

The SuperCDMS Experiment at SNOLAB

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On behalf of the SuperCDMS collaboration

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THE UNIVERSITY
OF BRITISH COLUMBIA

SuperCDMS collaboration



@SuperCDMS

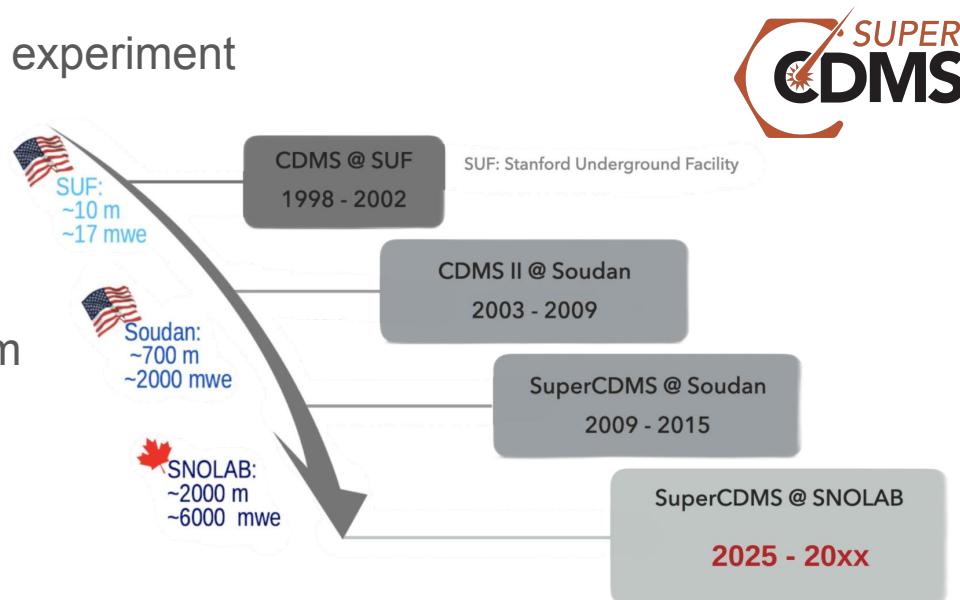
<https://www.snlab.ca/experiment/supercdms/>

SuperCDMS experiment

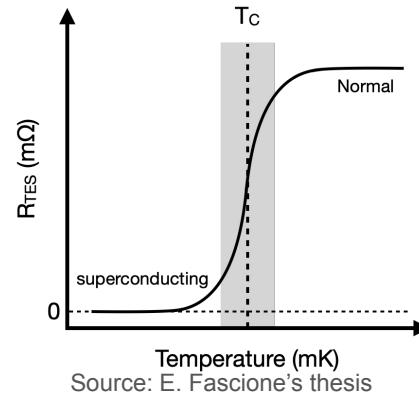
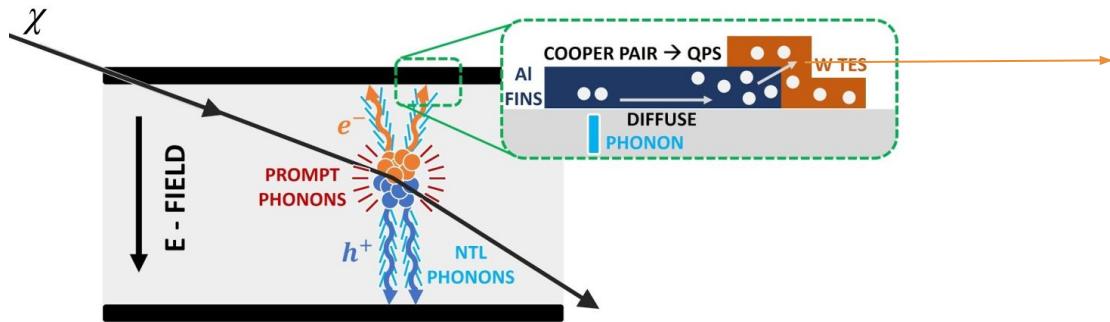
- Super Cryogenic Dark Matter Search experiment
- Use semiconductor crystal detectors
Si and Ge
- Operate at cryogenic temperature
- Measure energy deposition in the form of phonons and ionization
- Search for sub-GeV DM
- Recent publications:

- Light dark matter constraints from SuperCDMS HVeV detectors operated underground with an anticoincidence event selection
- G4CMP: Condensed matter physics simulation using the GEANT4 toolkit
- Investigating the sources of low-energy events in a SuperCDMS-HVeV detector
- Constraints on dark photons and axionlike particles from the SuperCDMS Soudan experiment

Full publication list at: <https://supercdms.slac.stanford.edu/science-results/publications>



Detection principle



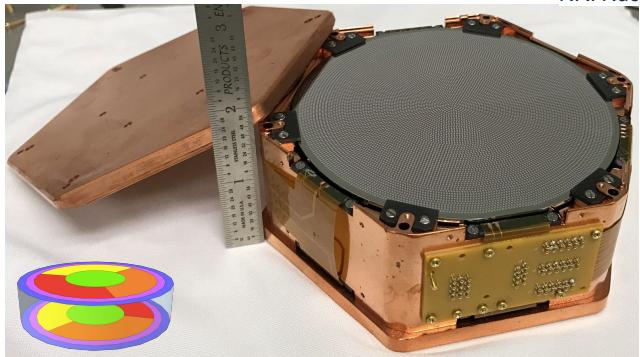
- Particle interactions produce electron-hole (eh) pairs and prompt phonons
- Voltage bias across detector allows eh pairs to drift, creating Neganov-Trofimov-Luke (NTL) phonons $E_{\text{total}} = E_{\text{recoil}} + (N_{\text{eh}} \cdot e \cdot V_{\text{bias}})$
- Energy from prompt and NTL phonons \rightarrow phonon sensor (QETs)
- Drifting eh pairs \rightarrow charge sensor (HEMTs)

QET: Quasiparticle Trap Assisted Electrothermal Feedback Transition Edge Sensor
HEMT: High Electron Mobility Transistor

SuperCDMS detectors

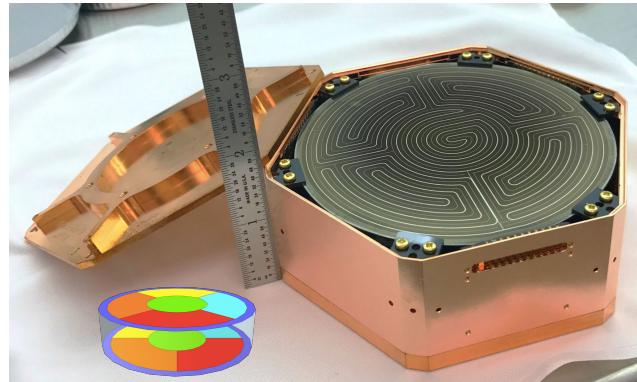
High Voltage (HV)

- 12 phonon channels
- Large bias voltage (~ 100 V)
- NTL amplification
 - low threshold \rightarrow low mass DM
 - no ER/NR discrimination
- Energy resolution $O(10$ eV)



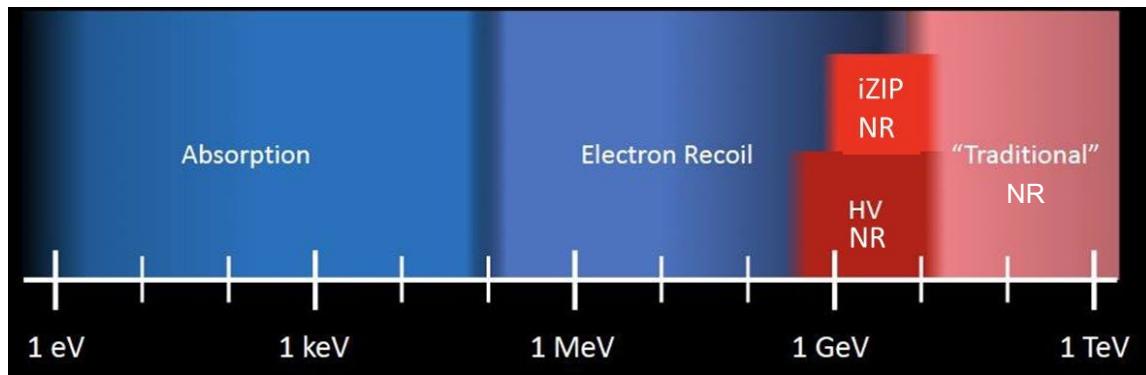
Interleaved Z-sensitive ionization phonon (iZIP)

- 12 phonon channels + 4 charge channels
- Small bias voltage (< 10 V)
- Measuring ionization + phonon
 - ER/NR discrimination
 - better background rejection



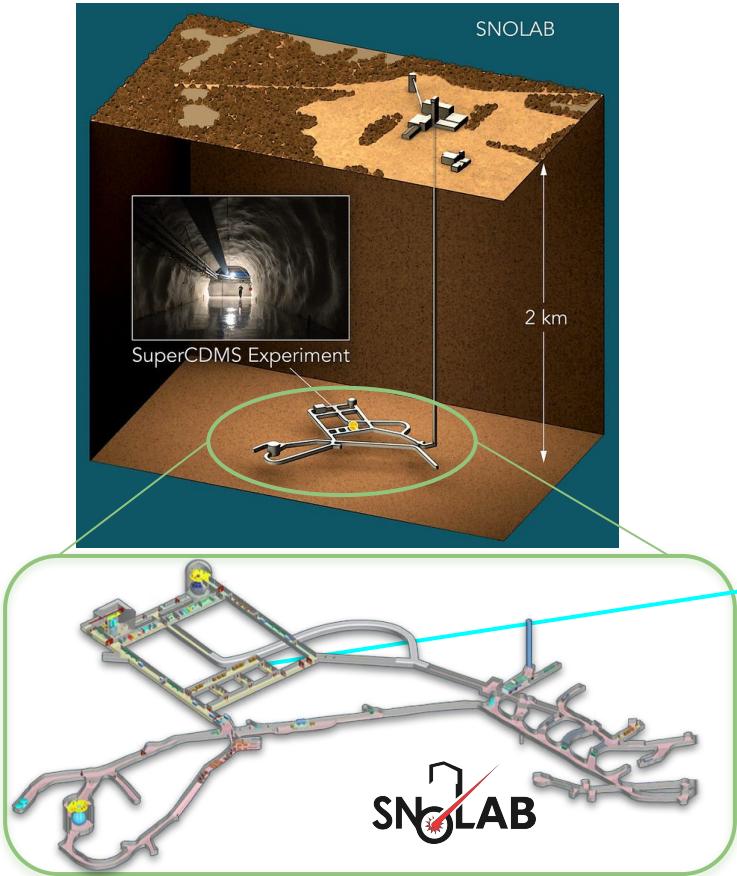
SuperCDMS science search

Nuclear recoil	iZIP, ER background free	$>5 \text{ GeV}/c^2$
	iZIP, limited ER/NR discrimination	$>1 \text{ GeV}/c^2$
	HV, no discrimination	$\sim 0.3 - 10 \text{ GeV}/c^2$
Electron recoil	HV, no discrimination	$\sim 0.5 \text{ MeV}/c^2 - 10 \text{ GeV}/c^2$
Absorption (dark photons, ALPs)	HV, no discrimination	$\sim 1 \text{ eV}/c^2 - 500 \text{ keV}/c^2$ (peak search)

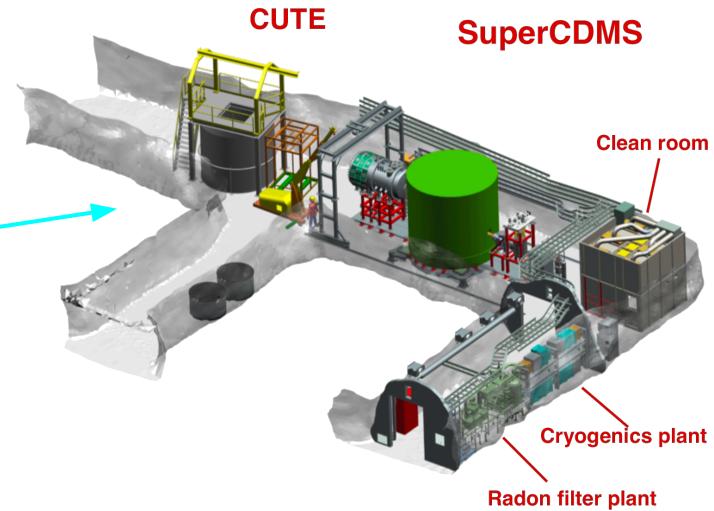


Refer to doi:10.1139/cjp-2024-0127 (Snowmass) for sensitivity forecast

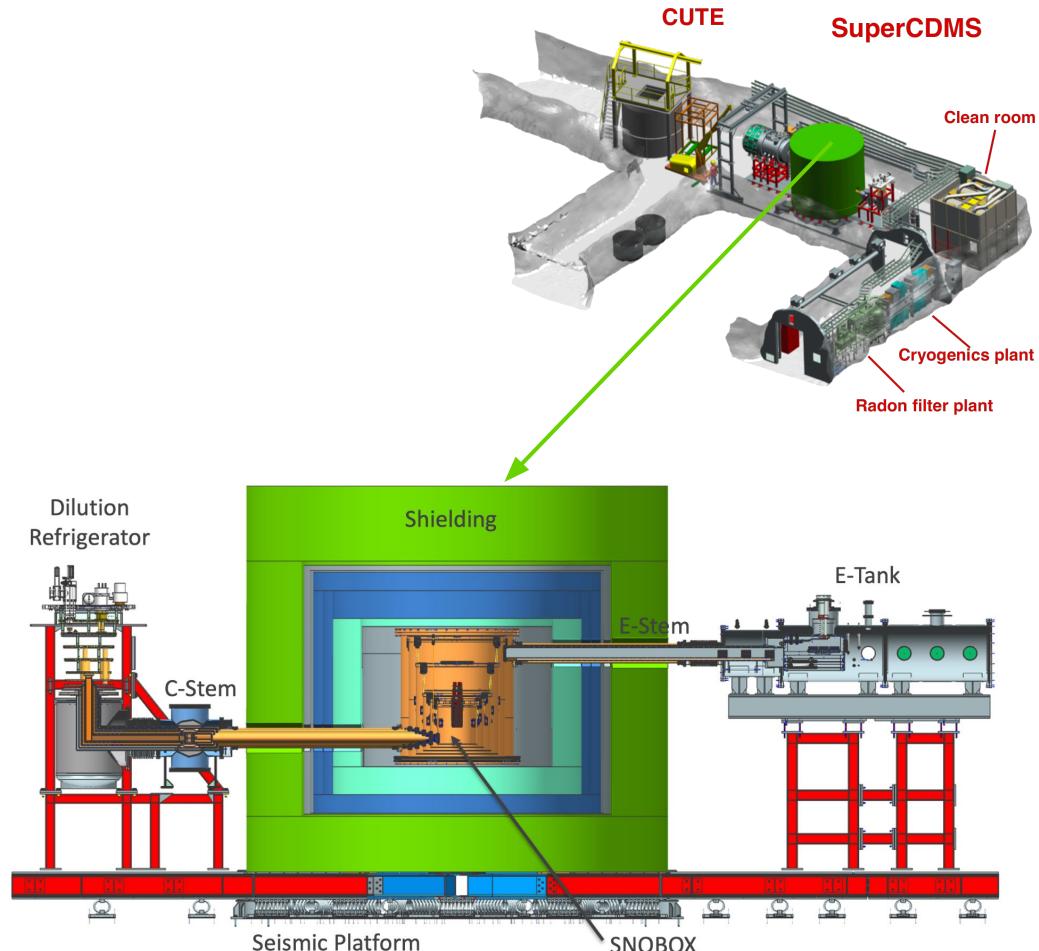
SNOLAB underground laboratory



- Located in Sudbury, Canada
- Has 2 km rock overburden / ~6000 mwe
- Hosts low background experiments
- Class 2000 clean room

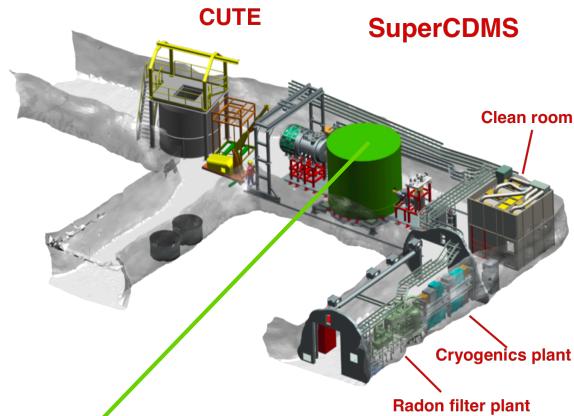
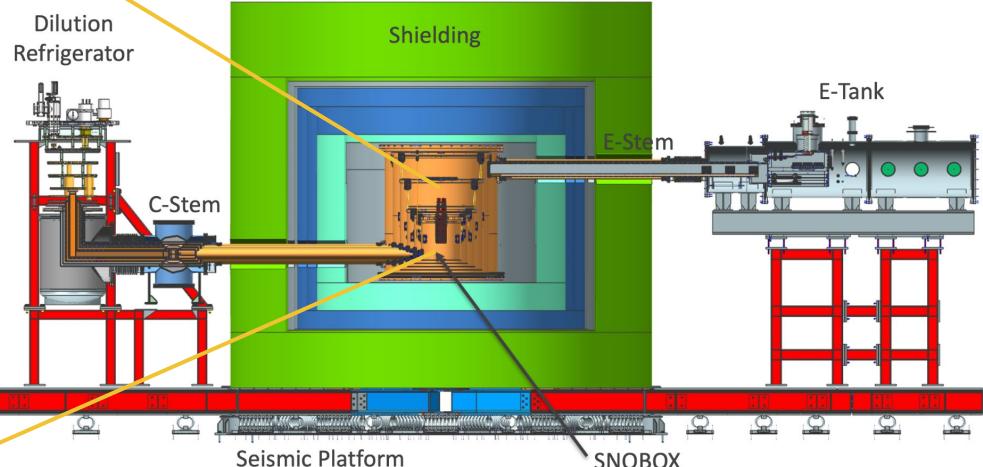
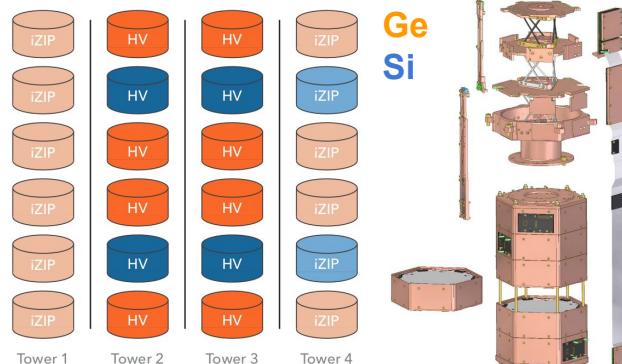


SuperCDMS SNOLAB



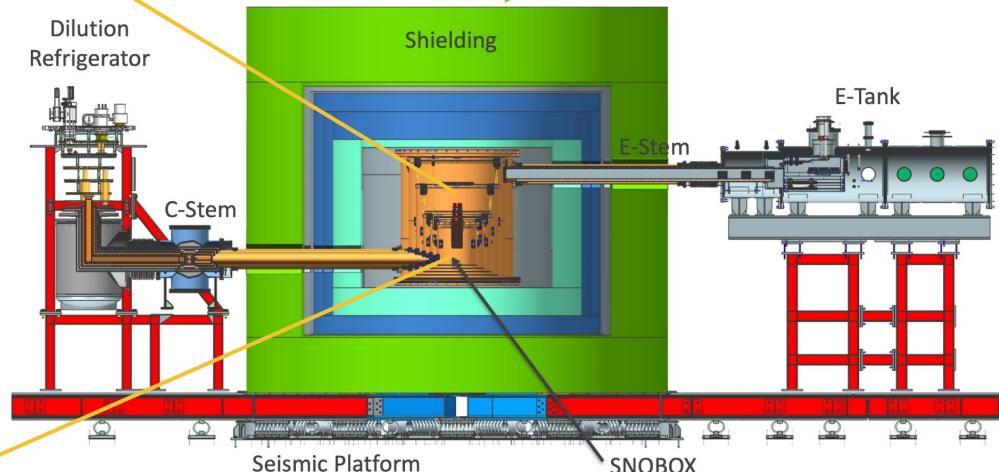
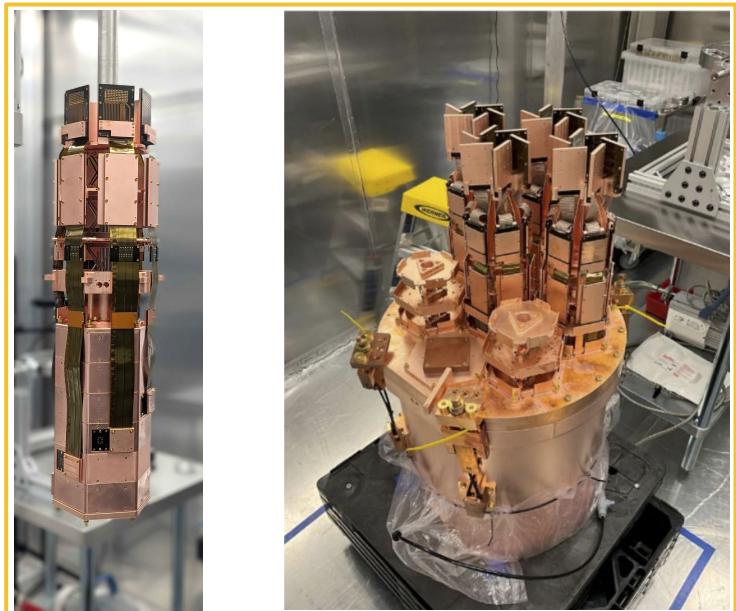
SuperCDMS SNOLAB

- Payload:
 - 4 towers with 6 detectors each
 - 2 iZIP and 2 HV towers
 - ~30kg in total



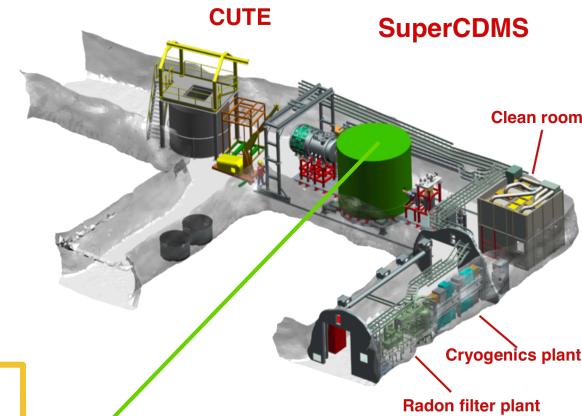
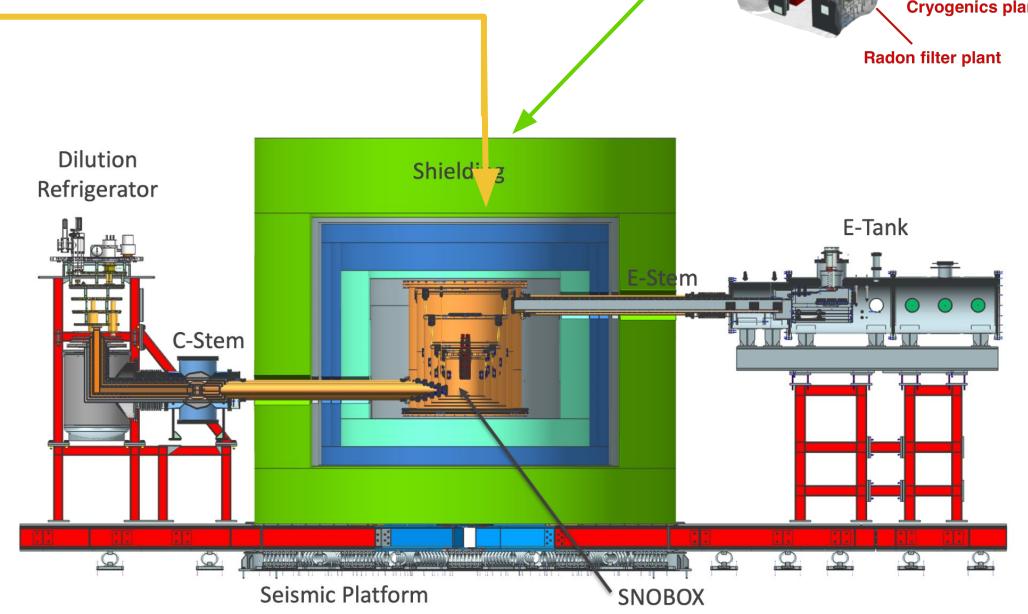
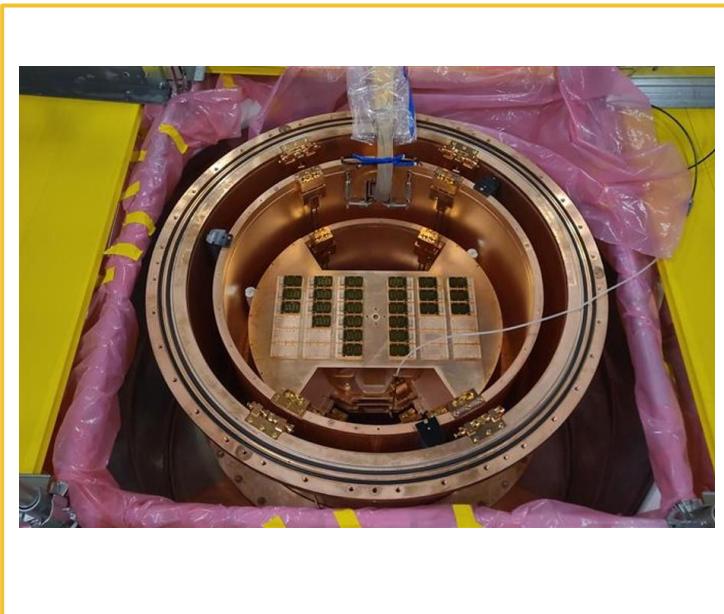
SuperCDMS SNOLAB installation status

- All towers in cryostat



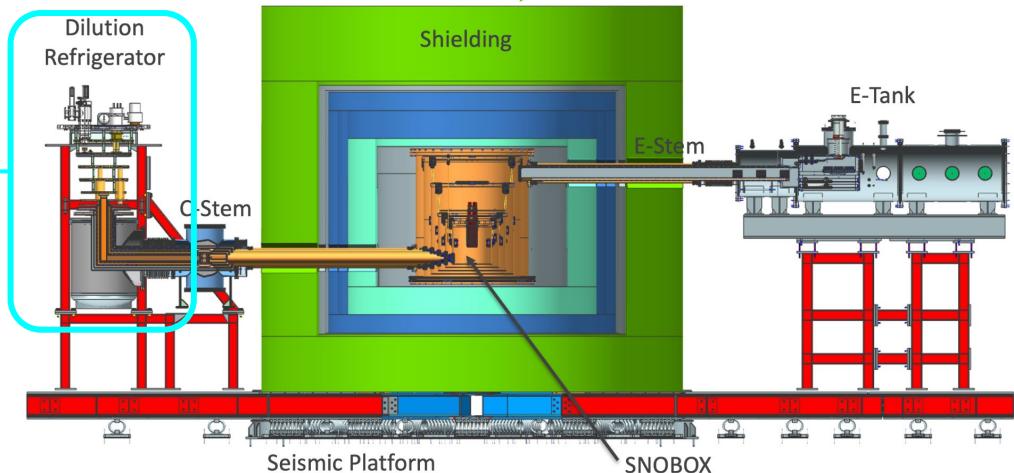
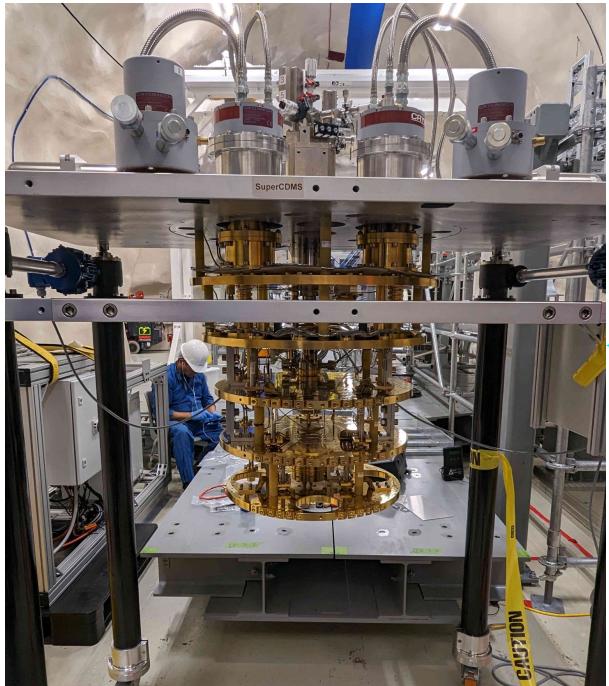
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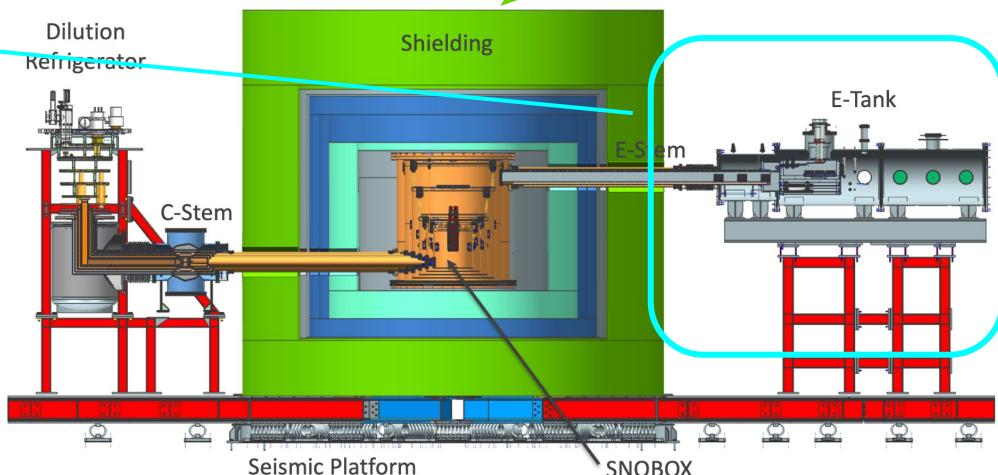
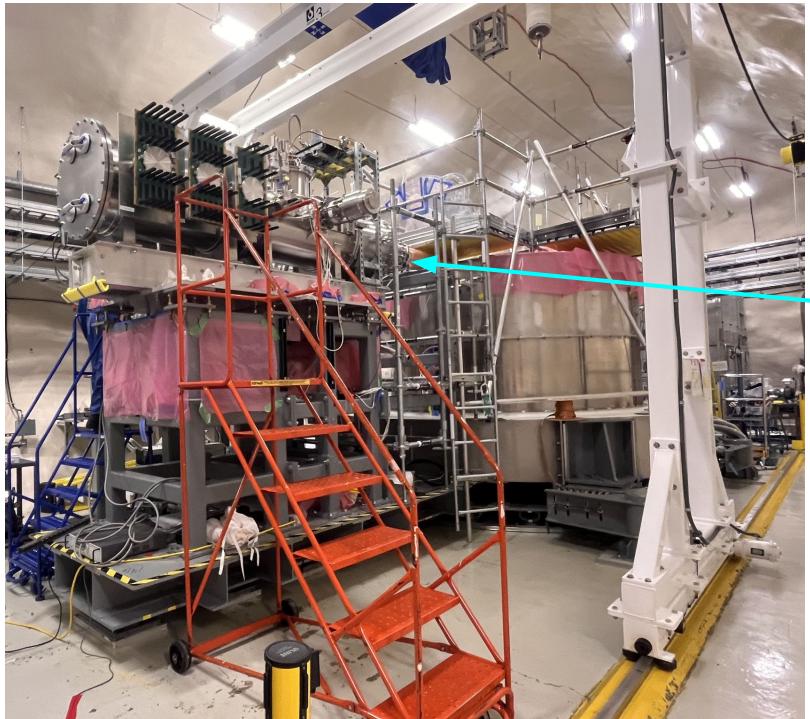


SuperCDMS SNOLAB installation status

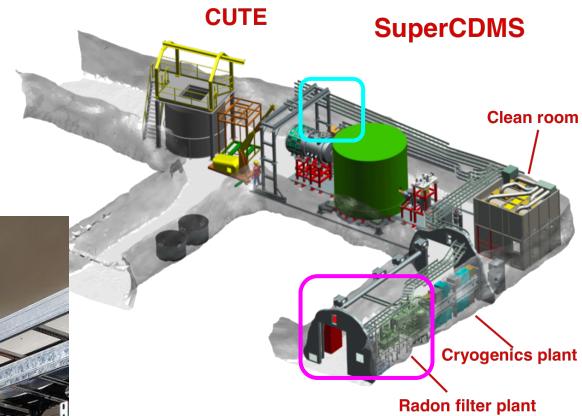
- Dilution refrigerator (~15mK) installed via C-stem



SuperCDMS SNOLAB installation status



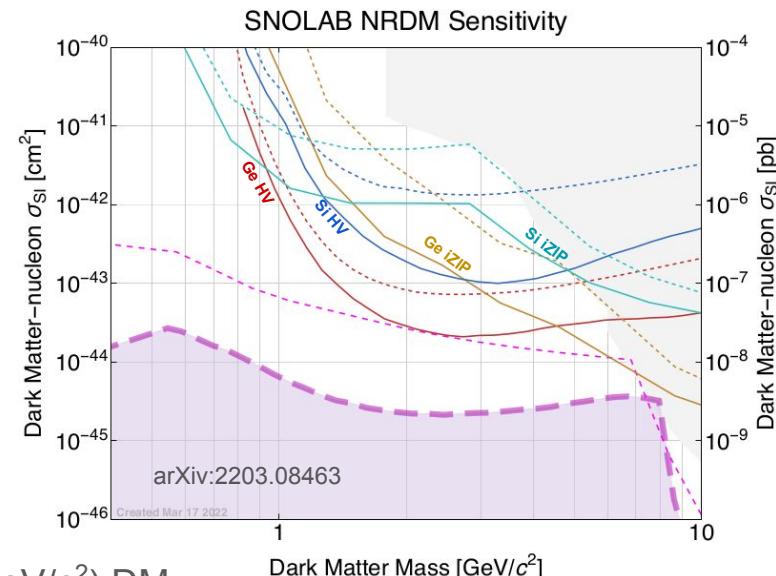
SuperCDMS SNOLAB installation status



SuperCDMS SNOLAB projected sensitivity for NRDM

- The projected sensitivities for Nuclear Recoil
 - Optimum interval - dashed
 - Profile likelihood - solid

	Germanium	Silicon
HV	Low threshold for low mass DM, larger exposure Limited by tritium betas	Lowest threshold for low mass DM, sensitive for lowest DM masses Limited by ^{32}Si (and tritium) betas
iZIP	ER/NR discrimination, understand Ge background Limited by exposure	ER/NR discrimination, understand Si background Limited by exposure

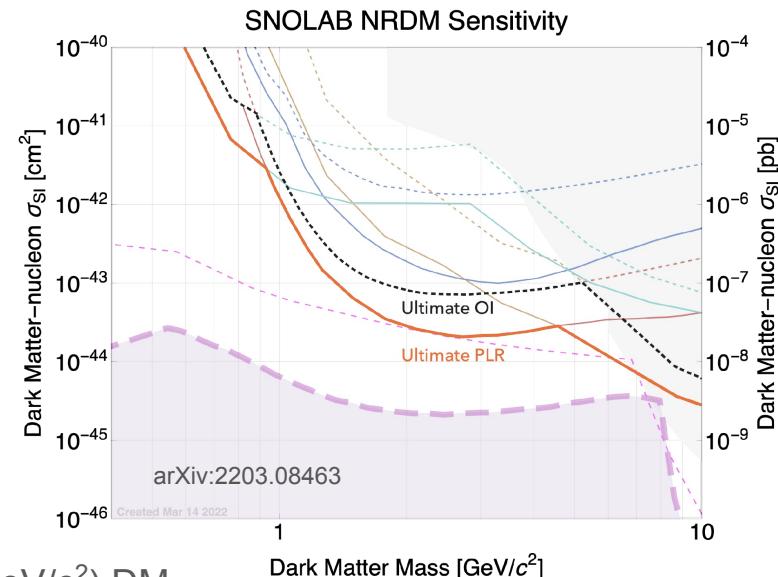


- Aiming for world leading sensitivity to low mass ($\lesssim 10 \text{ GeV}/c^2$) DM
- Challenges: Understanding and modeling detector response, low-energy calibration

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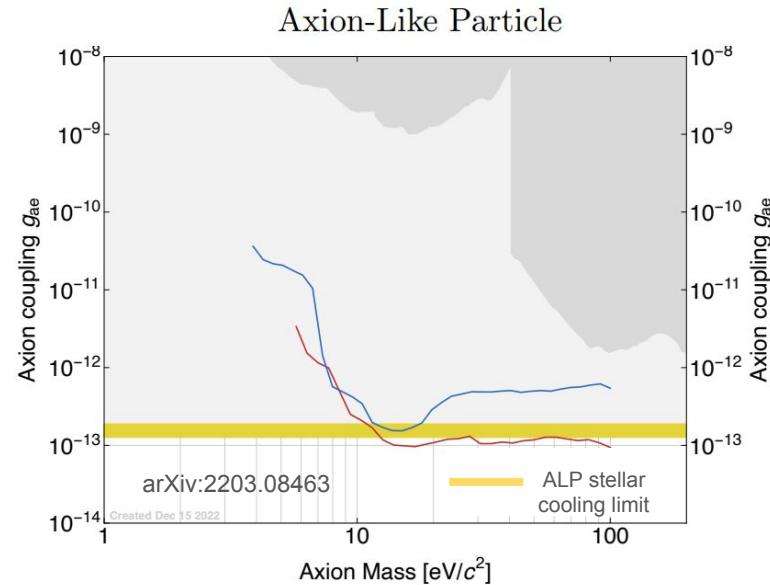
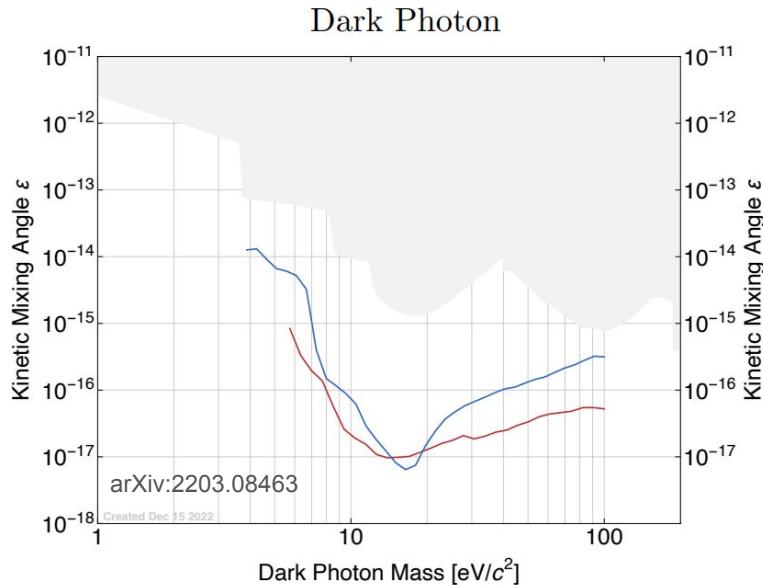
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SuperCDMS SNOLAB projected sensitivity for ERDM

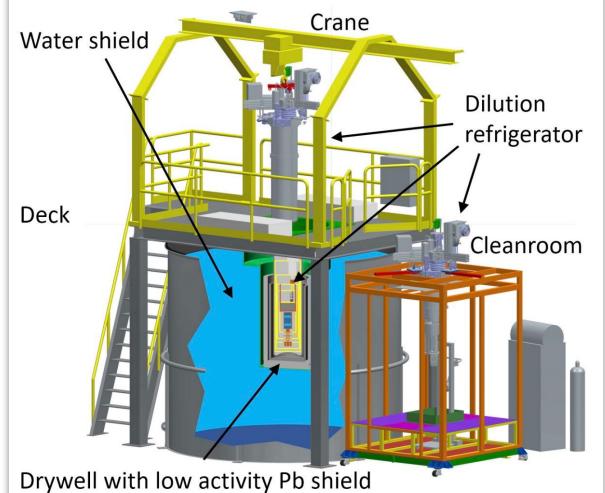
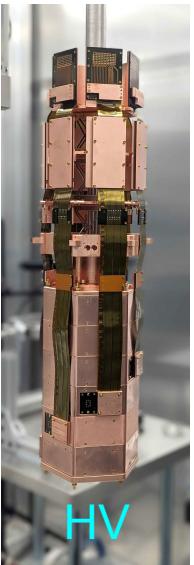
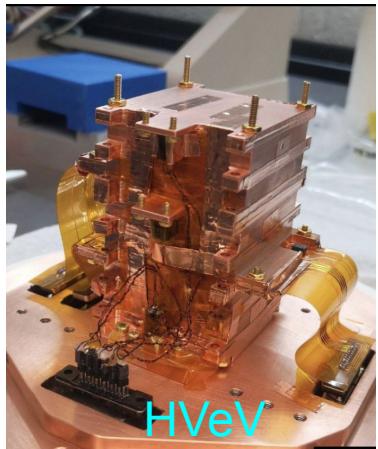
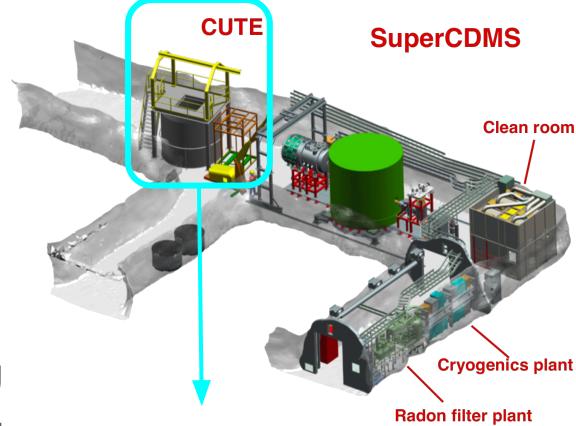
- PLR-based 90% CL projected sensitivities for Electron Recoil
- HV detectors: **Si** / **Ge**



Testing program at CUTE, SNOLAB

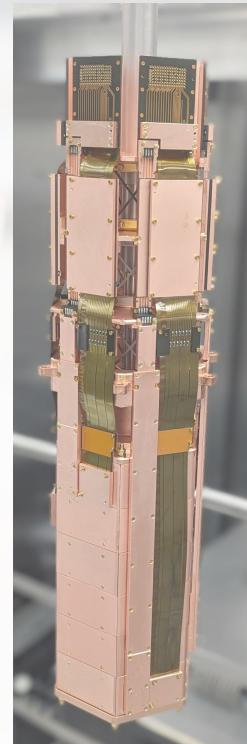
Cryogenic Underground Test facility

- SuperCDMS HV tower testing to understand detector characteristics - completed in Mar'24
- HVeV: 1 cm², gram-size detectors with eV baseline resolution: data-taking with 6 detectors to study Low Energy Excess (LEE) - completed in Sep'24



Summary

- SuperCDMS experiment: direct DM detection, targeting sub-GeV DM masses
- Cryogenic Si/Ge detectors
 - iZIP: background rejection
 - HV: low threshold
- Testing program at CUTE facility at SNOLAB
 - HV detector tower testing
 - HVeV testing
- Primary installation complete at SNOLAB and ready for cooldown
- Experiment commissioning at end of 2025



Stay Tuned!



Backup

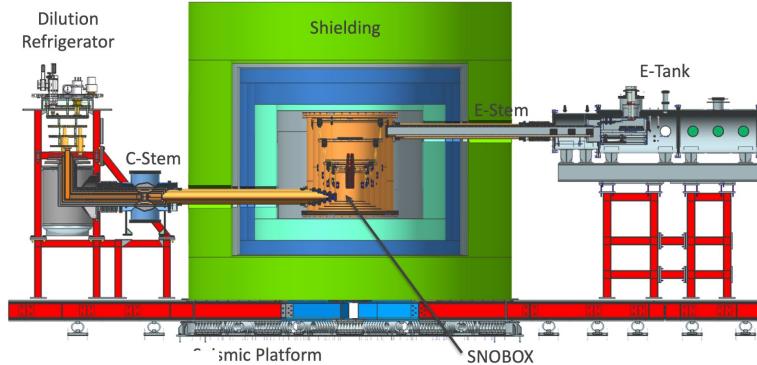
SuperCDMS SNOLAB: Backgrounds

Cosmogenic

- Cosmic ray muons
- Spallation neutrons
- Activated materials

Environmental

- Airborne radon & daughters
- Radio-impurities in materials



“Singles” Background Rates (counts/kg/keV/year)	Electron Recoil				Nuclear Recoil ($\times 10^{-6}$)	
	Ge HV	Si HV	Ge iZIP	Si iZIP	Ge iZIP	Si iZIP
Coherent Neutrinos					2300.	1600.
Detector-Bulk Contamination	21.	290.	8.5	260.		
Material Activation	1.0	2.5	1.9	15.		
Non-Line-of-Sight Surfaces	0.00	0.03	0.01	0.07	—	—
Bulk Material Contamination	5.4	14.	12.	88.	440.	660.
Cavern Environment	—	—	—	—	510.	530.
Cosmogenic Neutrons					73.	77.
Total	27.	300.	22.	370.	3300.	2900.

Phys. Rev. D 95, 082002

SuperCDMS SNOLAB: Shielding

▶ Outer neutron shield:

- ▶ Reduce MeV neutrons from cavern wall by $\sim 10^6$, ~ 20 for μ -induced neutrons (GeV)

▶ Gamma-ray shield:

- ▶ Reduces MeV gammas from cavern wall by $\sim 10^5$
- ▶ Inner layer is 1cm ancient lead, $^{210}\text{Pb} < 1 \text{ Bq/kg}$
- ▶ Reduce Bremsstrahlung ^{210}Bi in shield by factor of ~ 20

▶ Radon barrier:

- ▶ Al sheet banded around the lead shield with taped joints
- ▶ Reduce $\sim 100 \text{ Bq/m}^3$ to $< 0.1 \text{ Bq/m}^3$ (mine air)

▶ Inner neutron shield:

- ▶ Stacked 2" thick poly sheets
- ▶ Reduces neutron rate by $> 10 \mu$ -induced neutron from rock
- ▶ Neutrons from lead
- ▶ Absorbs backscatter from SNOBOX vacuum cans
- ▶ No neutron "reflection" from lead
- ▶ Copper cans of Cryostat System:
 - ▶ 6.7 cm of copper in nested cans (SNOBOX)
 - ▶ Copper is radiopure compared to lead
 - ▶ Final gamma shield
 - ▶ Reduce residual gammas by ~ 25
 - ▶ Reduces Bremsstrahlung from ^{210}Bi β -decay by ~ 20

61 cm outer neutron water shield in modular stainless steel tanks



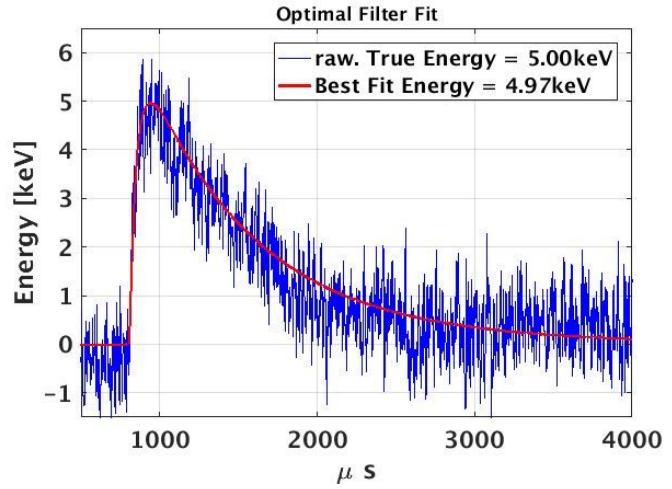
Slide credit: M. J. Wilson

Energy estimation (credit: W. A. Page)

$$S(t) = aA(t) + n(t)$$

↓ ↓ ↓ ↓
signal pulse amplitude known template gaussian noise

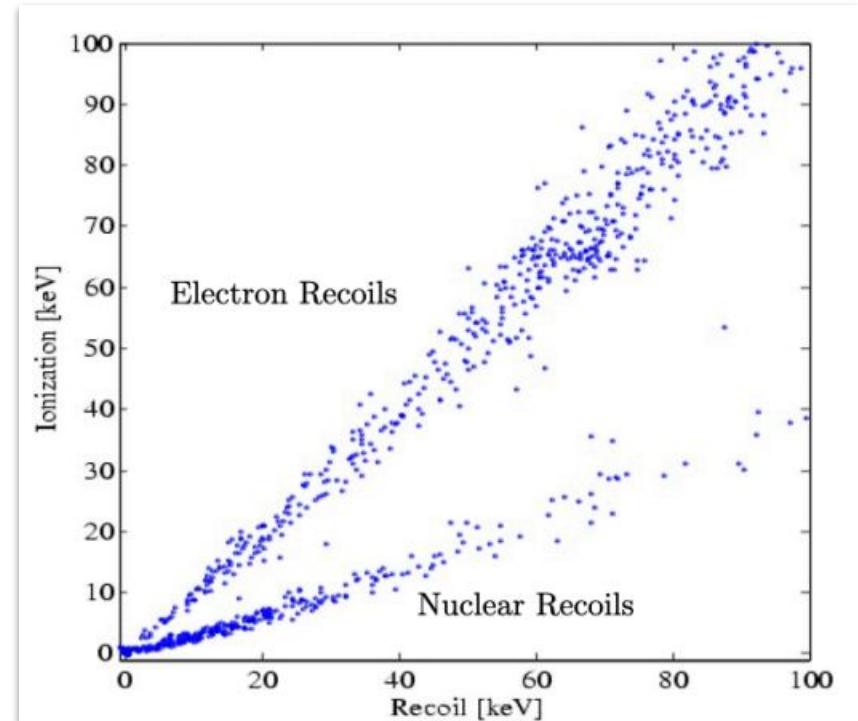
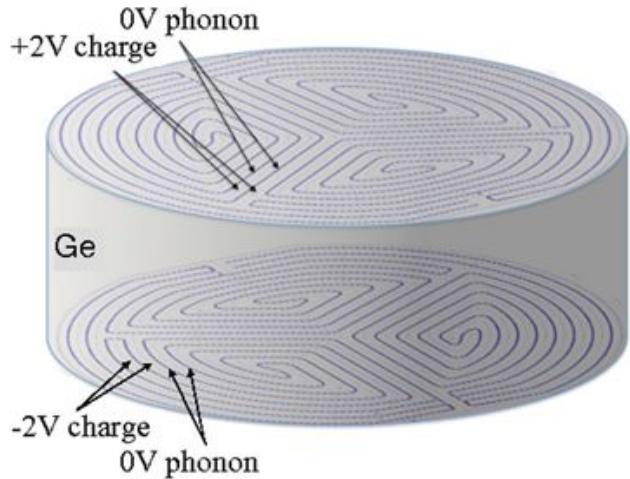
Calculate a frequency-domain χ^2 : $\chi^2(a) = \sum_n \frac{|\tilde{S}_n - a\tilde{A}_n|^2}{J_n}$



SuperCDMS SNOLAB: ER/NR

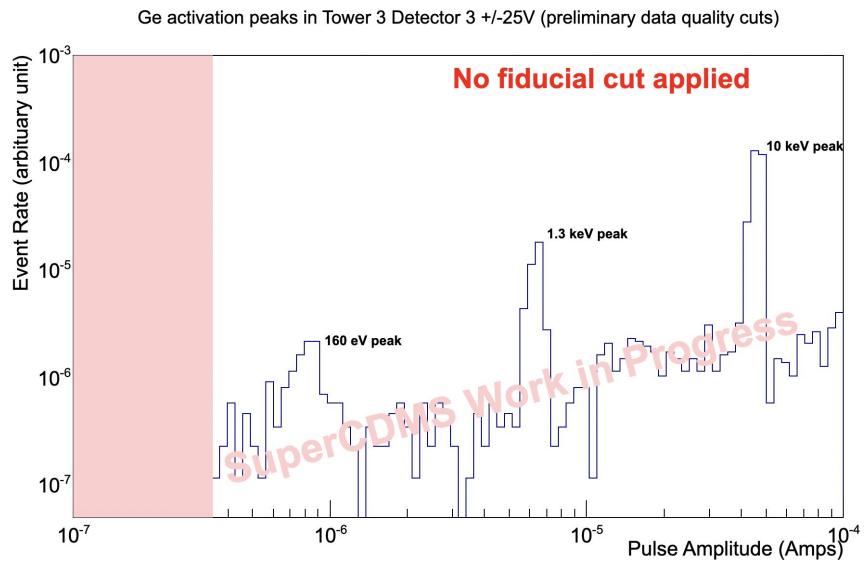
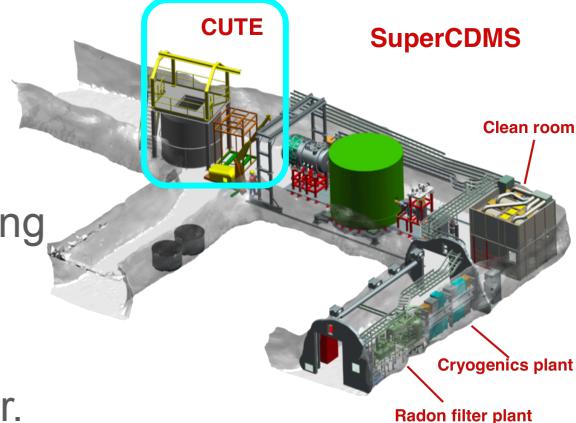
$$E_{ph} = E_R + \frac{Y(E_R)E_R}{\epsilon}e\Delta V$$

$$Y(E_R) = 1 \rightarrow \text{ER}$$
$$Y(E_R) < 1 \rightarrow \text{NR}$$



Tower testing at CUTE, SNOLAB

- SuperCDMS HV tower hosting 4 Ge and 2 Si detectors testing completed in Mar'24 ✓
- First time testing in low background environment
- Ongoing analysis efforts, possibility of early results next year.



SuperCDMS SNOLAB projected sensitivity for ER-LDM

- PLR-based 90% CL projected sensitivities for Electron Recoil
 - for heavy mediator (form factor $F(q) = 1$, left)
 - a light mediator (form factor $F(q) = 1/q^2$, right)
- HV detectors: Si: blue; Ge: red

