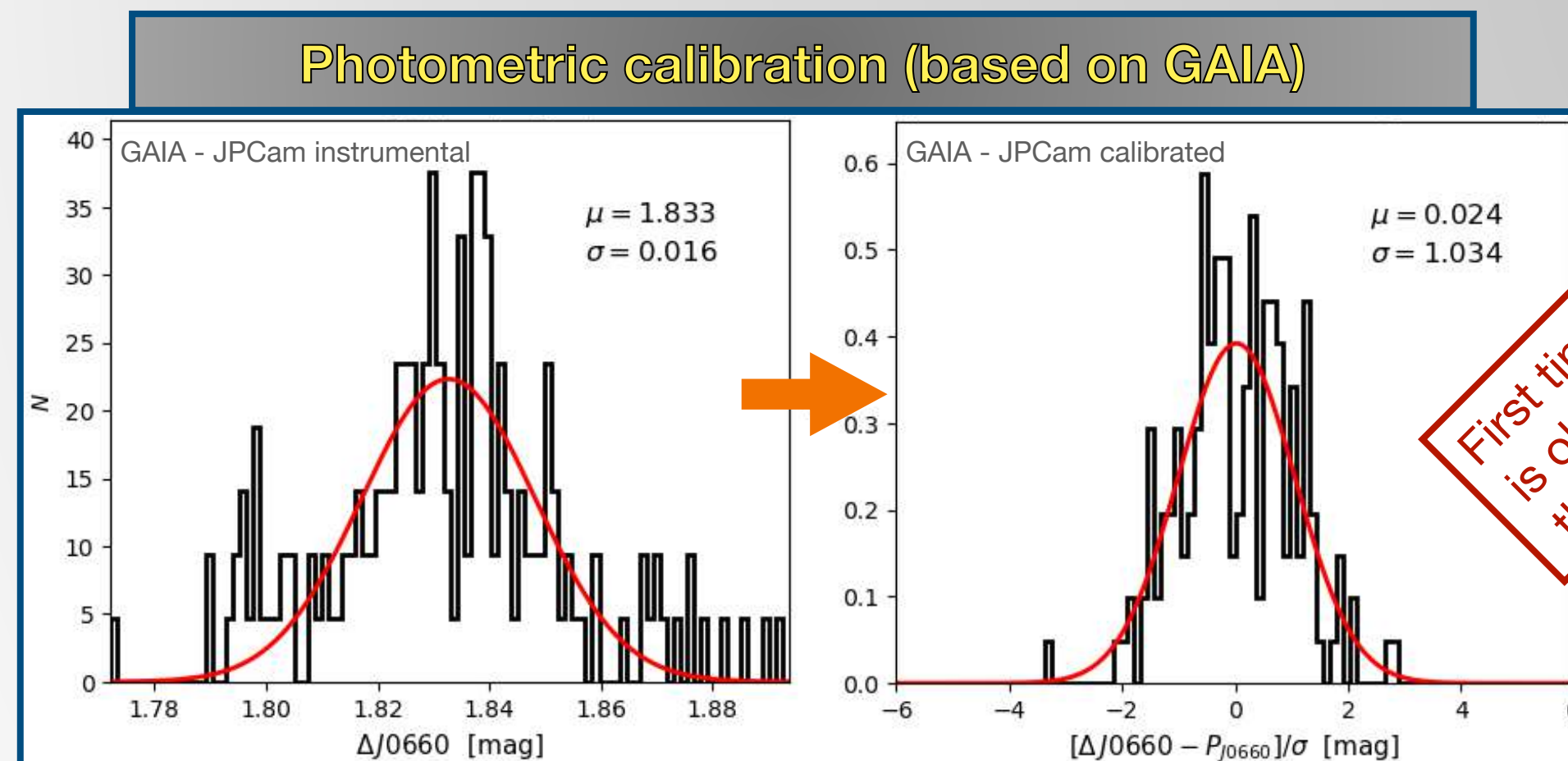
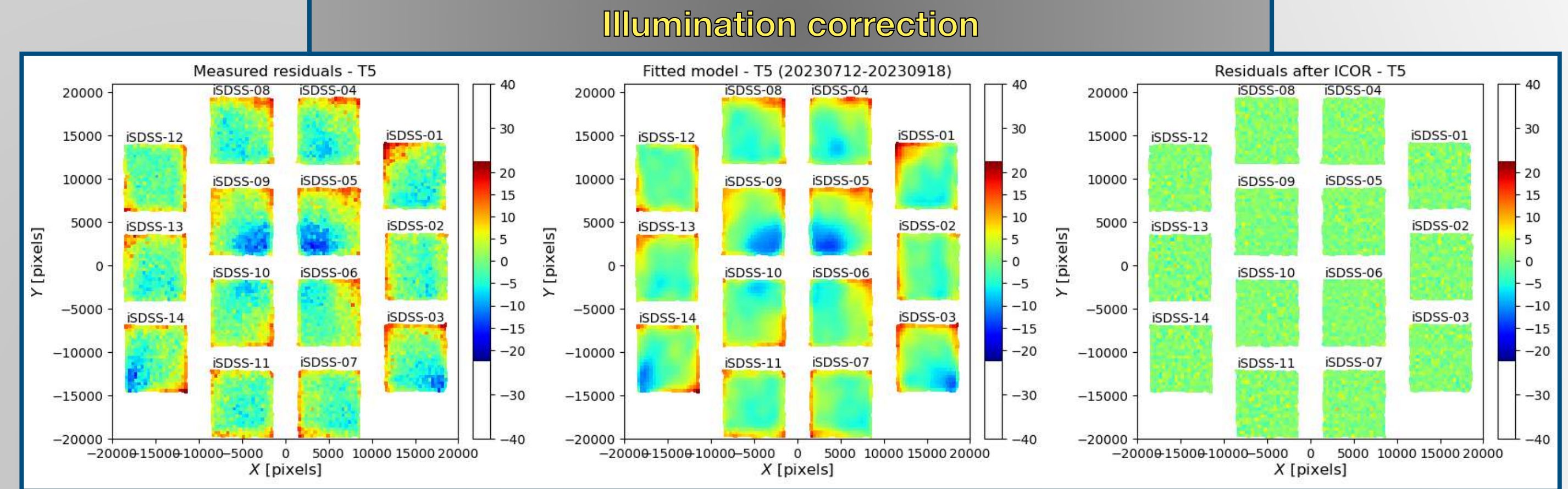
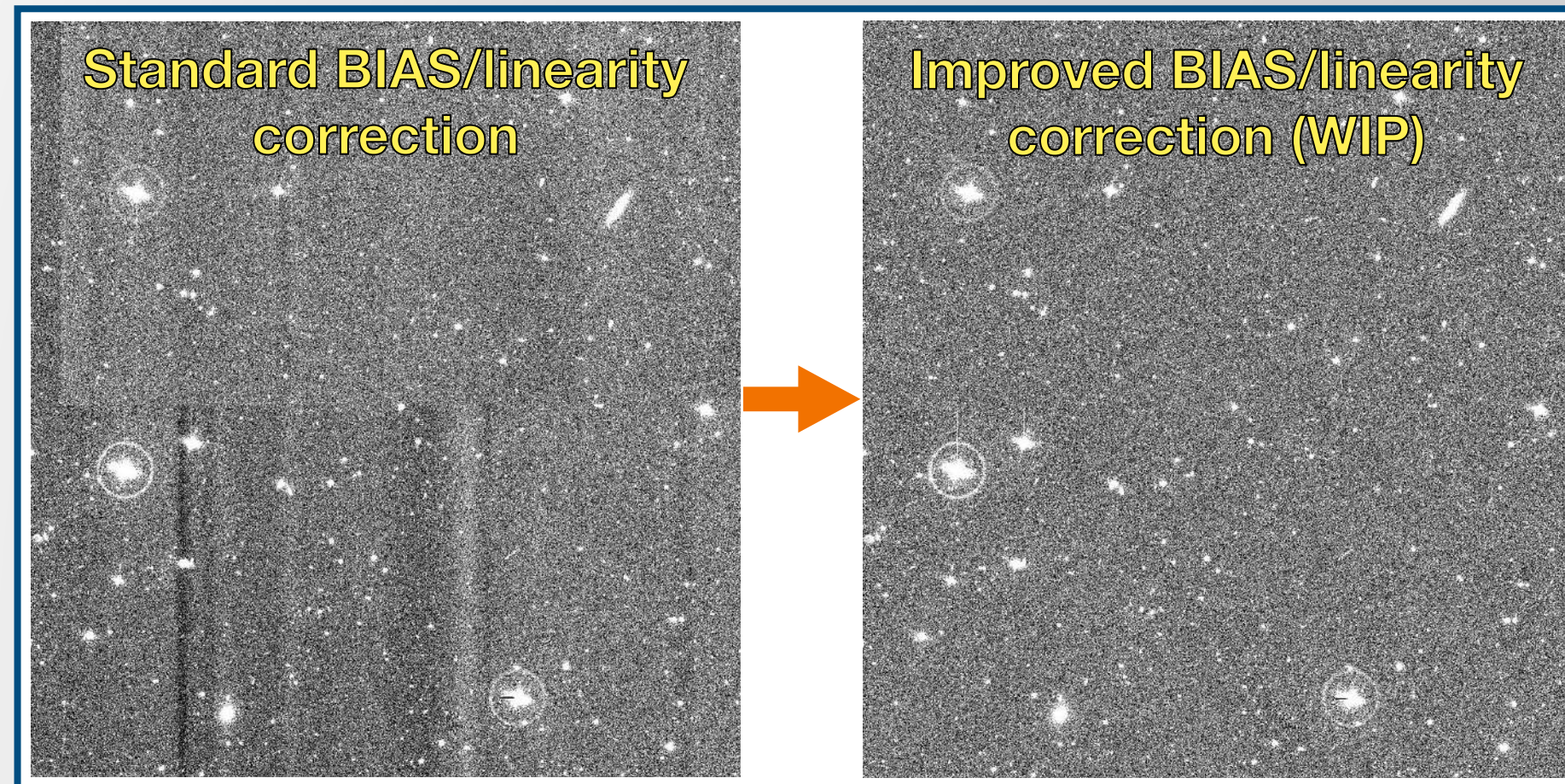
- Temperature set value: $-112,00^{\circ}\text{C}$
- Center-to-border variation: $1,27^{\circ}\text{C}$
- T stability over long periods of time: $1,06^{\circ}\text{C}$ p-t-v.
- Chamber pressure is maintained at $(1,20 \pm 0,10) \cdot 10^{-9}$ Bar

Read-out noise



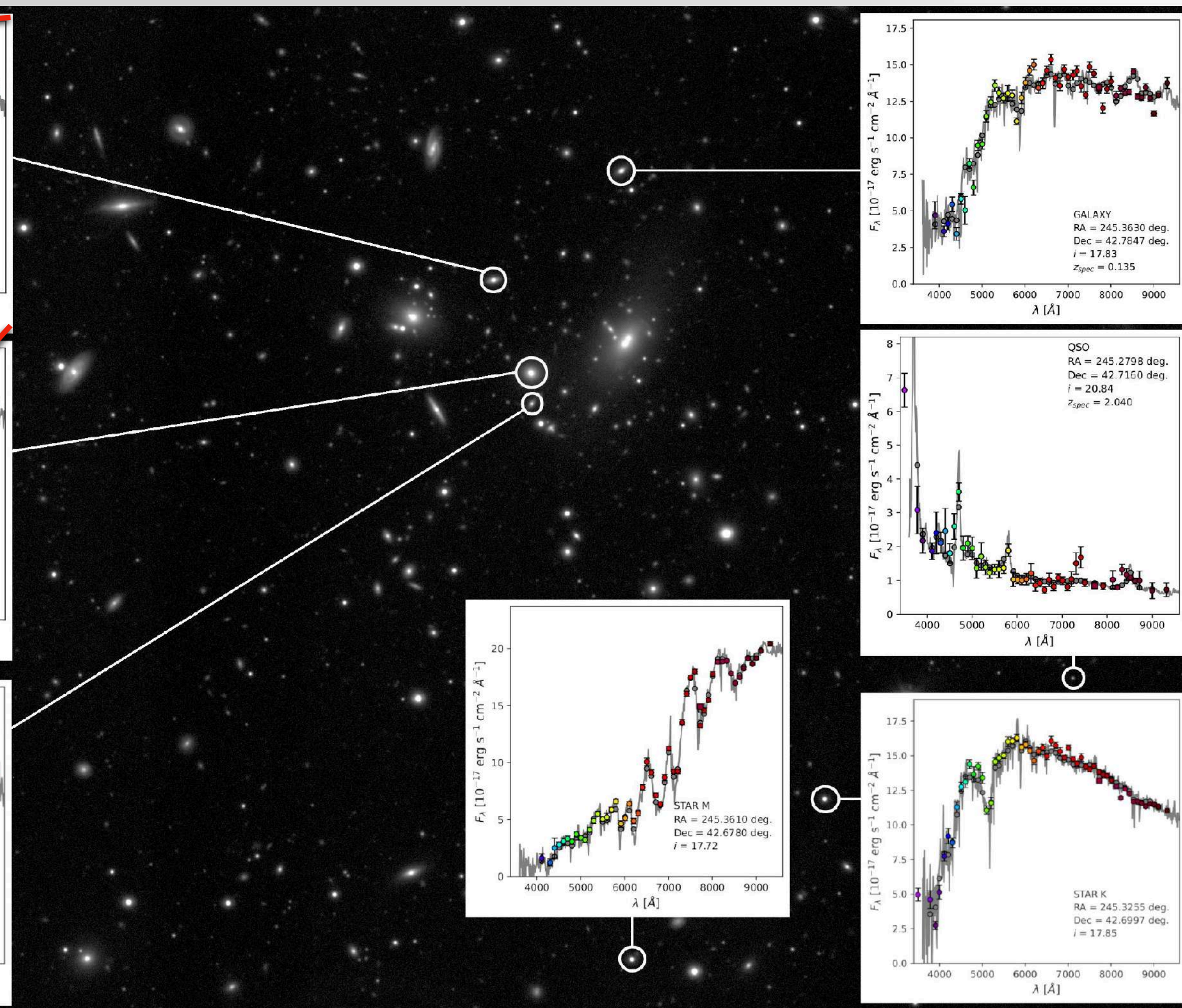
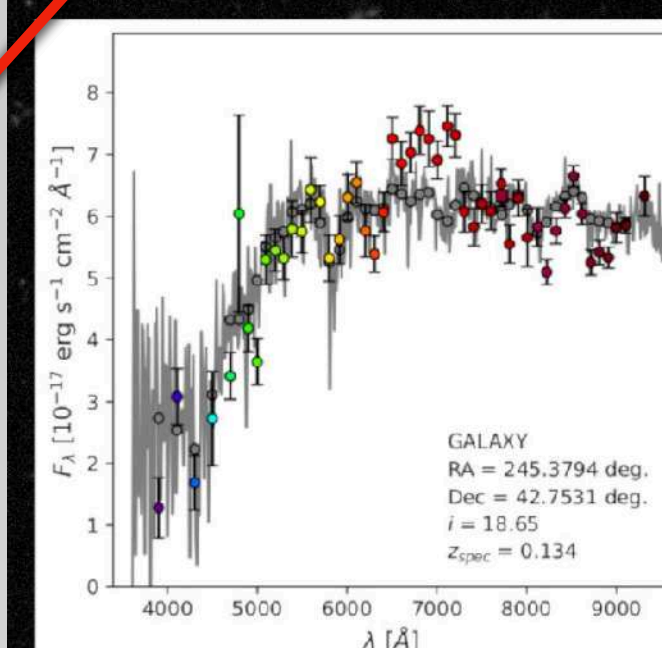
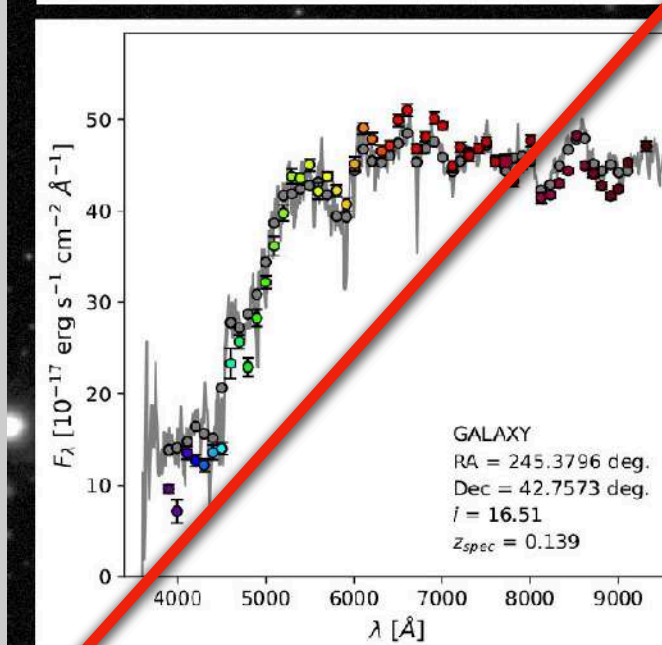
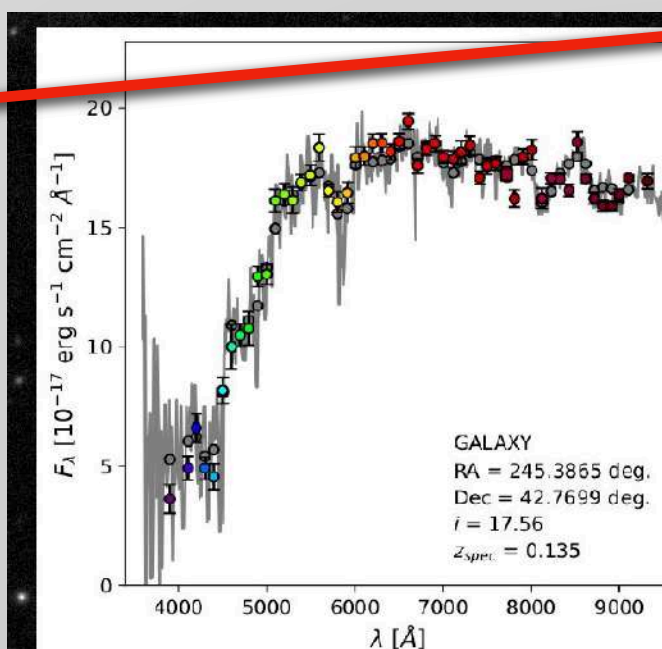
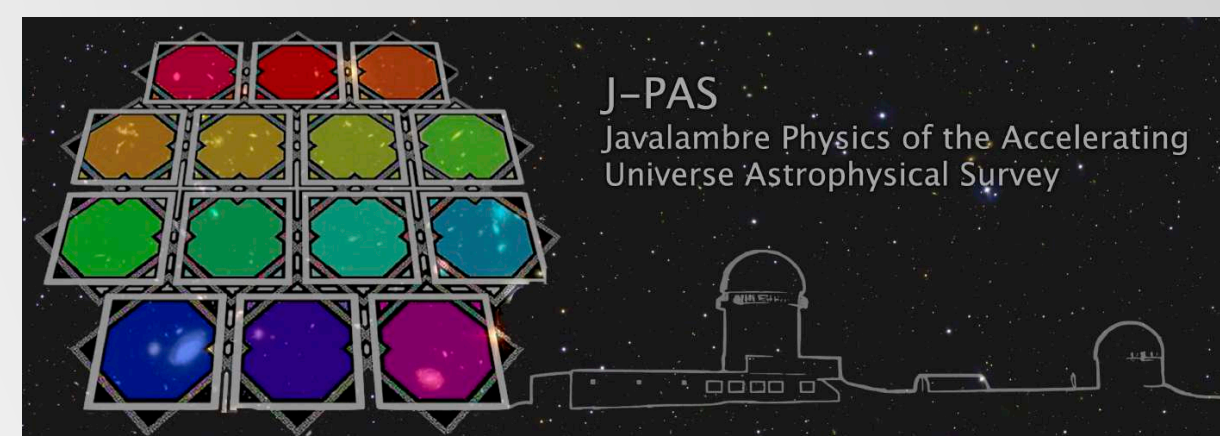
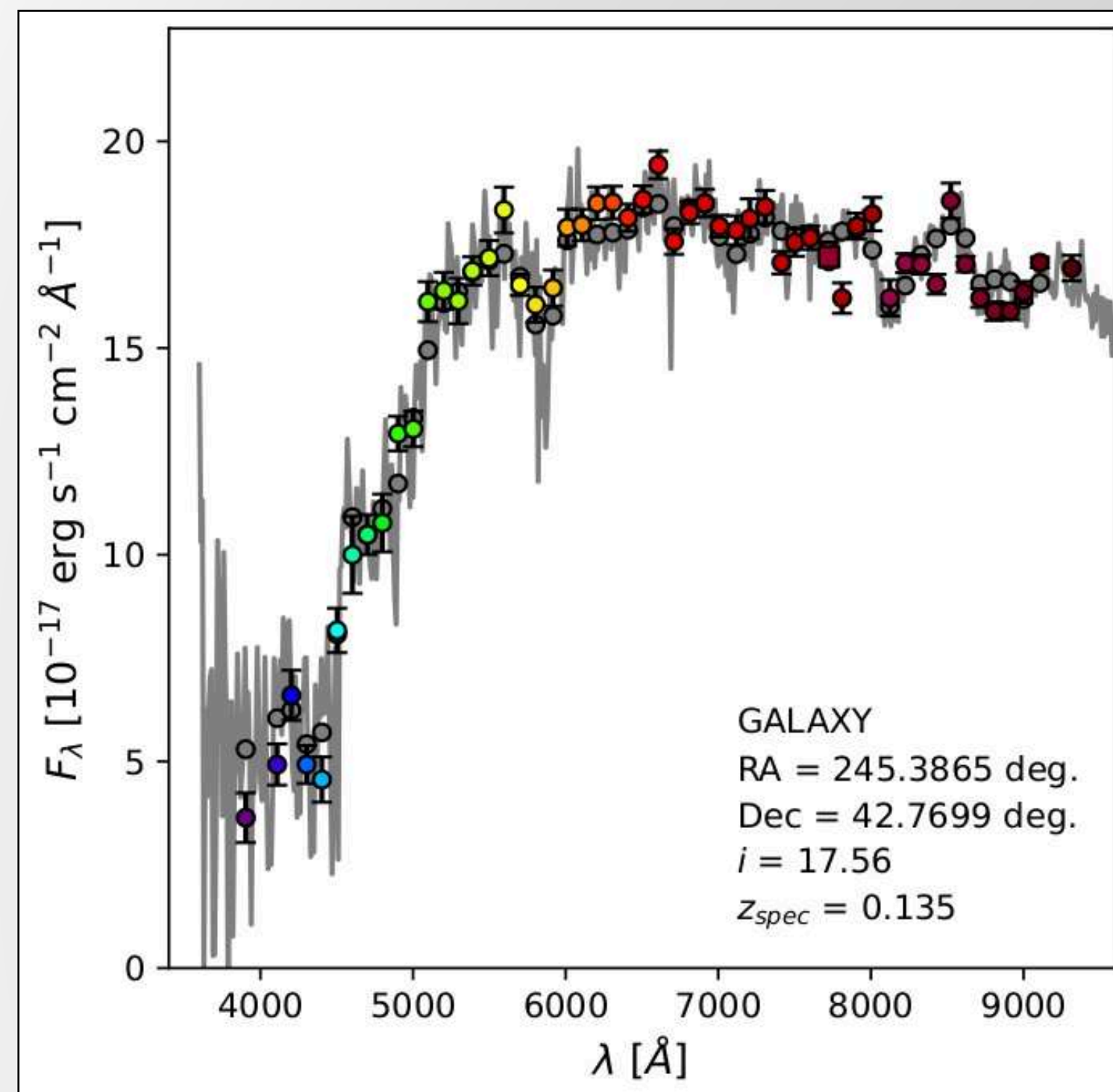
JPCam commissioning results

UPAD is developing special routines to optimize the pipeline to JPCam systematics, maximizing the scientific quality of delivered data.



Photometric error: 2 mmag precision at $\lambda > 400$ nm, 4 mmag for bluer passbands

Photo-z precision: $\sigma = 0.003$ after a quality cut in the probability distribution function of the redshift. This selected 39% of the galaxies (LRGs).



Data releases:

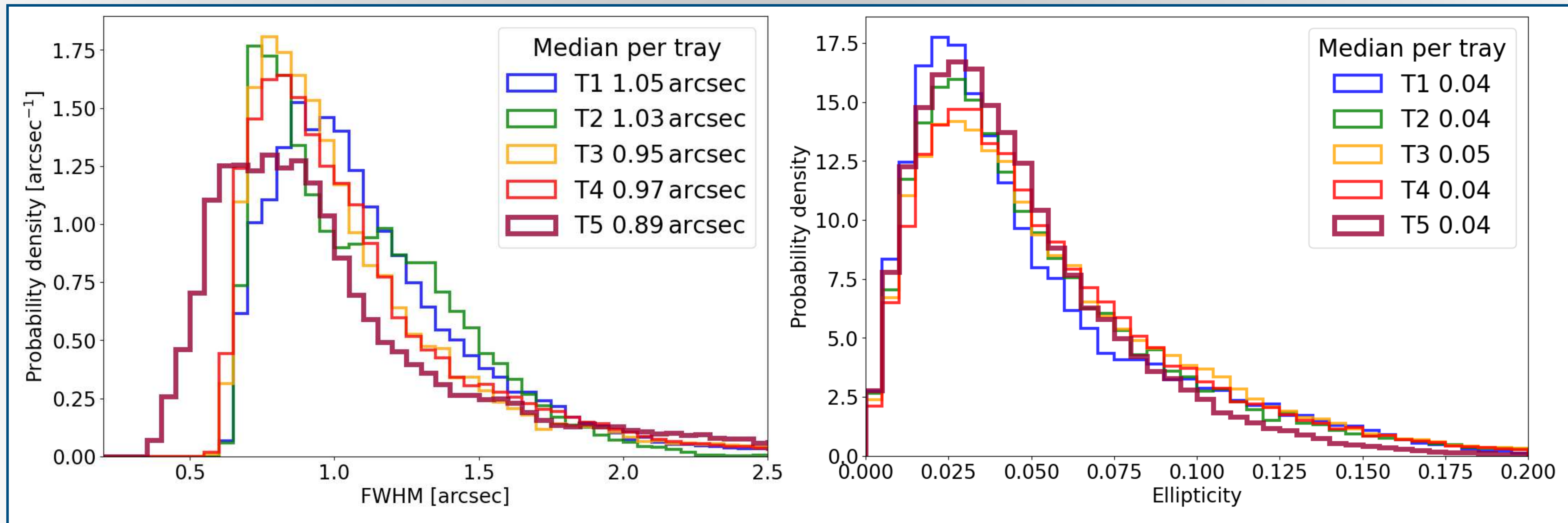
Previous DRs - Pathfinder instrument:

- miniJ-PAS (**public**) — Dec'19
- J-NEP (internal) — Jul'21
- J-NEP (**public**) — Jul'24

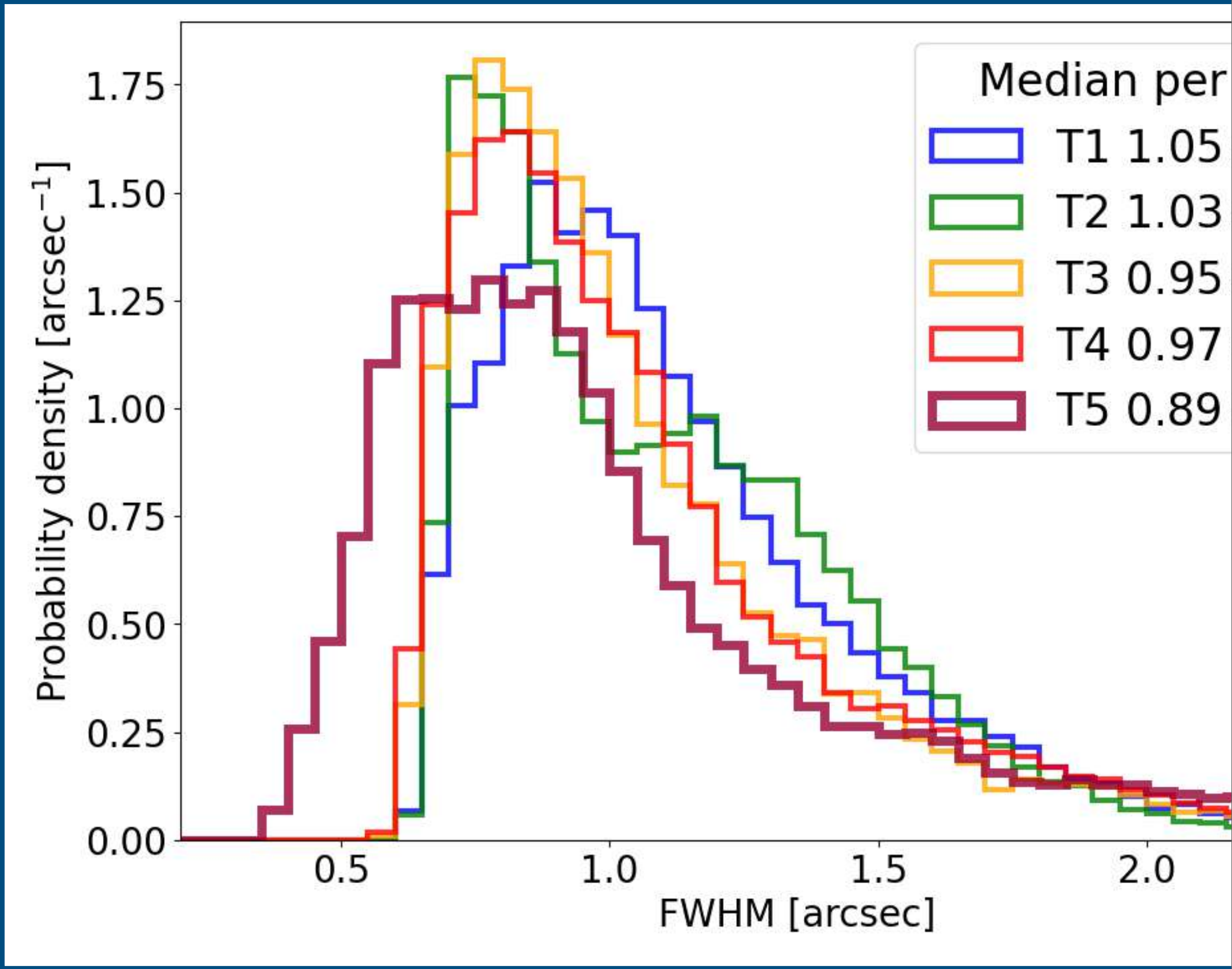
J-PAS DRs - JPCam

- JPSV (internal) — Oct'23
- IDR1 (internal) — Jun'24
- EDR1 (**public**) — Nov'24.

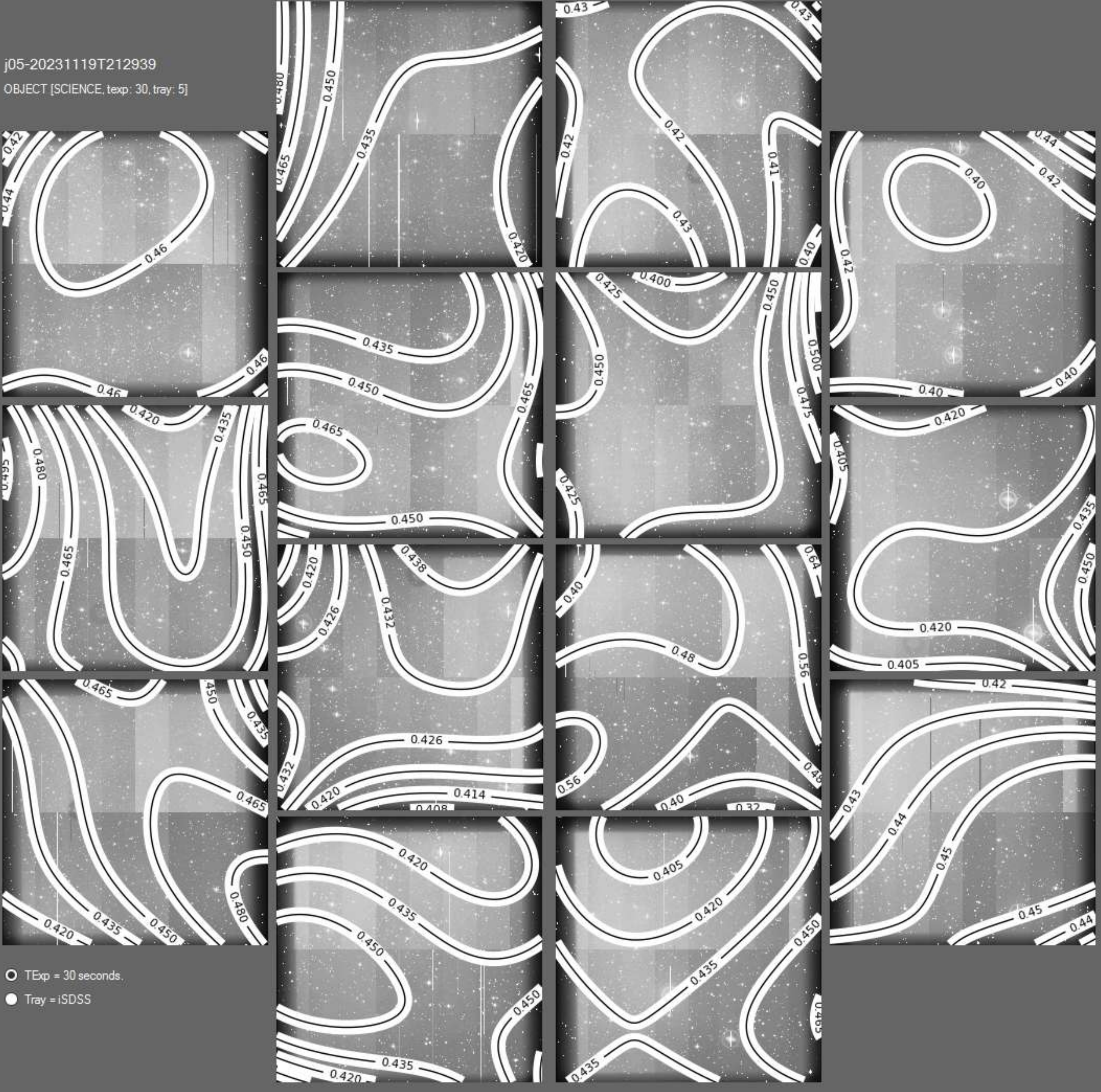
JPCam **image quality**. Statistics of the measured FWHM and ellipticity of the PSF for each of the five J-PAS filter trays. Observations gathered after ~ 1 year of scientific operation.



JPCam **image quality**. Statistics of the five J-PAS filter trays. Observations

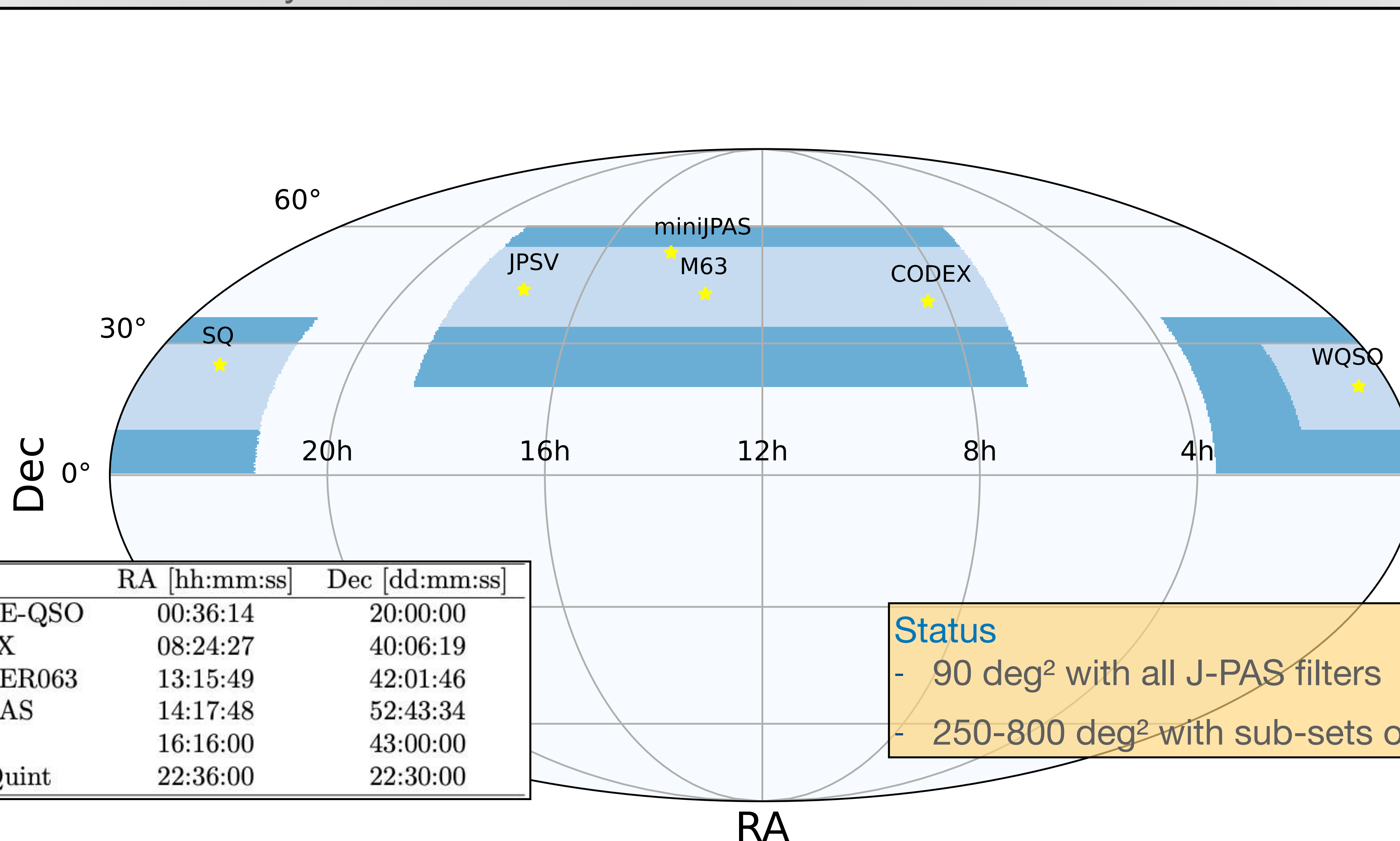


j05-20231119T212939
OBJECT [SCIENCE, texp: 30, tray: 5]

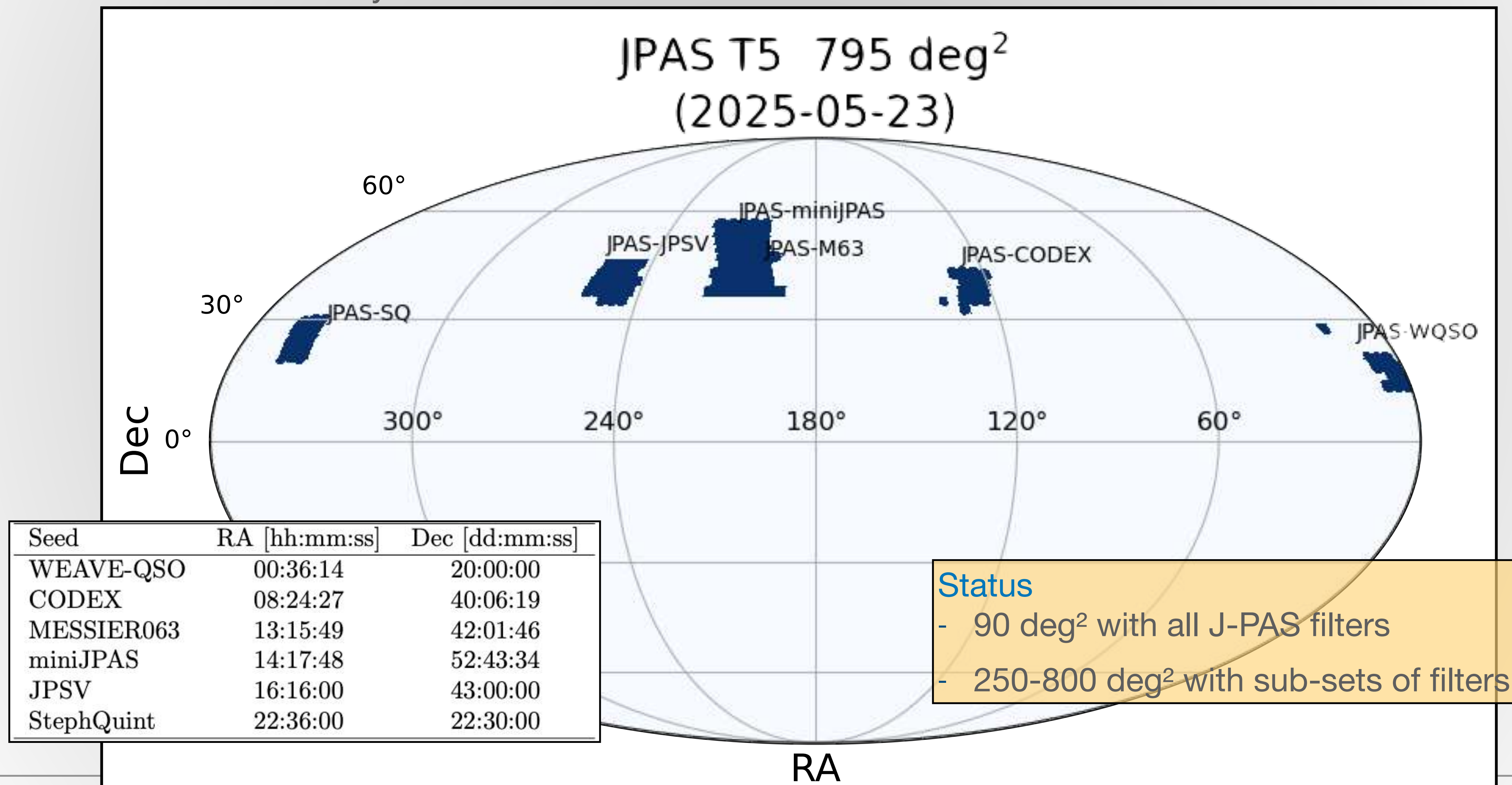


○ TExp = 30 seconds.
● Tray = iSDSS

J-PAS Observations started centered in 6 “seeds” chosen based on specific scientific interests and visibility.

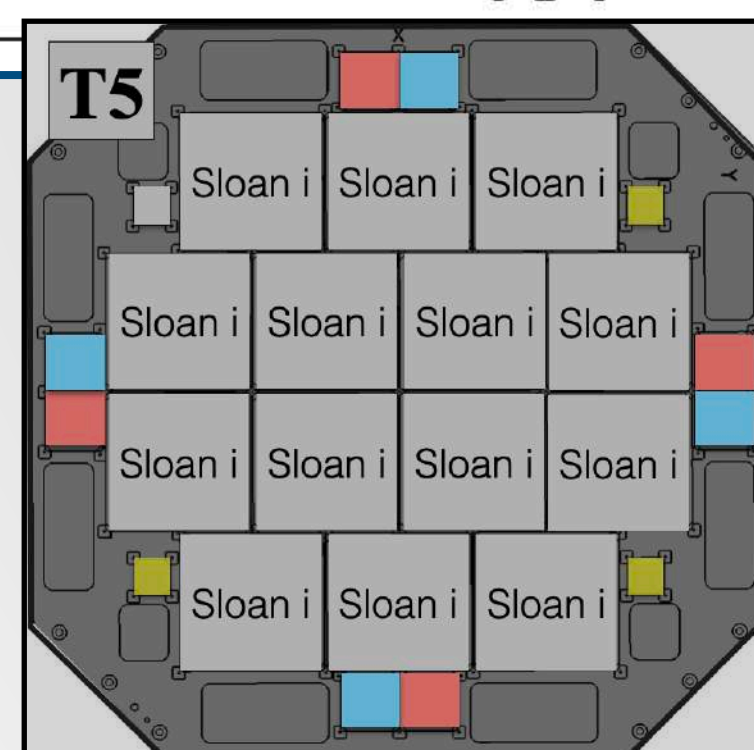
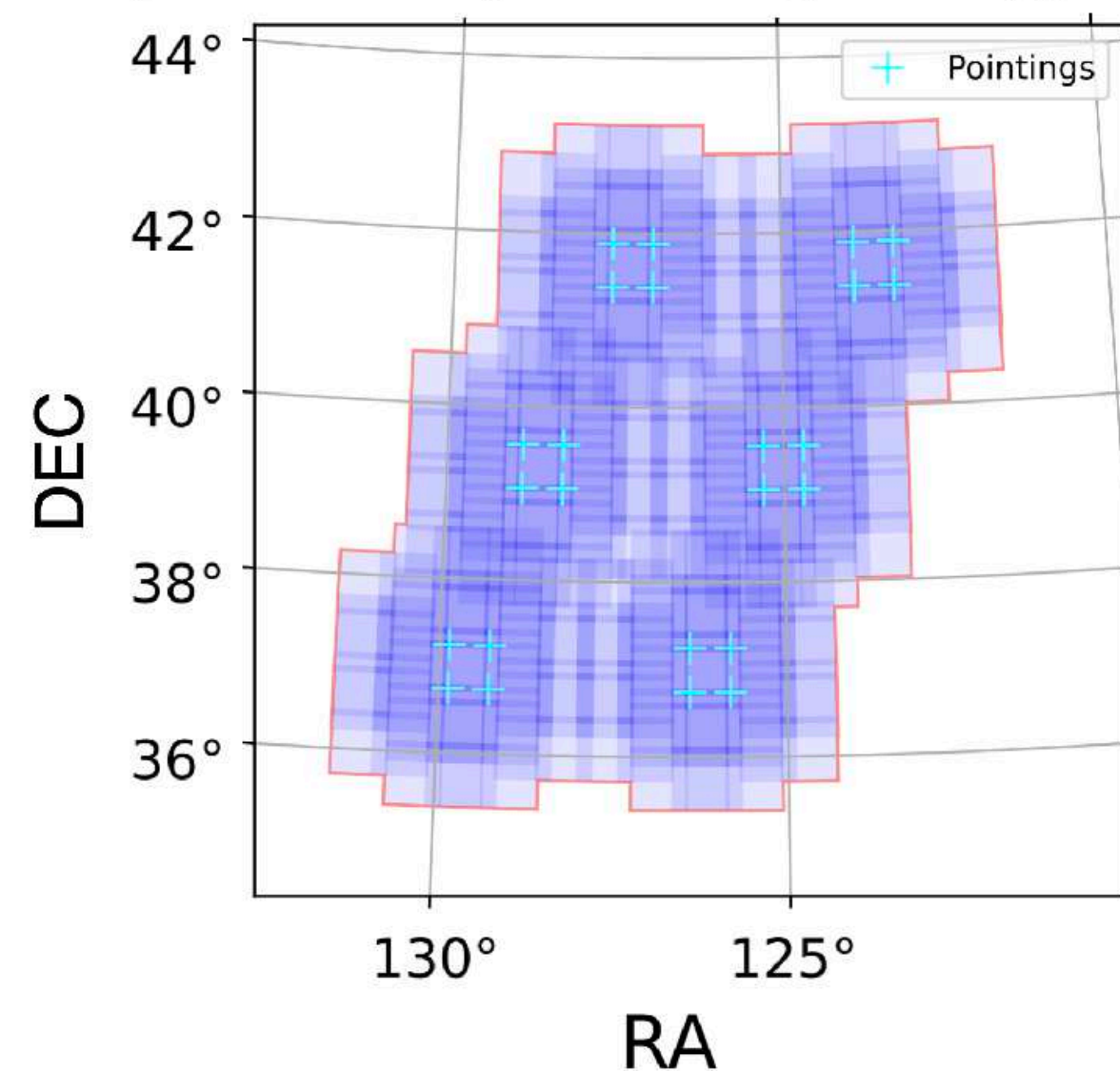


J-PAS Observations started centered in 6 “seeds” chosen based on specific scientific interests and visibility.

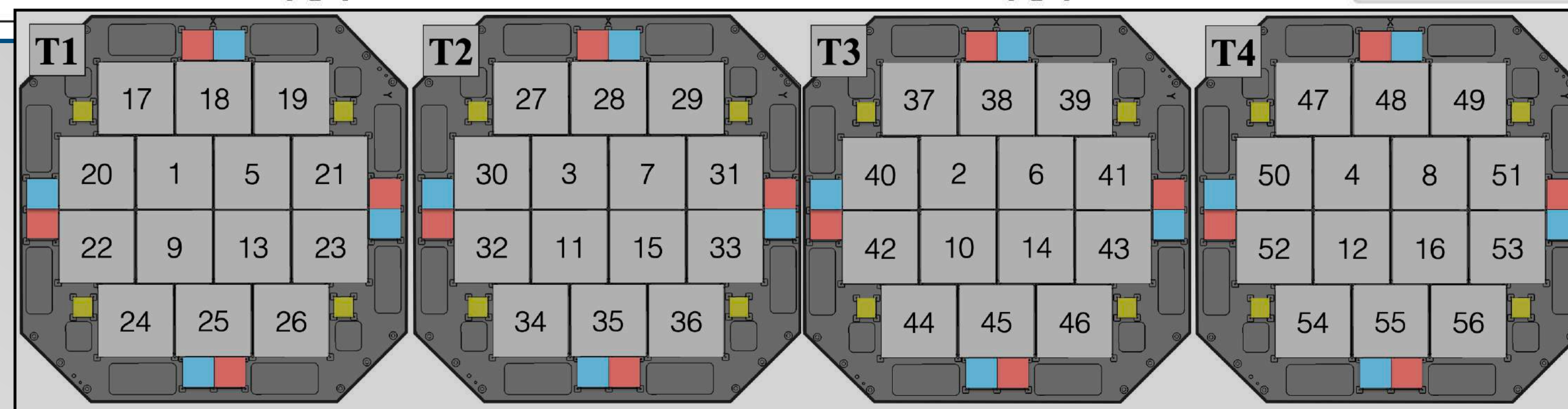
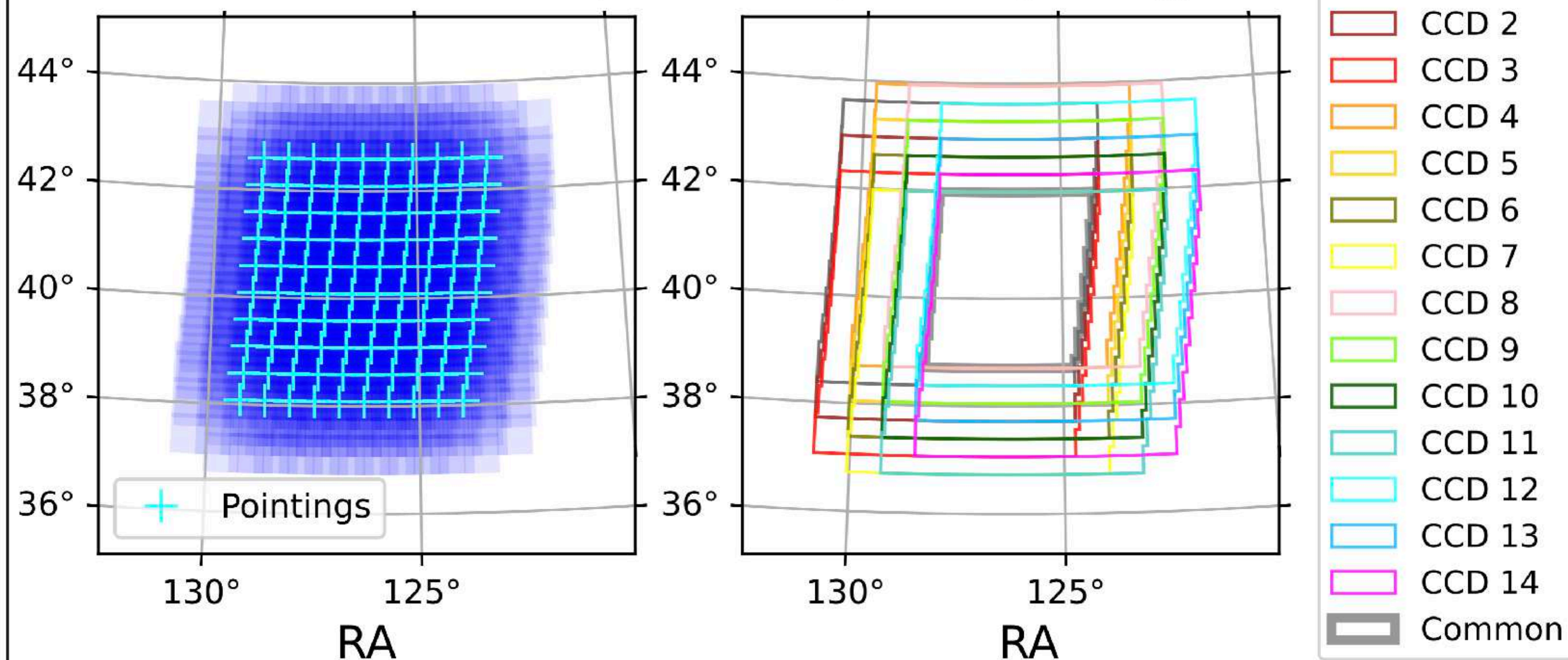


J-PAS Observational strategy is optimized for the different filter trays.

T5 (broad band) observing strategy

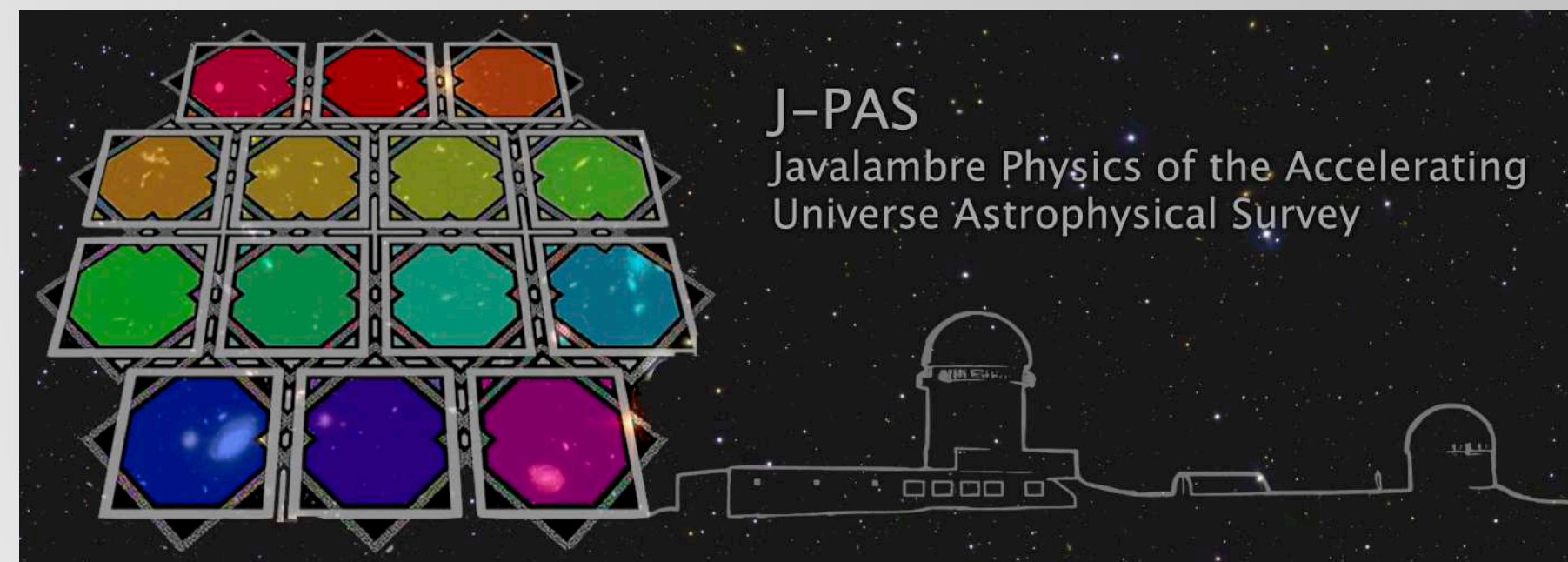


T1 to T4 (narrow band) observing strategy



Summary

- The **commissioning of JPCam**, the 1.2Gpix, large FoV J-PAS Camera, has been successfully completed at the JST250 telescope.
- During its first on-sky operation a number of improvements have been implemented both at instrument level (opto-mechanics and light baffling, control system, electronics, IQ monitoring and control...) and data reduction level.
- Main **J-PAS scientific requirements have been achieved** and the main survey has started.
- In November'24 a **J-PAS first data release** became public, making J-PAS data available to the whole scientific community.
- In 2025, the tandem JST250-JPCam has been offered to the scientific community through a competitive, **open call for proposals**.



CCD format	14 × 9216 × 9232 pix, 10 μm pix ⁻¹ 1.2 Gpix camera
Pixel scale	0.2265'' pix ⁻¹
Unvignetted FoV	3.4deg ² – (14×) 0.48deg × 0.51deg
Read out time (633kHz)	10.9 s (full frame) – 6.1 s (2x2 binning)
Read out noise (633kHz)	5.5 e ⁻ (RMS)
Read out time (400kHz)	16.4 s (full frame) – 8.9 s (2x2 binning)
Read out noise (400kHz)	4.3 e ⁻ (RMS)
Gain	2.274 e ⁻ ADU ⁻¹
Minimum exposure time	0.1 s
Exposure homogeneity	1 ms
Full well	> 125 000 e ⁻
Dark current	0.001 e ⁻ pix ⁻¹ s ⁻¹