

# XVII CPAN DAYS

Centro Nacional de Física de Partículas,  
Astropartículas y Nuclear.

**First full  $\beta$ -strength  
measurement with DTAS across  
N=126 at FAIR phase-0**

David Rodríguez García

*On behalf of DESPEC collaboration*

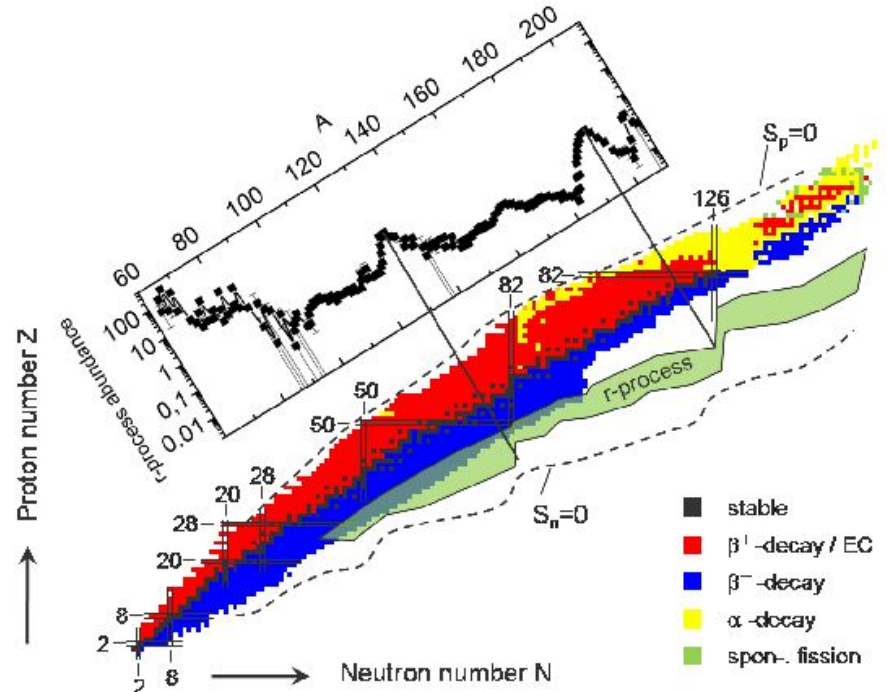
**19–21 Noviembre 2025  
Valencia**

david.rodriguez@ific.uv.es



# Aim of the experiment

- Site for the astrophysical r-process is still uncertain
- Nuclear input data is also uncertain: lack of experiments
- In particular data for the 3<sup>rd</sup> r-process peak (~N=126)
- Benchmarking  $\beta$ -strength theoretical models used for  $T_{1/2}$  and  $P_n$  predictions on r-process calculations
- Using Total Absorption Gamma-ray Spectroscopy (TAGS) to benchmark directly the  $\beta$ -strength (and not the  $T_{1/2}$  and  $P_n$ )



!  $T_{1/2}$  and  $P_n$  derived from  $S_\beta$  ( $\beta$ -strength) calculations:

$$\frac{1}{T_{1/2}} = \int_0^{Q_\beta} S_\beta(E_x) \cdot f(Q_\beta - E_x) dE_x$$

$$S_\beta^{th}(E_x) = \frac{1}{D} \frac{g_A^2}{g_V^2} \frac{1}{2I_i + 1} \left| \langle f \| M_{\lambda\pi}^\beta \| i \rangle \right|^2$$

$$P_n = T_{1/2} \int_{S_n}^{Q_\beta} \frac{\Gamma^n(E_x)}{\Gamma_{tot}(E_x)} S_\beta(E_x) \cdot f(Q_\beta - E_x) \cdot dE_x$$

$$S_\beta^{exp}(E_x) = \frac{I_\beta(E_x)}{T_{1/2} f(Q_\beta - E_x)}$$

TAGS

# S505 - Experimental Set-up at S4

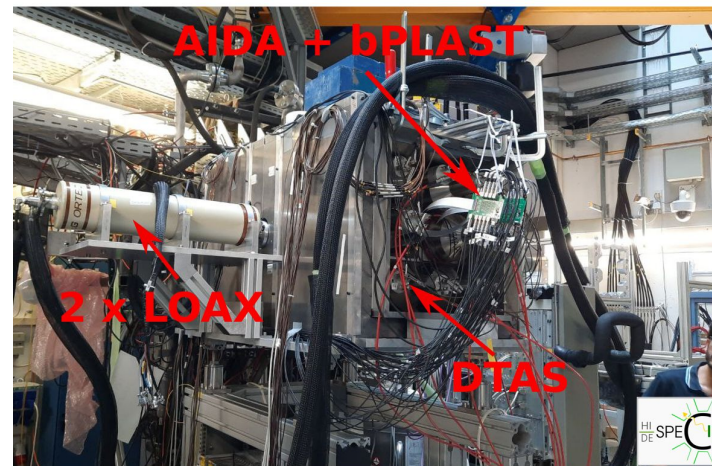
Spokesperson: Jose Luis Tain, Ana Isabel Morales, Enrique Nacher

Performed: June 21-28, 2022

Beam/Target:  $\leq 4.5 \times 10^8$  ppb (1.6s/2.2s), 1GeV/u

$^{208}\text{Pb}$  on  $1.6 \text{ g/cm}^2$  Be

FRS Settings:  $^{204}\text{Pt}$  (~125h),  $^{207}\text{Hg}$  (~15h)



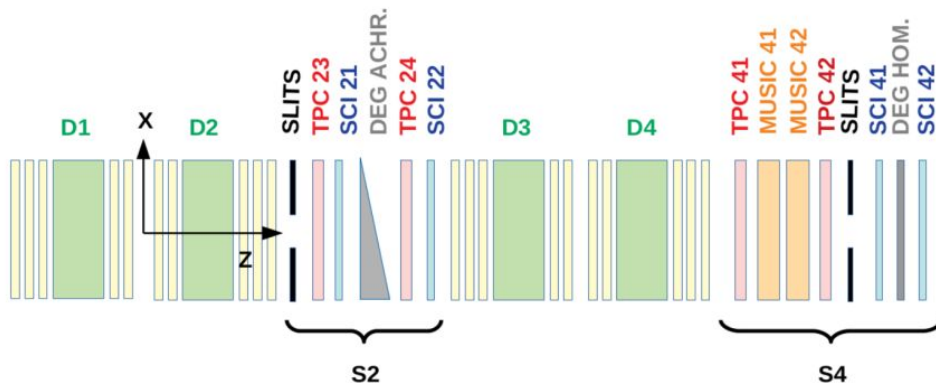
**DTAS** (16xNaI(Tl) det.)

**AIDA** (2xDSSDs-80mmx80mm)

2xb**Past** detectors

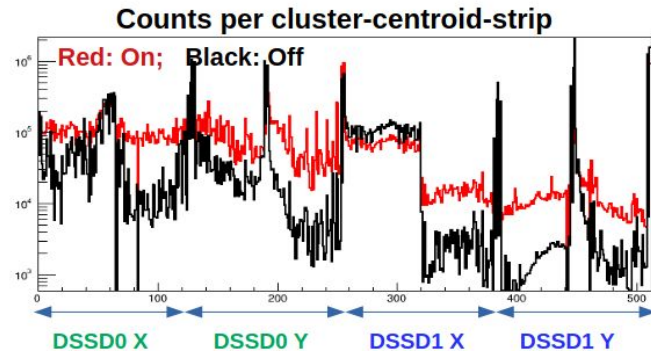
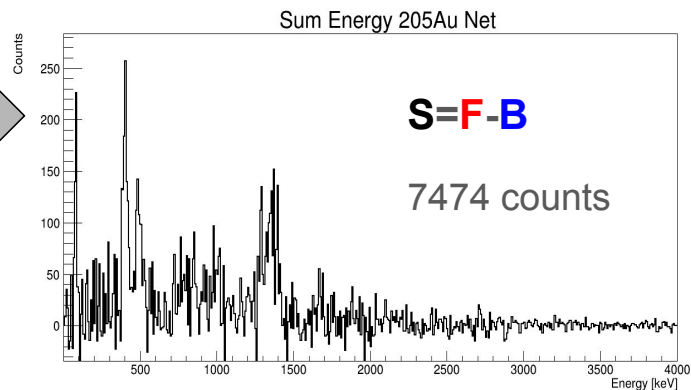
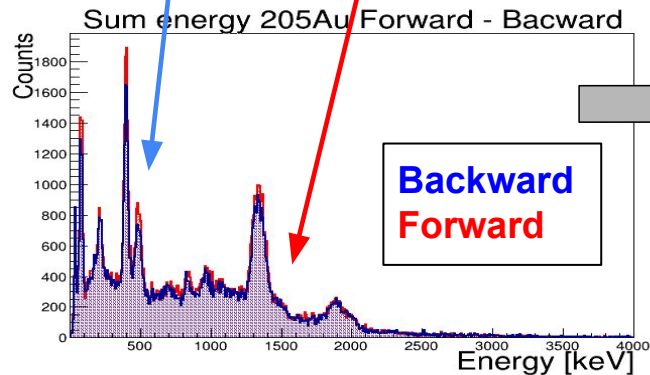
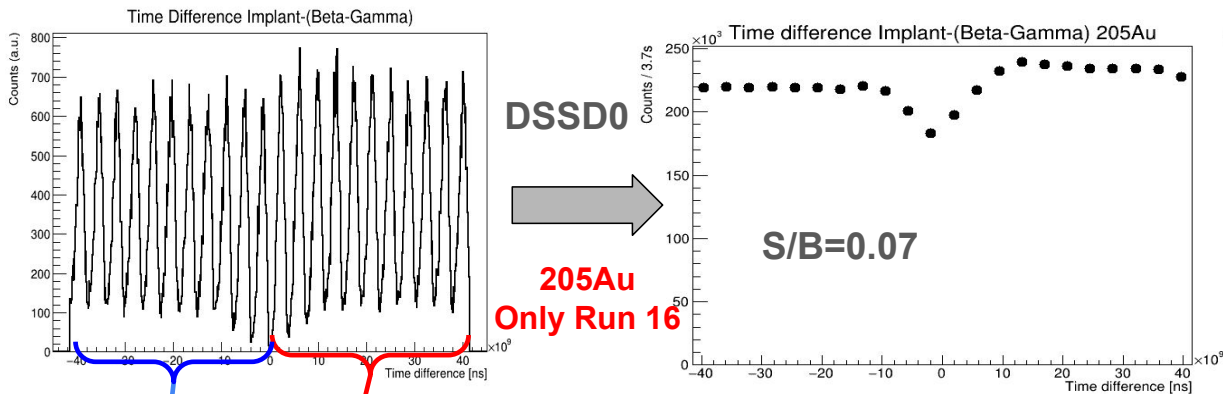
2x**LOAX** HPGe (1 not operational)

		$^{207}\text{Hg}$ 2.9m 4.6MeV (9/2+)
$^{204}\text{Au}$ 39.8s 4.0MeV (2-)	$^{205,205\text{mAu}}$ 32.5/6s 3.5/4.4MeV (3/2+, 11/2-)	$^{206}\text{Au}$ 45s 6.7MeV (?)
$^{203,203\text{mPt}}$ 22s/12s 3.5MeV (1/2-, ?)	$^{204}\text{Pt}$ 16s 2.7MeV 0+	





- Implant-beta(gamma) time correlation: forward and backward
- Gamma spectra from forward-backward correlation differences of identified implants



- Huge rate in decay branch, in particular for DSSD0 during spill
- Large noise on most FEE64s
- Noisy strips can be identified
- Large number of strips firing in one decay event

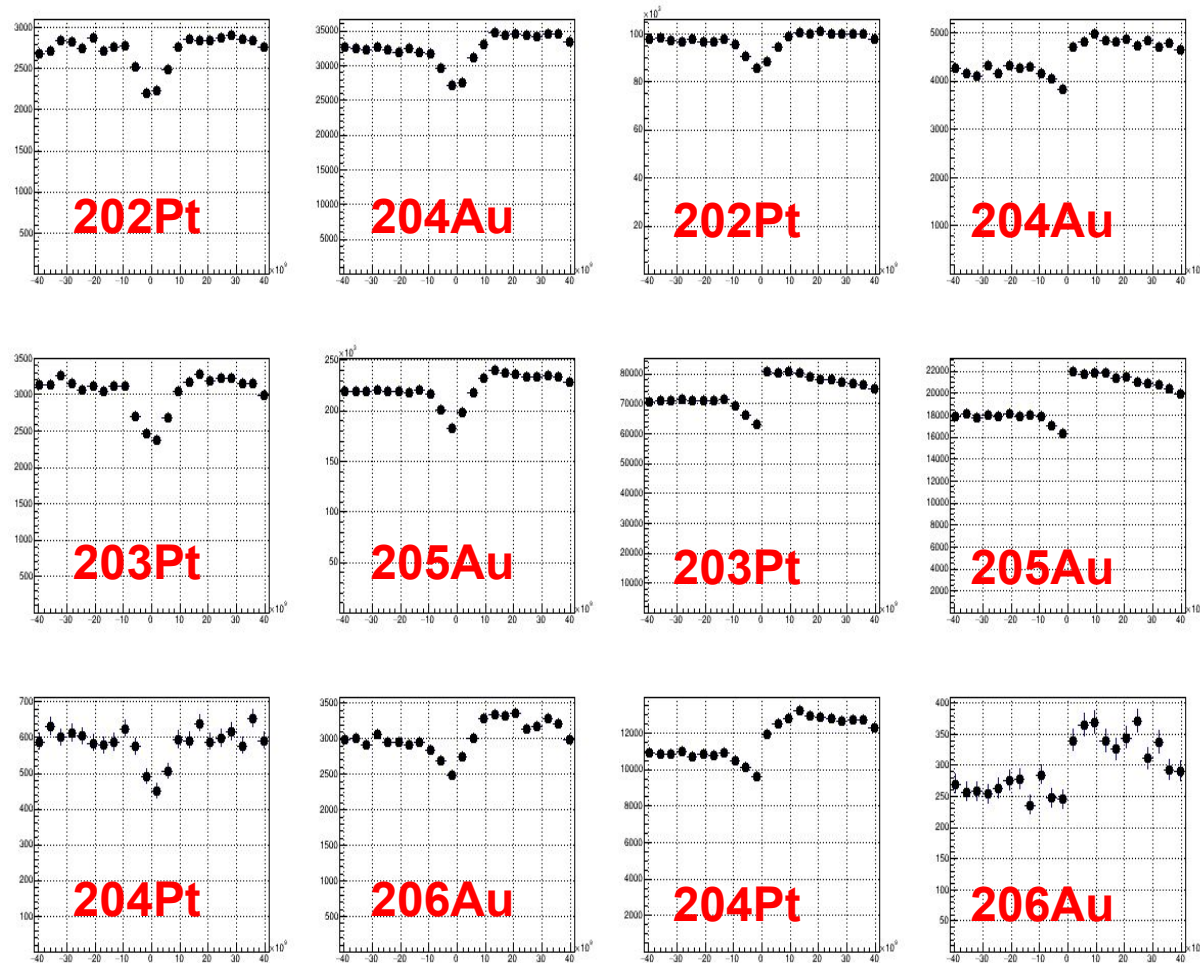
**Issue: very poor S/B**

## DSSD0 (S/B~0.06)

## DSSD1 (S/B~0.15-0.20)

Implant-beta(gamma) time correlation window:  $\pm 41.5$ s

Large difference between DSSD0 and DSSD1, in signal-to-background and shape.



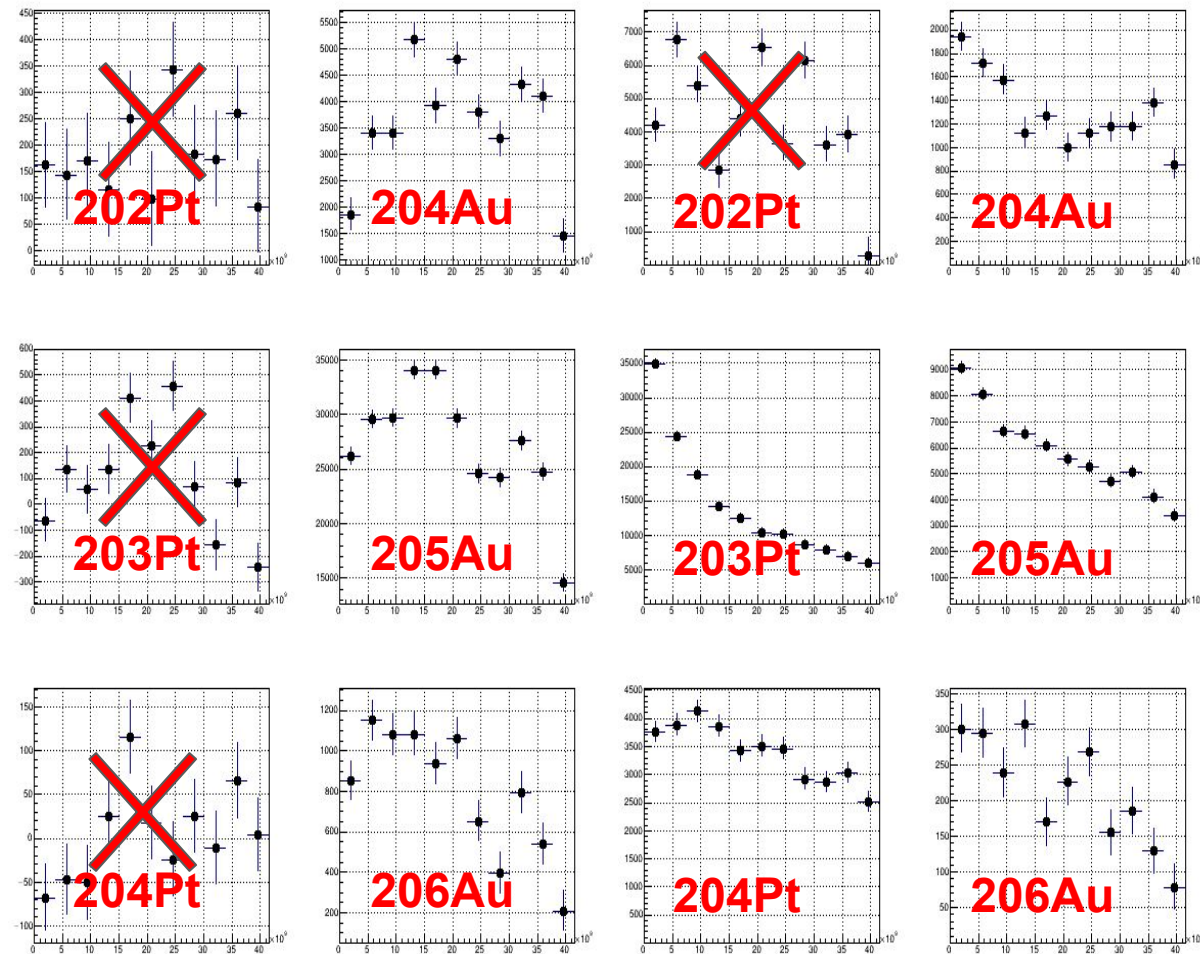
Reminder: We implanted Pt in DSSD1 and Au in both detectors  
80% Au implants in DSSD0

## DSSD0 (S/B~0.06)

## DSSD1 (S/B~0.15-0.20)

Difference forward - backward in  
the time correlation  
implant-beta(gamma).

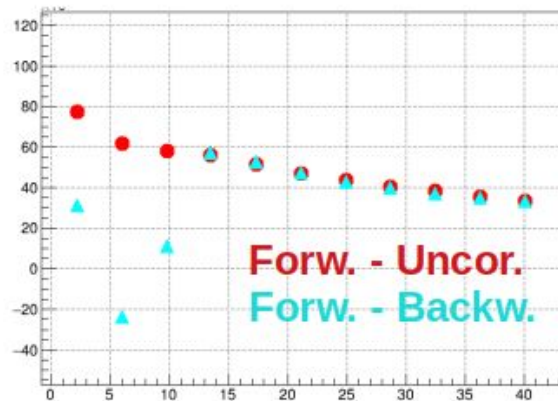
Strange behaviour in DSSD0  
DSSD1 has a better shape



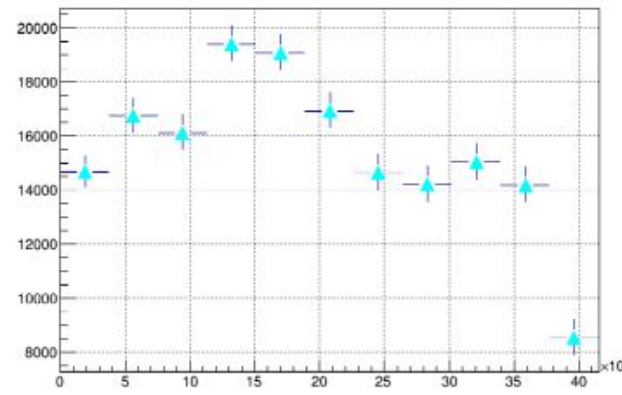
Reminder:  
We implanted Pt in DSSD1 and Au in both detectors.  
 $T_{1/2}$  202Pt = 45(15)h  
80% Au implants in DSSD0

## 205Au case (10% 205mAu)

DSSD0: MC



EXP



We made a realistic MC simulation of implant-beta correlation trying to understand this behaviour.

Experimental information used in the simulation:

Spill sequence (1.6s/2.2s), 18 delivered, 3 not delivered.

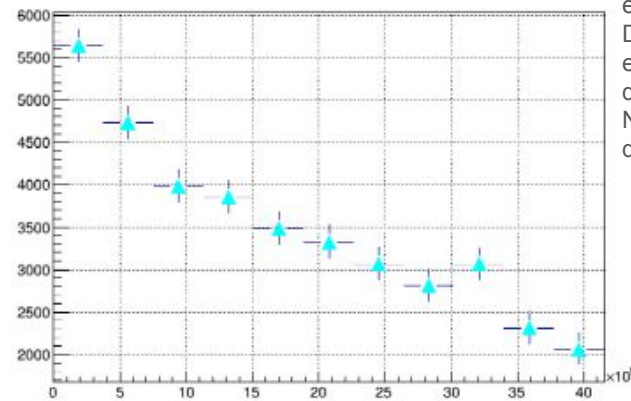
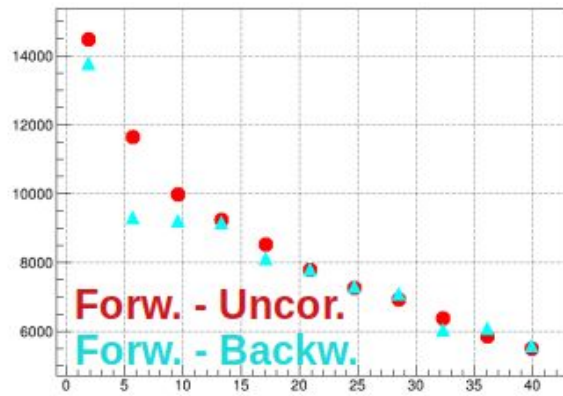
Implants: time distribution, experimental efficiency and rate XY distribution.

Decay: sequence to stability, experimental beta-gamma efficiency, dead-time and XY distribution.

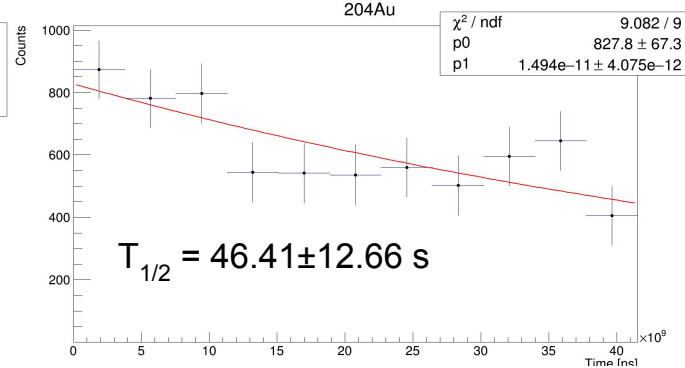
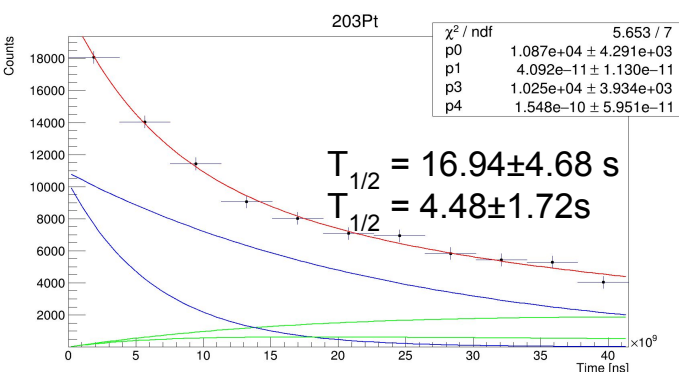
Noise: rate variation over time and XY distribution.

## 205Au case (10% 205mAu)

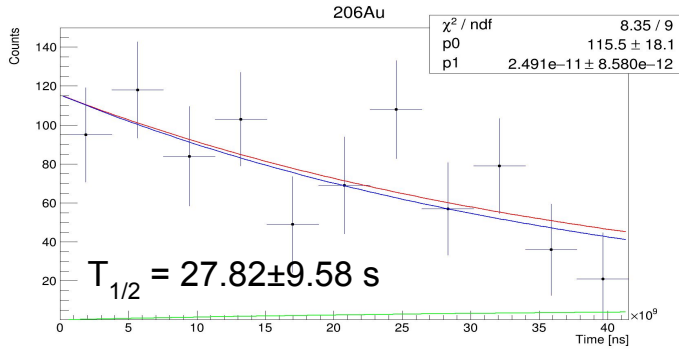
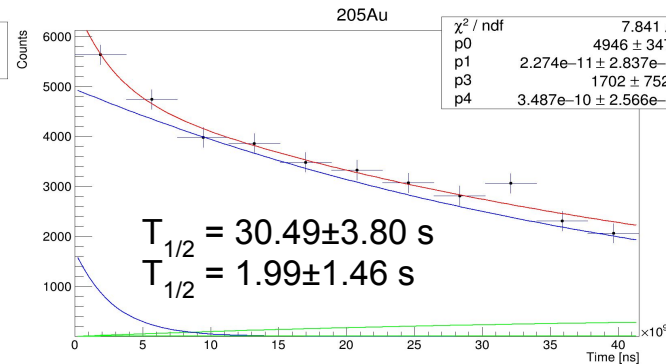
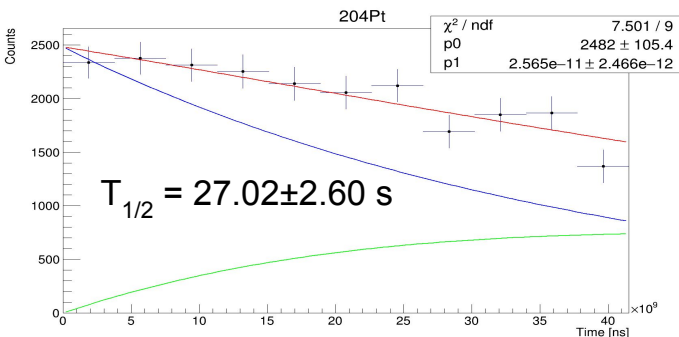
DSSD1: MC







**PRELIMINARY**



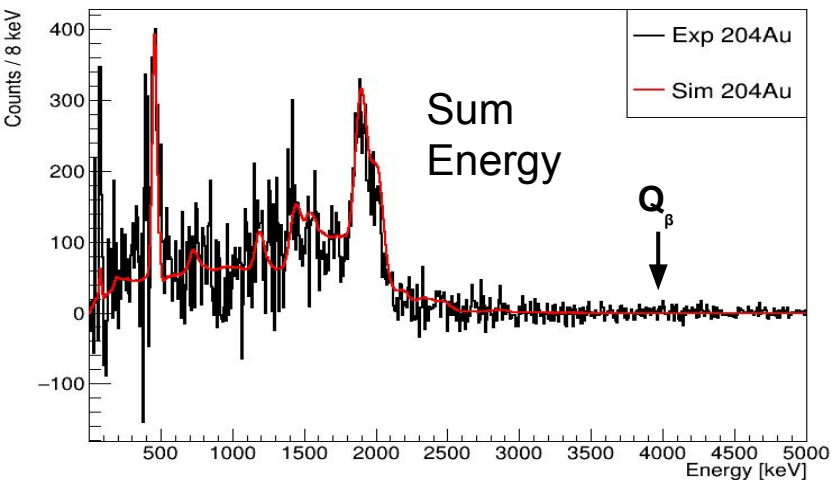
Short (isomer) half-lives affected by background subtraction (!?)

Isotope	Previous (ENSDF)	This work
<b>203Pt</b> <b>203mPt</b>	22(4); 12(5)	17(5); 4.5(17)
<b>204Pt</b>	10.3(13)	27(3)
<b>204Au</b>	39.8(9)	46(13)
<b>205Au</b> <b>205mAu</b>	32.0(14); 6(2)	30.5(38); 2.0(15)
<b>206Au</b>	40(15)	28(10)

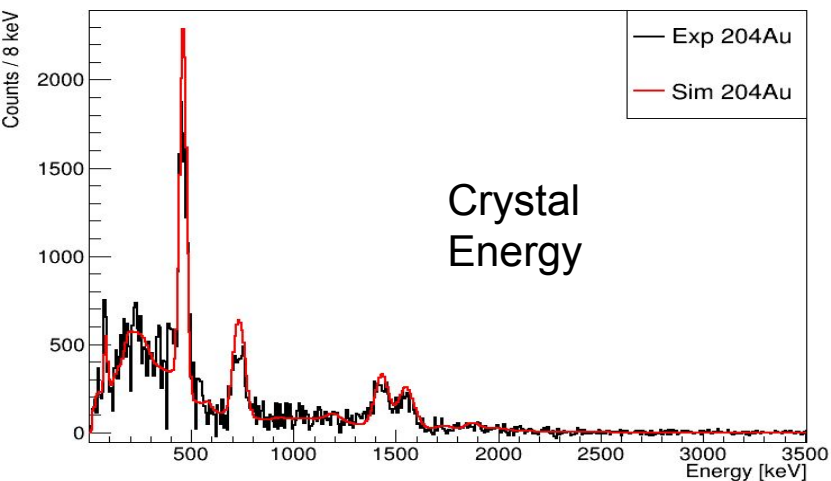


# Preliminary analysis of beta intensity distribution of $^{204}\text{Au}$ and $^{205}\text{Au}$

204Au Sum Energy



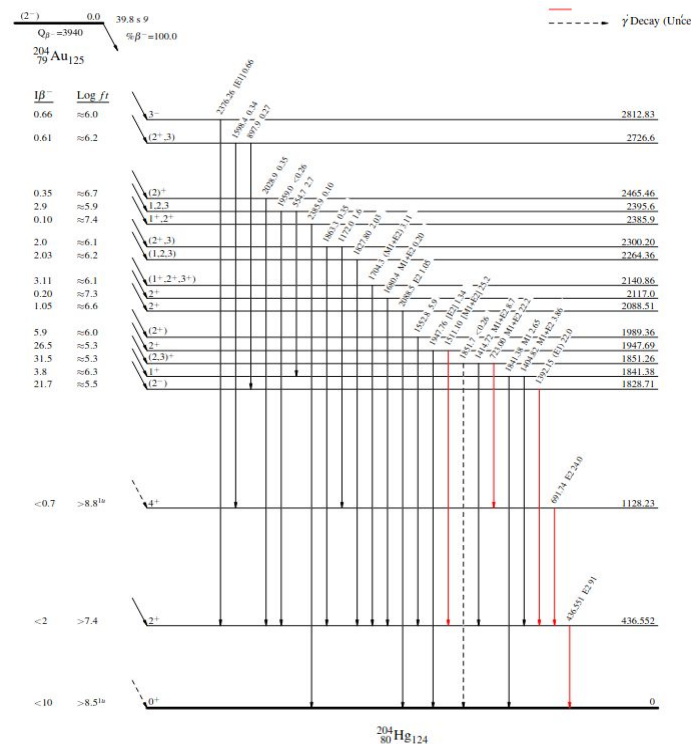
204Au Cry Energy



204Au

Full  
statistics:  
26.8k  
counts

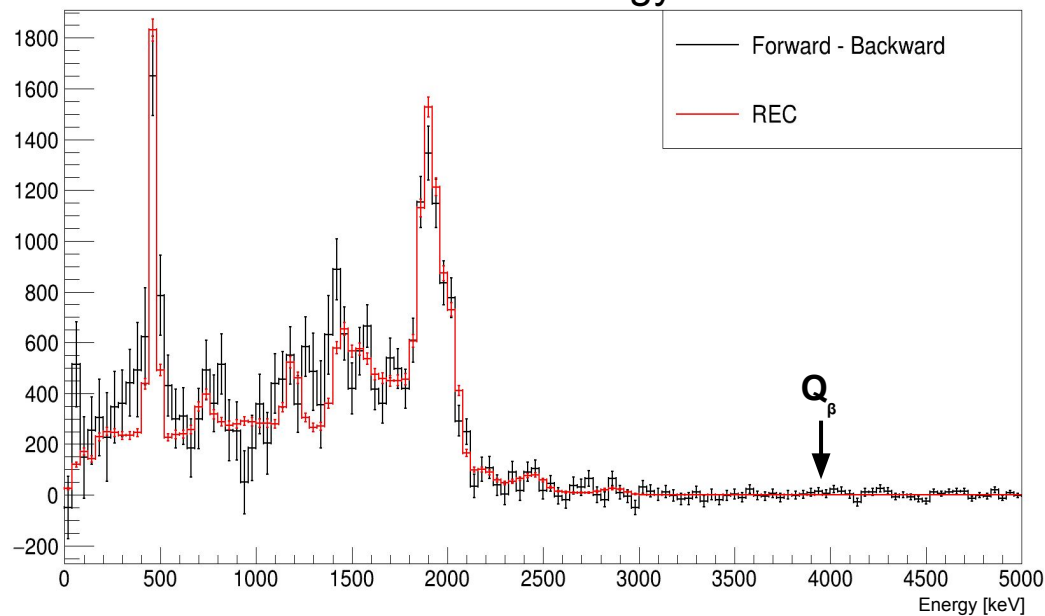
Comparison  
with previous  
knowledge



D.A. Craig and H.W. Taylor 1984 *J. Phys. G: Nucl. Phys.* **10** 1133

Within our sensitivity, no additional  
beta intensity

## 204Au Sum Energy



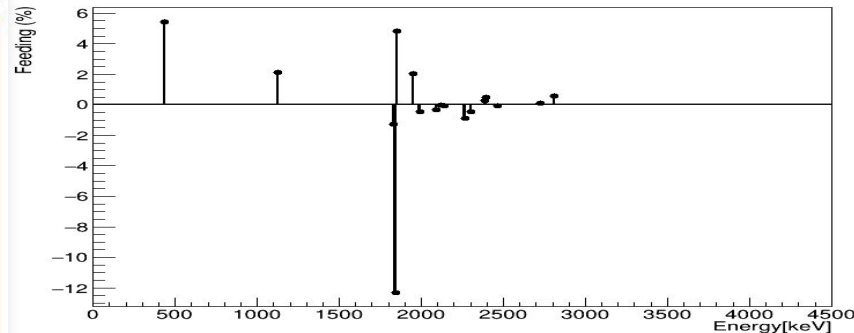
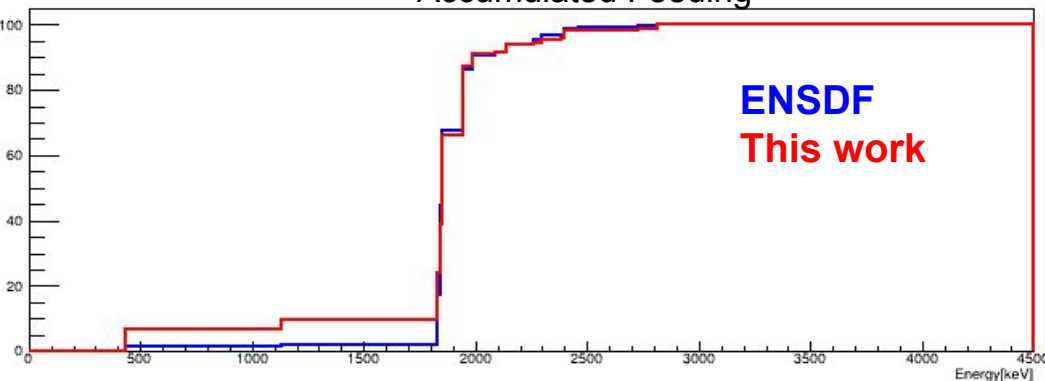
Preliminary beta intensity distribution results for 204Au.

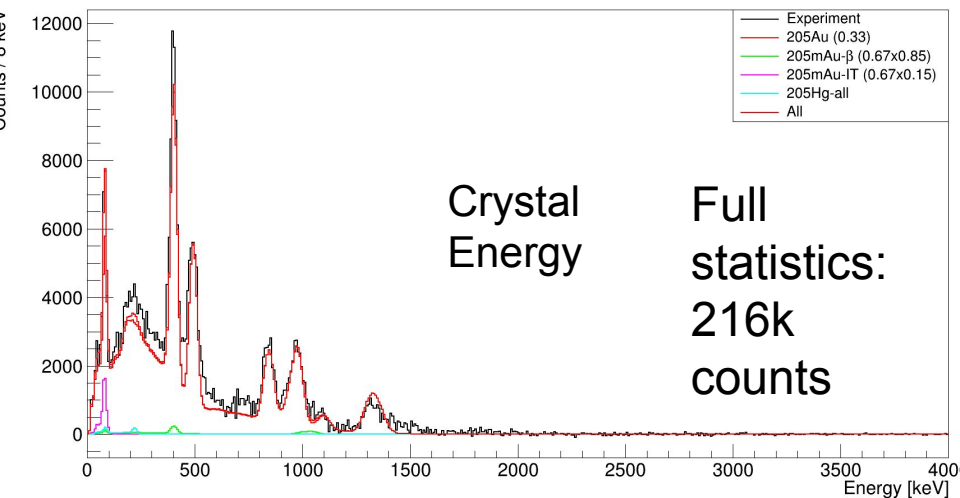
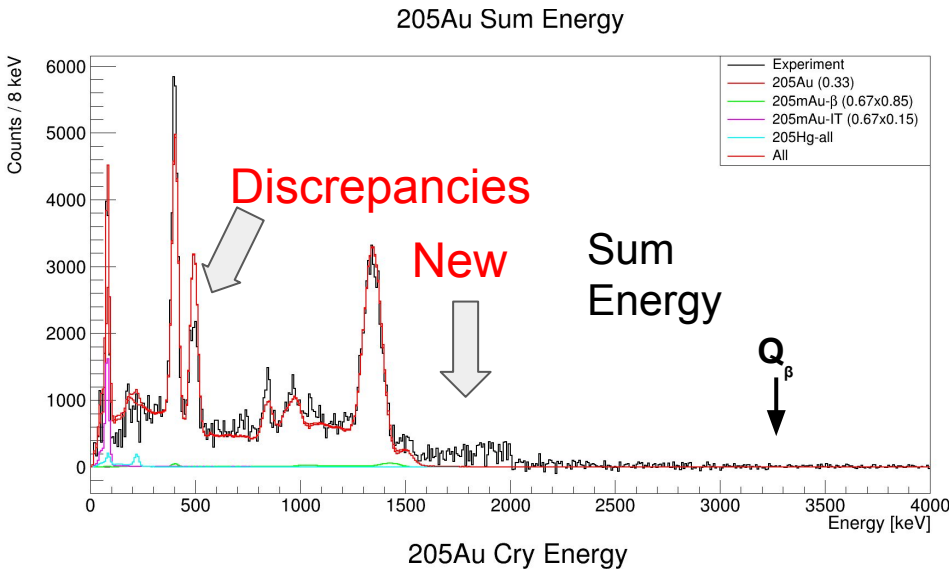
The feeding to g.s. fixed to 0 (Craig et al.)

In good agreement with Craig et al., but some feeding to first and second excited states

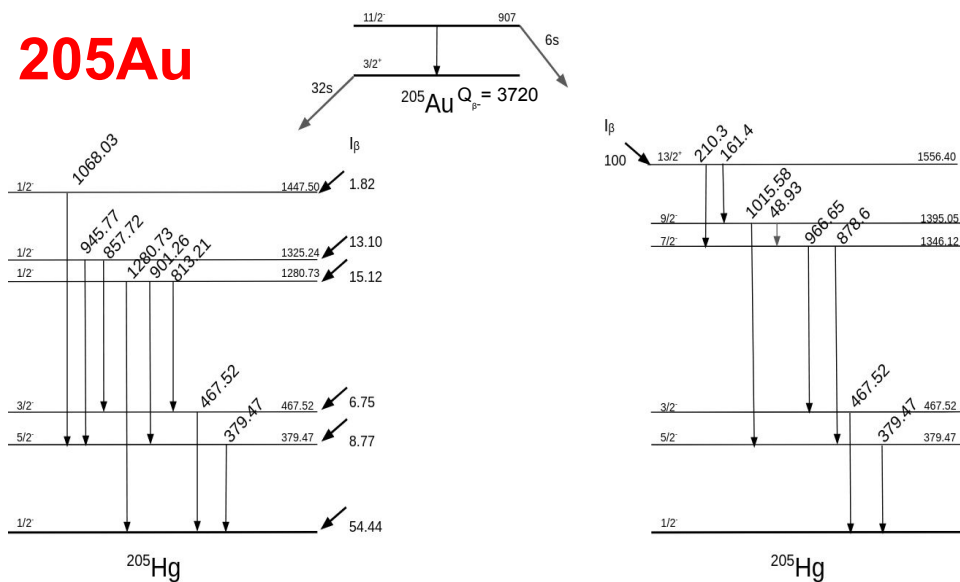
**PRELIMINARY**

## Accumulated Feeding





205Au



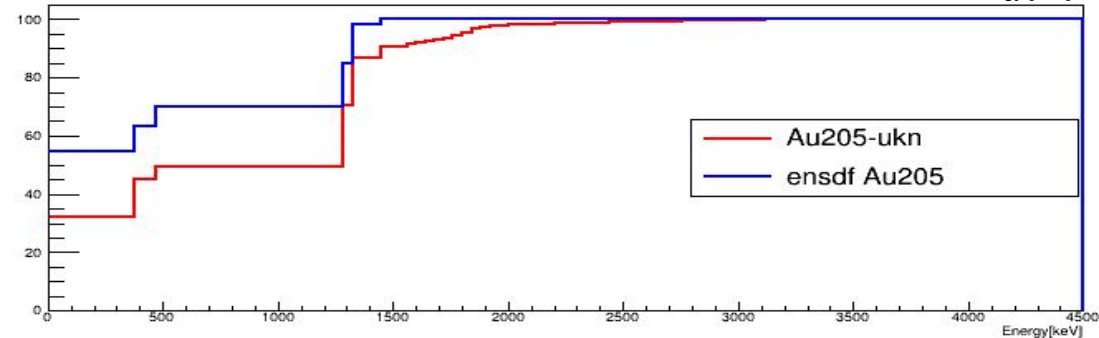
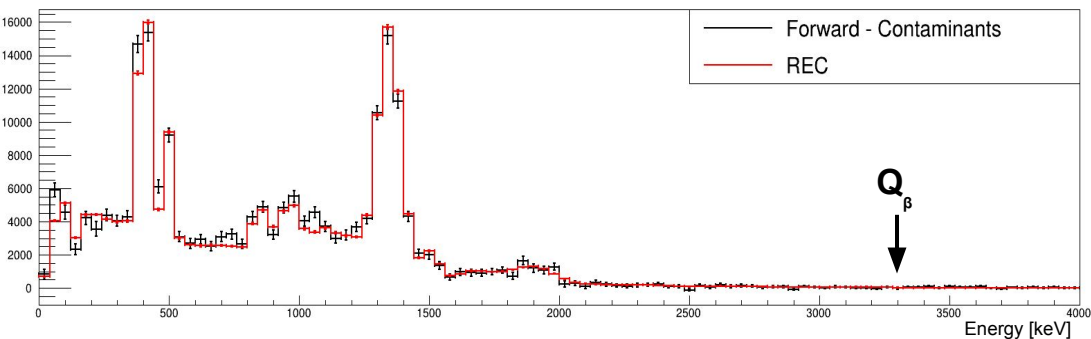
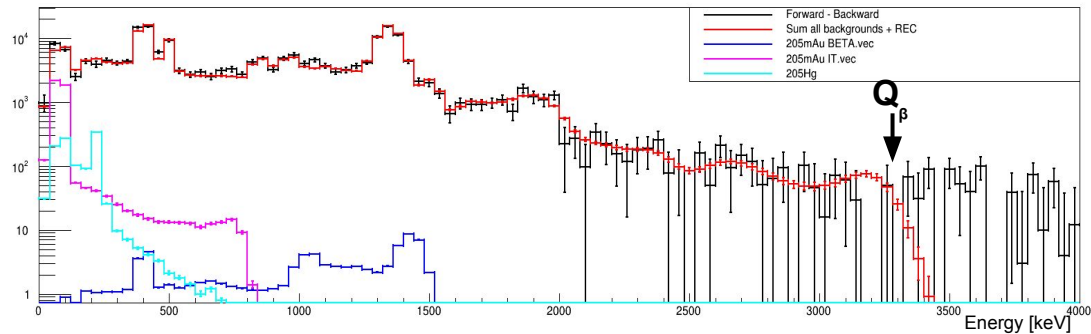
Comparison  
with previous  
knowledge

C. Wennemann et al. Z.Phys. A 347, 185-189 (1994)

Z. Podolyák et al. PLB. 672 (2009) 116-119

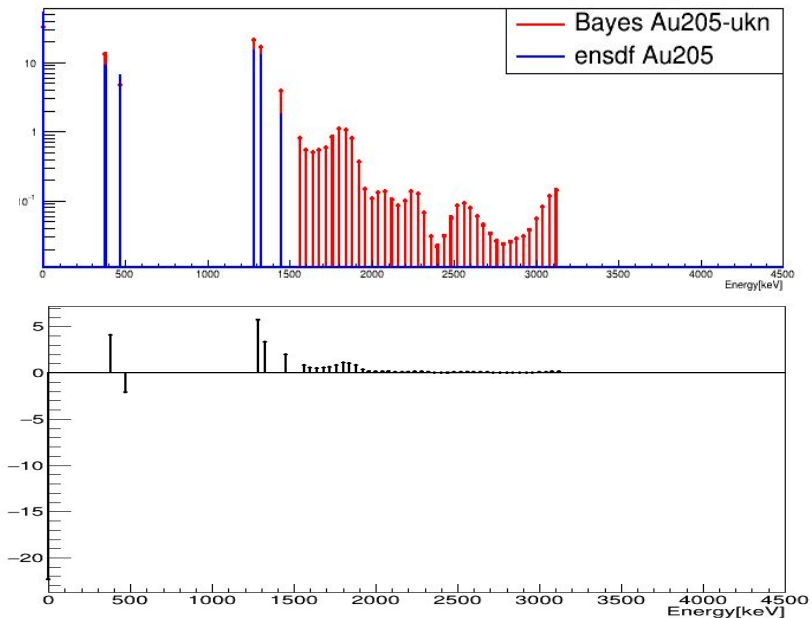
We see additional beta  
intensity





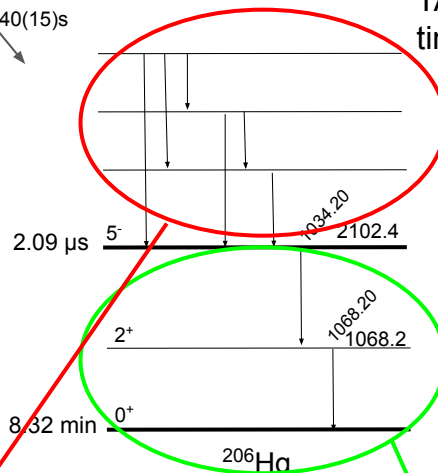
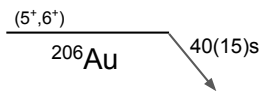
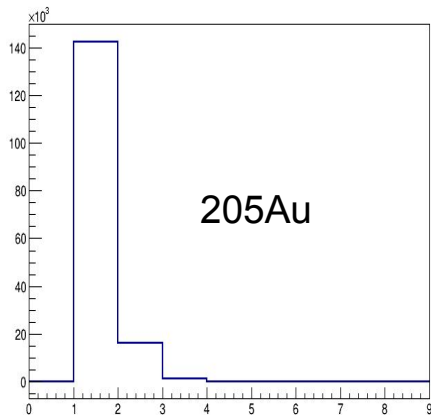
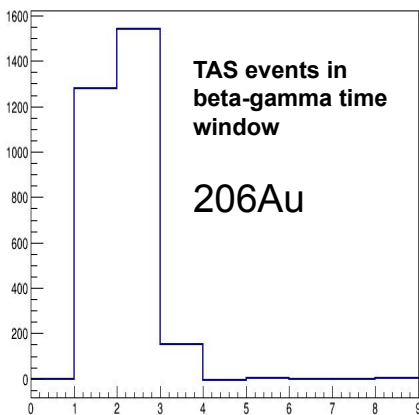
Preliminary beta intensity distribution results for  $^{205}\text{Au}$ .

Additional beta intensity above 1.55 MeV, lower g.s. feeding intensity



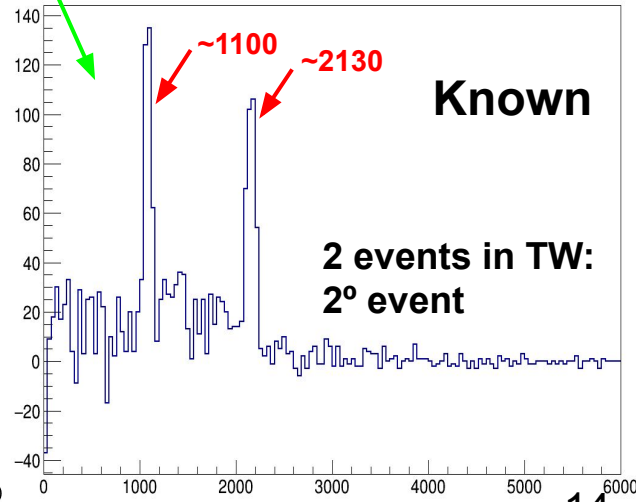
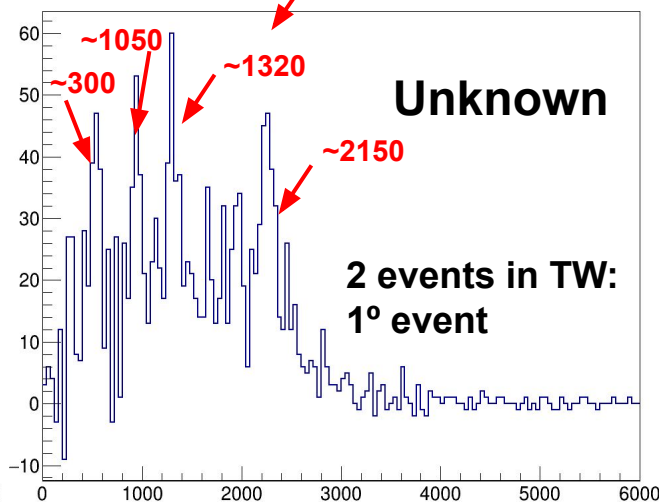
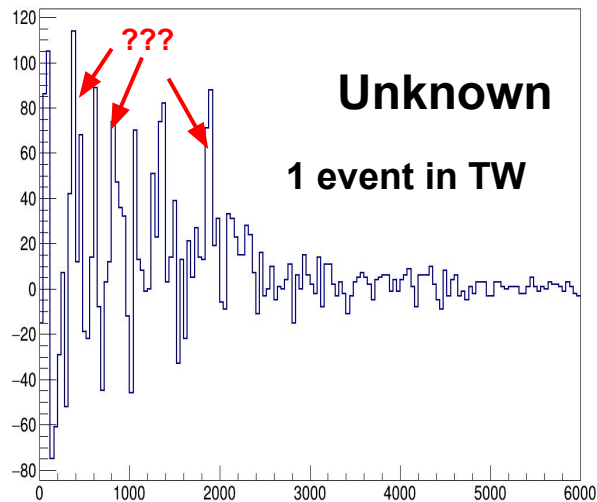
**PRELIMINARY**

# 206Au



TAS events in a 16  $\mu s$  beta-gamma time window.

- Only 1 TAS event: Feeds levels that decays to gs avoiding the isomer level
- 2 TAS events: 1° TAS event feed level that decay into the isomer level, and 2° event is the decay of the isomer level



# Outlook:

- Refining analysis of Au isotopes
- Working on preliminary analysis of  $^{203,204}\text{Pt}$
- In contact with theoreticians to obtain calculated  $\beta$ -strength functions

# S505 Collaboration:

*D.Rodriguez-Garcia<sup>1,\*</sup>, J. L.Tain<sup>1</sup>, A. I.Morales<sup>1</sup>, G.Aggez<sup>2,3</sup>, J.Agramunt<sup>1</sup>, M.Alaqeel<sup>4,5</sup>, B.Alayed<sup>5,6</sup>, H. M.Albers<sup>2</sup>, G.Alcala<sup>1</sup>, A.Algora<sup>1</sup>, A.Alharbi<sup>5</sup>, S.Alhomaidhi<sup>2,7</sup>, F.Amjad<sup>2</sup>, T.Arici<sup>2</sup>, M.Armstrong<sup>8</sup>, M.Bajzek<sup>2</sup>, A.Banerjee<sup>2</sup>, G.Bartram<sup>2</sup>, G.Benzoni<sup>9</sup>, Z.Chen<sup>2</sup>, B.Das<sup>2</sup>, T.Davinson<sup>10</sup>, T.Dickel<sup>2</sup>, I.Dillmann<sup>11</sup>, C.Domingo-Pardo<sup>1</sup>, H.Ekawa<sup>12</sup>, Z.Ge<sup>2</sup>, W.Gelletly<sup>13</sup>, J.Gerl<sup>2</sup>, M.Gorska<sup>2</sup>, E.Haettner<sup>2</sup>, O.Hall<sup>10</sup>, P.Herrmann<sup>2</sup>, C.Hornung<sup>2</sup>, N.Hubbard<sup>2,7</sup>, C.Jones<sup>2</sup>, E.Kazantseva<sup>2</sup>, R.Knoebel<sup>2</sup>, I.Kojouharov<sup>2</sup>, G.Kosir<sup>2,1</sup>, D.Kostyleva<sup>2</sup>, T.Kurtukian-Nieto<sup>15</sup>, N.Kurz<sup>2</sup>, M.Labiche<sup>16</sup>, A.Mccarter<sup>5</sup>, M.Mikolajczuk<sup>2</sup>, A. K.Mistry<sup>2,7</sup>, I.Mukha<sup>2</sup>, E.Nacher<sup>1</sup>, M.Nakagawa<sup>12</sup>, B.S.Nara-Singh<sup>20</sup>, S.Nishimura<sup>12</sup>, S. E. A. Orrigo<sup>1</sup>, P.Papadakis<sup>13</sup>, S.Pietri<sup>2</sup>, W.Plass<sup>2</sup>, Z.Podolyak<sup>13</sup>, M.Polettini<sup>9,17</sup>, R.Prajapat<sup>2,18</sup>, E.Rocco<sup>2</sup>, B.Rubio<sup>1</sup>, E.Sahin<sup>2,7</sup>, M.Satrazani<sup>5</sup>, H.Schaffner<sup>2</sup>, C.Scheidenberger<sup>2</sup>, A.Sharma<sup>19</sup>, Y.K.Tanaka<sup>12</sup>, A.Tolosa-Delgado<sup>21</sup>, M.Vencelj<sup>14</sup>, J.Vesic<sup>14</sup>, P.Vi<sup>12</sup>, J. A.Victoria<sup>1</sup>, H.Weick<sup>2</sup>, K.Wimmer<sup>2</sup>, H. J.Wollersheim<sup>2</sup>, A.Yaneva<sup>2,8</sup>, and J.Zhao<sup>2,22</sup>*

<sup>1</sup>Instituto de Fisica Corpuscular, CSIC-Universidad de Valencia, E-46980 Paterna, Spain

<sup>2</sup>GSF Helmholtzzentrum für Schwerionenforschung, 64291 Darmstadt, Germany

<sup>3</sup>Institute of Graduate Studies in Sciences, Istanbul University, 34452 Istanbul, Turkey

<sup>4</sup>Physics Department, Imam Mohammad Ibn Saud Islamic University (IMISU), P.O. Box 90950, Riyadh, 11623, Saudi Arabia

<sup>5</sup>Department of Physics, University of Liverpool, Liverpool, L69 7ZE, UK

<sup>6</sup>Department of Physics, ArRass College of Sciences and Art, Qassim University, Saudi Arabia

<sup>7</sup>Institut für Kernphysik, Technische Universität Darmstadt, D-64289 Darmstadt, Germany

<sup>8</sup>IKP, University of Cologne, D-50937 Cologne, Germany

<sup>9</sup>INFN Sezione di Milano, I-20133 Milano, Italy

<sup>10</sup>School of Physics and Astronomy, University of Edinburgh, Edinburgh, EH9 3FD, UK

<sup>11</sup>TRIUMF, 4004 Wesbrook Mall, Vancouver, BC V6T 2A3, Canada

<sup>12</sup>RIKEN Cluster for Pioneering Research, RIKEN, Saitama 351-0198 Japan

<sup>13</sup>Department of Physics, University of Surrey, Guildford, GU2 7XH, UK

<sup>14</sup>Faculty of mathematics and physics of the University of Ljubljana, SI-1000 Ljubljana, Slovenia

<sup>15</sup>Instituto de Estructura de la Materia, CSIC, E-28006, Madrid, Spain

<sup>16</sup>Science and Technology Facilities Council, Daresbury Laboratory, Daresbury, WA4 4AD, UK

<sup>17</sup>Dipartimento di Fisica, Università degli Studi di Milano - Milano, Italy

<sup>18</sup>Department of Physics, Indian Institute of Technology Roorkee, Roorkee-247667, Uttarakhand, India

<sup>19</sup>Department of Physics, Indian Institute of Technology Ropar, Rupnagar - 140001, India

<sup>20</sup>Paisley, University of the West of Scotland, High St, Paisley PA1 2BE, UK

<sup>21</sup>Faculty of Mathematics and Science, University of Jyväskylä, FI-40014 Jyväskylä, Finlandia

<sup>22</sup>School of Physics, Peking University, Beijing 100871, China

*Thank you very much for your attention*

Thanks to:

PID2019-104714GB-C21, PID2022-138297NB-C21,

PROMETEO/2019/007, CIPROM/2022/9, CISEJI/2022/25



# XVII CPAN DAYS

Centro Nacional de Física de Partículas,  
Astropartículas y Nuclear.

## Questions?

19-21 Noviembre 2025  
Valencia



Extra Slides

# FRS: PID using $B\rho$ - ToF - $\Delta E$ method

- Ion identification combining information for each ion, on magnetic rigidity  $B\rho$  in the dipoles, the time-of-flight between detectors in the spectrometer flight path, and the energy loss in suitable “thin” detectors

## A/Q calculations:

$$\beta = \frac{L}{c \cdot \text{ToF}}$$

$$B\rho = B\rho_0 \cdot \left( 1 + \frac{x_4}{D_x^{24}} + \frac{x_2}{D_x^{02}} \right)$$

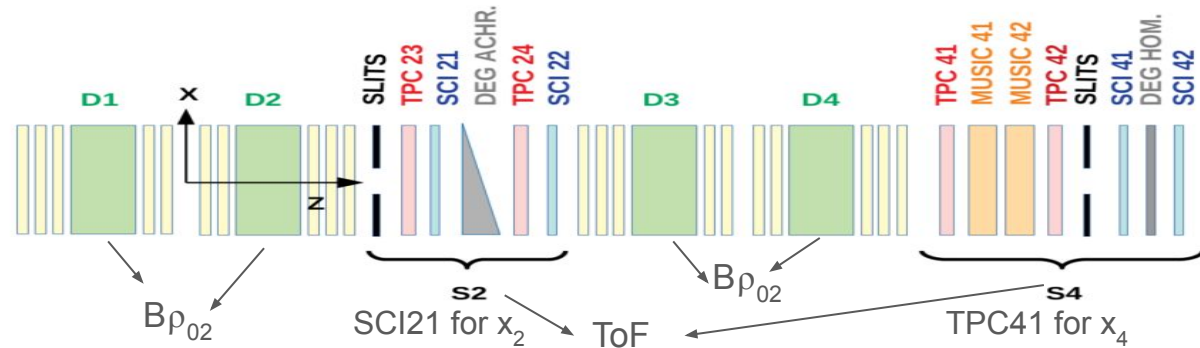
$$\frac{A}{Q} = \frac{e}{u \cdot c} \frac{B\rho}{\beta \gamma}$$

## Z identification: Using the MUSIC

$$\Delta E_{\text{MUSIC}} \propto Z^2$$

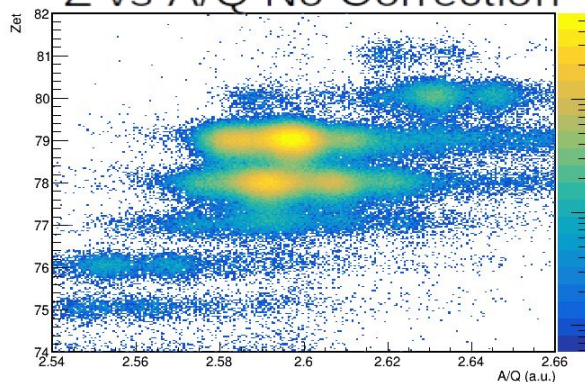
## Loss energy in the degrader:

$$\Delta E_{\text{deg}} = c \cdot e \cdot Q \cdot \left( \frac{B\rho_{02}}{\beta_{02}} - \frac{B\rho_{24}}{\beta_{24}} \right)$$

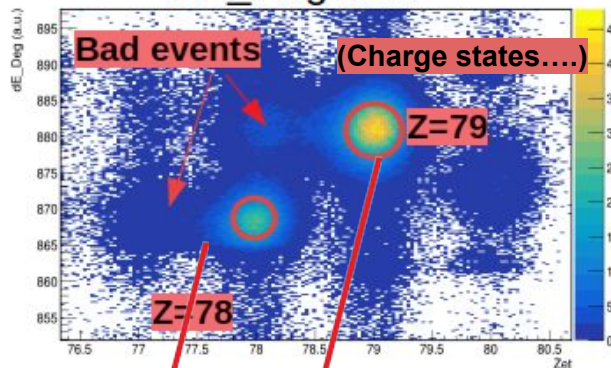


## FRS: PID corrections A/Q and $\Delta E_{\text{deg}}$

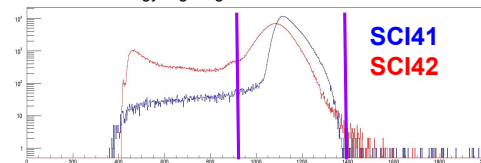
Z vs A/Q No Correction



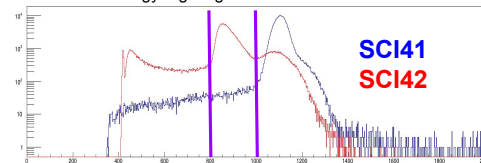
$\Delta E_{\text{Deg}}$  vs Z



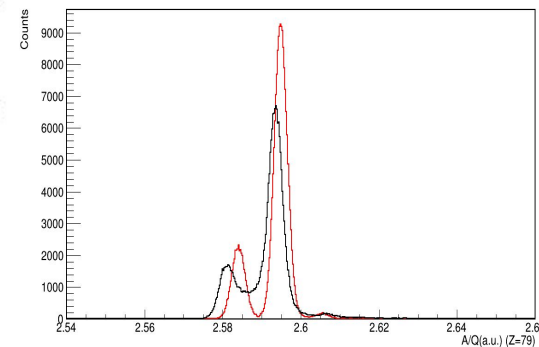
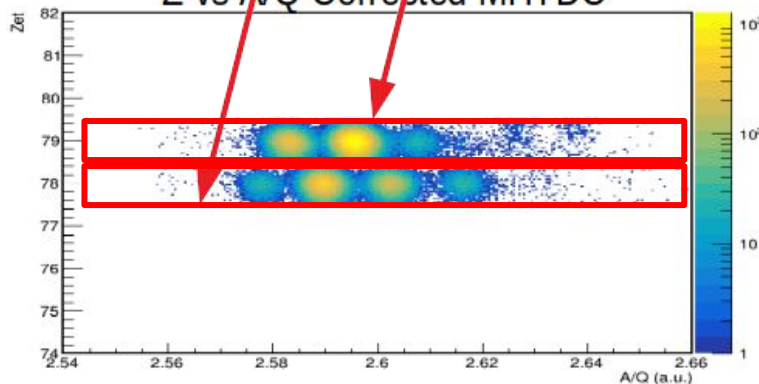
Energy Signal gated in Au



Energy Signal gated in Pt

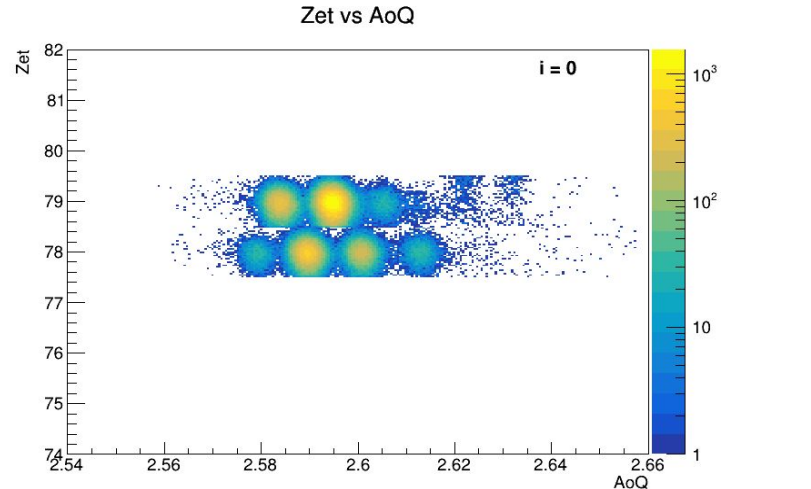
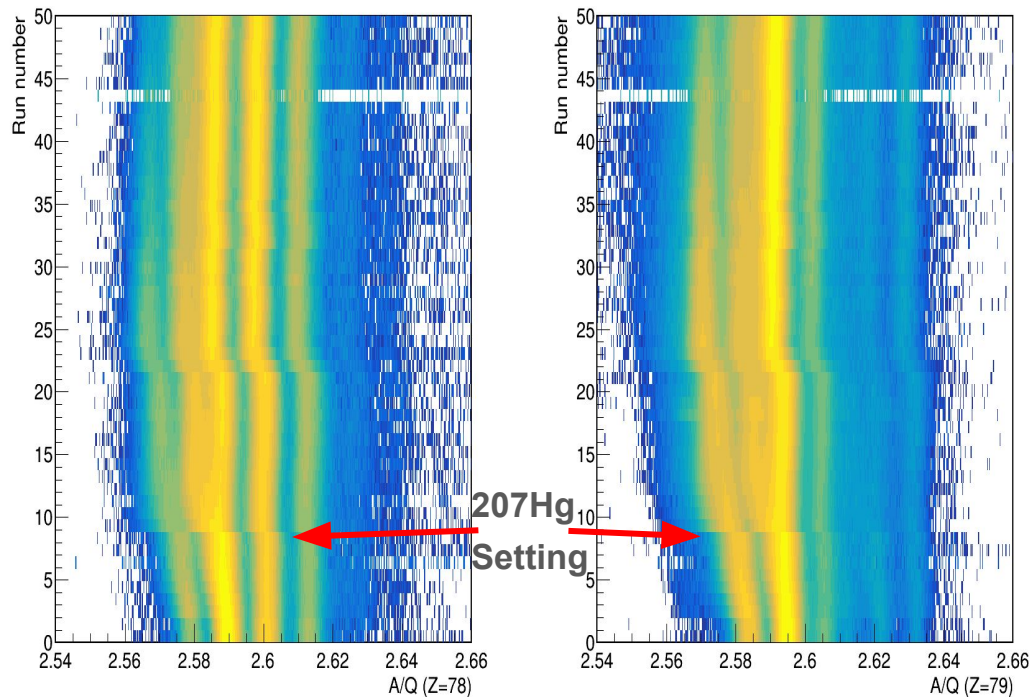


Z vs A/Q Corrected MHTDC

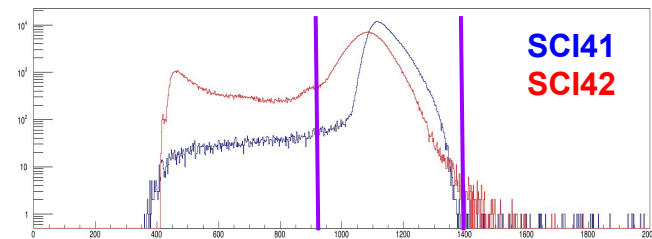




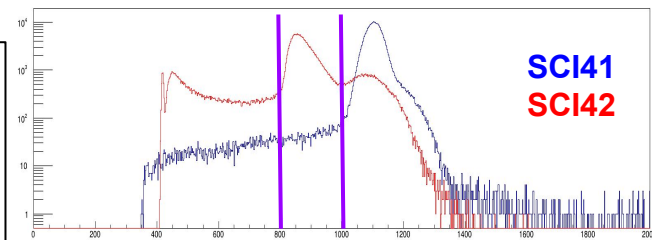
# FRS: A/Q drift along experiment



Energy Signal gated in Au



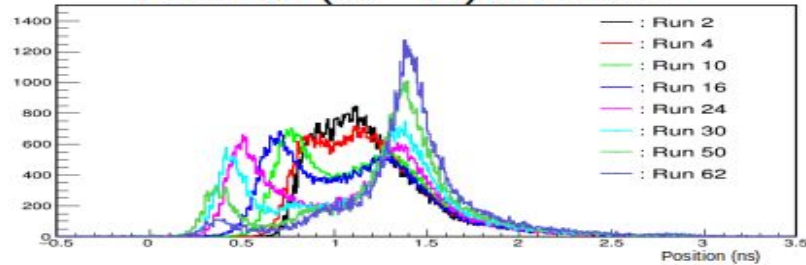
Energy Signal gated in Pt



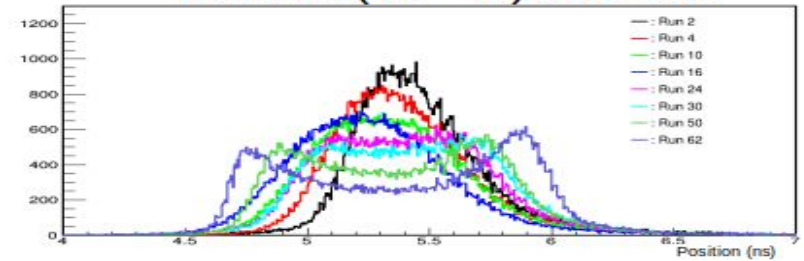
- Drift along the experiment due to damage in SCI21 and SCI22
- We added a energy condition in SCI42 to minimize the effect of secondary reactions in last degrader

# FRS: MHTDC: S2 SCI Position and Energy

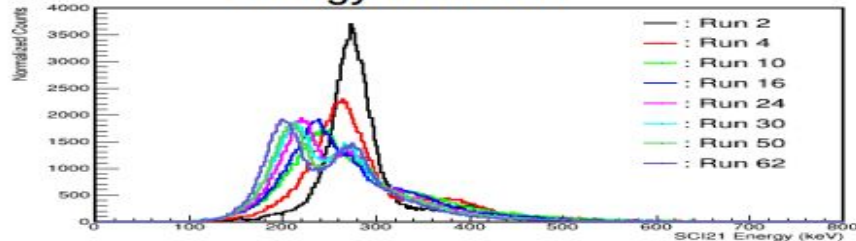
Position ( $T_L - T_R$ ) in SCI21



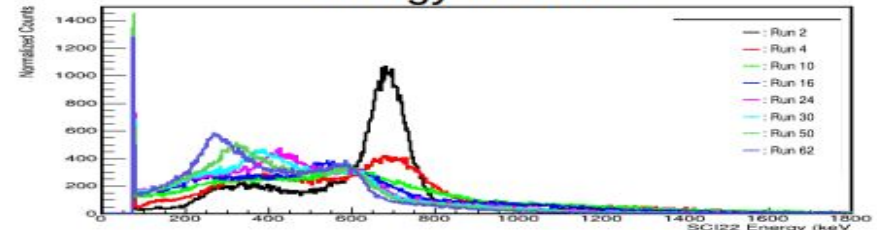
Position ( $T_L - T_R$ ) in SCI22



Energy in SCI21



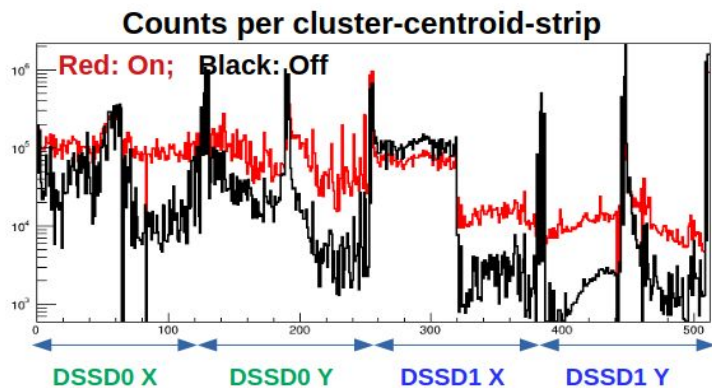
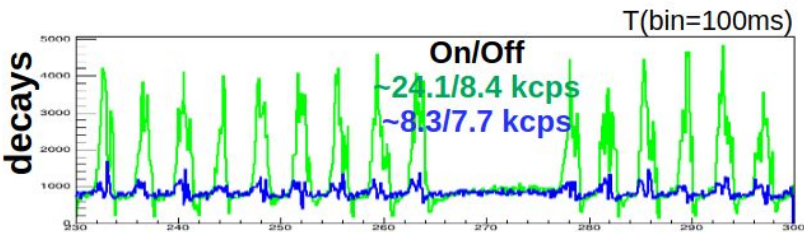
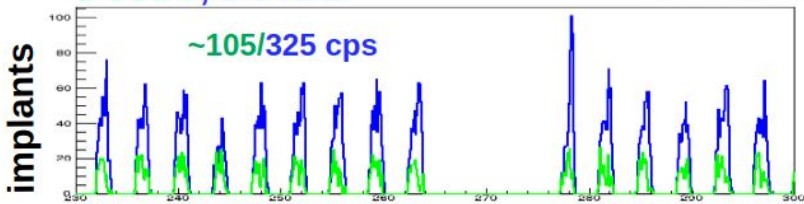
Energy in SCI22



Position distortion: gradual development of side wings  
Energy spectrum distortion and light reduction  
Consequence of detector damage

DSSD0, DSSD1

(FRS Run 16)

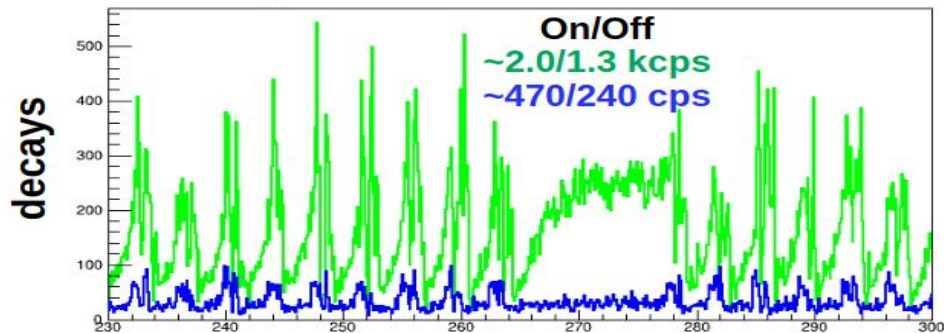


Strange effect on the rate during spill and first few seconds after spill for DSSD0

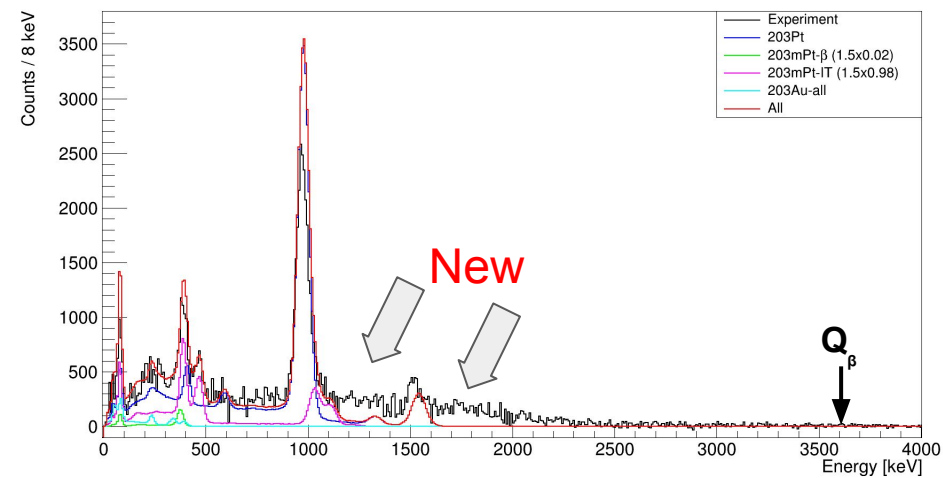
- Huge rate in decay branch, in particular for DSSD0 during spill
- Large noise on most FEE64s
- Noisy strips can be identified
- Large number of strips firing in one decay event

1. Increase threshold (currently 150keV)
2. Limit event strip multiplicity ( $n_x, n_y < 6$ )
3. Only analyze off-spill data
4.  $\gamma$ -gating

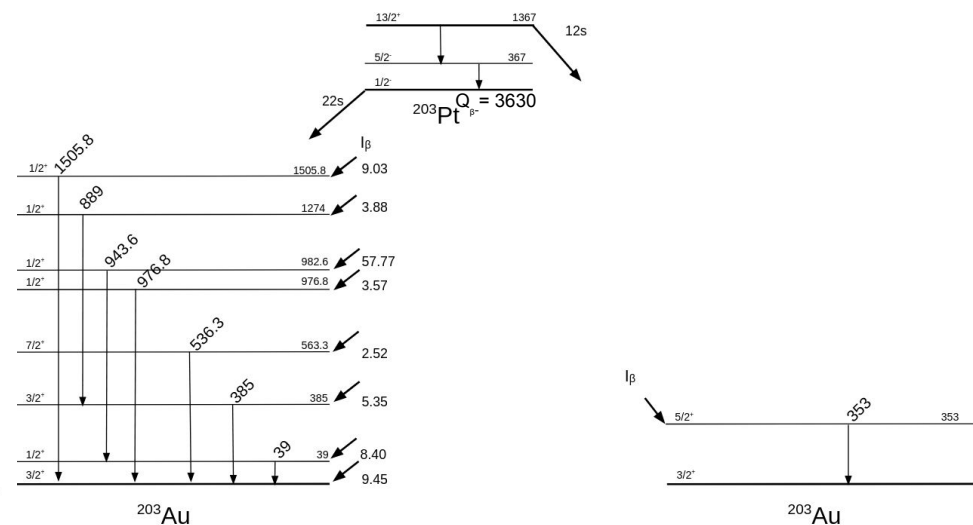
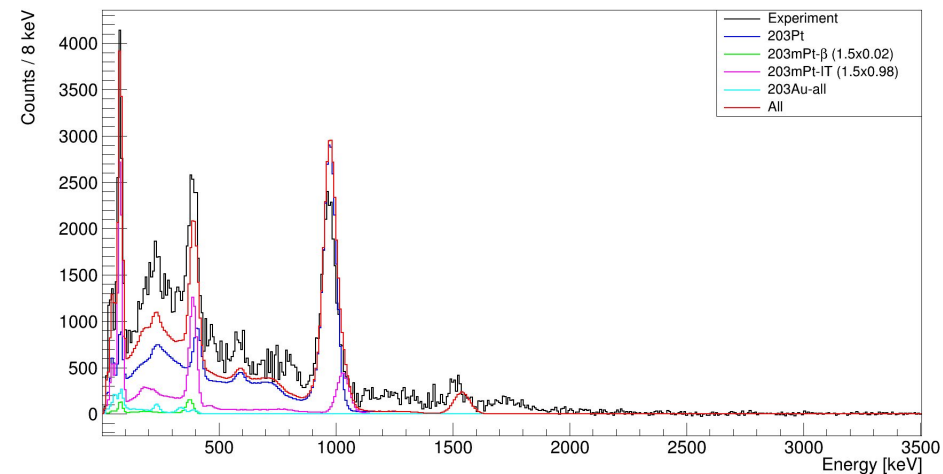
DSSD0, DSSD1



203Pt Sum Energy



203Pt Cry Energy

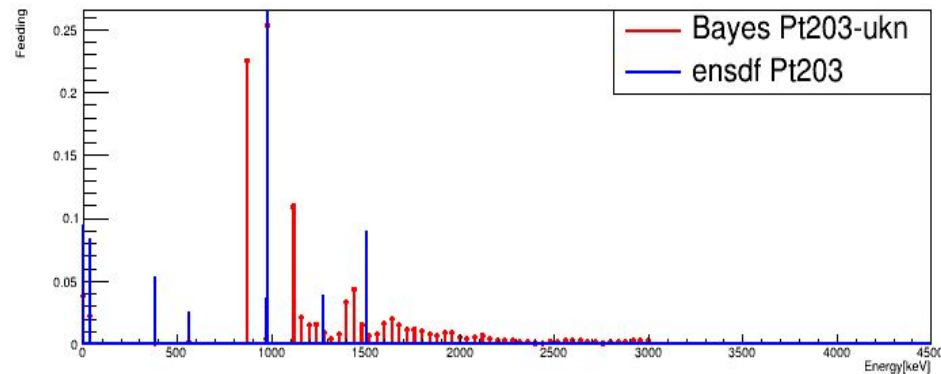
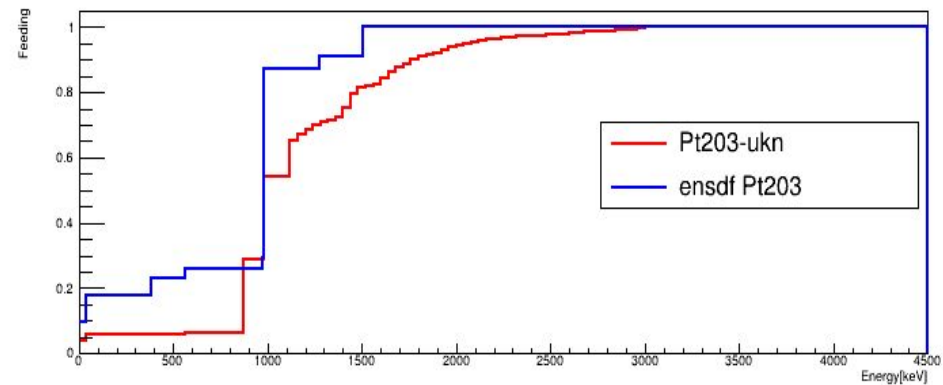
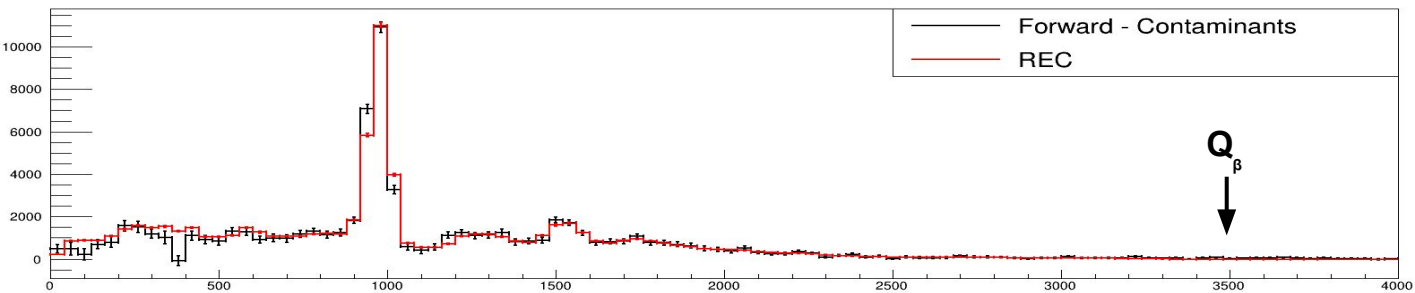
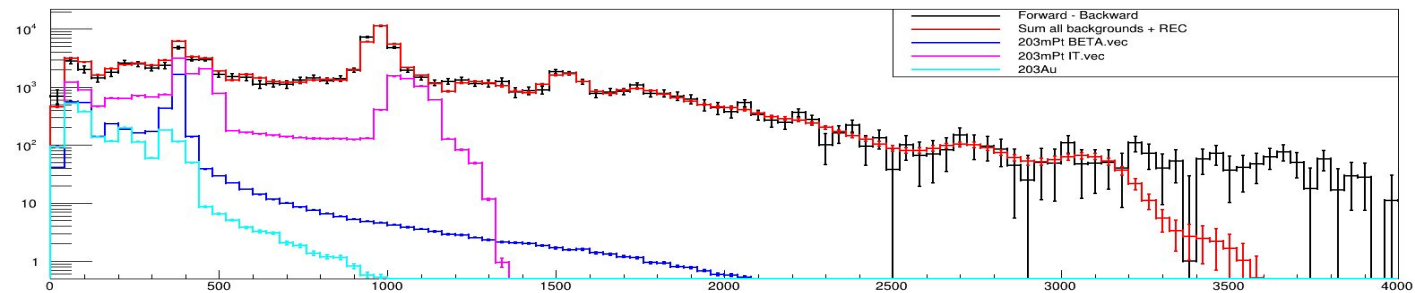


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Plot with all the experiment statistics

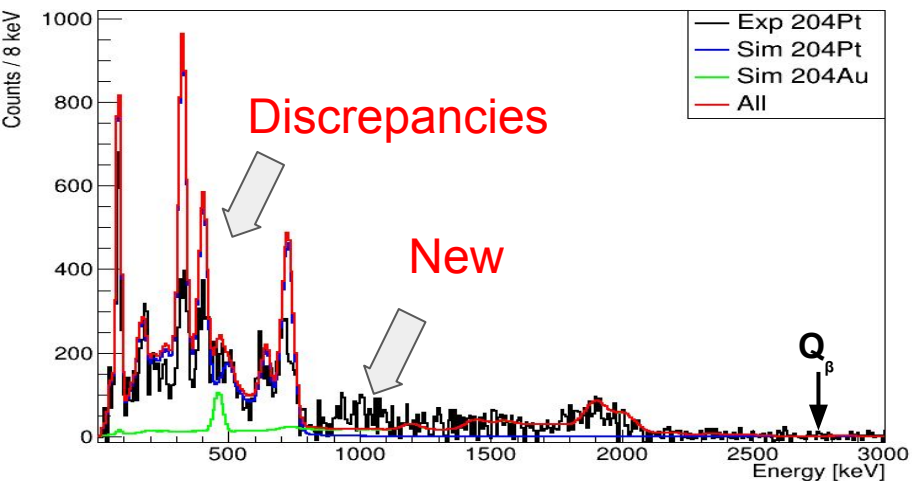
Total statistics in Sum spectrum: 98.7k

**203Pt**

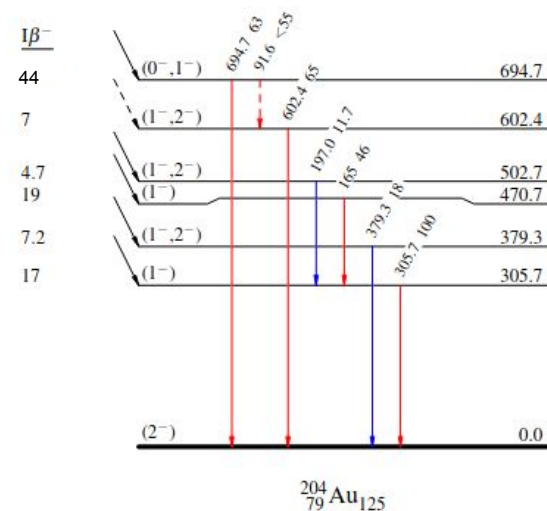
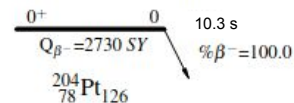
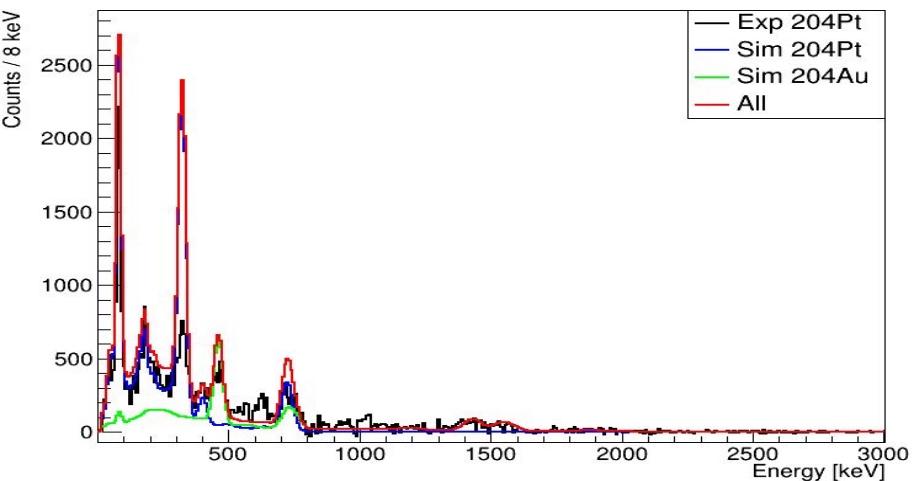




204Pt Sum Energy



204Pt Cry Energy

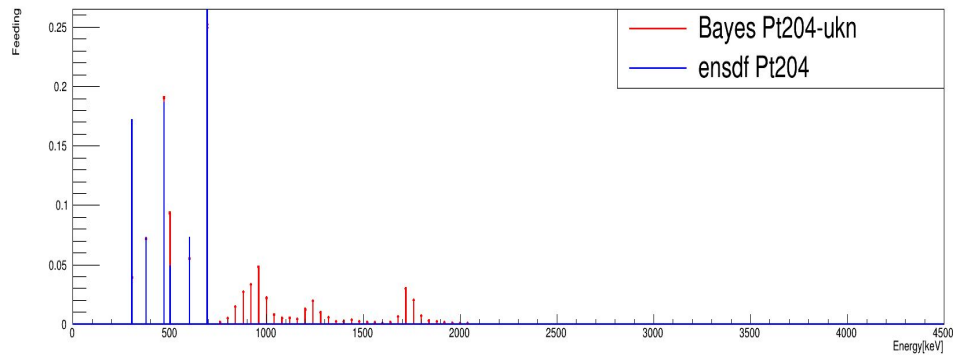
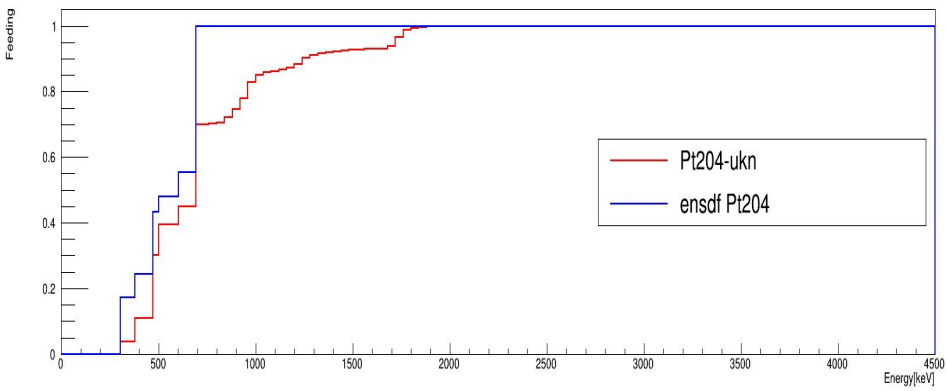
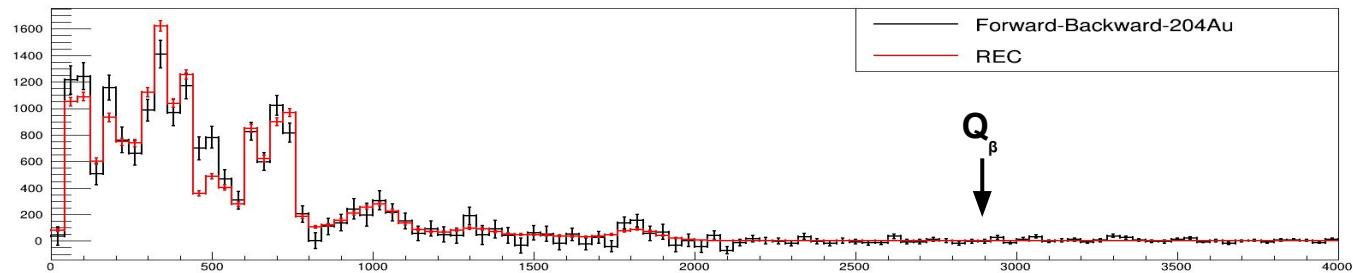
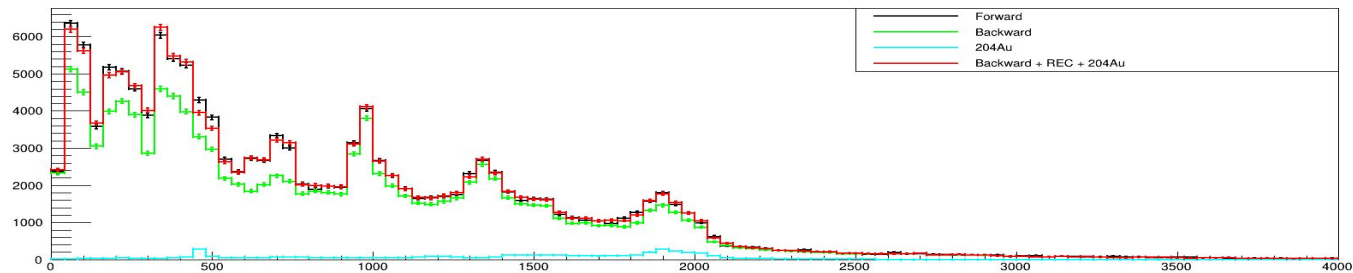


- A.I. Morales et al. Phys. Rev. C **88**, 014319 (2013)

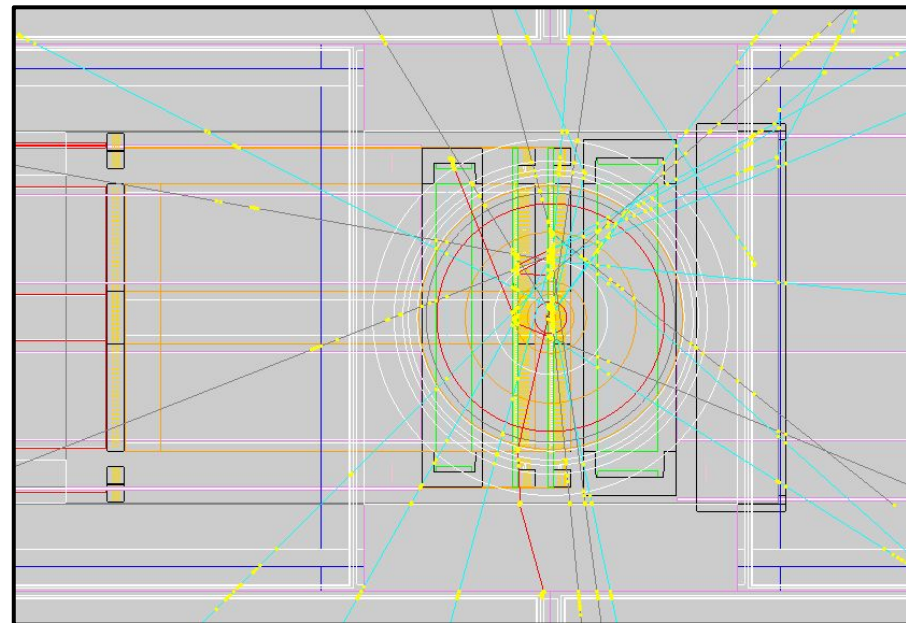
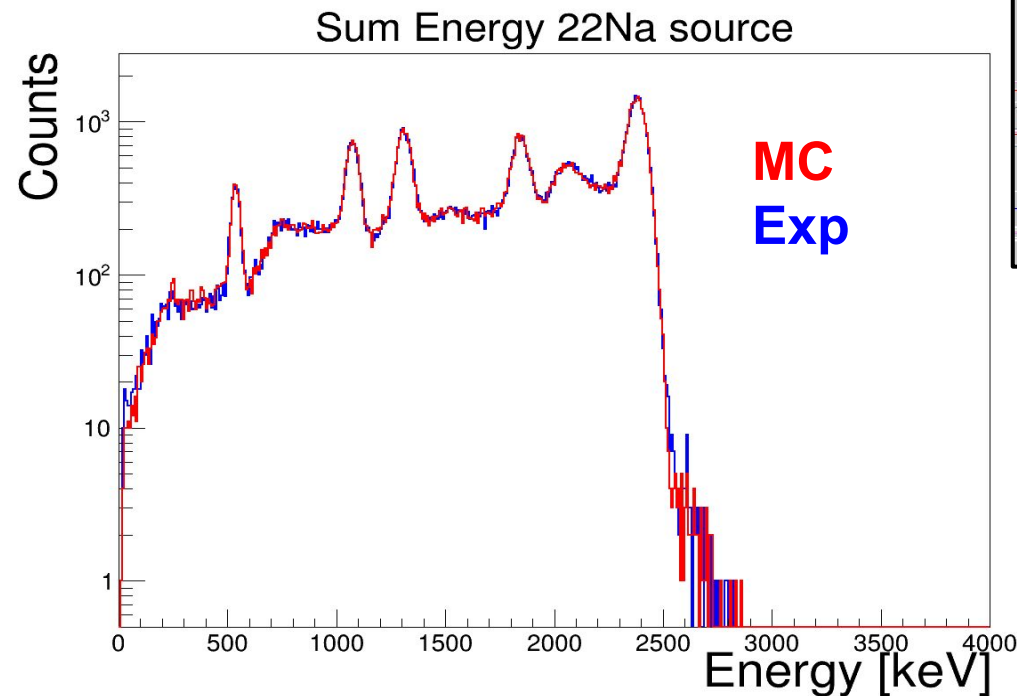
Plot with all the experiment statistics

Total statistics in Sum spectrum: 23k

**204Pt**

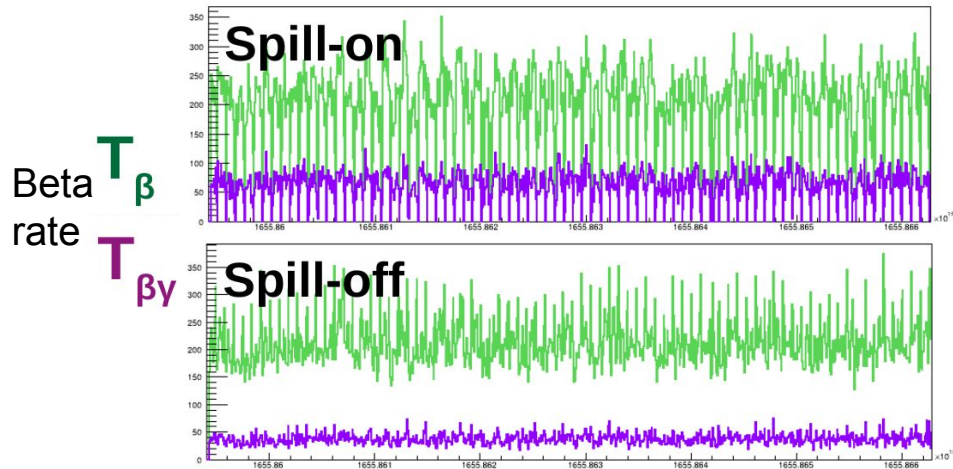
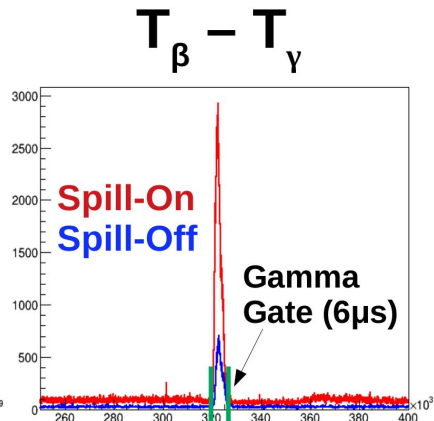
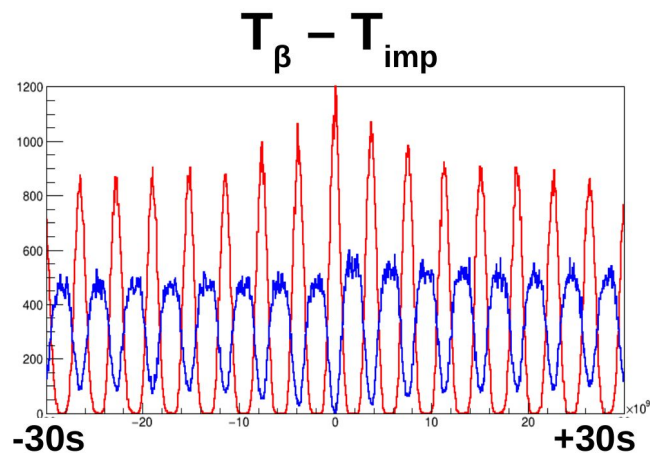


Benchmark of MC simulations:  
different geometries (DTAS, DTAS+LOAX,  
DTAS+AIDA, DTAS+AIDA+LOAX) with  
sources ( $^{137}\text{Cs}$ ,  $^{60}\text{Co}$ ,  $^{22}\text{Na}$ ,  $^{24}\text{Na}$ )



New tree with DTAS event  
information: TS, number of crystals  
and crystal energy (time event  
window: 250ns)

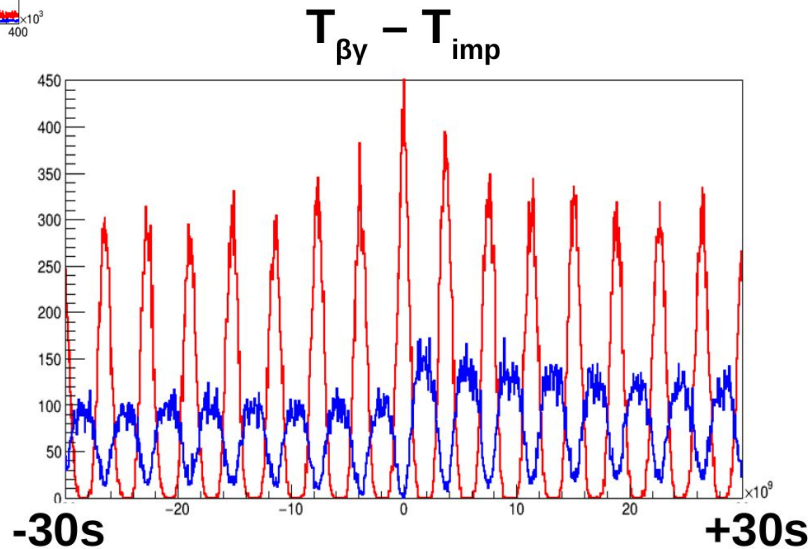
## Low intensity run



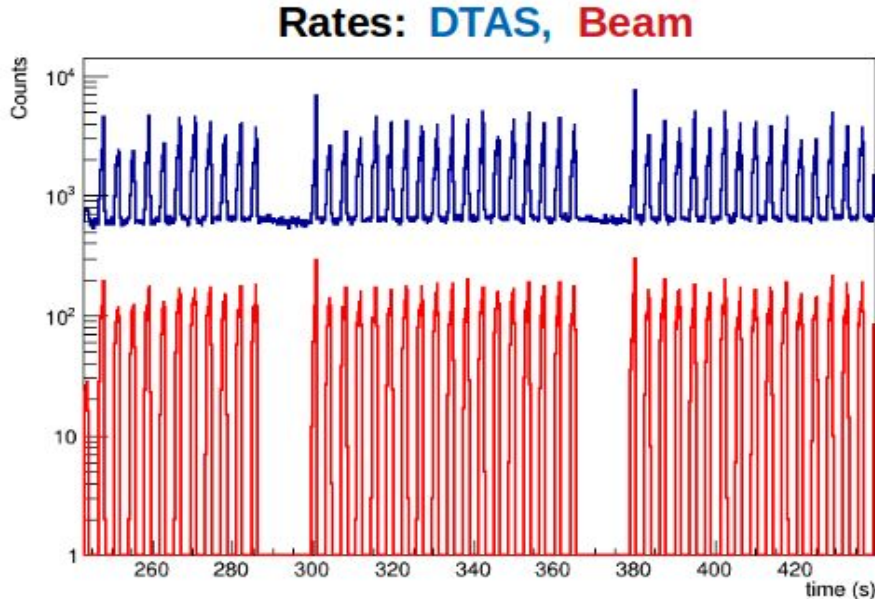
## MERGER:

Combines DTAS+AIDA+FRS data  
(with presorted time-correlations)  
Allows the extraction of analyzable data

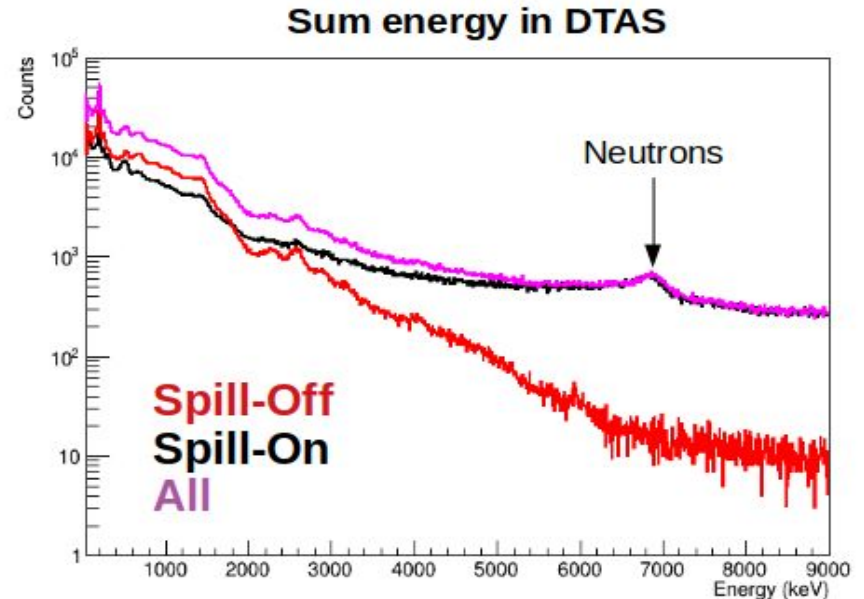
True  $\beta\gamma/\beta$ -implant correlations are clearly seen in the off-spill data (not so clear on-spill)



# DTAS: Rates and comparison: Spill-Off, Spill-On



Beam (spill) → ~1000 pps  
DTAS spill-off → ~8000 cps  
DTAS spill-on → ~42000 cps



- Huge rate on spill, too high rate off spill
- Particles (neutrons, ...) and EM radiation coming with the beam disturbing the spectra
- Possibly only spill-off data is useful