

RPC readout optimization: miniTRASGO as a path to a single-plane multiplexed timing system

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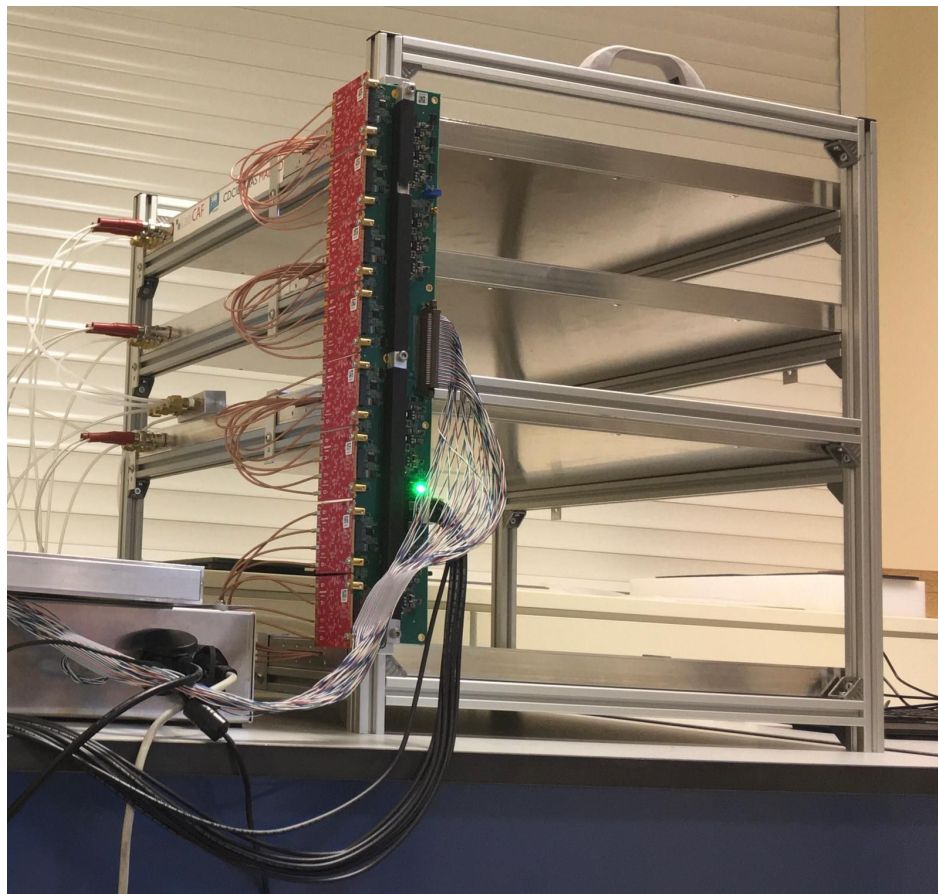
going for the single plane

Introduction

Resistive Plate Chambers in miniTRASGO

miniTRASGO: RPC-based muon telescope

- 4 parallel 30x30 cm² **Multigap glass Resistive Plate Chambers (MRPCs)**
 - **~400 ps** timing
 - **~3 cm** spatial resolution
- **Gas:** R134a, flux of <1 kg/month



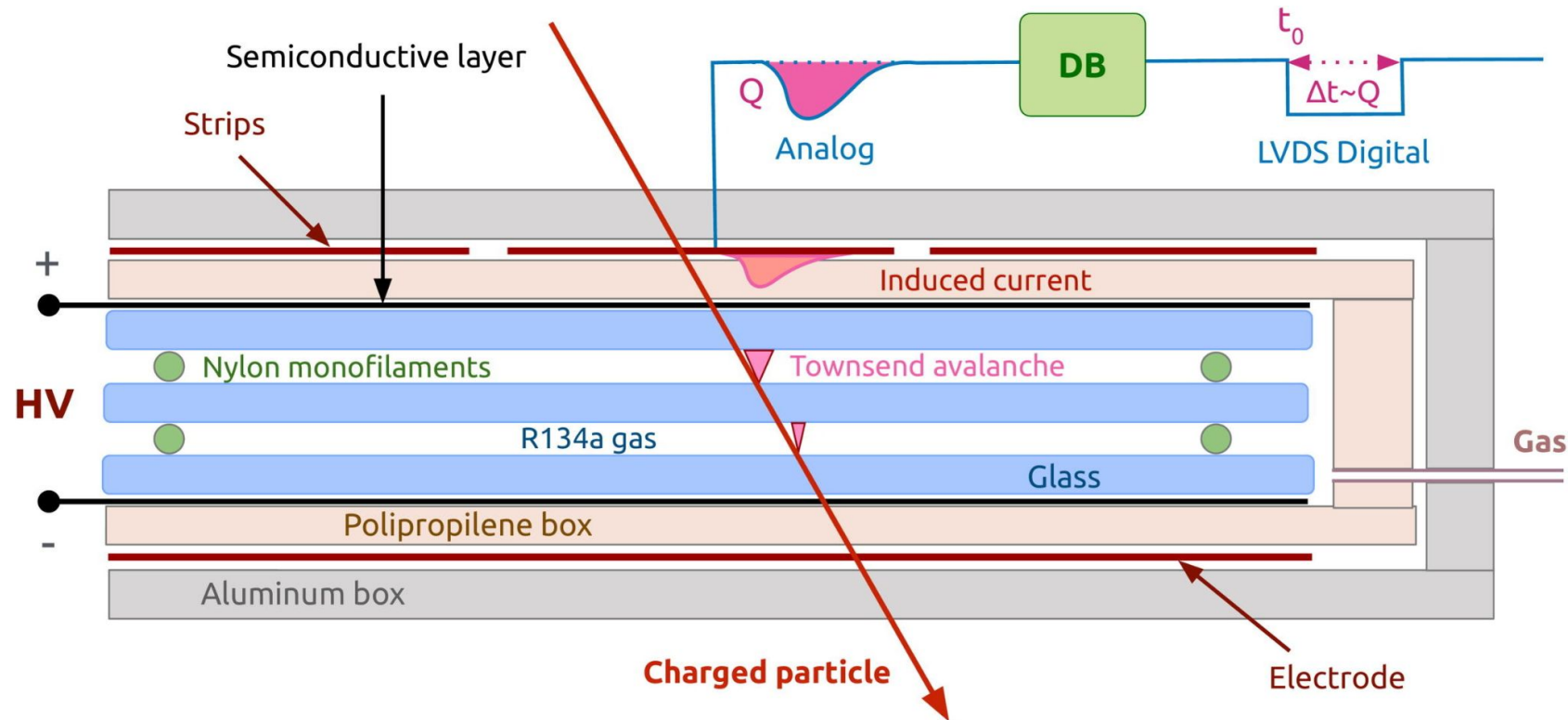
C. Soneira-Landín et al., NIM A, 2025, <https://doi.org/10.1016/j.nima.2025.170511>

C. Soneira-Landín et al., Adv. Space Res., 2025, <https://doi.org/10.1016/j.asr.2025.07.096>

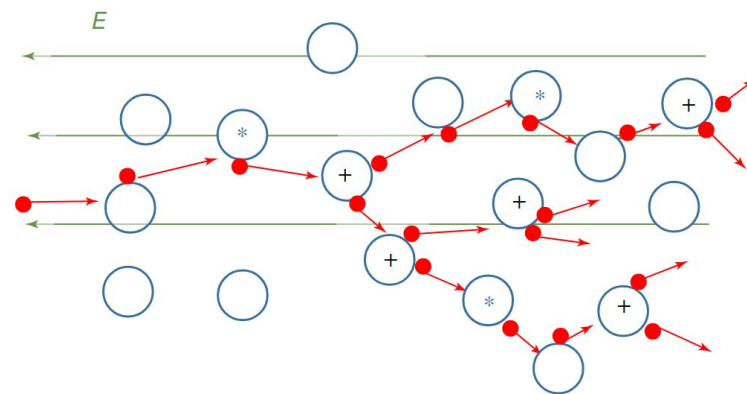
C. Soneira-Landín et al., PoS(ICRC2025)1368

Resistive Plate Chamber layout

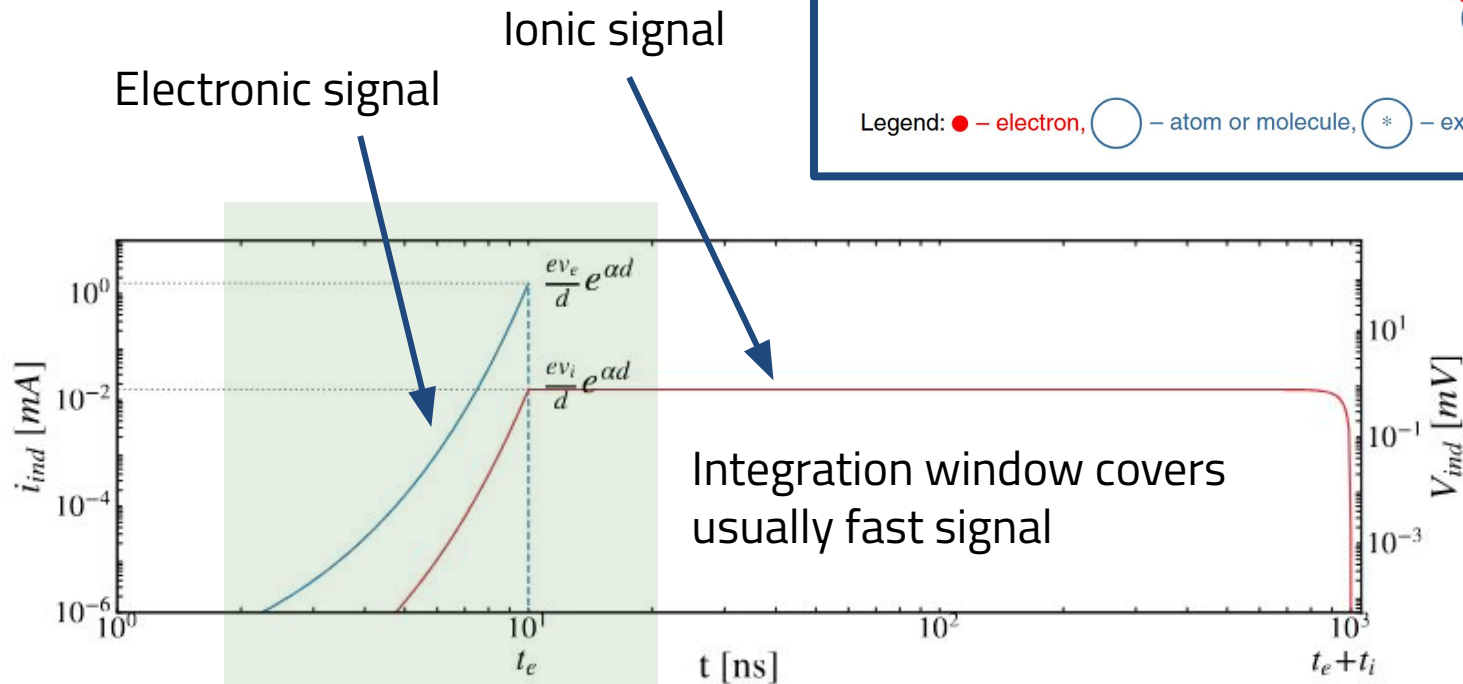
- Every **Multigap glass RPC** has:
 - **2 gas gaps** → 1 mm/gap
 - **3 glass electrodes** → 2 mm/glass
- **Readout:** *wide strips* (6-9 cm)



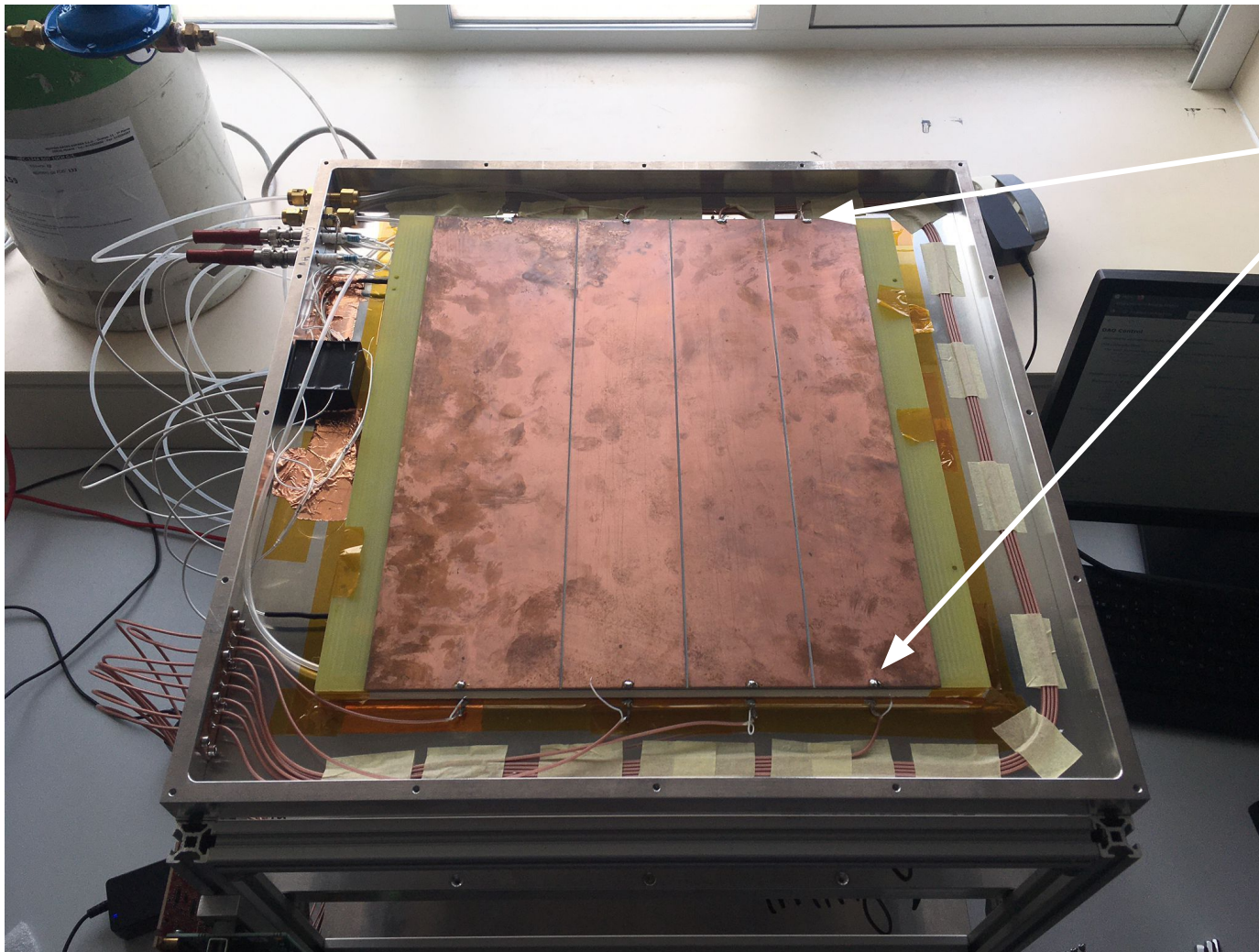
Signal induction in RPCs



Legend: ● – electron, ○ – atom or molecule, (*) – excited atom or molecule, (+) – ion



Townsend avalanche
(from Abbrescia et al., 2018)



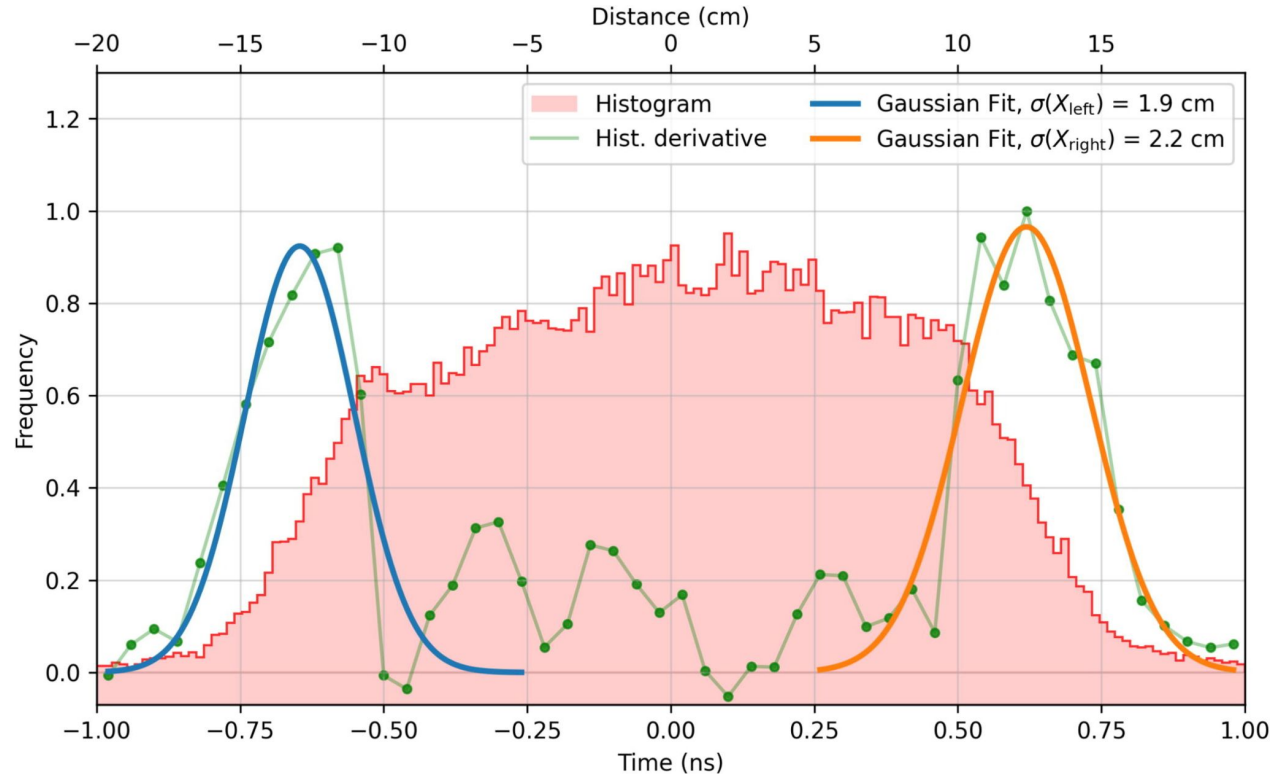
Asymmetric
parallel **Cu**
strips read in
Front and Back
(32 channels)



Time and
charge
obtained

Estimated position uncertainty

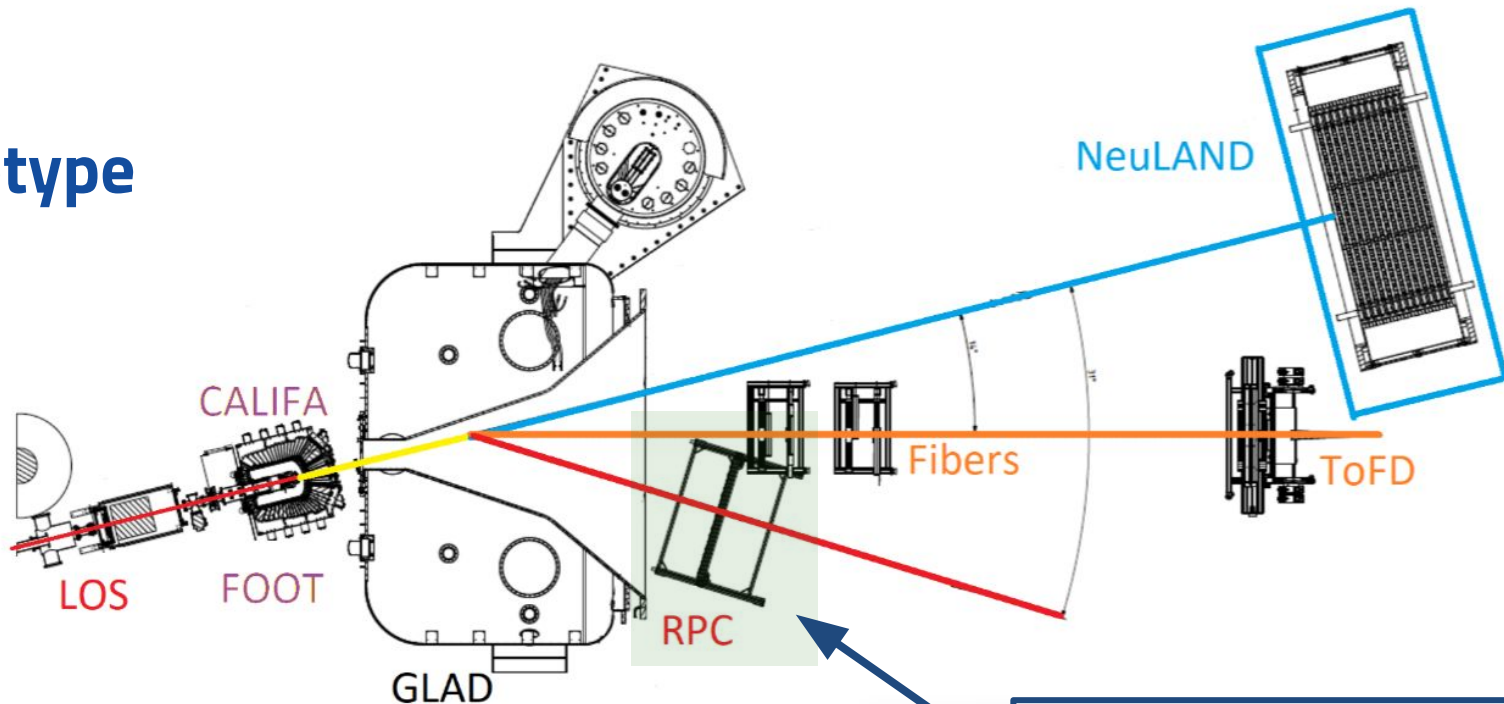
Using the **time difference** in both sides of the **strip** the position of the event can be determined



More advanced versions

R3B prototype

R3B prototype



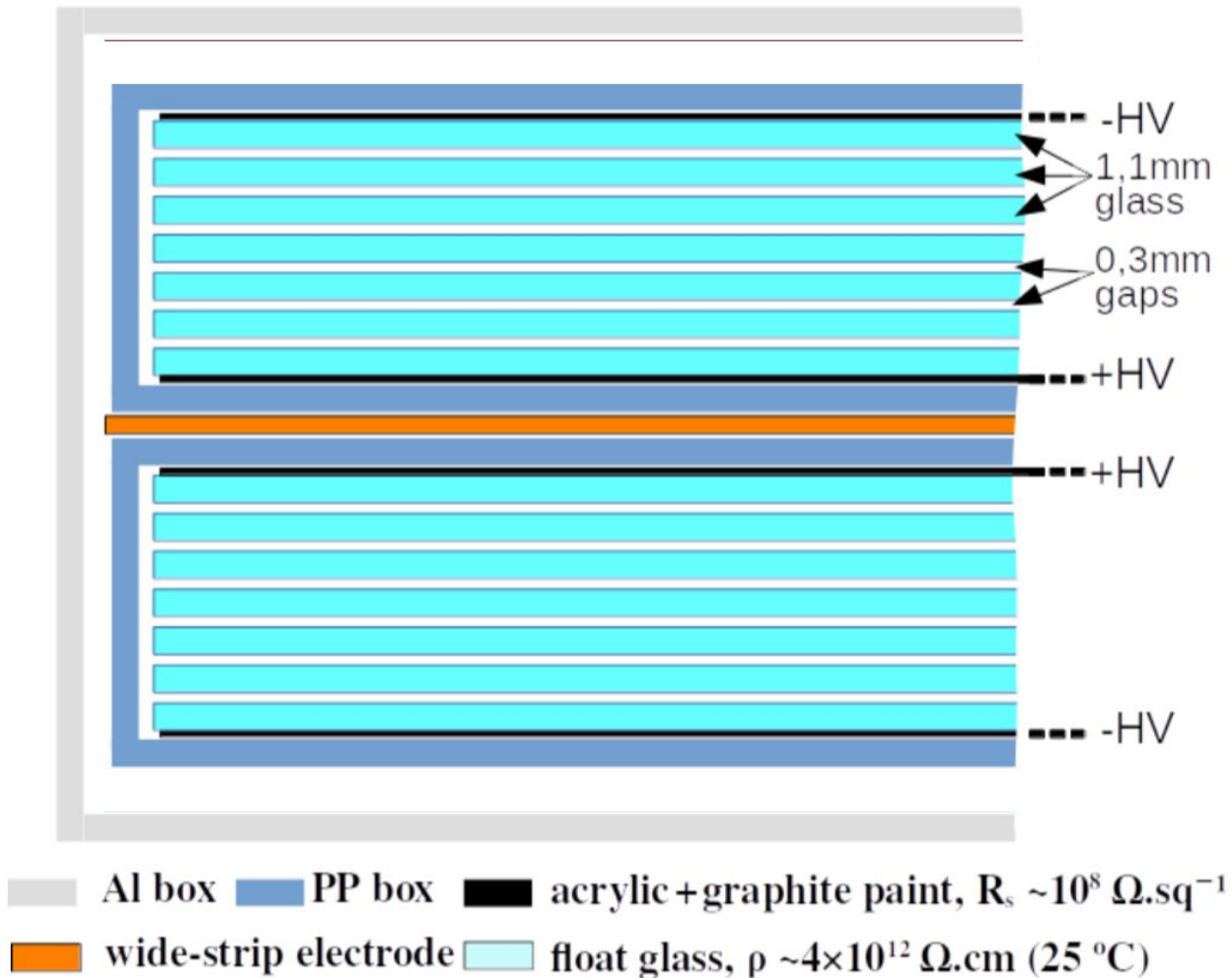
*The next step in the detector development
is being used in R3B experiment at FAIR*

2 m²
6 gaps x 2 chambers
3 cm pitch strips

R3B prototype layout

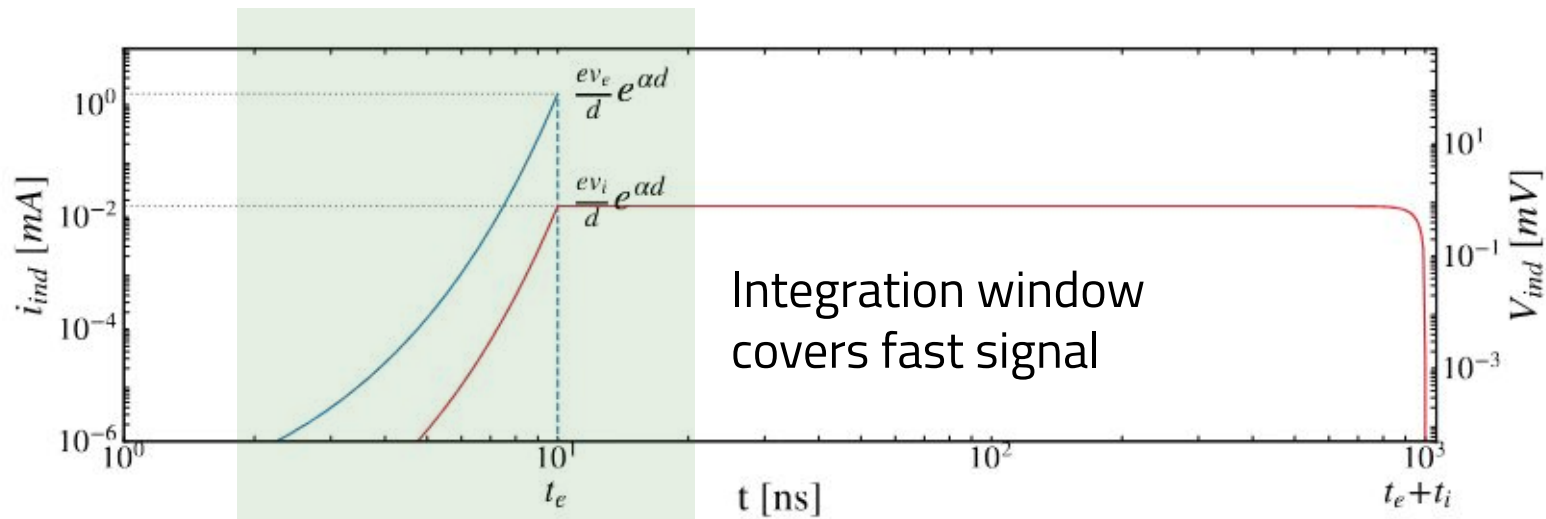
2 multigap
RPCs **stacked**

Thick readout strips
are in the **middle**



Still measuring the fast signal

→ **Faint** (less efficient) and **ballistic deficit** (losing information)

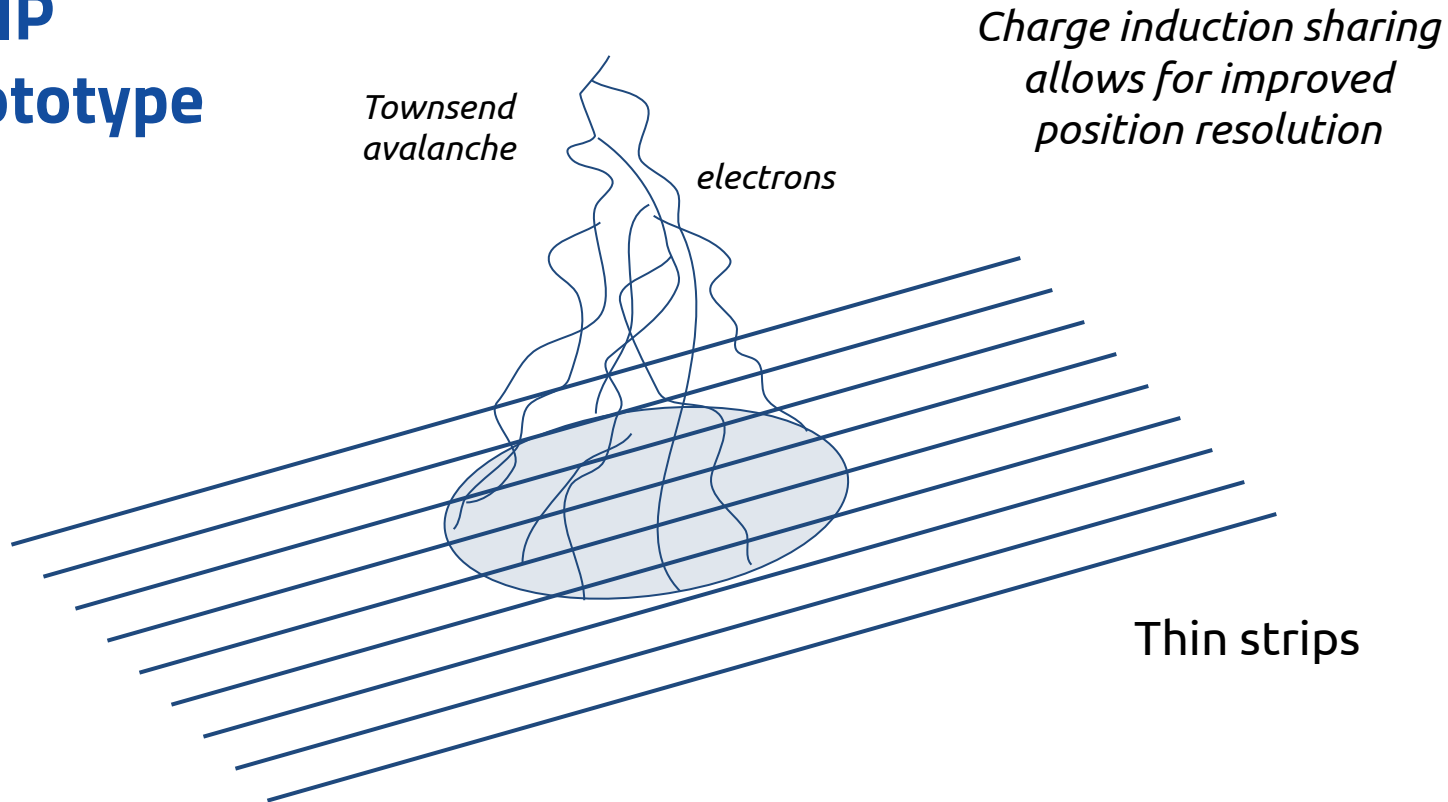


Still, using a double stack allows for **98% efficiency for MIPs**

More advanced versions

Multiplexing: SHIP prototype

SHIP prototype

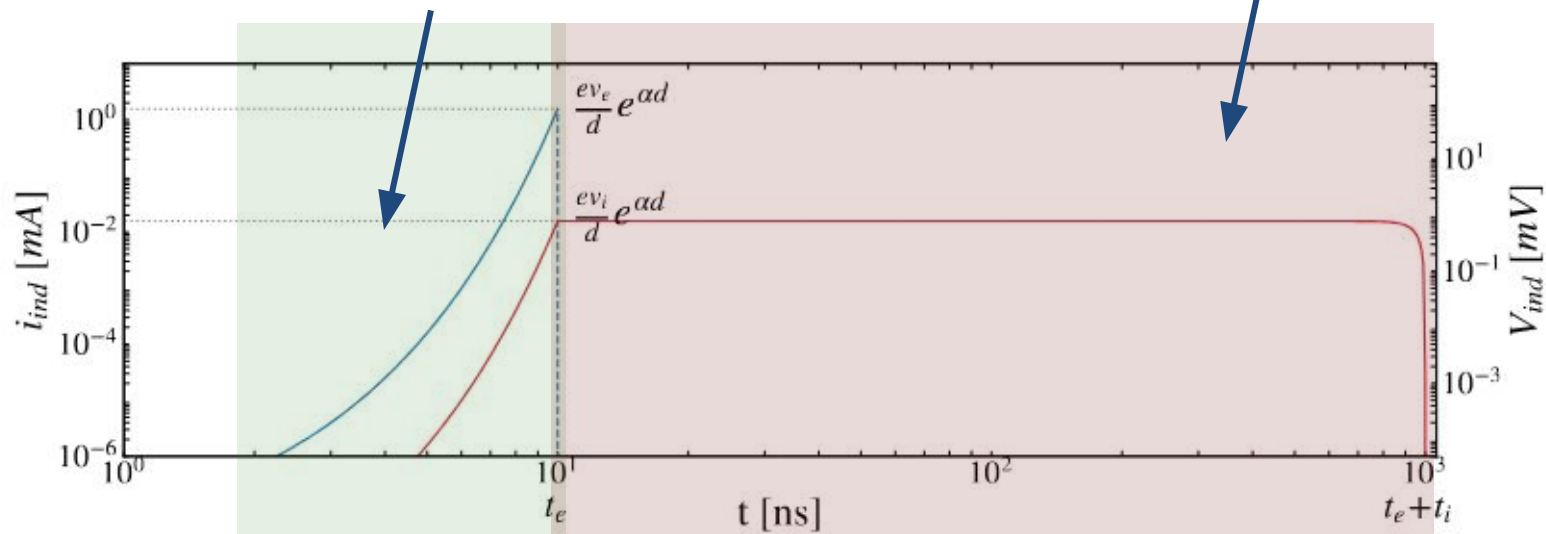


Let's take profit on
the long signal

But combining readouts

Thin strips for slow signal → position

Thick strips for fast signal → timing



→ **Robust** (more efficient) and **total charge collection**

Readout electrodes

Pitch 2.54 mm
Interstrip 1 mm



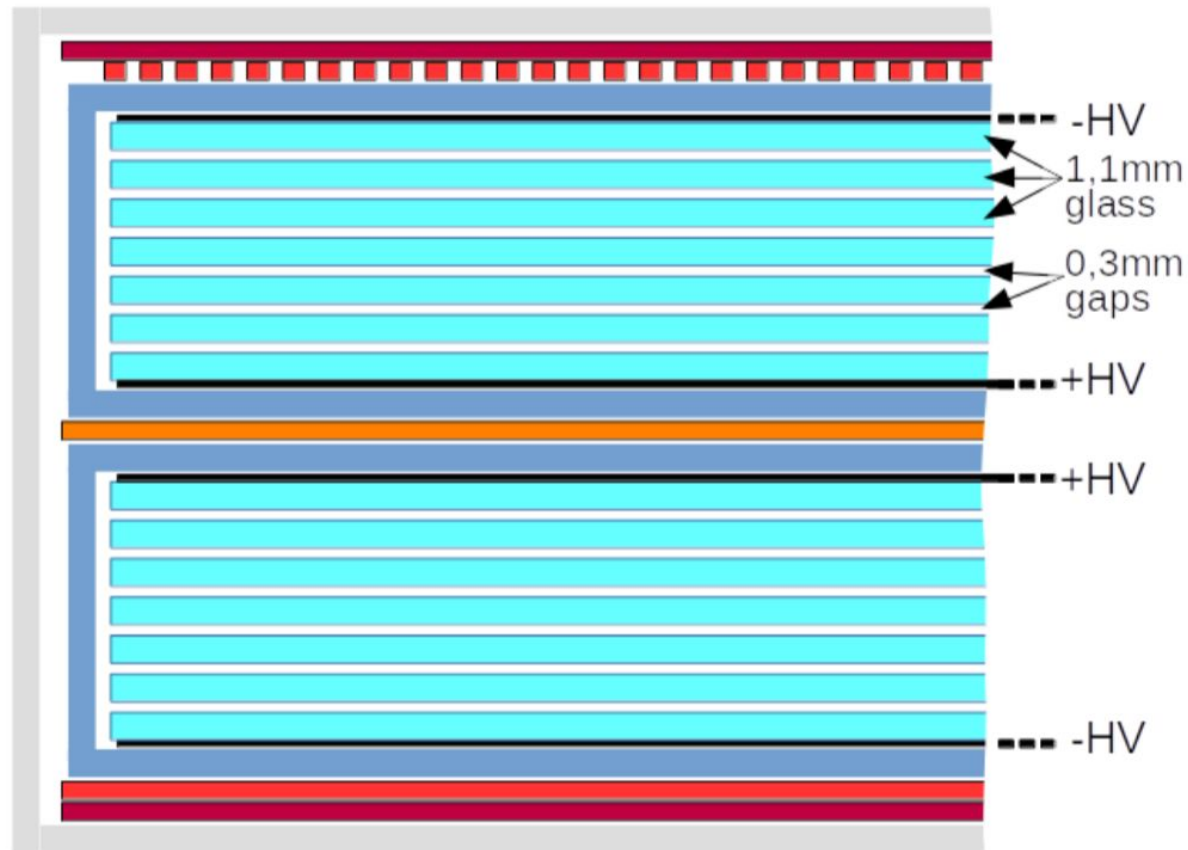
(a) Thin-strip PCB

Pitch 61.0 mm
Interstrip 2 mm



(b) Wide-strip PCB

SHIP prototype

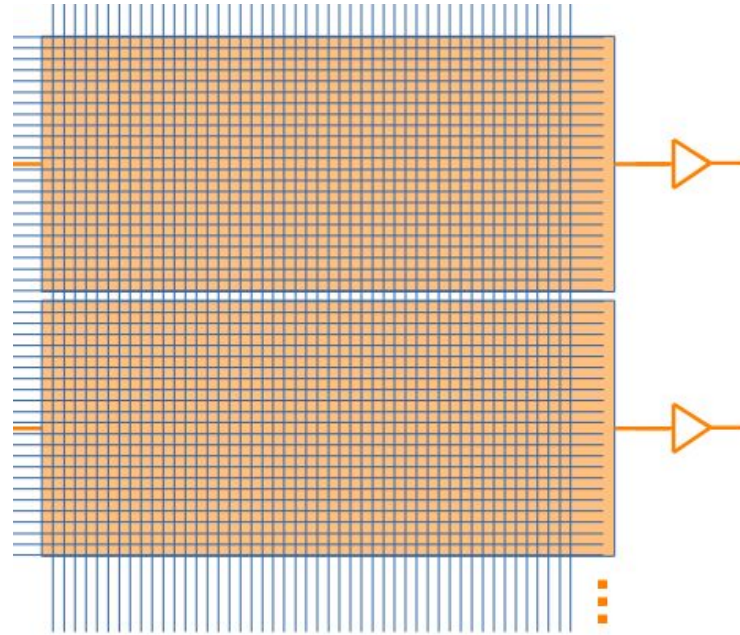


- Al box
 PP box
 acrylic+graphite paint, $R_s \sim 10^8 \Omega \cdot \text{sq}^{-1}$
- thin-strip electrode
 float glass, $\rho \sim 4 \times 10^{12} \Omega \cdot \text{cm}$ (25 °C)
- wide-strip electrode
 ground plane

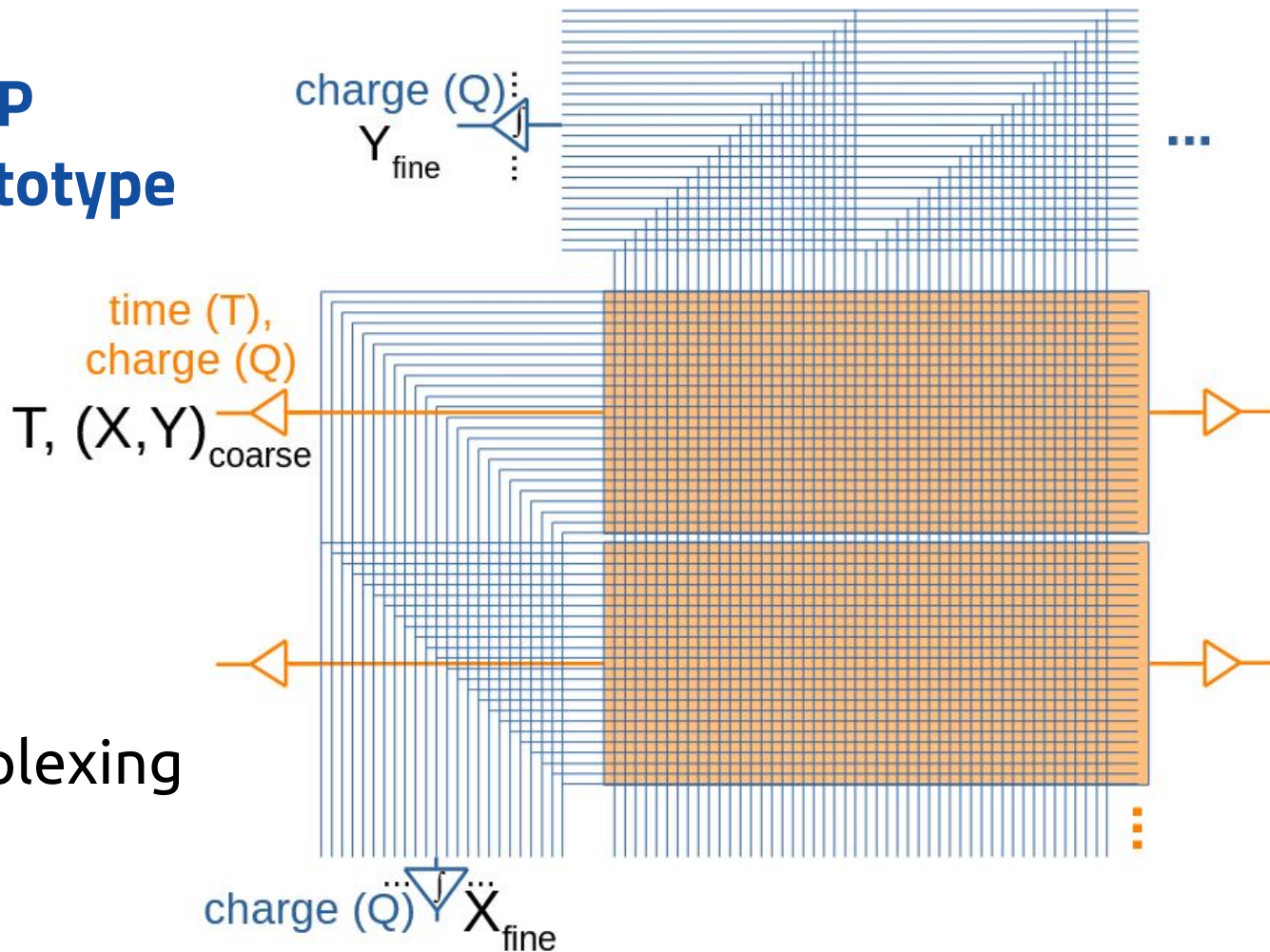
SHIP prototype

But this makes the channel count massive → high cost

Solution?



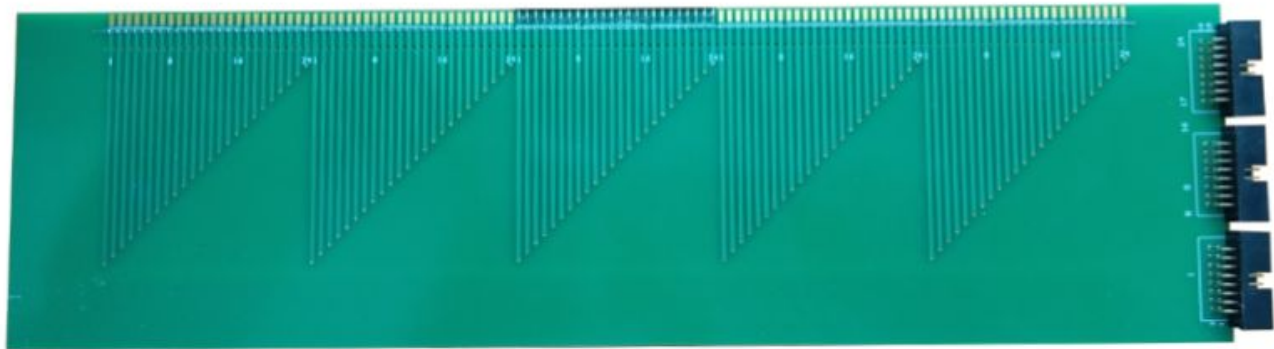
SHIP prototype



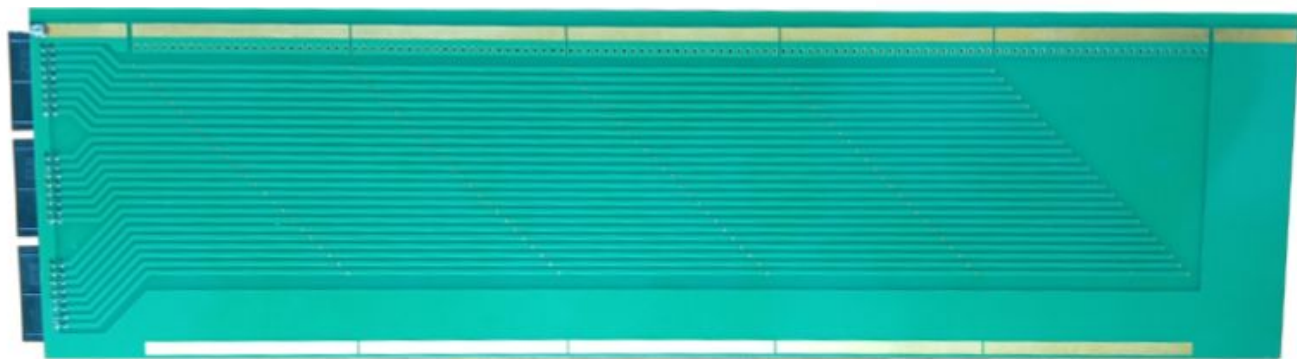
Multiplexing

SHIP prototype

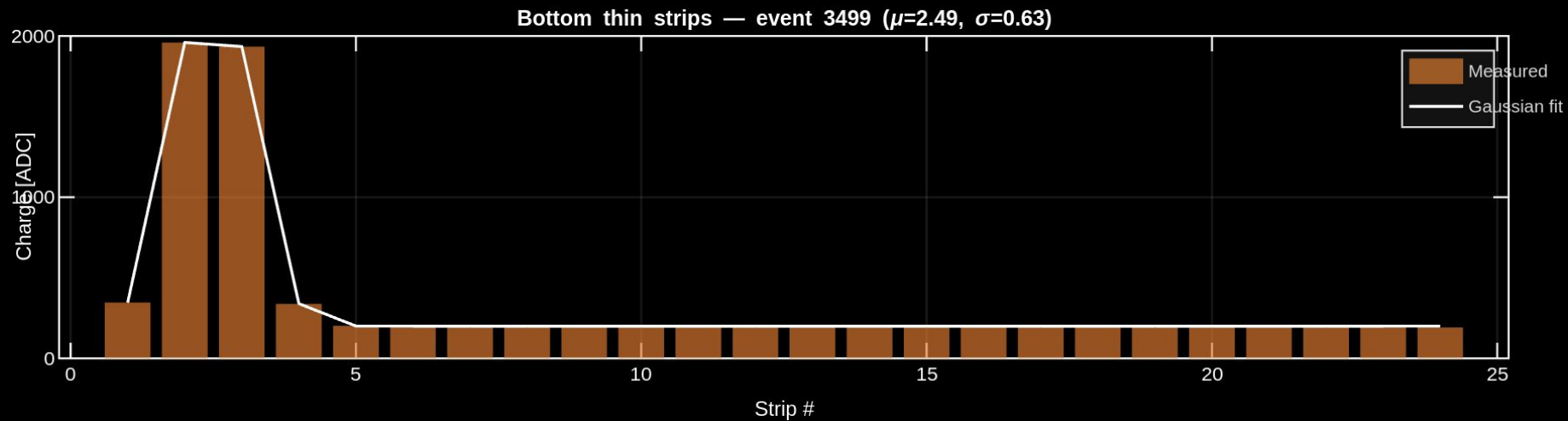
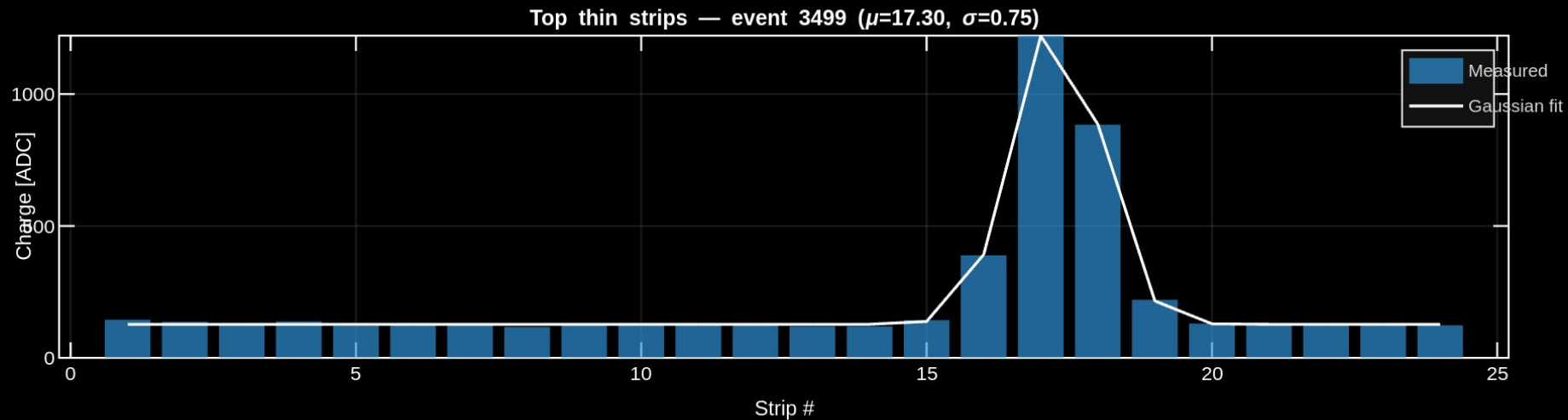
Multiplexing PCB



Front side

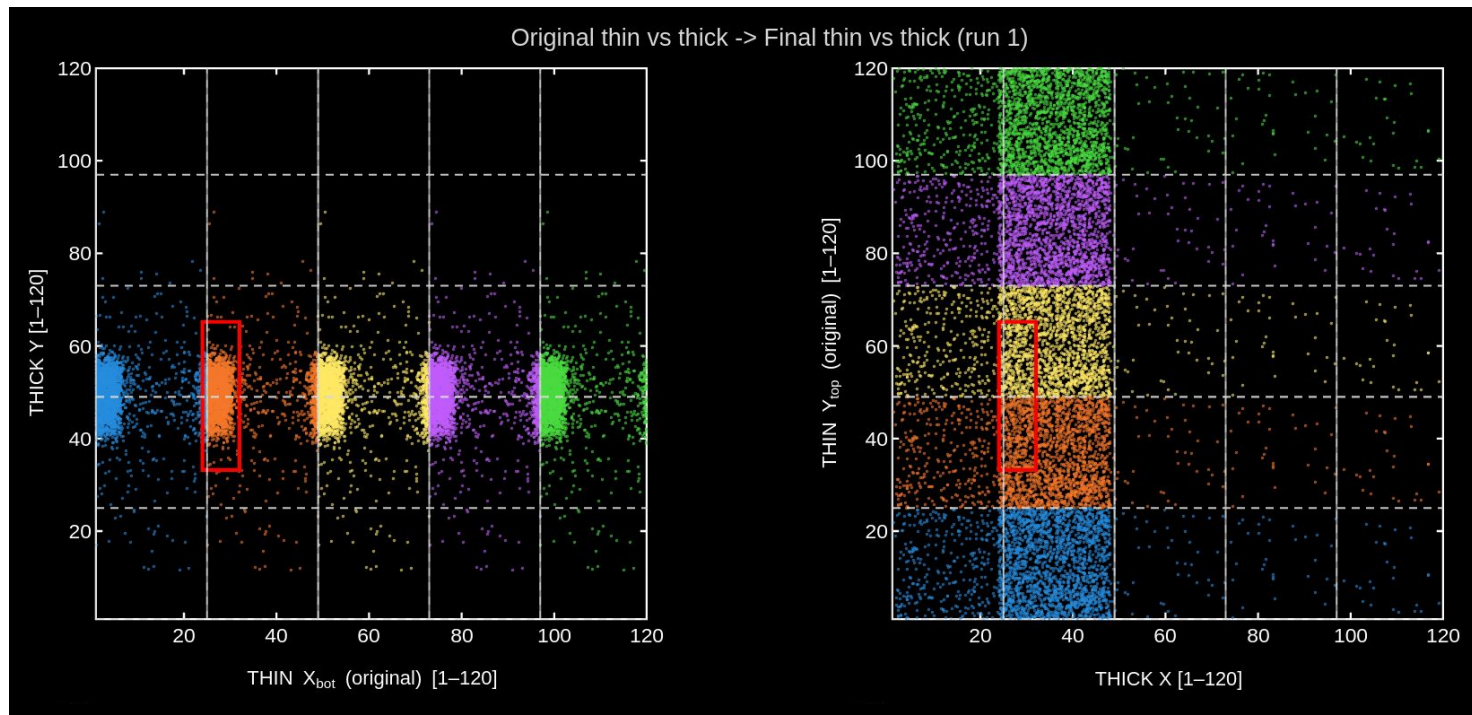


Rear side



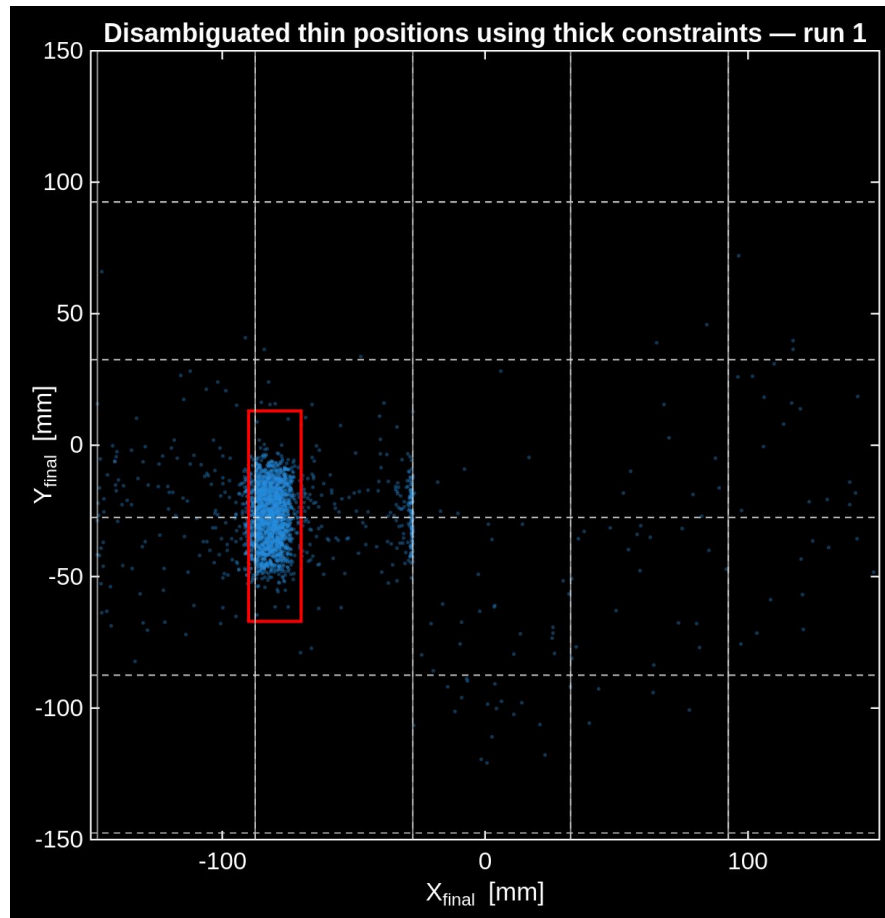
SHIP prototype

Ambiguities can be eliminated by combining information from different detector components



SHIP prototype

Unwrapped
positions
(preliminary
method, yet to
be refined)



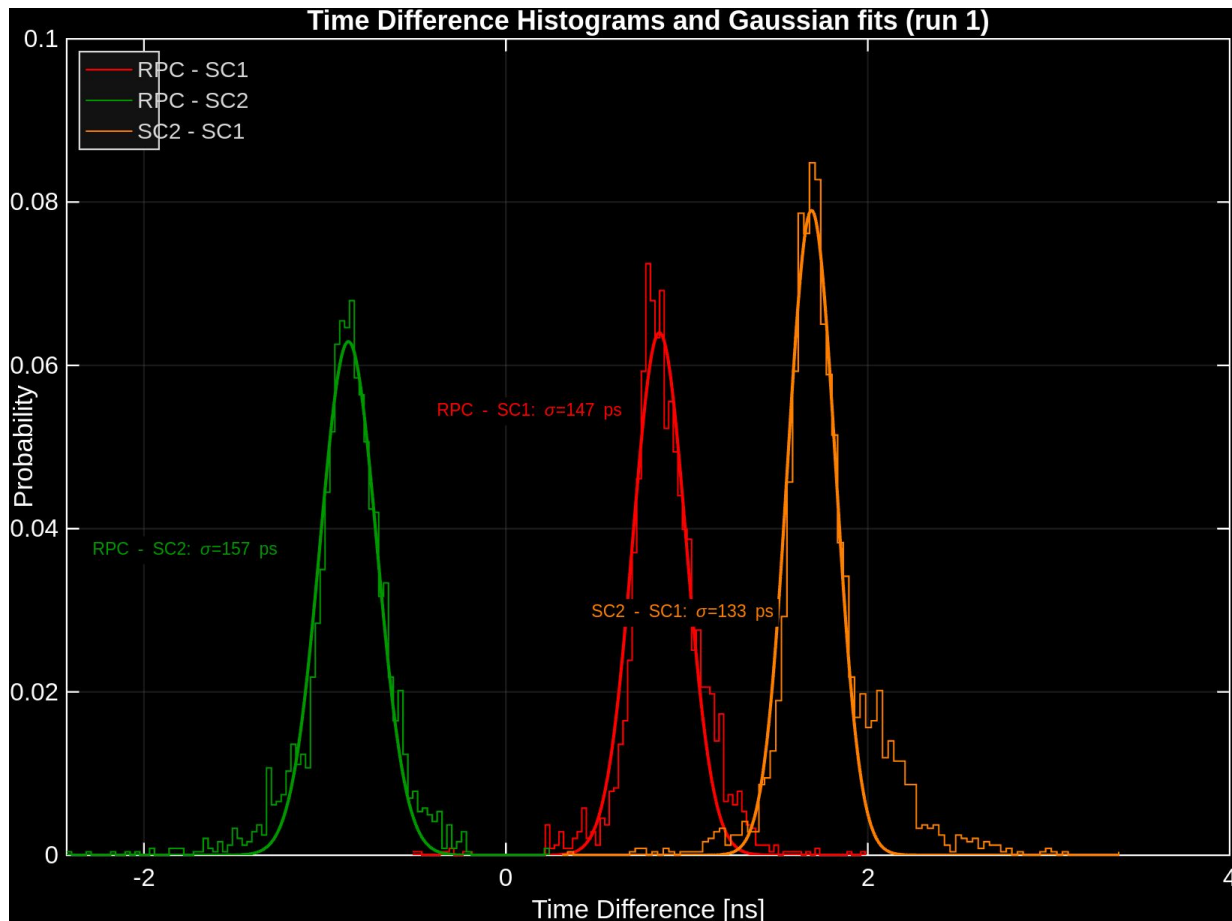
SHIP prototype

Time resolution (preliminary)

$$\begin{bmatrix} 1 & 1 & 0 \\ 1 & 0 & 1 \\ 0 & 1 & 1 \end{bmatrix} \begin{bmatrix} \sigma_{RPC}^2 \\ \sigma_{SC_1}^2 \\ \sigma_{SC_2}^2 \end{bmatrix} = \begin{bmatrix} \sigma_{RPC-SC_1}^2 \\ \sigma_{RPC-SC_2}^2 \\ \sigma_{SC_1-SC_2}^2 \end{bmatrix}$$

Time res. (equal-PMT assumption)

- RPC → **119 ps (preliminary)**
- SC → 94 ps

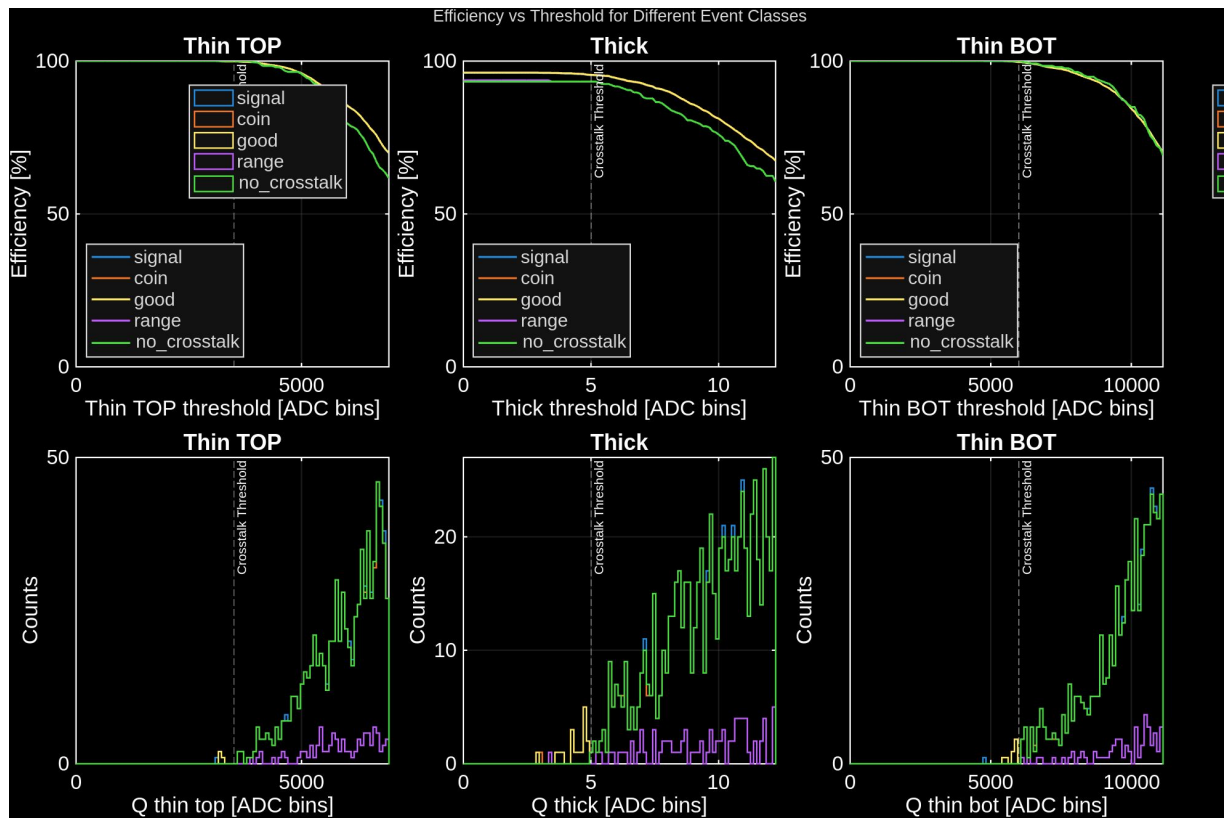


SHIP prototype

Efficiency (preliminary)

*Using scintillators
as reference*

RPC	Efficiency [%]
Thin_top	100.0 (2.8)
Thick_center	96.2 (2.7)
Thin_bottom	100.0 (2.8)



Next step

going for the single plane

Single-plane prototype

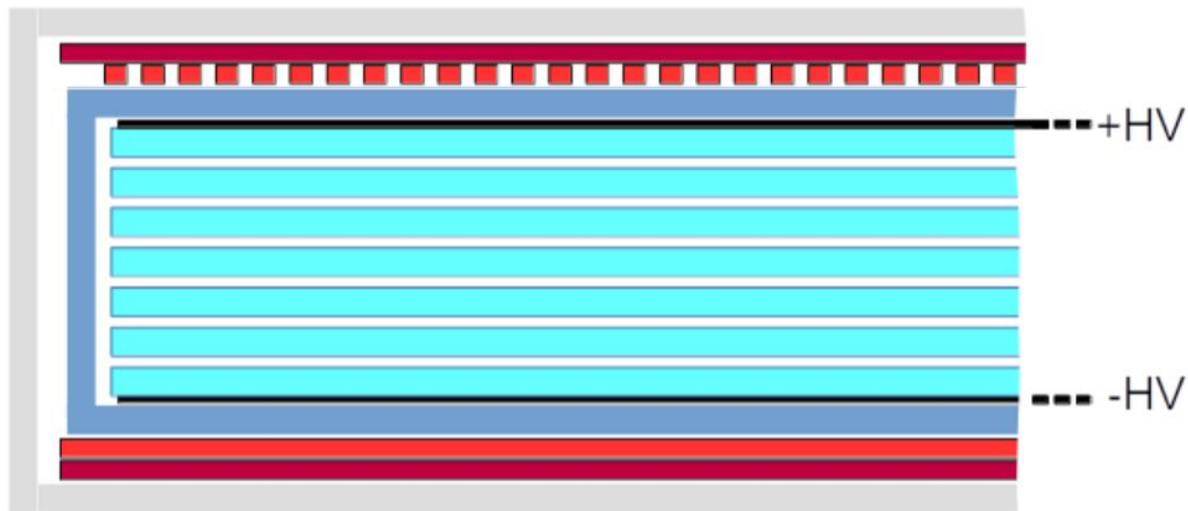
→ Fewer channels, less electronics, less material

Problem

Which readout for timing, which one for position?

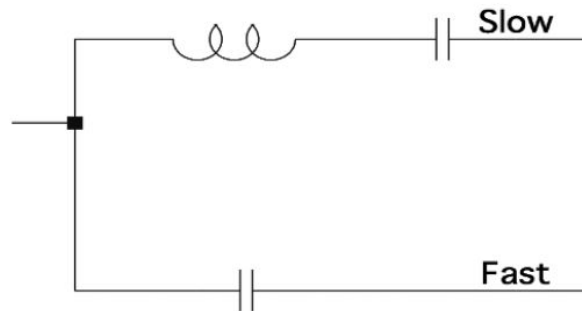
Solution

Both of them!



- Al box
- PP box
- acrylic+graphite paint, $R_s \sim 10^8 \Omega.\text{sq}^{-1}$
- thin-strip electrode
- float glass, $\rho \sim 4 \times 10^{12} \Omega.\text{cm}$ (25 °C)
- ground plane

Single-plane prototype: signal separation



Signal collected fully on thin strips

- Dedicated RC network to split fast and slow components
- Tests (Joana Pinto @ LIP):
 - Emulation of a typical RPC pulse.
 - Capacitors scanned and optimal ratio found
 - i. $C_{\text{fast}} : C_{\text{slow}} \approx 1 : 10$, with final values 560 pF (fast) and 5.6 nF (slow), avoiding saturation and keeping amplitudes.

→ *Next step is plugging real RPC signals*

Concluding

a summary of the story

Conclusions - *A detector story*

- **miniTRASGO**

a good workspace for detector characterization

- **R3B prototype**

→ double mRPC read in the middle

- **SHIP prototype**

→ introducing thin strips and multiplexing

- **Single plane**

→ thin strips only and fast/slow filter

We are here.



Thank you! *Questions?*

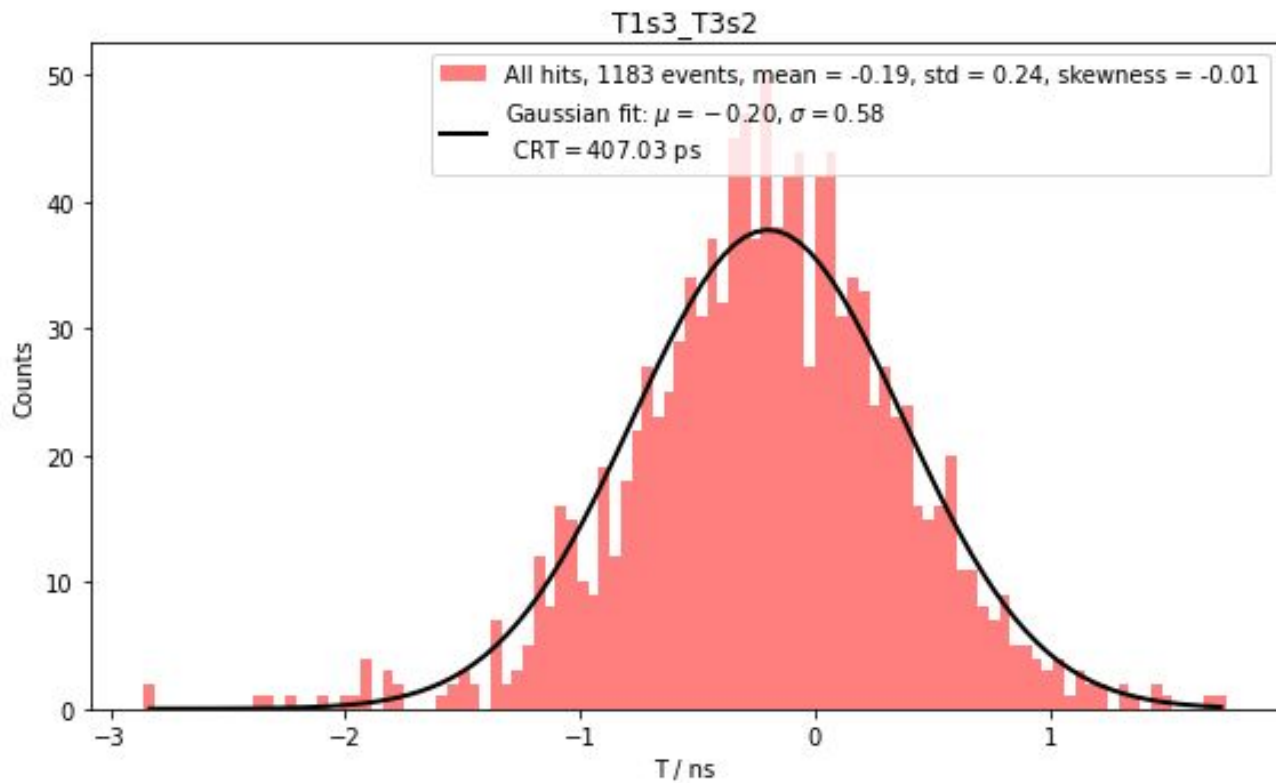
*RPC readout optimization:
miniTRASGO as a path to a
single-plane multiplexed
timing system*

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C. Soneira-Landín*



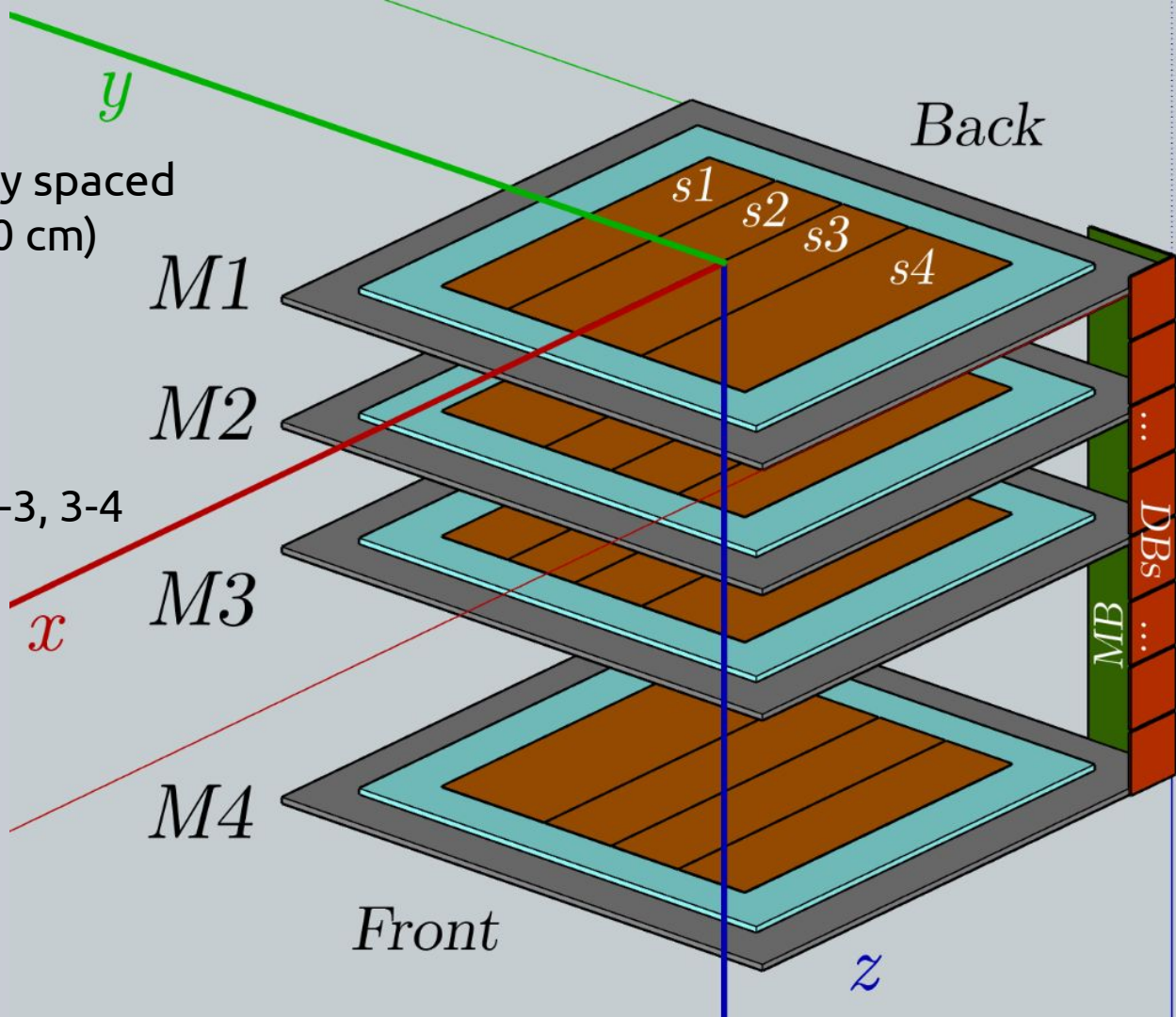
Back-up slides

for those curious



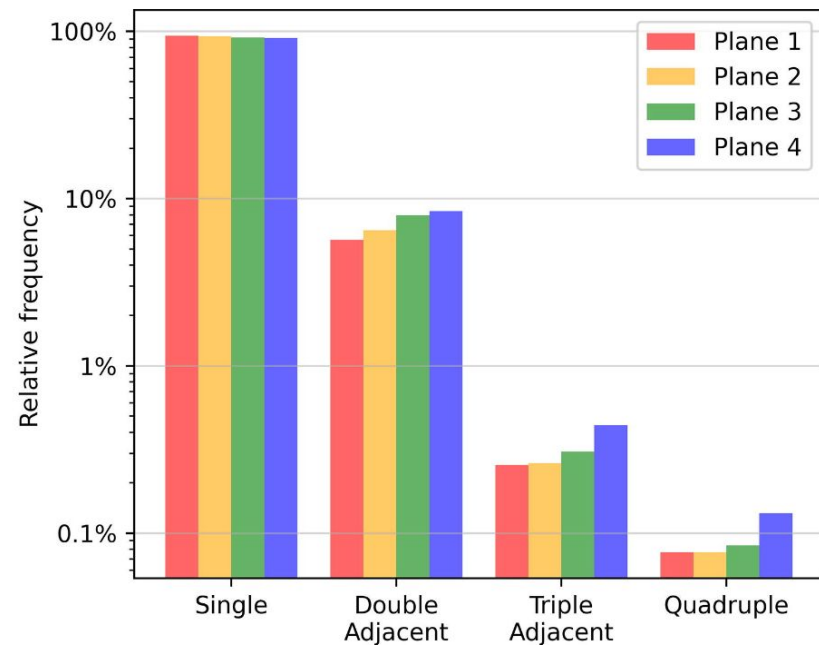
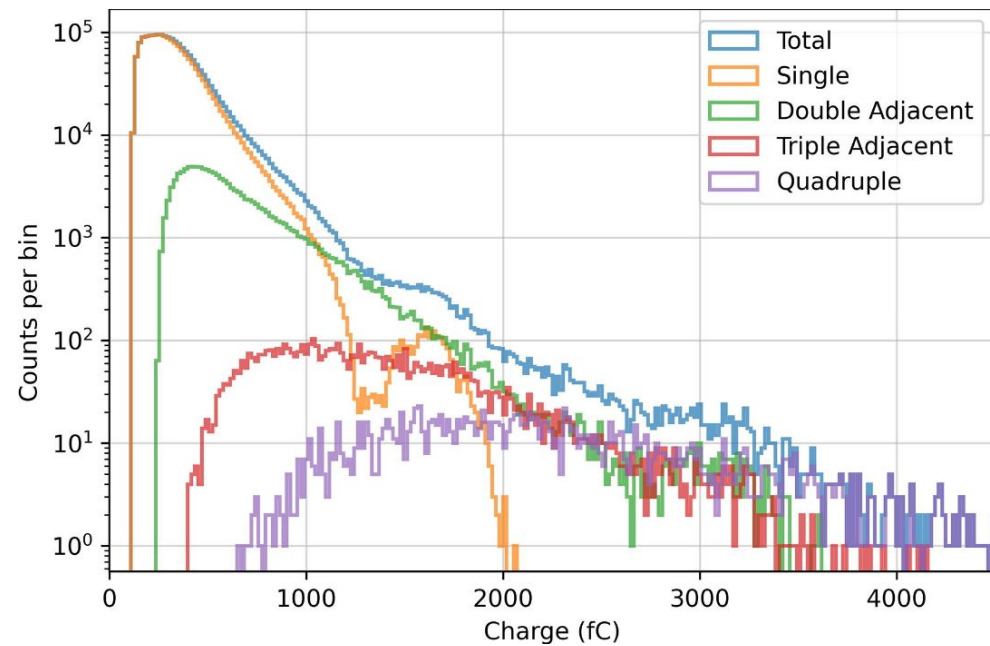
Not evenly spaced
(10 and 20 cm)

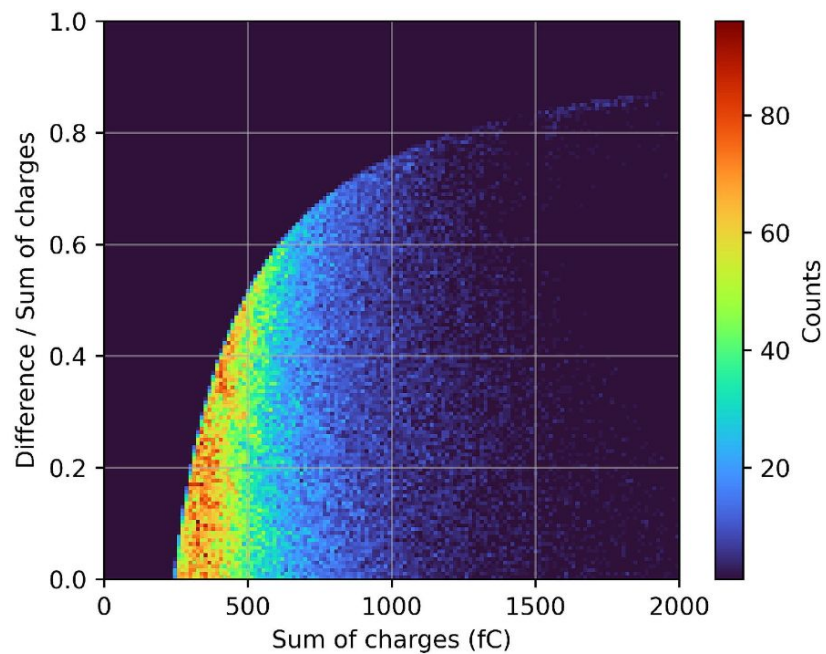
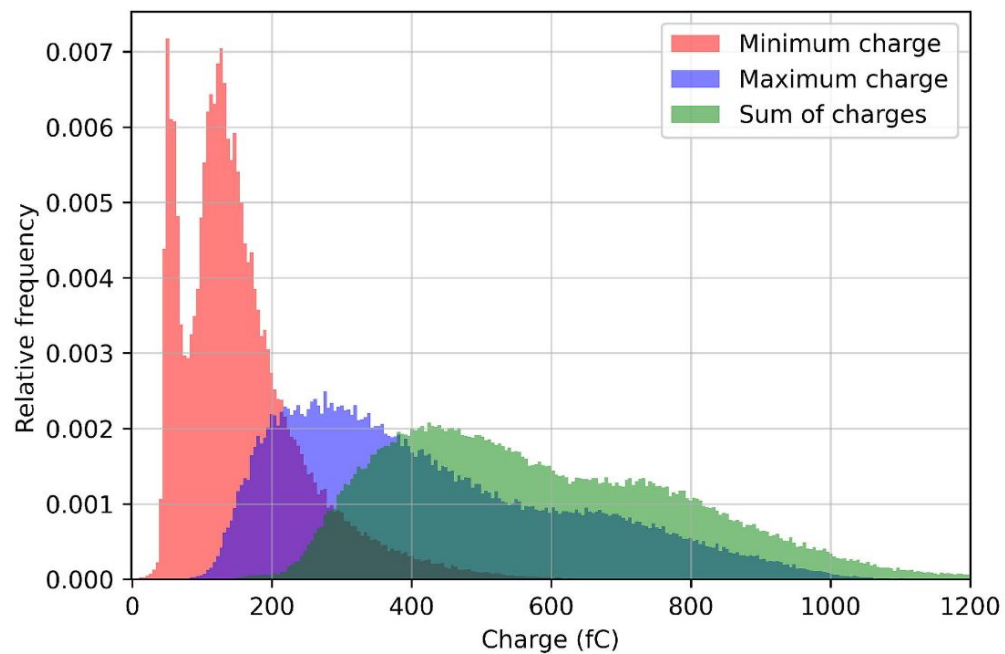
Trigger:
1-2, 1-3, 2-3, 3-4

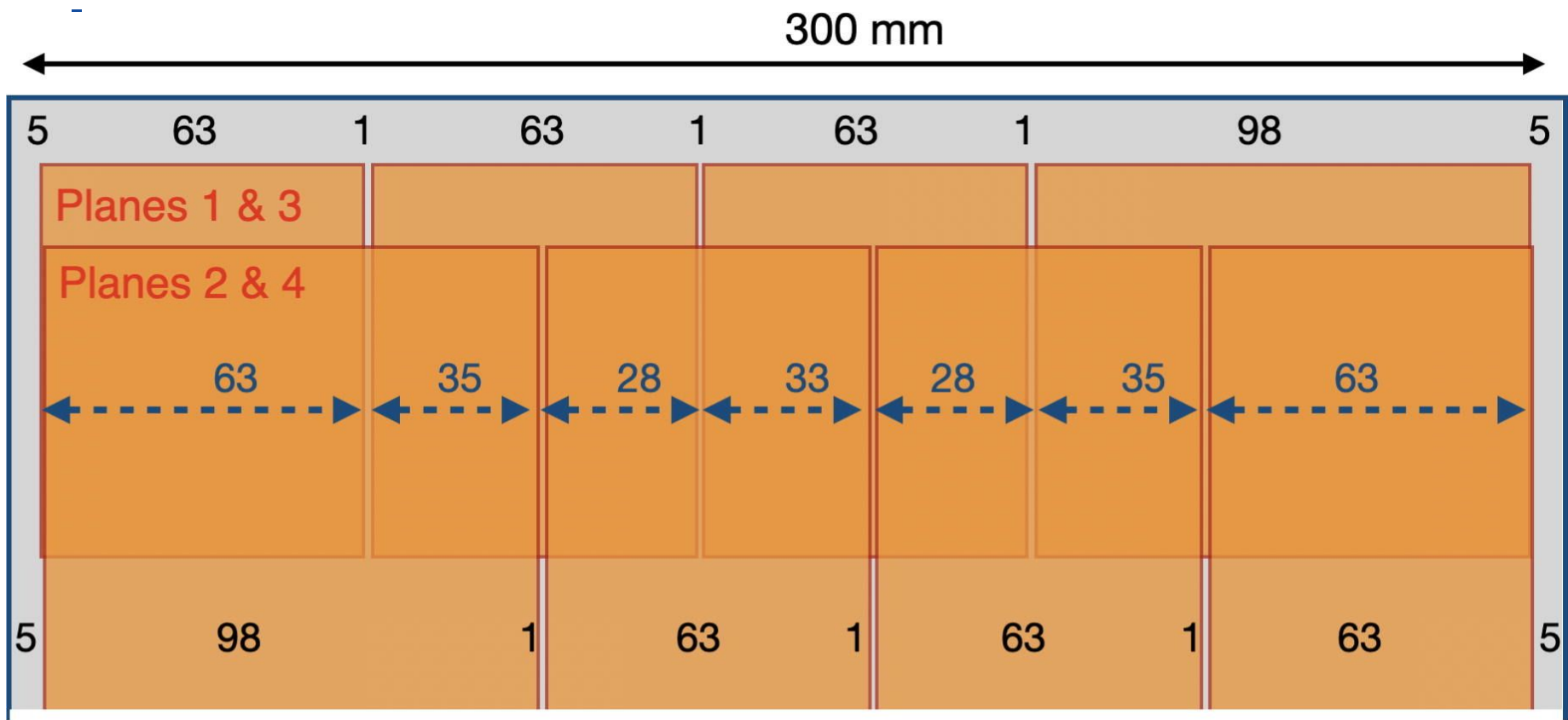


Asymmetric
parallel **Cu**
strips read in
Front and Back

32 channels







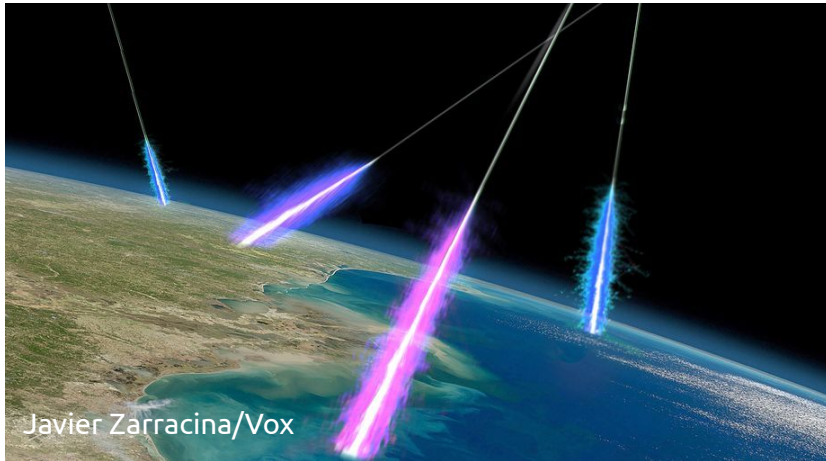
Nuclear reactions in the upper atmosphere

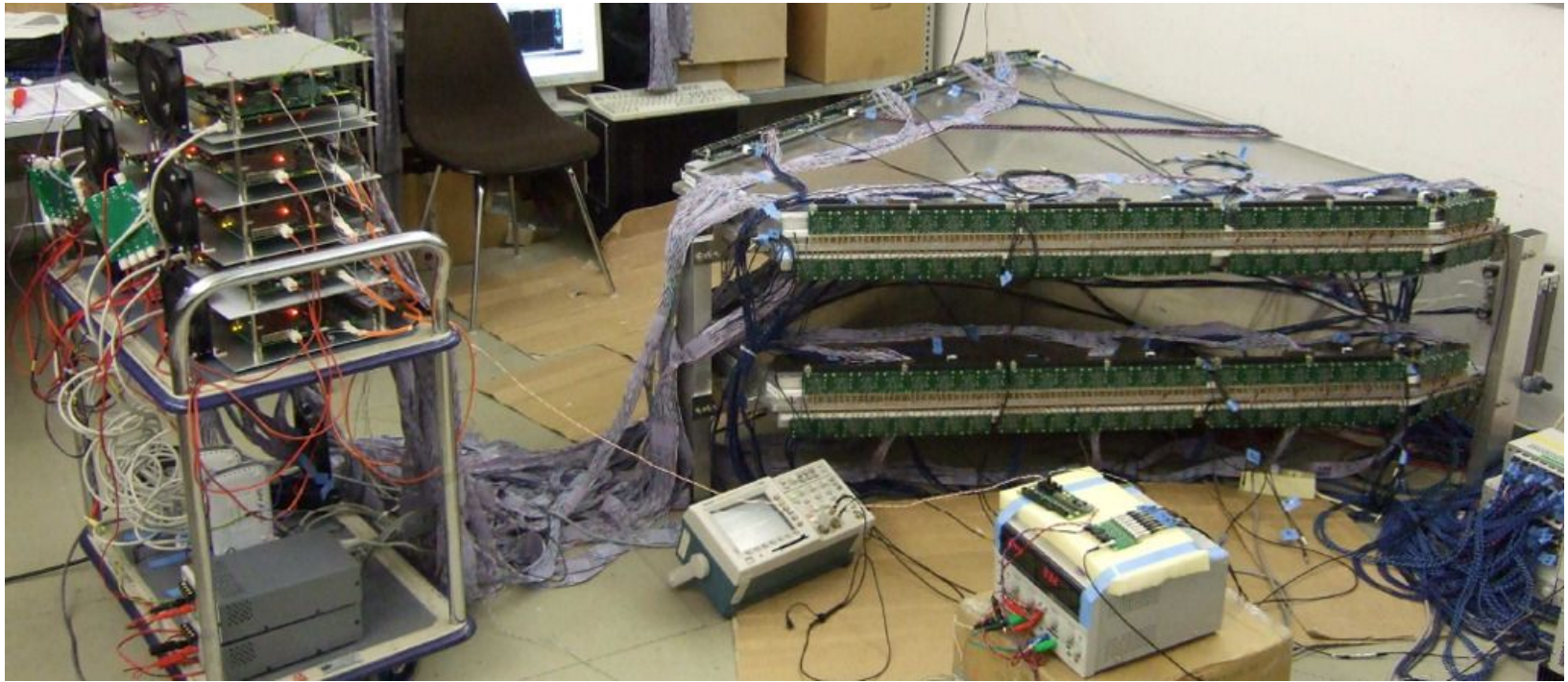
$$p + A \rightarrow p + n + \pi^0 + \pi^\pm + \dots$$

$$\pi^\pm + A \rightarrow \pi^0 + \pi^\pm + \dots$$

$$\pi^+ \rightarrow \mu^+ + \nu_\mu$$

$$\pi^- \rightarrow \mu^- + \bar{\nu}_\mu$$





HADES Time-of-Flight Wall (commissioning)

- D. Belver et al., *TRASGO: A proposal for a timing RPCs based detector for analyzing cosmic ray air showers*, NIM A, 2012, <https://doi.org/10.1016/j.nima.2010.09.173>
- A. Blanco et al., *RPC HADES-TOF wall cosmic ray test performance*, NIM A, 2012, <https://doi.org/10.1016/j.nima.2010.08.068>
- D. Belver et al., *Analysis of the space-time microstructure of cosmic ray air showers using the HADES RPC TOF wall*, JINST, 2012, 10.1088/1748-0221/7/10/P10007



TRAGALDABAS

- A Blanco et al., *TRAGALDABAS: a new RPC based detector for the regular study of cosmic rays*, JINST, 2014, DOI 10.1088/1748-0221/9/09/C09027
- I. Riádigos et al., *Revisiting the limits of atmospheric temperature retrieval from cosmic-ray measurements*. Earth and Space Science, 2022, <https://doi.org/10.1029/2021EA001982>
- I. Riádigos et al., *Atmospheric temperature effect in secondary cosmic rays observed with a 2 m² ground-based tRPC detector*. Earth and Space Science, 2020, <https://doi.org/10.1029/2020EA001131>



- Logicmelt, *Stratos DS – Stratospheric temperature prediction using artificial intelligence*, Logicmelt – Use Cases, 2025, https://logicmelt.com/en/use-cases_eng/stratos-ds-prediction-of-the-stratosphere-temperature/
- NAC-Intercom, *Project STRATOS – Ground Station for Continuous Monitoring of the Stratosphere Temperature through Cosmic Ray Directional Flow*, NAC-Intercom – Projects, 2025, <https://www.nac-inter.com/en/content/30-project-stratos>

STRATOS



TRISTAN

- J.P. Saraiva, et al., *The TRISTAN detector—2018–2019 latitude survey of cosmic rays. Journal of Instrumentation*, JINST, 2020, <https://doi.org/10.1088/1748-0221/15/09/C09024>



miniTRASGO network

Warsaw 🇵🇱

Madrid 🇪🇸

Puebla 🇲🇽

Monterrey 🇲🇽

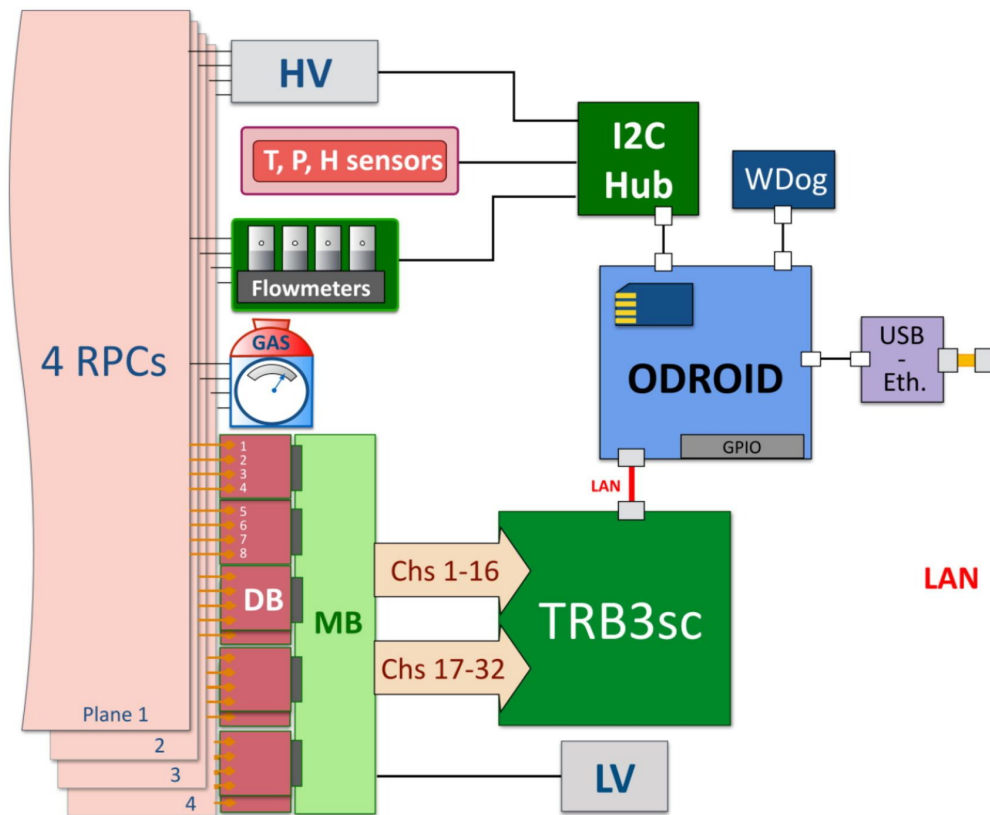
Compact and autonomous

Only three connections:

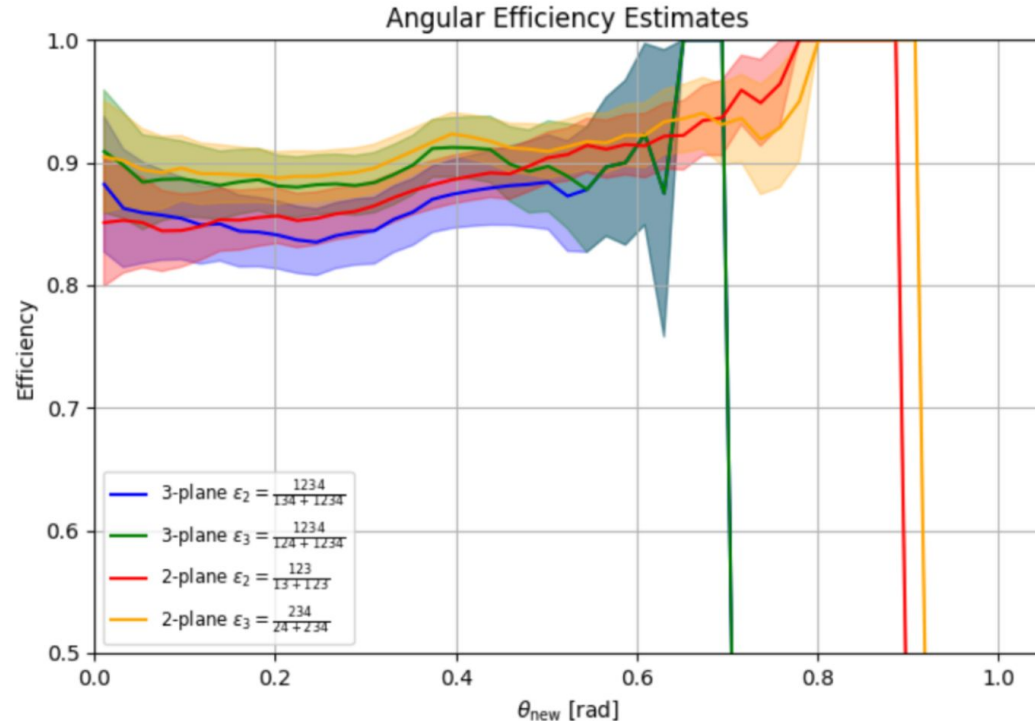
- **Power** (standard plug)
- **Gas** bottle (15 kg)
- **Ethernet** (not required)

Includes:

- environment sensors, flowmeters, DAQ and a PC for data analysis and storage



Efficiency vs. incident zenith muon angle



*Oblique tracks
spend more time in
the active volume*

Forbush Decrease

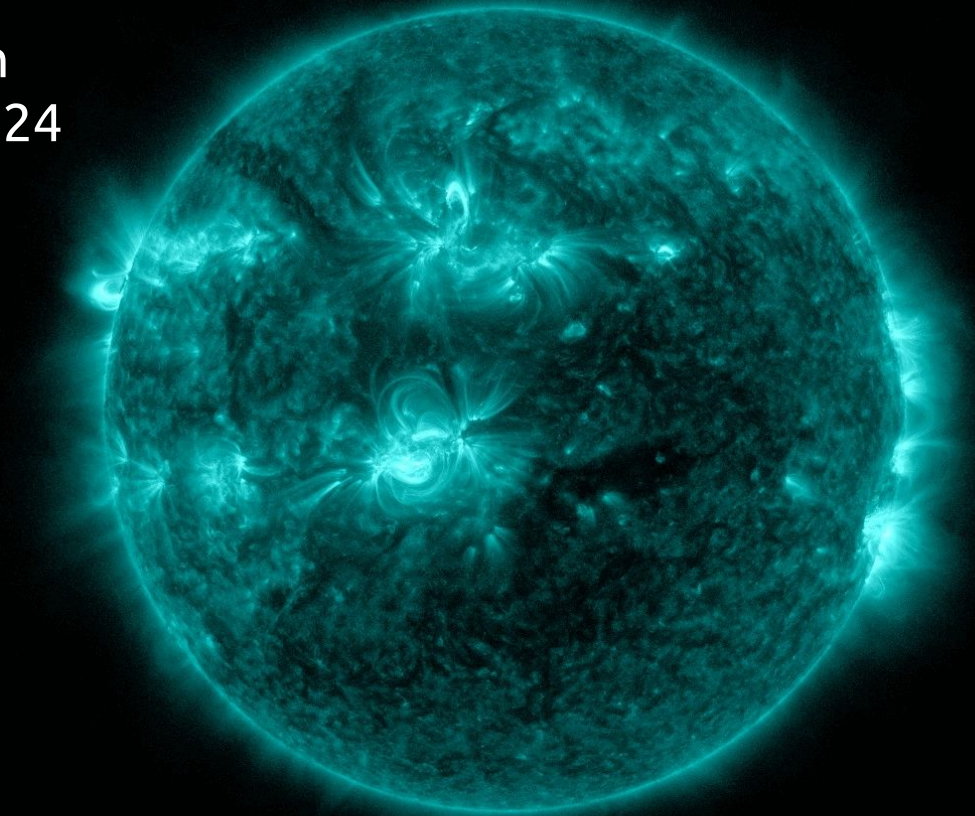
On March 2024

Two Sunspots exploded in
tandem on March 23rd 2024
(03:30 CET)

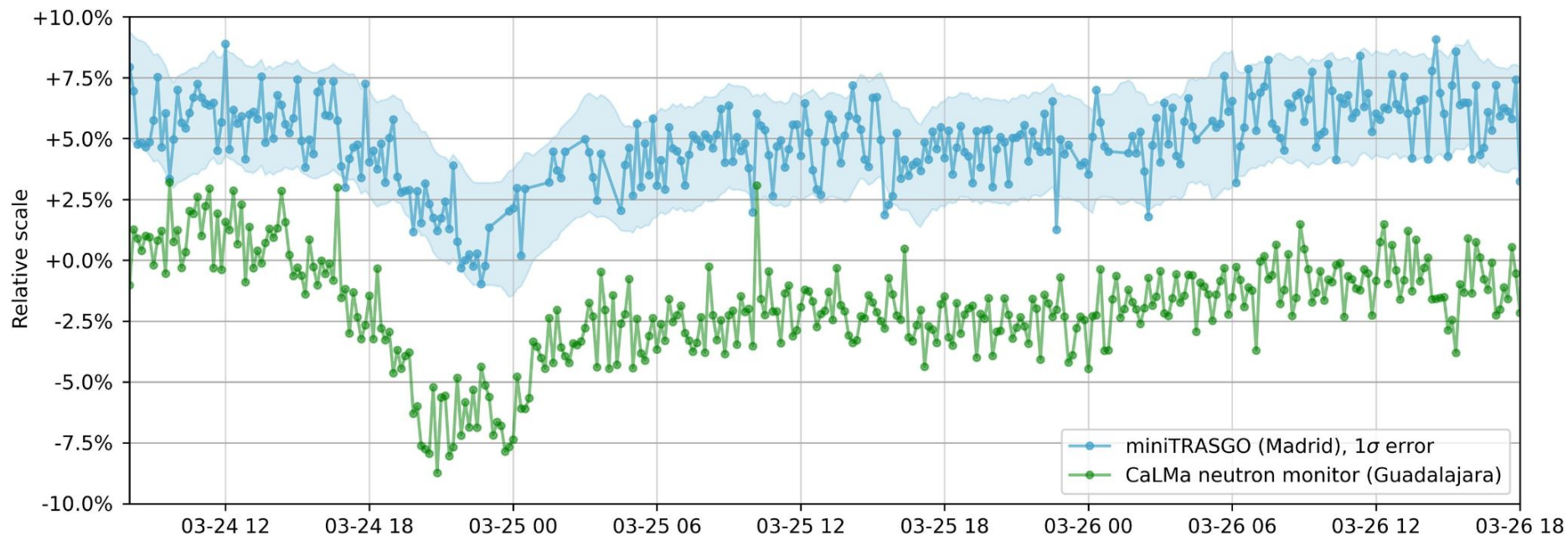


8 hours later...

From Solar Dynamics
Observatory, NASA



Eff. and pressure corrected rate
→ the Forbush Decrease can be seen



Measure: time and charge in F and B per strip



Time sum → Incidence time

Time diff → Position

Charge sum → Total charge

Charge diff → Monitor

