

Producing Light Dark Matter at Fixed Target Experiments

Taylor Rose Gray

Supervisor: Riccardo Catena

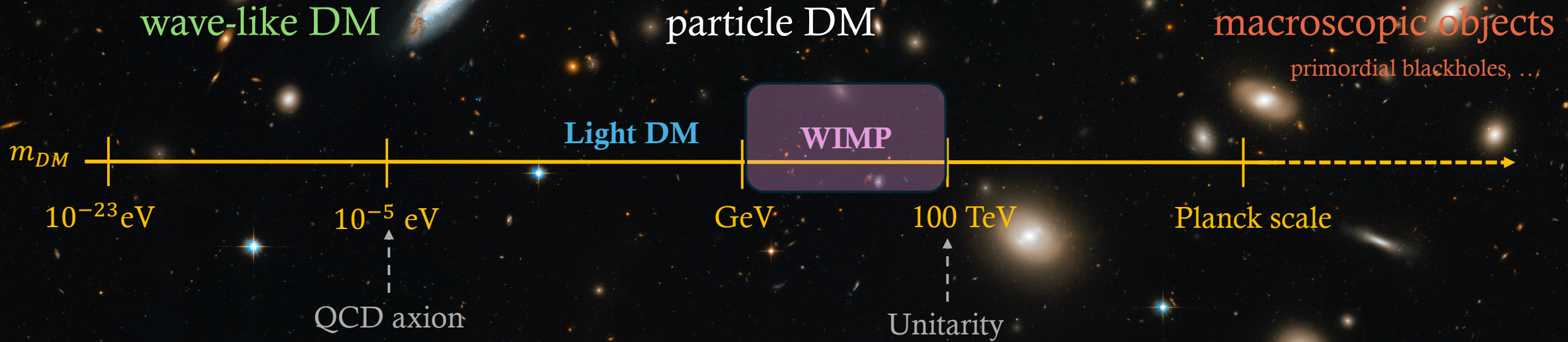
TeVPA 2025 Valencia

*Knut and Alice
Wallenberg
Foundation*

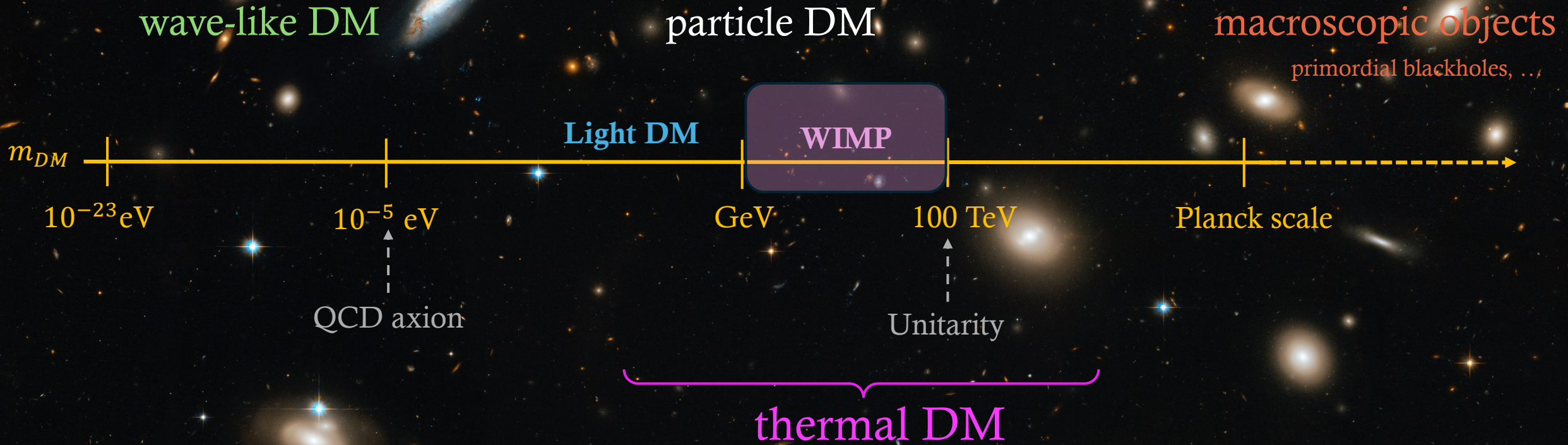


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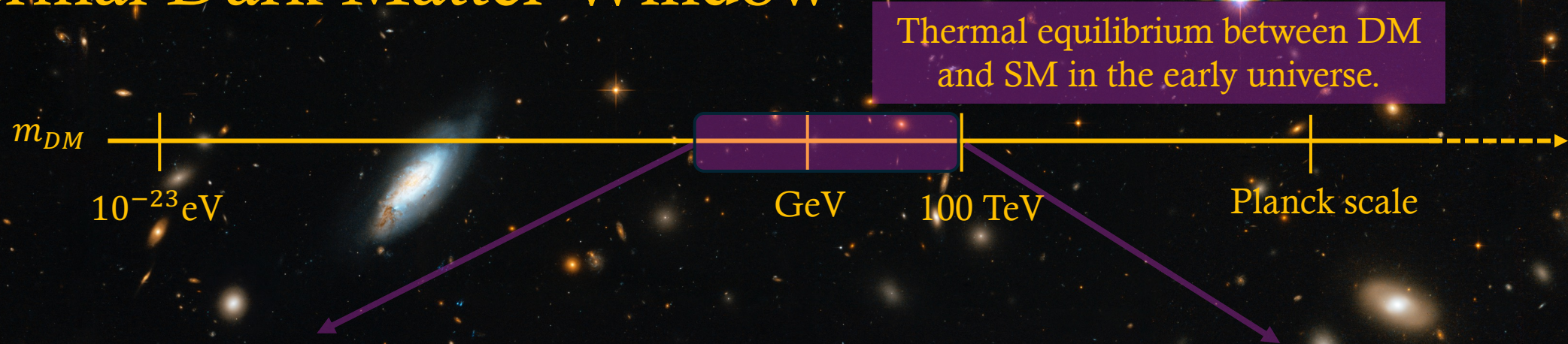
What is the *Dark Matter* made of?



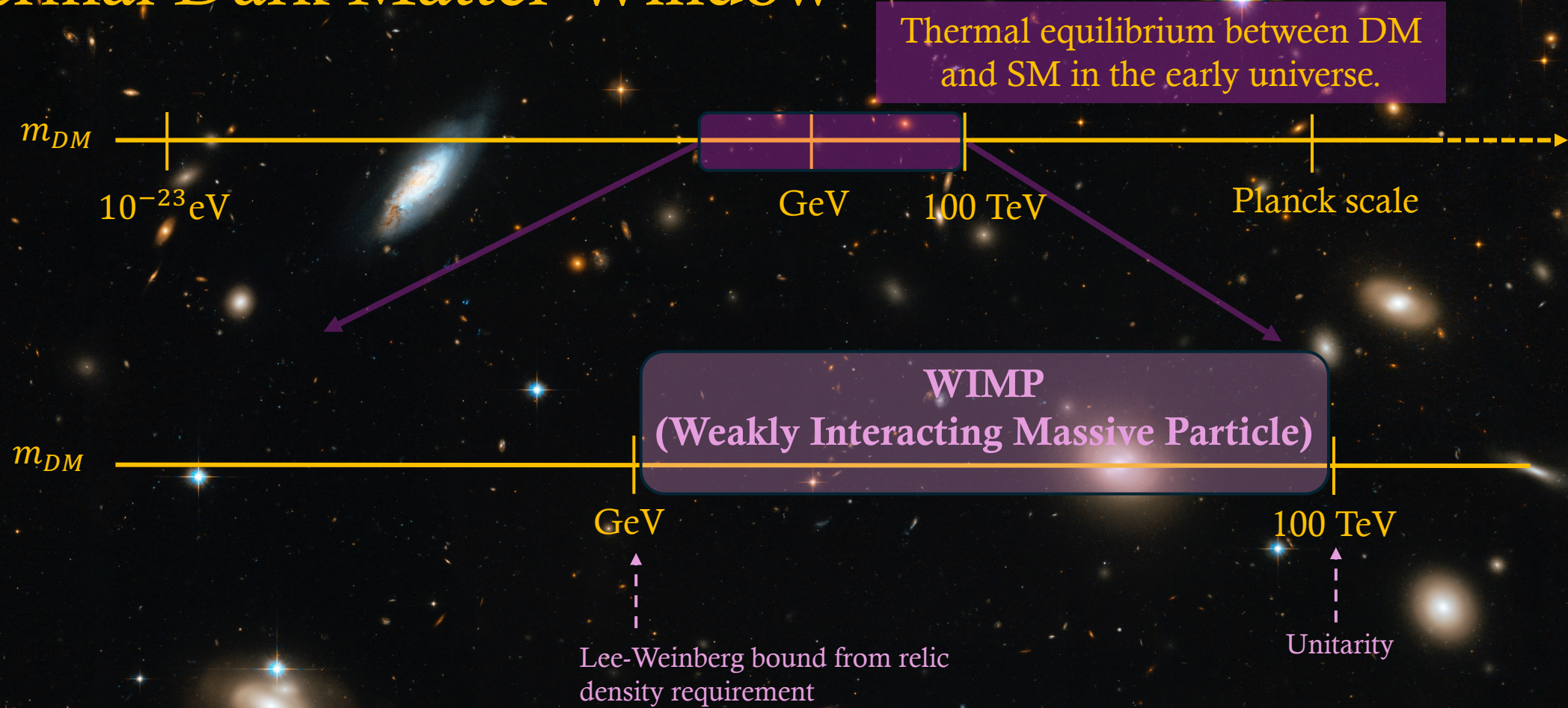
What is the *Dark Matter* made of?



Thermal Dark Matter Window

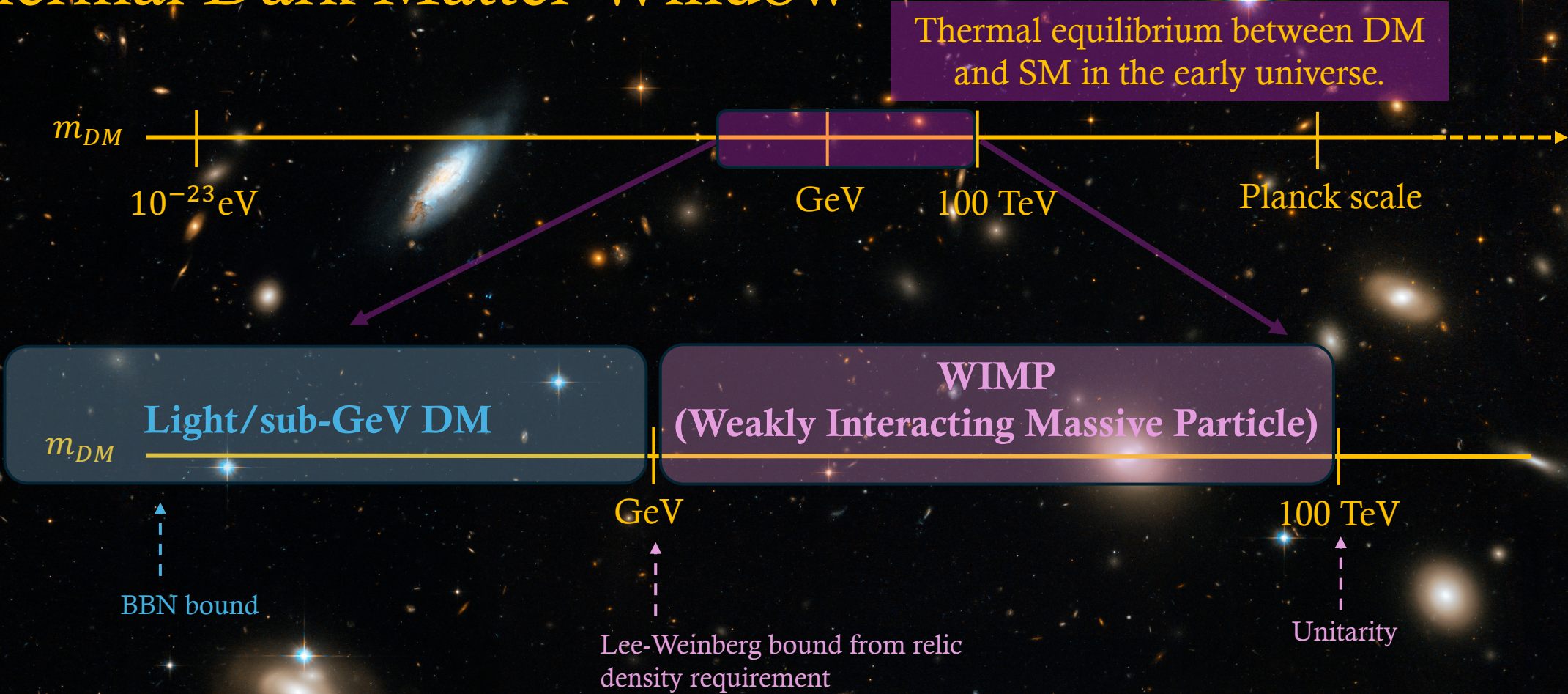


Thermal Dark Matter Window



- DM interacts via the weak force near the weak scale
- Widely tested by direct detection experiments

Thermal Dark Matter Window

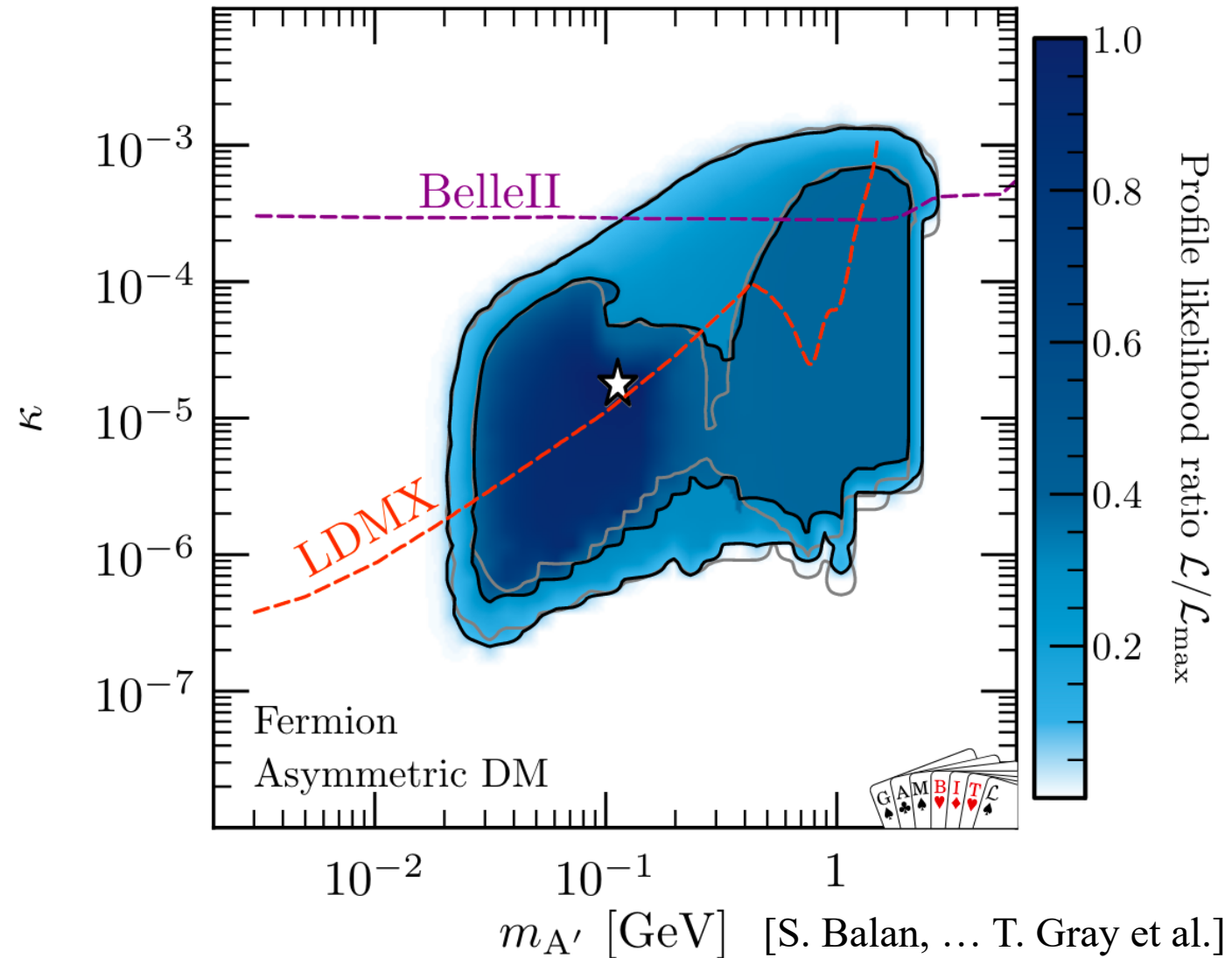


- Evade LW bound by introducing a **new mediator**
- DM too light to deposit sufficient energy in nuclear recoils

- DM interacts via the weak force near the weak scale
- Widely tested by direct detection experiments

Light Dark Matter Status

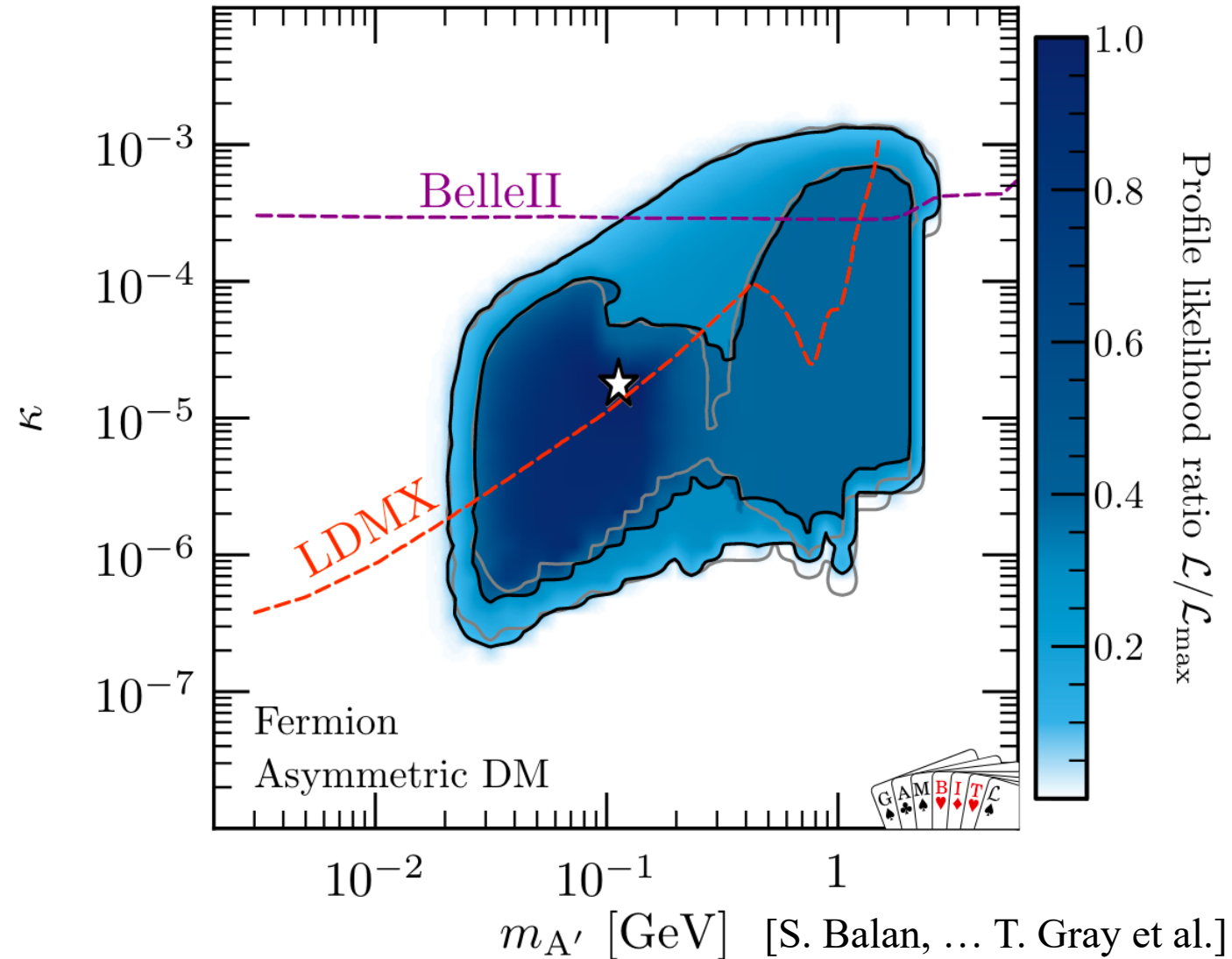
Frequentist Global Fits



Light Dark Matter Status

Frequentist Global Fits

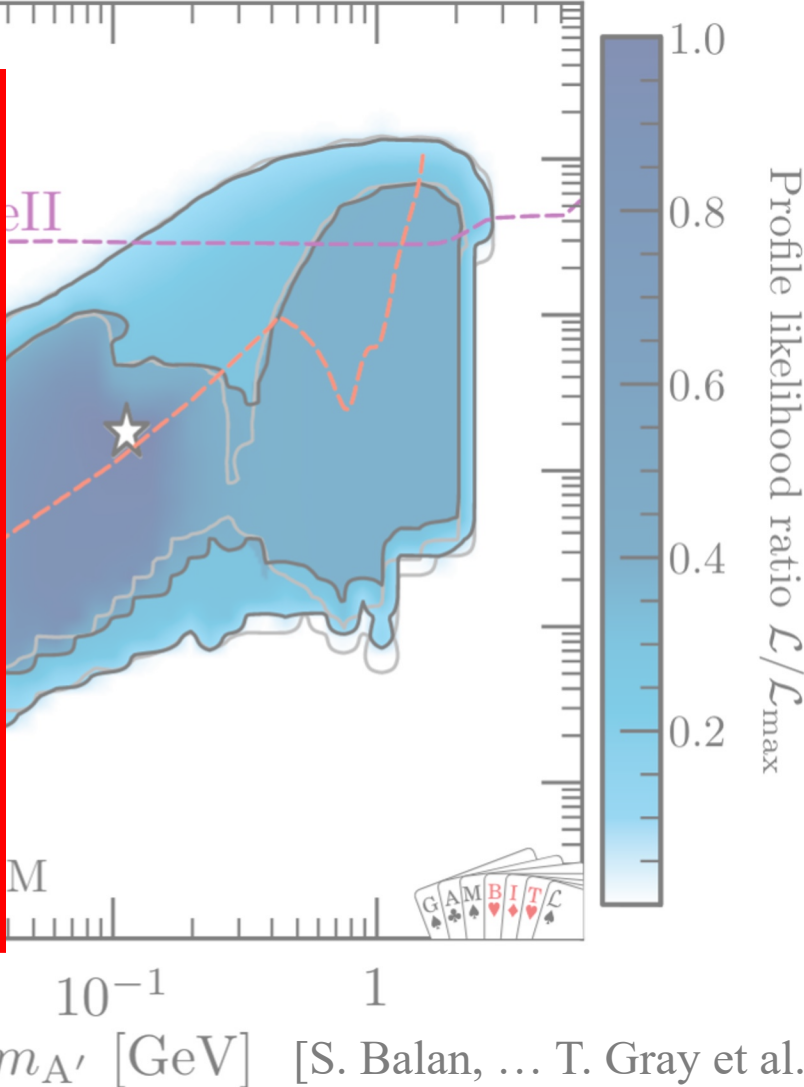
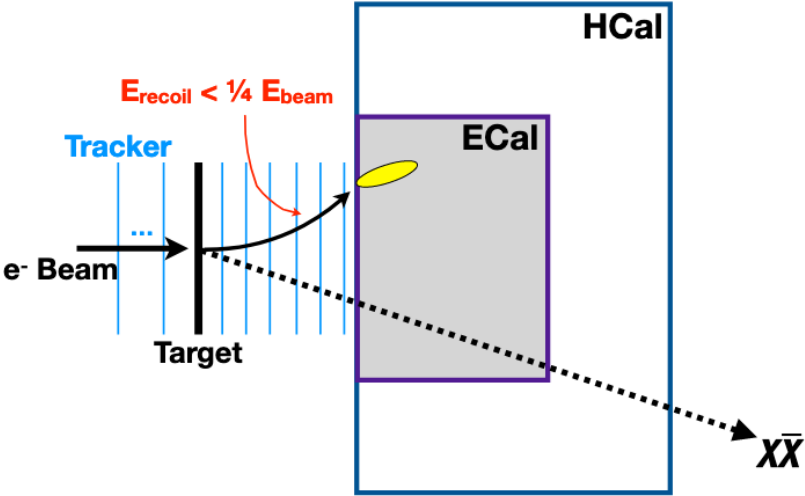
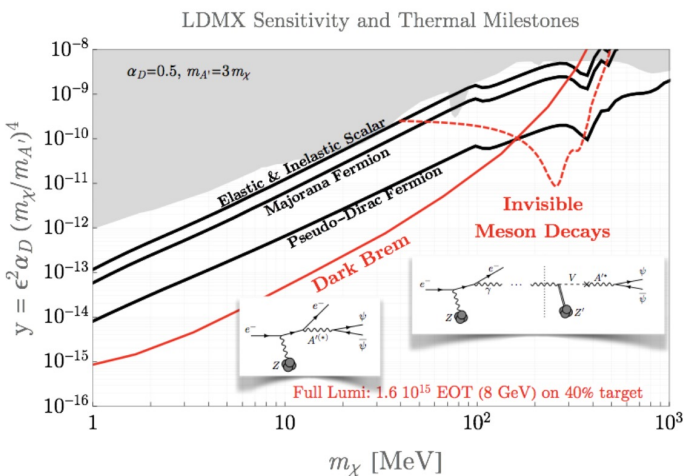
LDMX will probe **64%** of posterior volume!



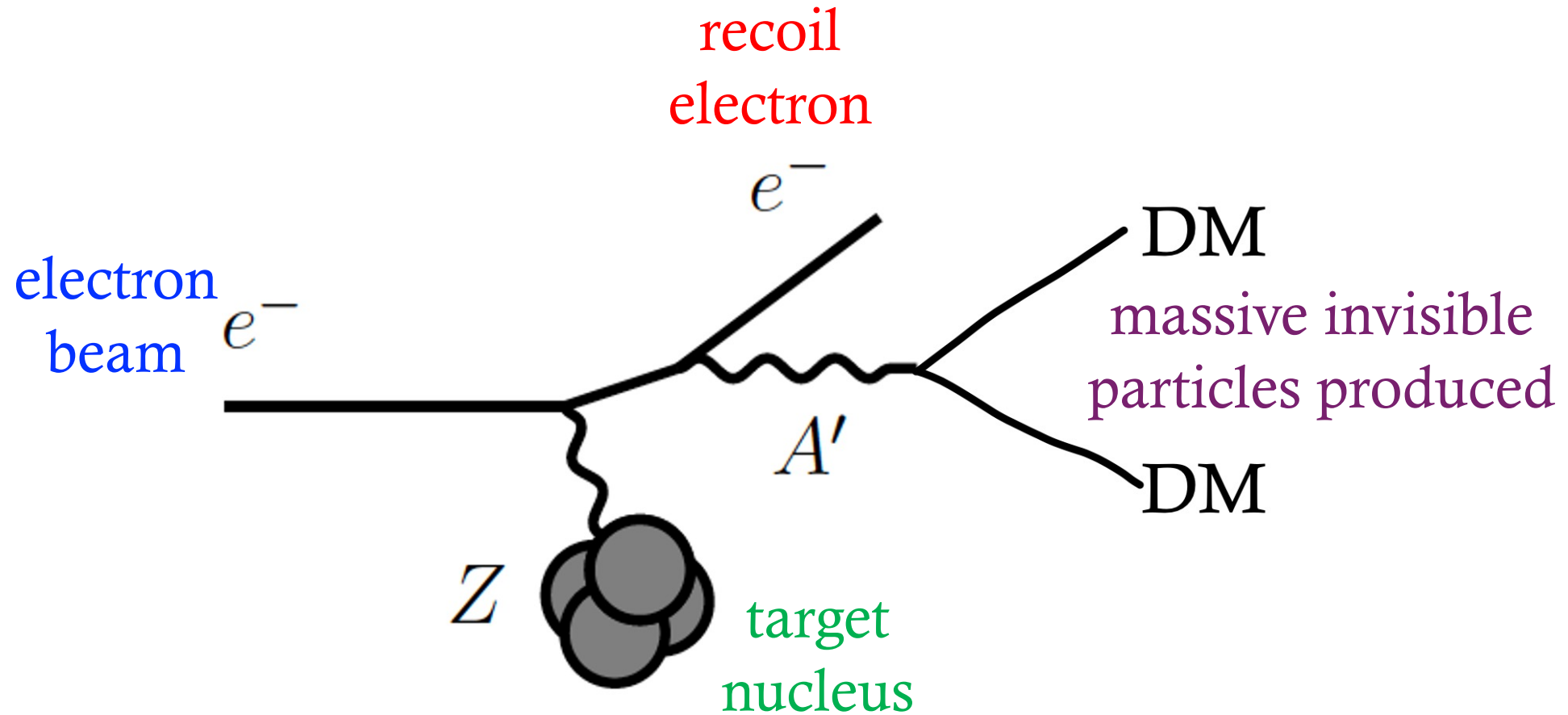
Light Dark Matter Status

LDMX

Future Fixed-Target Experiment
(Missing Energy/Momentum)

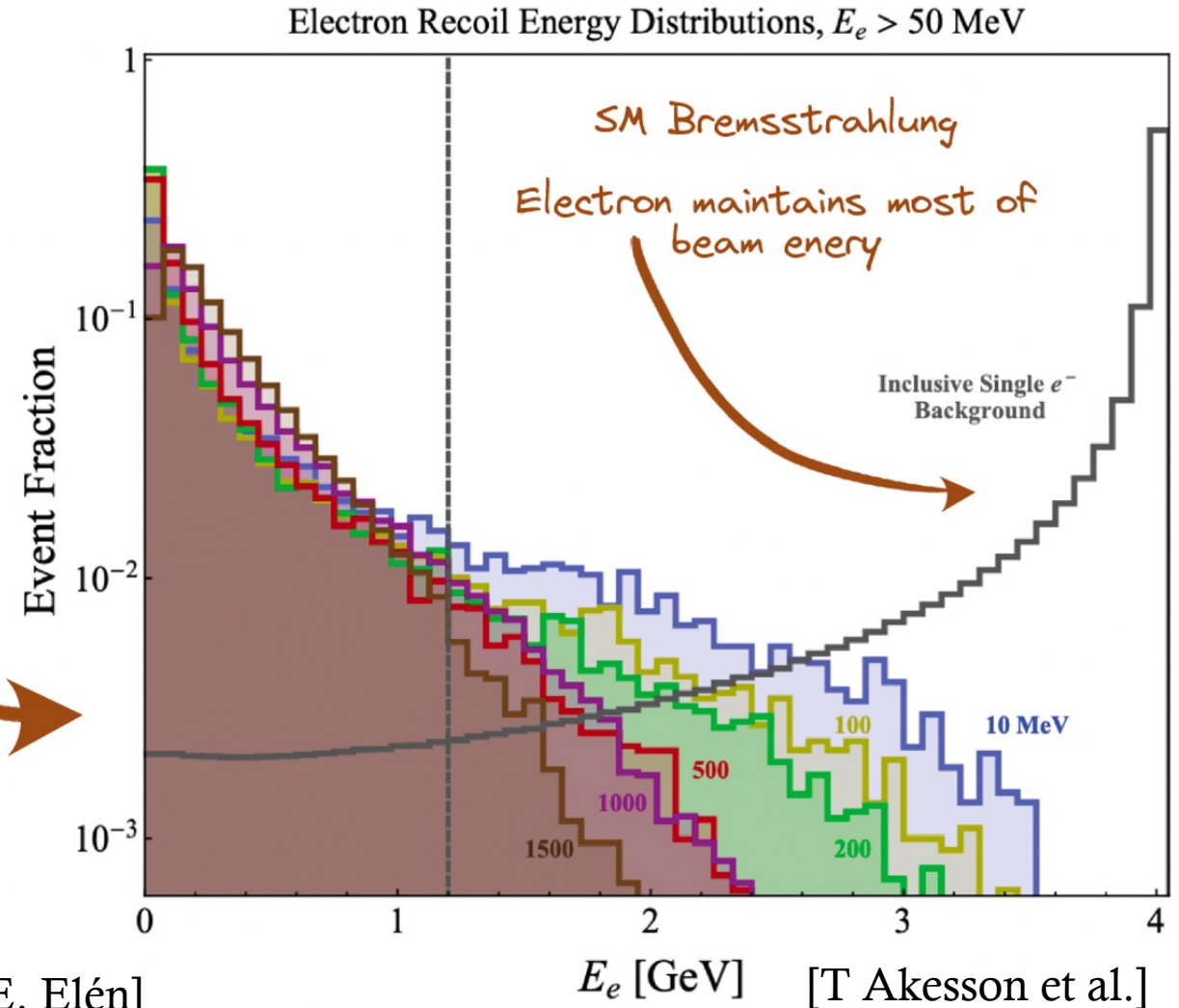
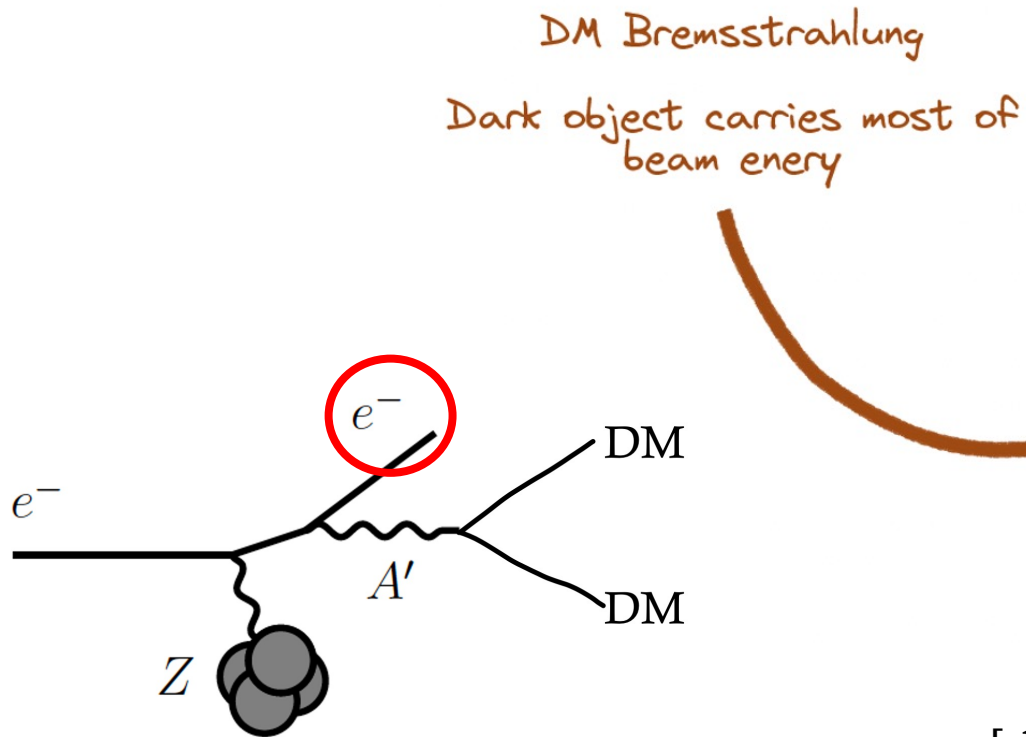


LDMX: Light Dark Matter EXperiment



LDMX: Light Dark Matter EXperiment

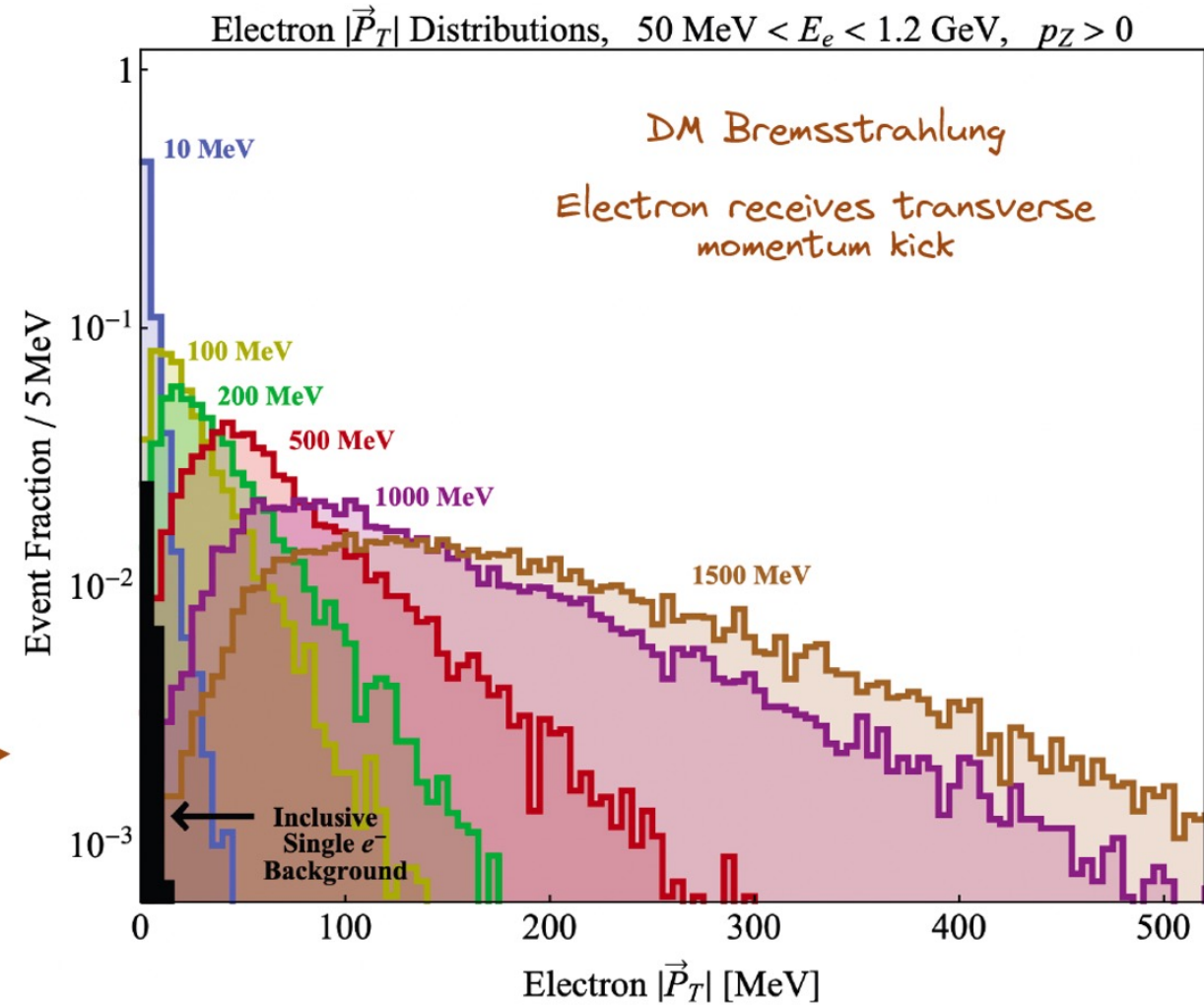
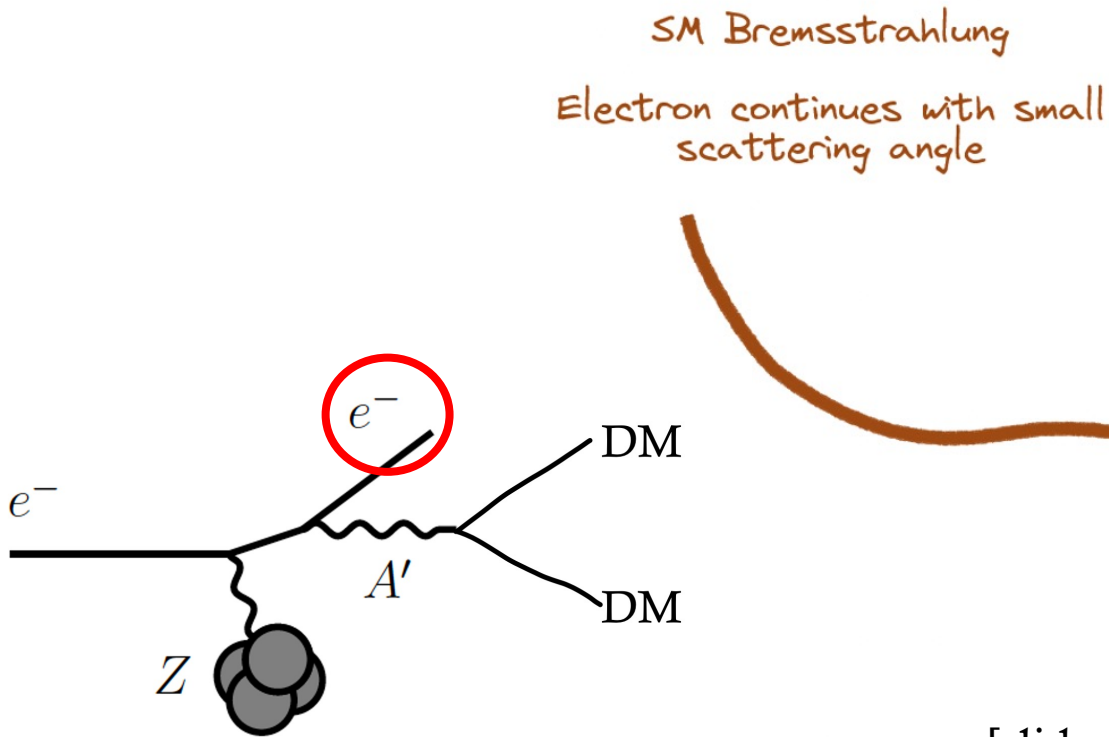
Recoil e^- Energy



[slide from E. Elén]

LDMX: Light Dark Matter EXperiment

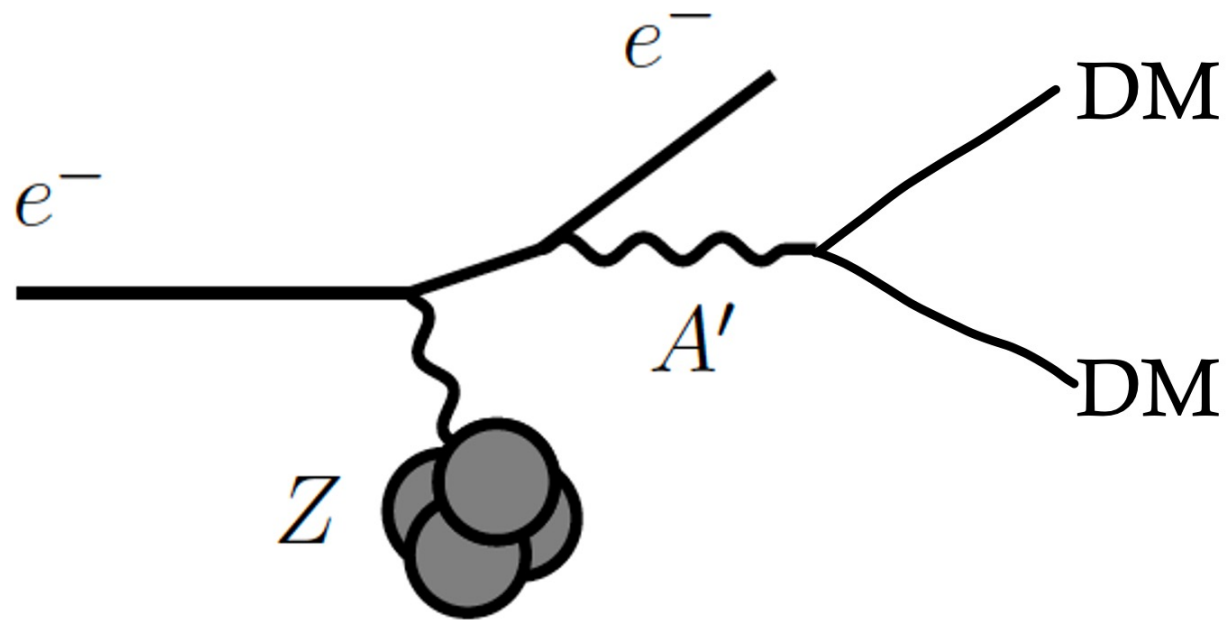
Recoil e^-
Transverse Momentum p_T



[slide from E. Elén]

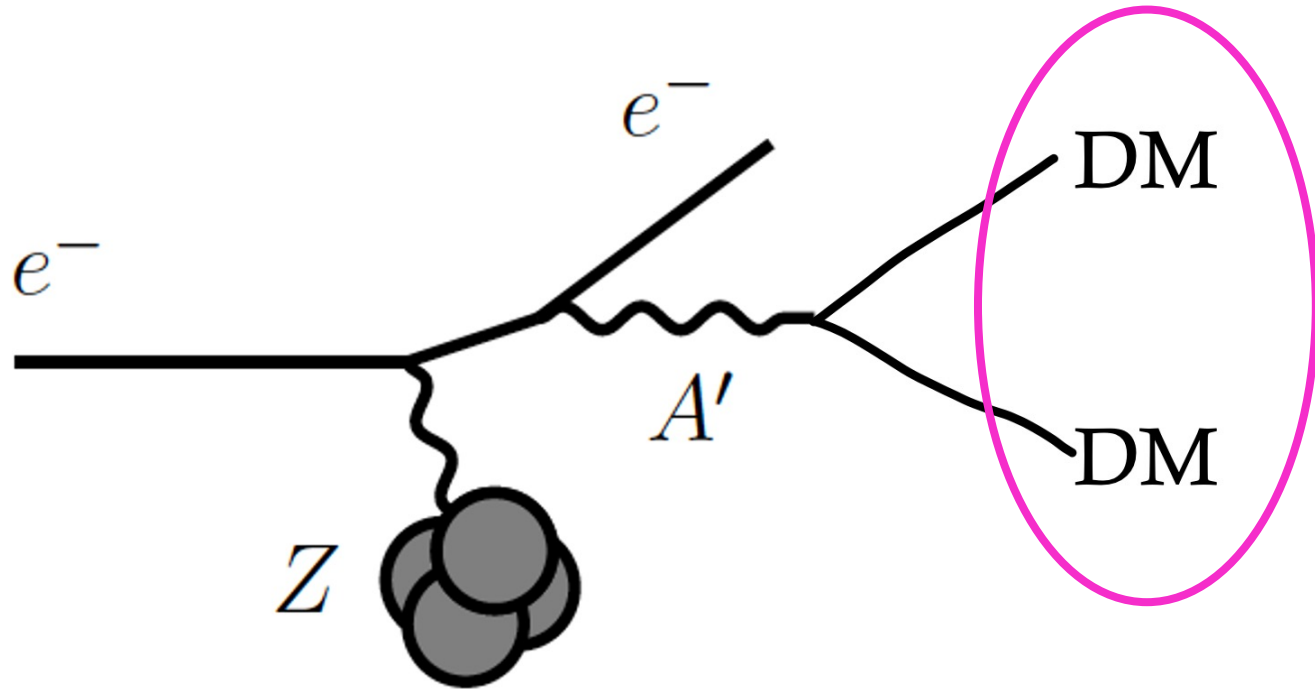
[T Akesson et al.]

Light Dark Matter Production at Fixed Targets



Light Dark Matter Production at Fixed Targets

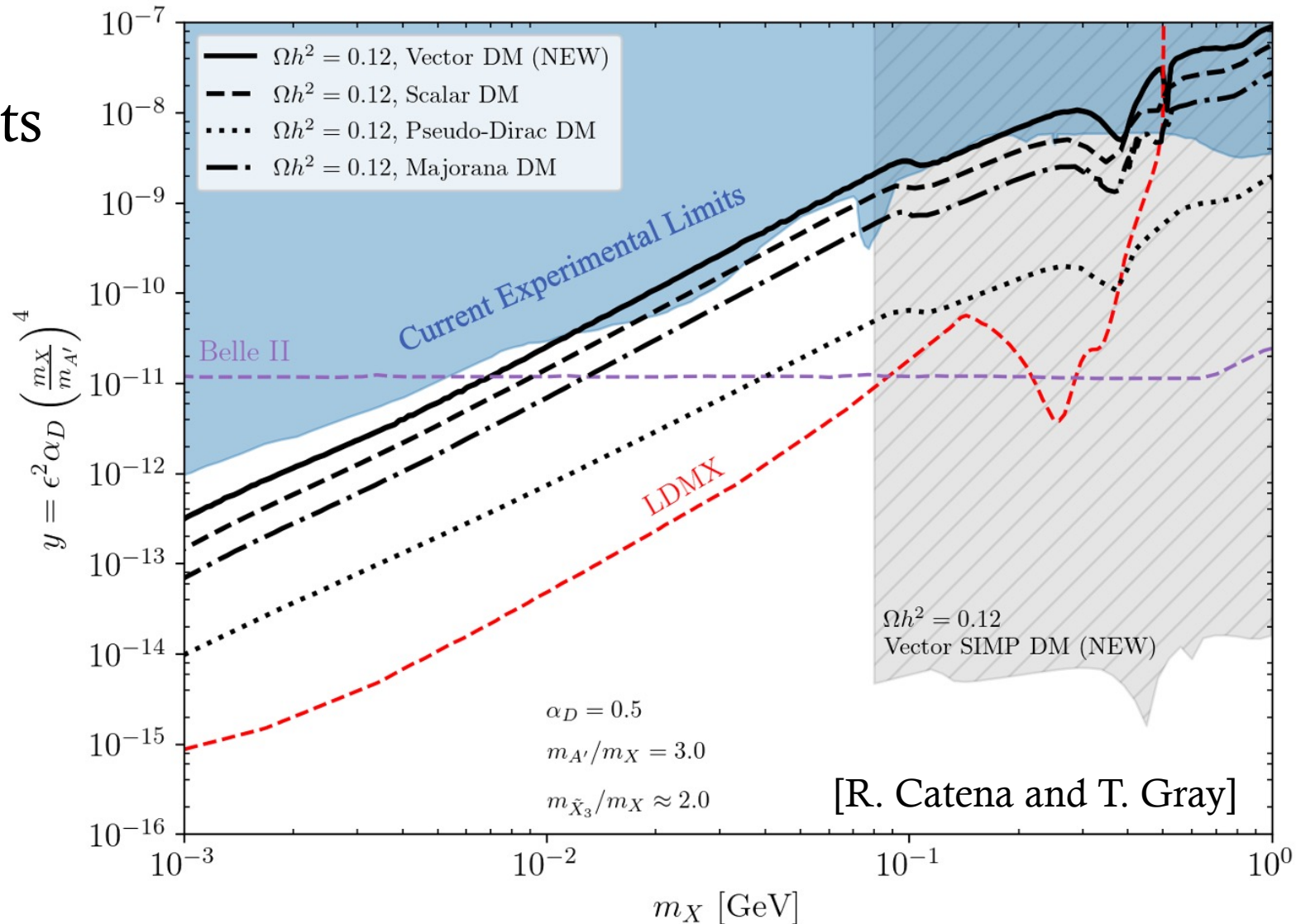
DM theory landscape



Light Dark Matter Theory Landscape

Summary of Thermal Targets

- including **vector DM**!



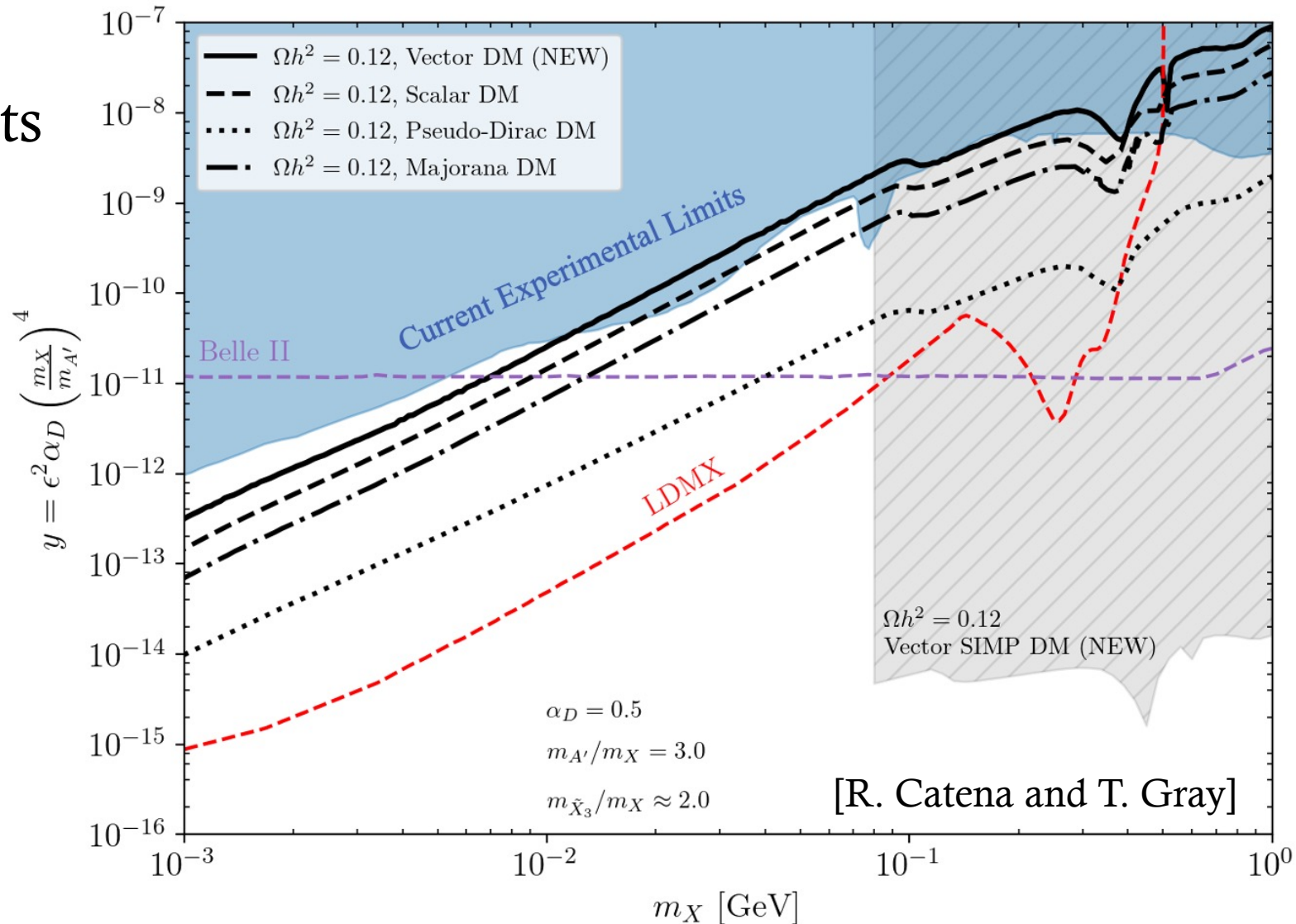
Light Dark Matter Theory Landscape

Summary of Thermal Targets

- including **vector DM**!

Mass Hierarchy:

$$m_{DM} < m_{med}/2$$



Light Vector Dark Matter Scenario

$$\text{Dark SU(2)} \left\{ \begin{array}{l} X_{\mu}^a \rightarrow \begin{array}{l} X^+ X^- \text{ Vector DM} \\ Z' \text{ "Dark Photon"} \end{array} \\ \Phi_D \text{ (doublet)} \text{ gives mass to dark gauge bosons} \\ \Sigma_D \text{ (triplet)} \text{ gives more mass to the DM} \end{array} \right.$$



Reversed Mass Hierarchy:

$$m_{DM} > m_{med}/2$$

Light Vector Dark Matter Scenario

$$\text{Dark SU(2)} \left\{ \begin{array}{l} X_{\mu}^a \rightarrow \begin{matrix} X^+ X^- \\ Z' \end{matrix} \text{ Vector DM "Dark Photon"} \\ \Phi_D \text{ (doublet)} \text{ gives mass to dark gauge bosons} \\ \Sigma_D \text{ (triplet)} \end{array} \right.$$

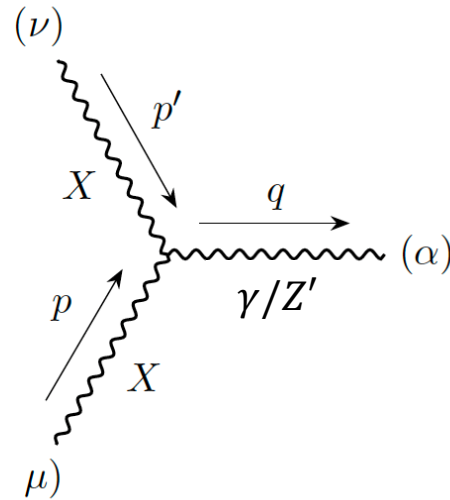
gives more mass to the DM



Reversed Mass Hierarchy:

$$m_{DM} > m_{med}/2$$

$$\mathcal{L}_{Portal} = -\frac{\sin \zeta}{v_{\Sigma_D}} \text{Tr} [\Sigma_D X_{\mu\nu}] B^{\mu\nu}$$



(Magnetic Dipole)

[A. Banerjee, R. Catena, T. Gray, arXiv: 2511.XXXX]

Light Vector Dark Matter Scenario

Dark SU(2)

$X_\mu^a \longrightarrow \begin{matrix} X^+ X^- \\ Z' \end{matrix}$

Vector DM
"Dark Photon"

Φ_D (doublet)
 Σ_D (triplet)

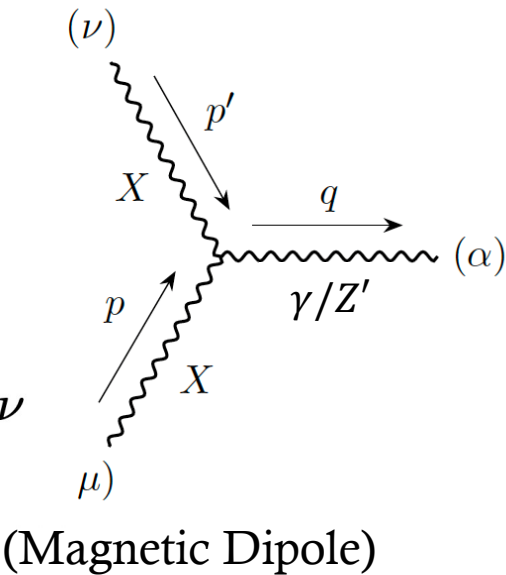
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Reversed Mass Hierarchy:
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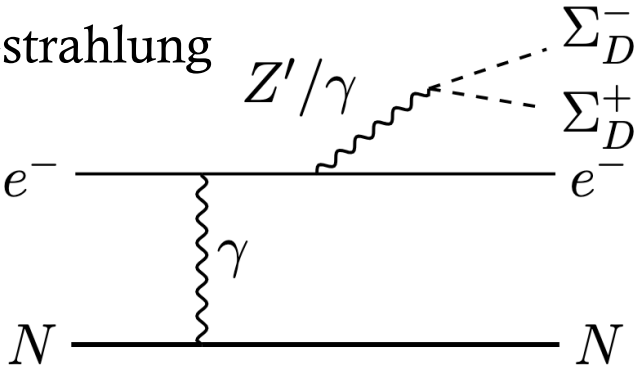
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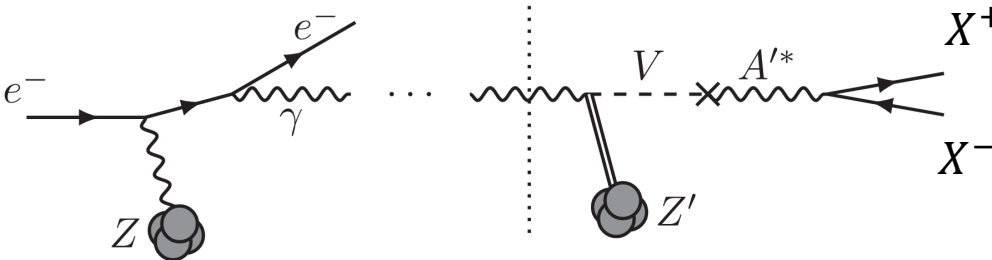


Phenomenology:

- Off-shell Dark Brem
- Dark Higgs-strahlung

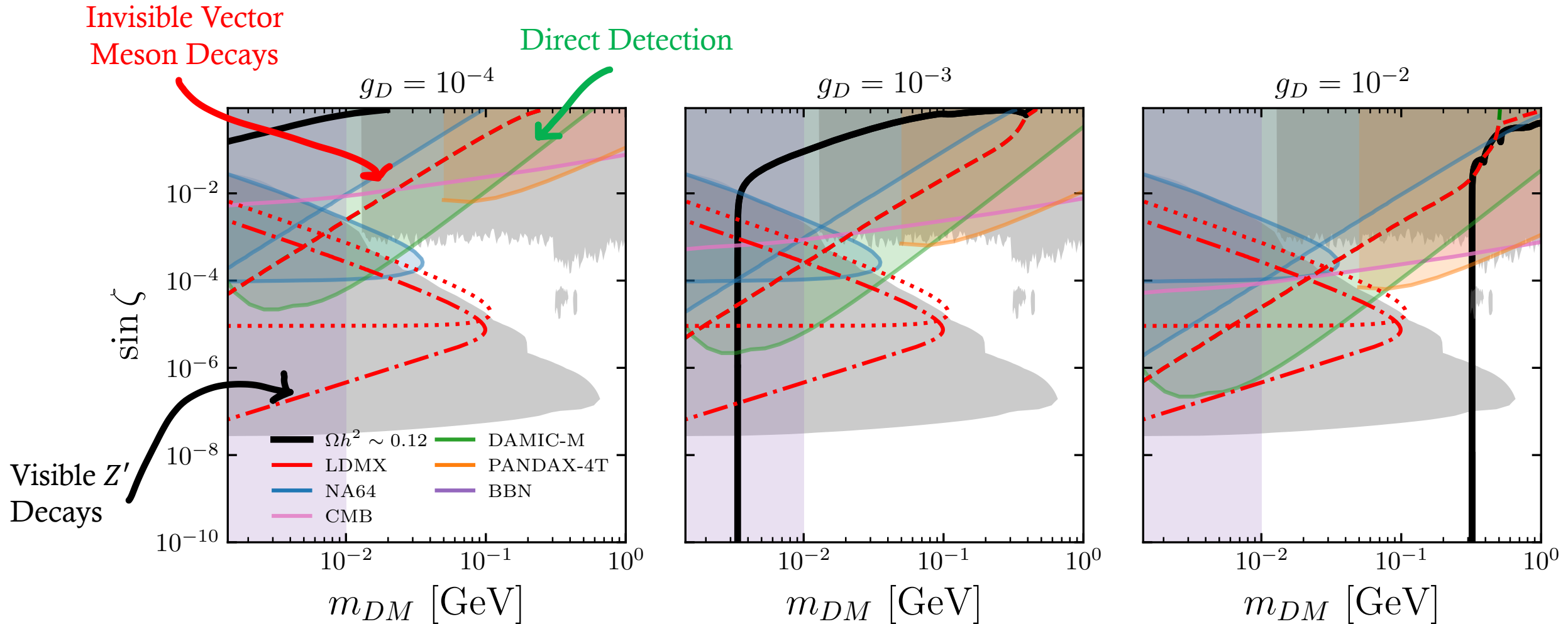


- Invisible Vector Meson Decay

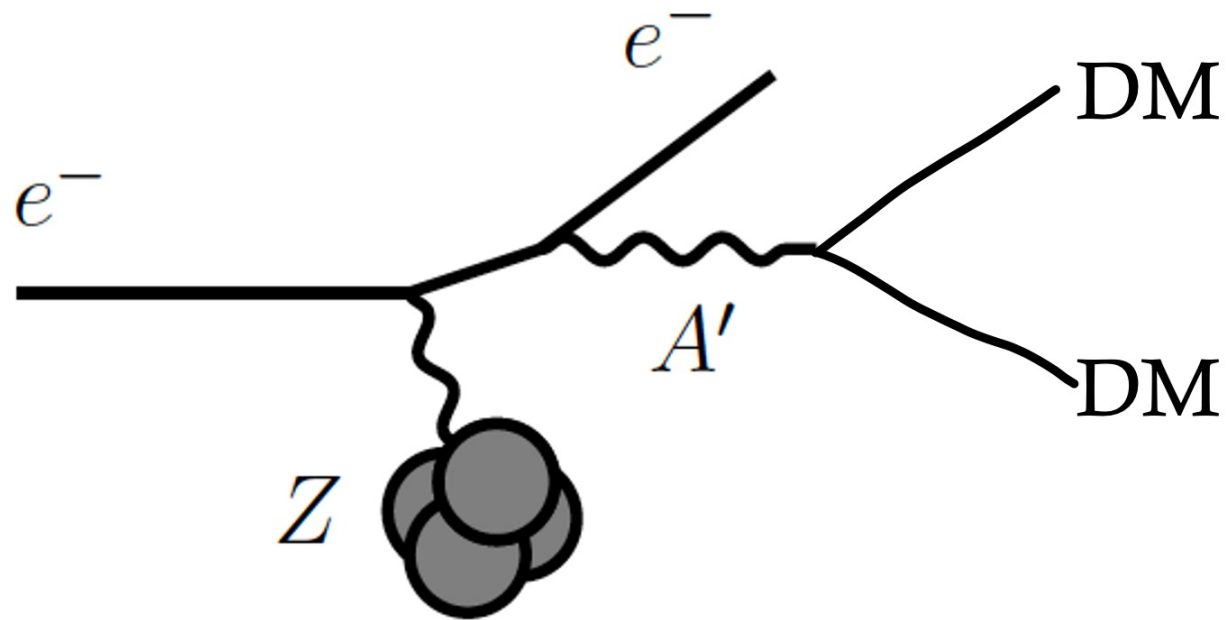


Light Vector Dark Matter Scenario

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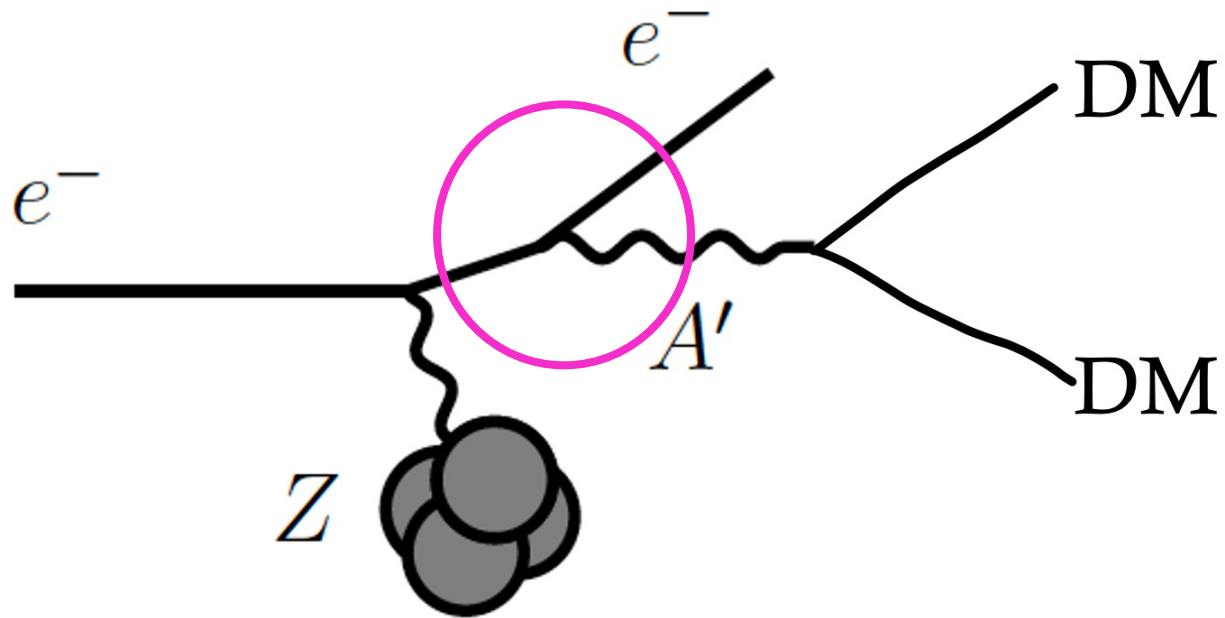


Light Dark Matter Production at Fixed Targets



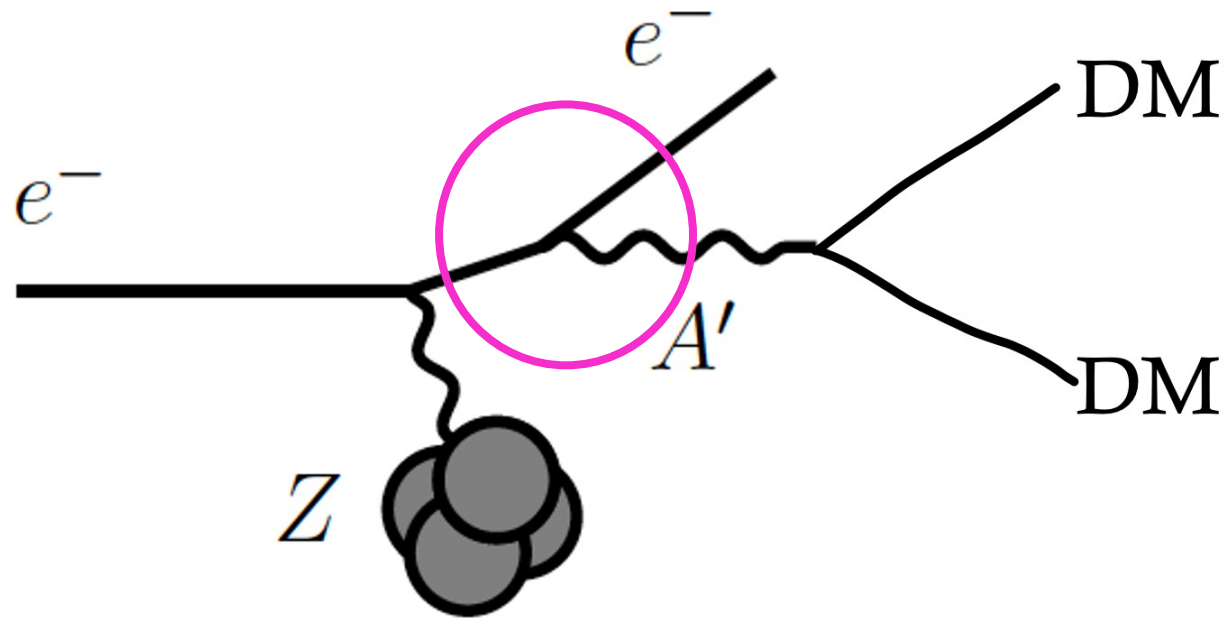
Light Dark Matter Production at Fixed Targets

DM theory landscape: Dark Photon Model



Light Dark Matter Production at Fixed Targets

DM theory landscape: Dark Photon Model



Beyond ordinary
kinetic mixing!

Production of Dark Photons through Higher Electromagnetic Moments at LDMX

- ordinary kinetic mixing (KM)
- magnetic dipole (M)
- electric dipole (E)
- charge radius (C)
- anapole moment (A)

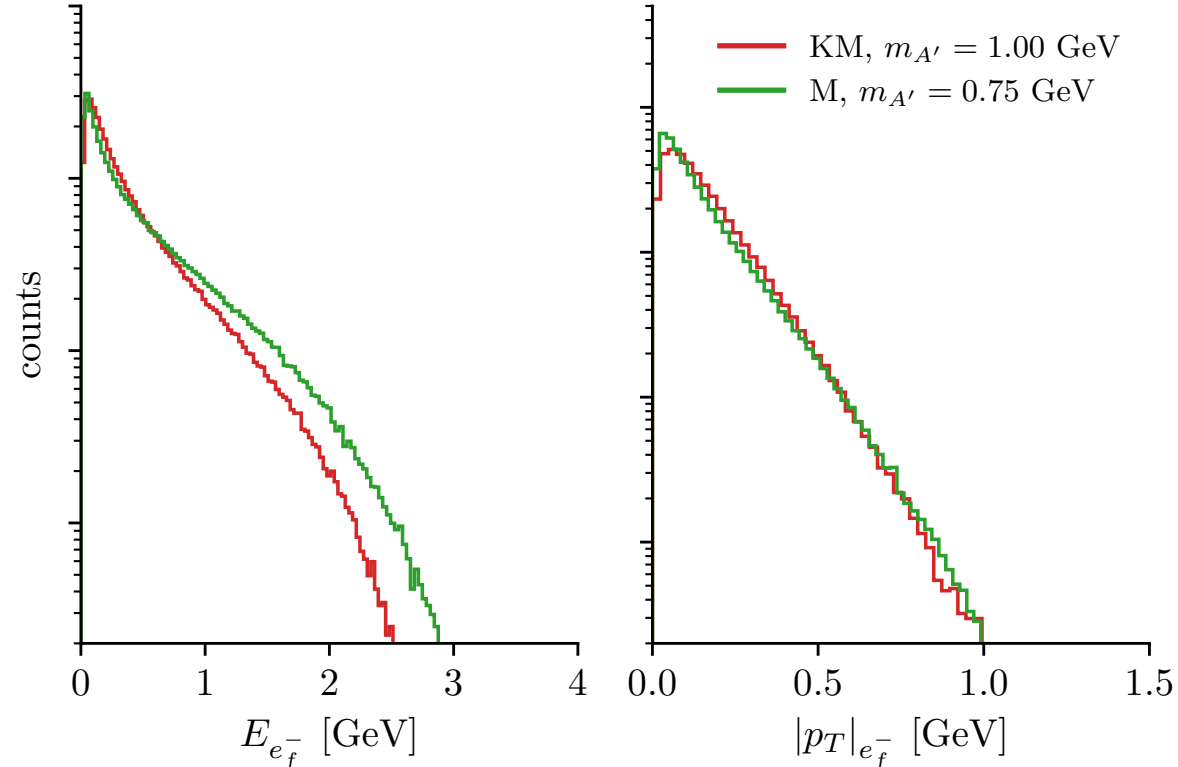
$$\begin{aligned}\mathcal{L}_{A'f} = & \boxed{-e\epsilon Q_f \bar{f} \gamma^\mu f A'_\mu} - \boxed{i\mu_f \partial_\nu (\bar{f} \sigma^{\mu\nu} f) A'_\mu} + d_f \partial_\nu (\bar{f} \sigma^{\mu\nu} \gamma^5 f) A'_\mu \\ & - b_f [\partial^\nu \partial_\nu (\bar{f} \gamma^\mu f) - \partial^\mu \partial_\nu (\bar{f} \gamma^\nu f)] A'_\mu \\ & + a_f [\partial^\nu \partial_\nu (\bar{f} \gamma^\mu \gamma^5 f) - \partial^\mu \partial_\nu (\bar{f} \gamma^\nu \gamma^5 f)] A'_\mu,\end{aligned}$$

[R. Catena, T. Gray, T. Jerkvall]

Production of Dark Photons through Higher Electromagnetic Moments at LDMX

- ordinary kinetic mixing (KM)
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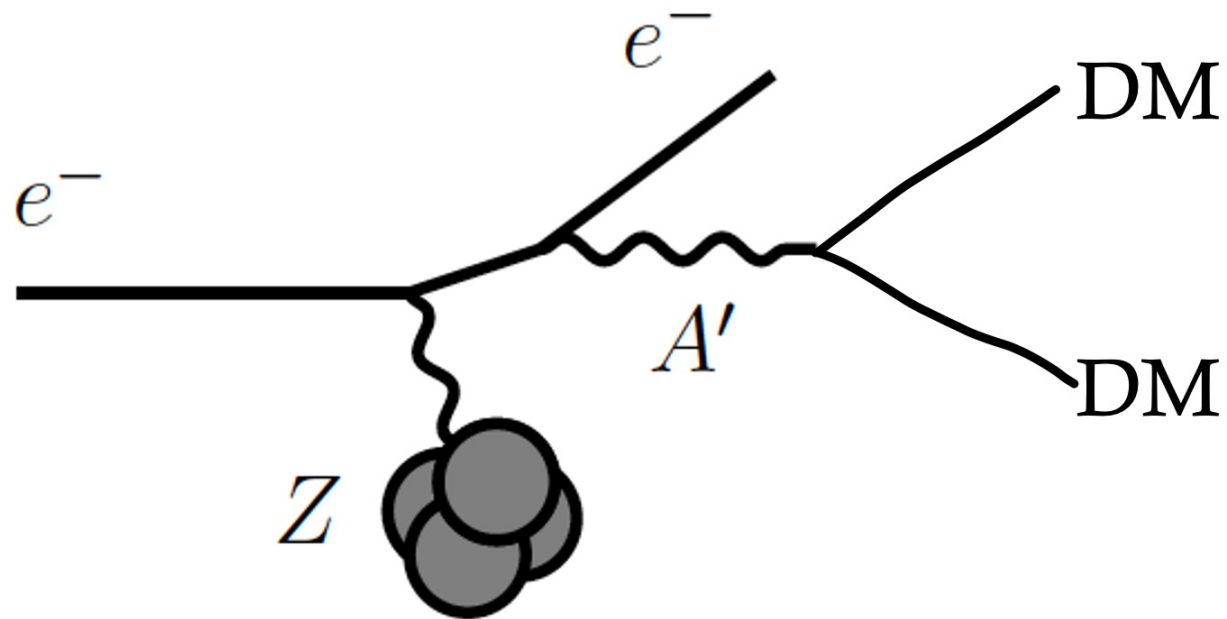
$$\mathcal{L}_{A'f} = \boxed{-e\varepsilon Q_f \bar{f} \gamma^\mu f A'_\mu} - \boxed{i\mu_f \partial_\nu (\bar{f} \sigma^{\mu\nu} f) A'_\mu} + d_f \partial_\nu (\bar{f} \sigma^{\mu\nu} \gamma^5 f) A'_\mu \\ - b_f [\partial^\nu \partial_\nu (\bar{f} \gamma^\mu f) - \partial^\mu \partial_\nu (\bar{f} \gamma^\nu f)] A'_\mu \\ + a_f [\partial^\nu \partial_\nu (\bar{f} \gamma^\mu \gamma^5 f) - \partial^\mu \partial_\nu (\bar{f} \gamma^\nu \gamma^5 f)] A'_\mu,$$



- In addition to **mass reconstruction**
→ DP **model reconstruction?**
- **Break degeneracy** between groups with both E_e and $|p_T|_E$

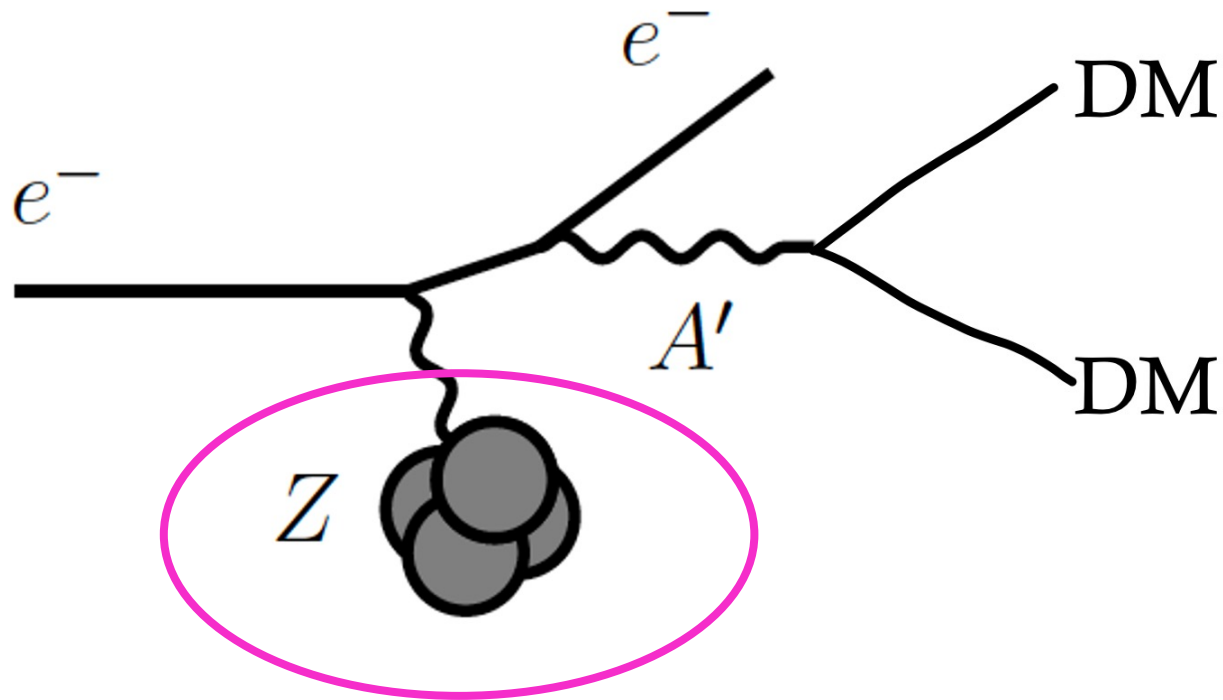
[R. Catena, T. Gray, T. Jerkvall]

Light Dark Matter Production at Fixed Targets



Light Dark Matter Production at Fixed Targets

Ab-Initio Modeling of the Target Nucleus
using Many-Body Methods

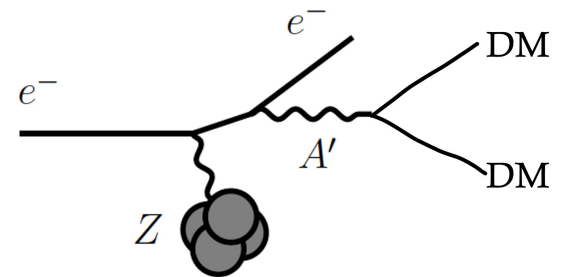


Ab-Initio Modeling of the Target Nucleus

using Many Body Methods

$$N_{\text{events}} \propto \int dq \frac{q^2 - q_{\text{min}}^2}{q^3} |F(q)|^2$$

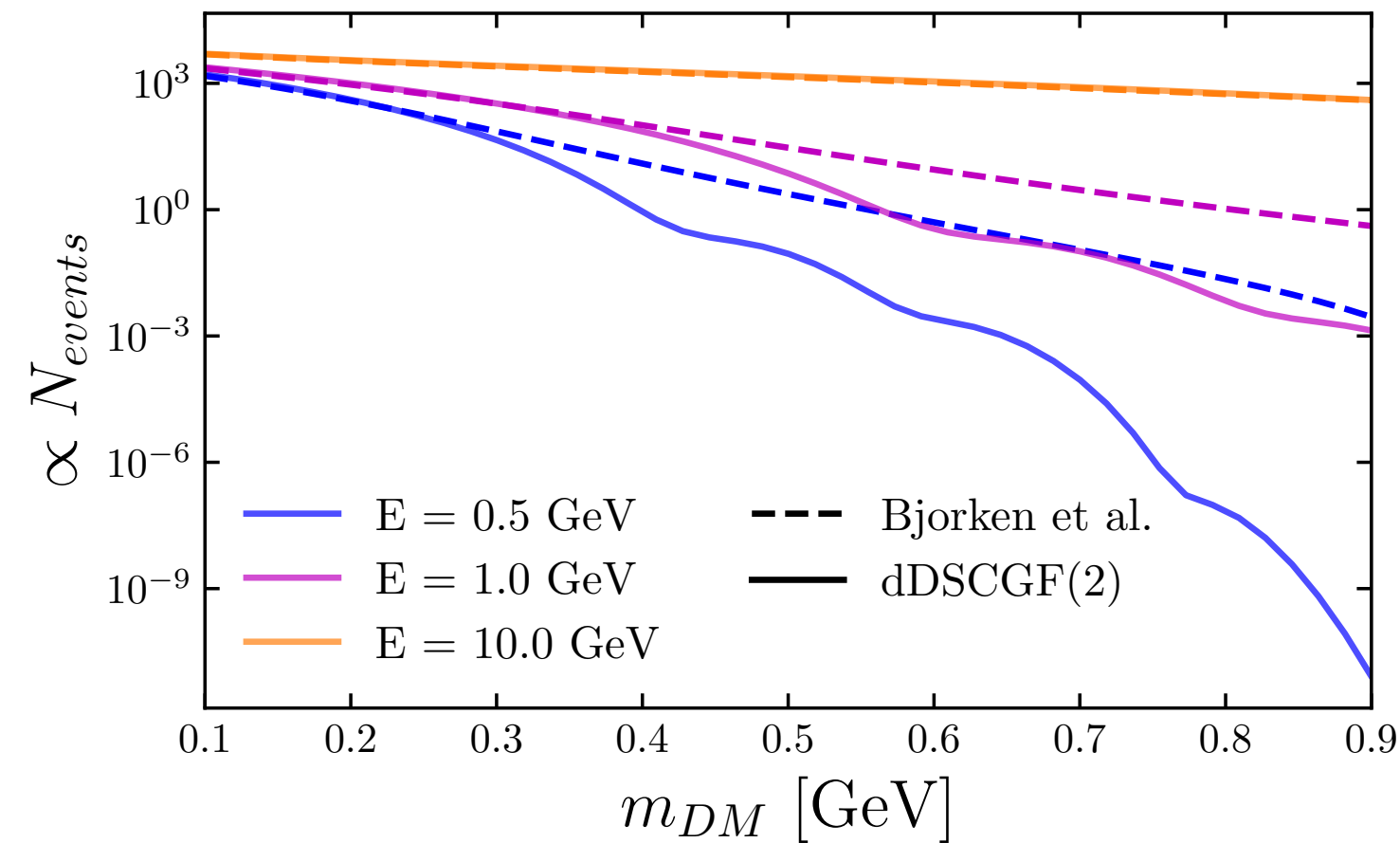
Form Factor



[T. Gray and A. Scalesi *in preparation*]

Ab-Initio Modeling of the Target Nucleus

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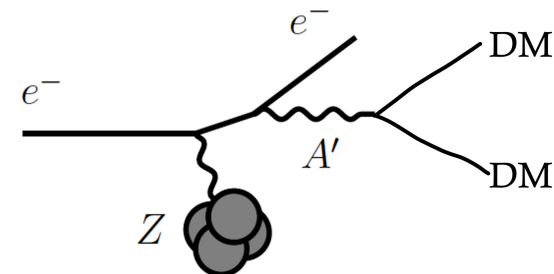


[T. Gray and A. Scalesi *in preparation*]

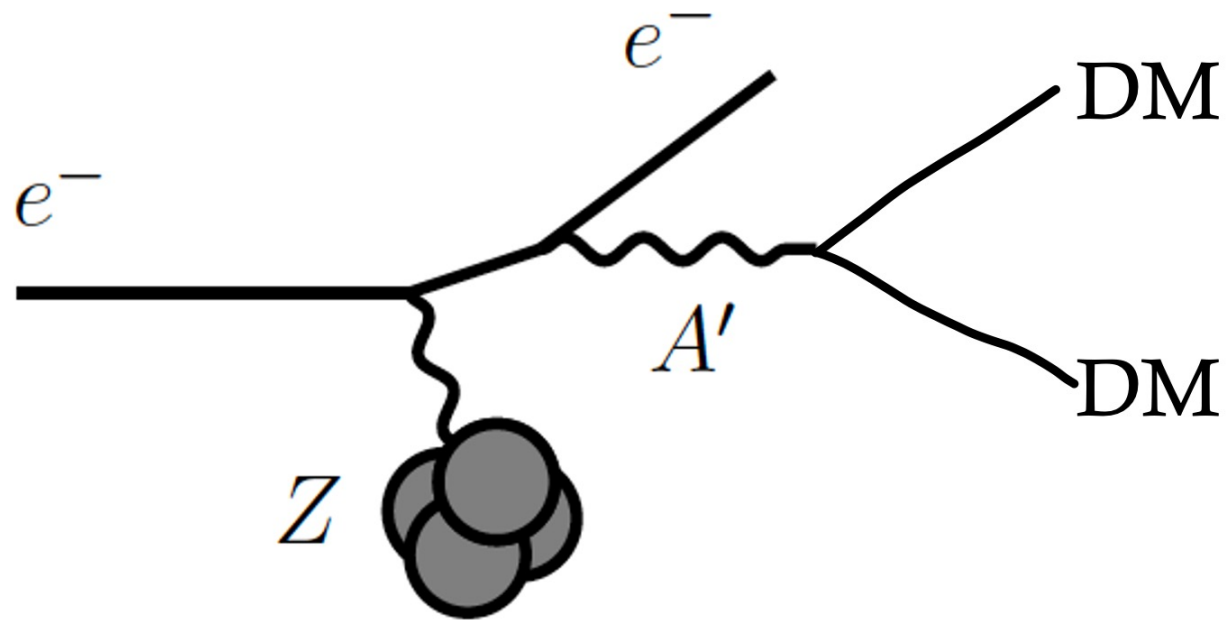
$$N_{events} \propto \int dq \frac{q^2 - q_{min}^2}{q^3} |F(q)|^2$$

Form Factor

Experimental sensitivity depends on nuclear structure!

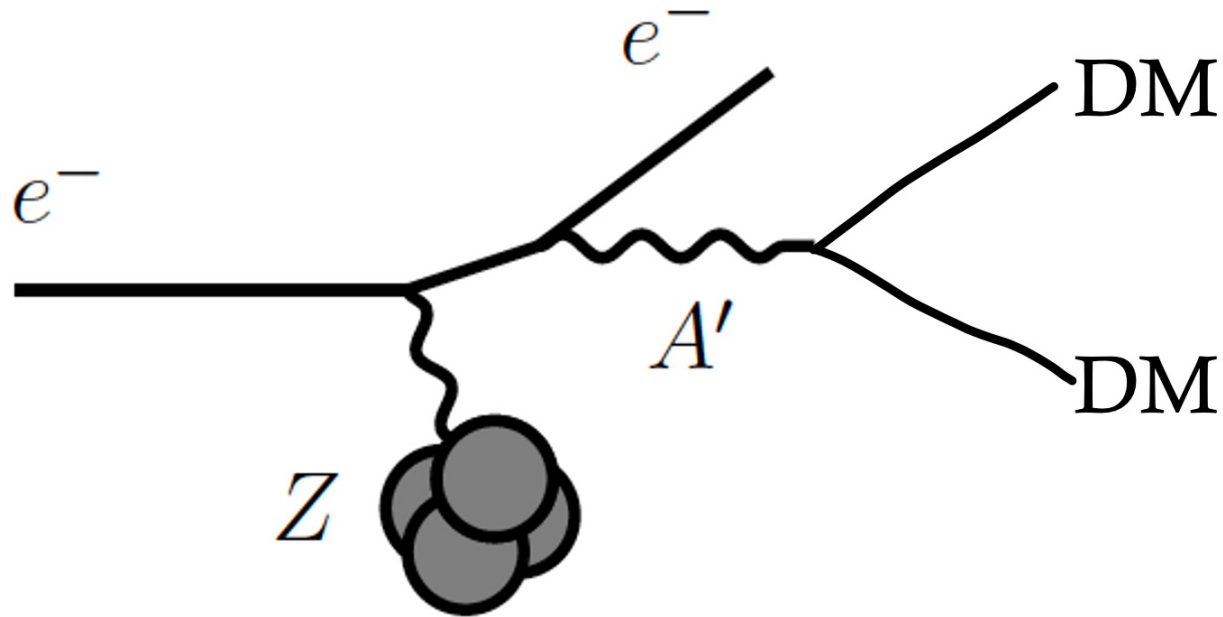


Light Dark Matter Production at Fixed Targets



Light Dark Matter Production at Fixed Targets

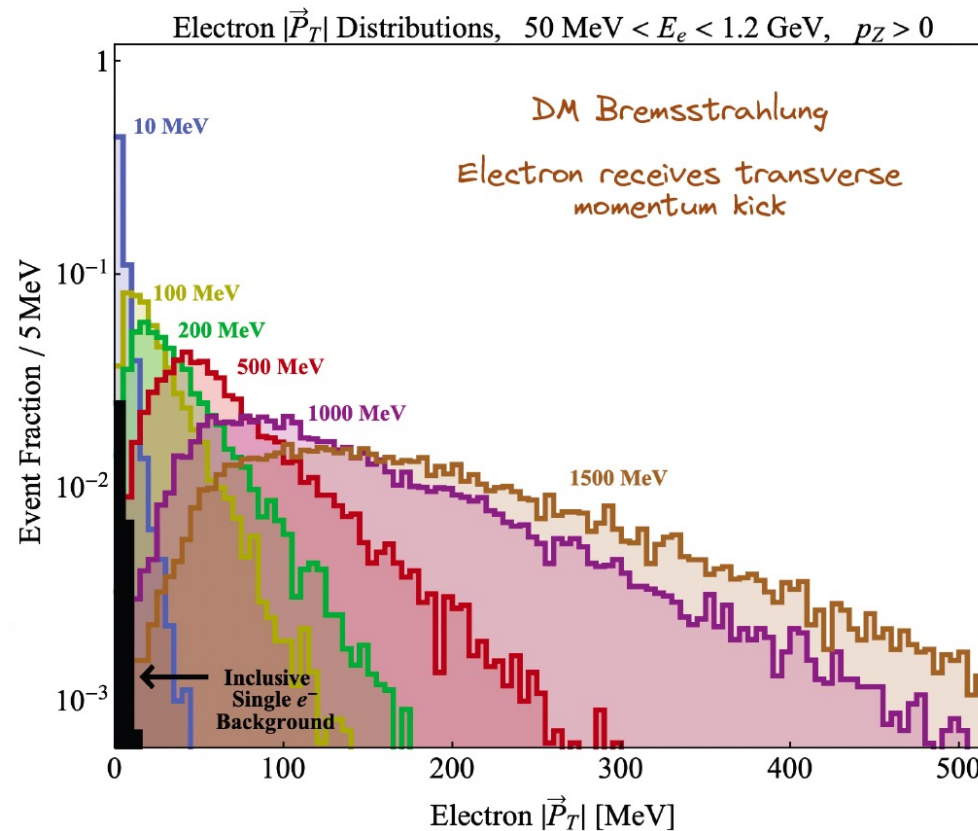
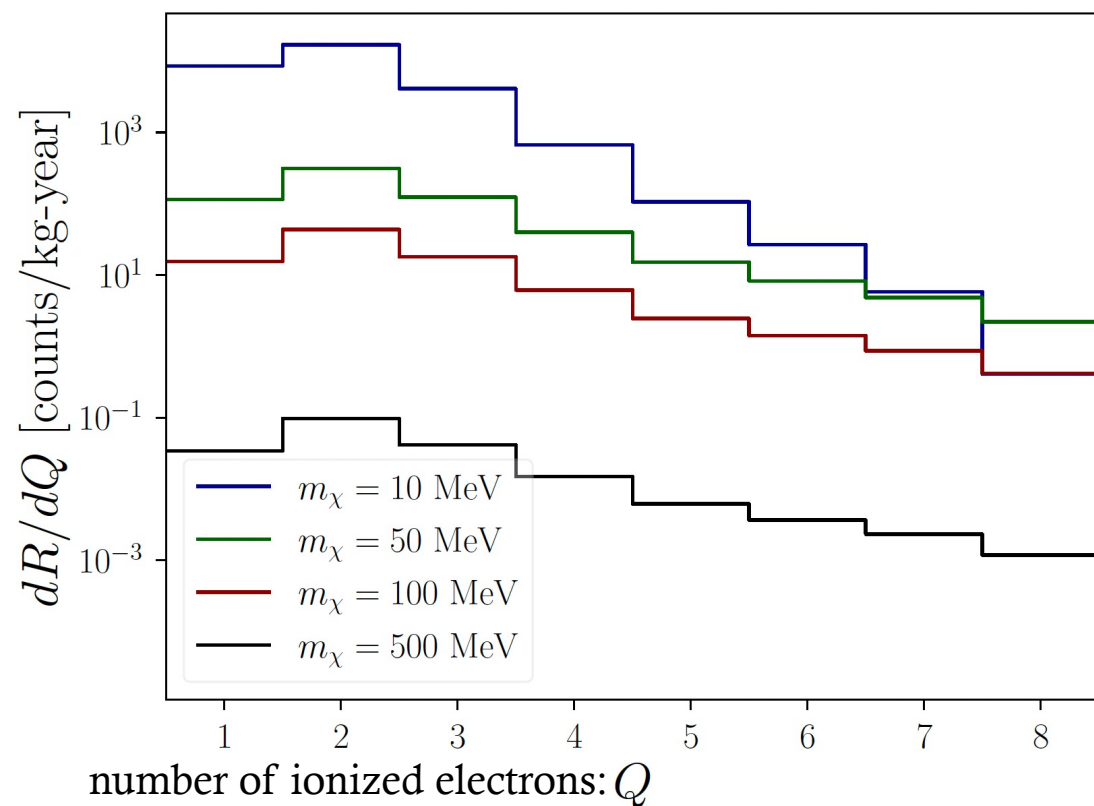
What if we measure a signal? :o



On the dark matter origin of an LDMX signal

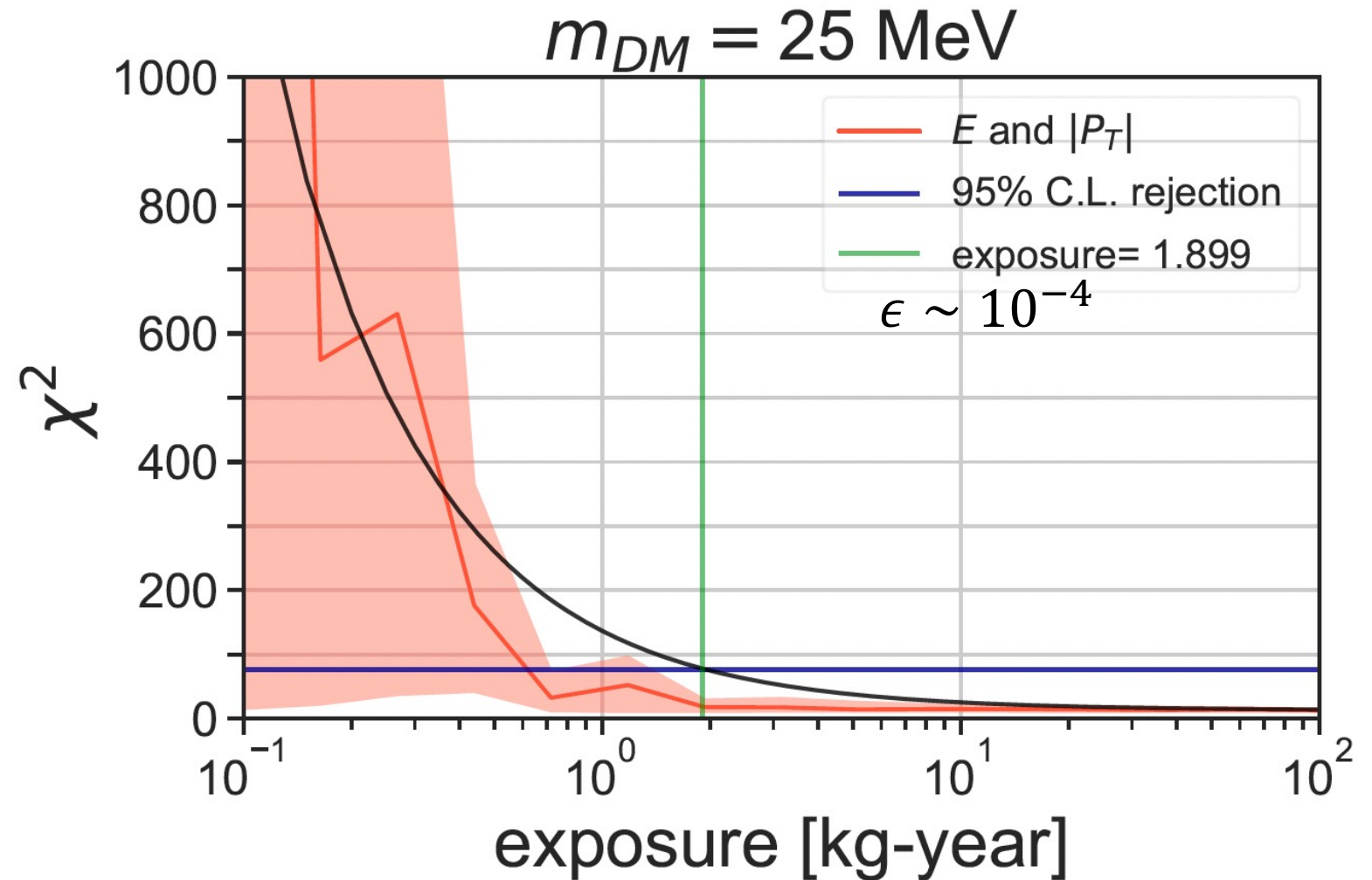
How can we validate that an LDMX signal is of DM nature?

→ a direct detection signal



[R. Catena, T. Gray, A. Lund]

On the dark matter origin of an LDMX signal



On the dark matter origin of an LDMX signal

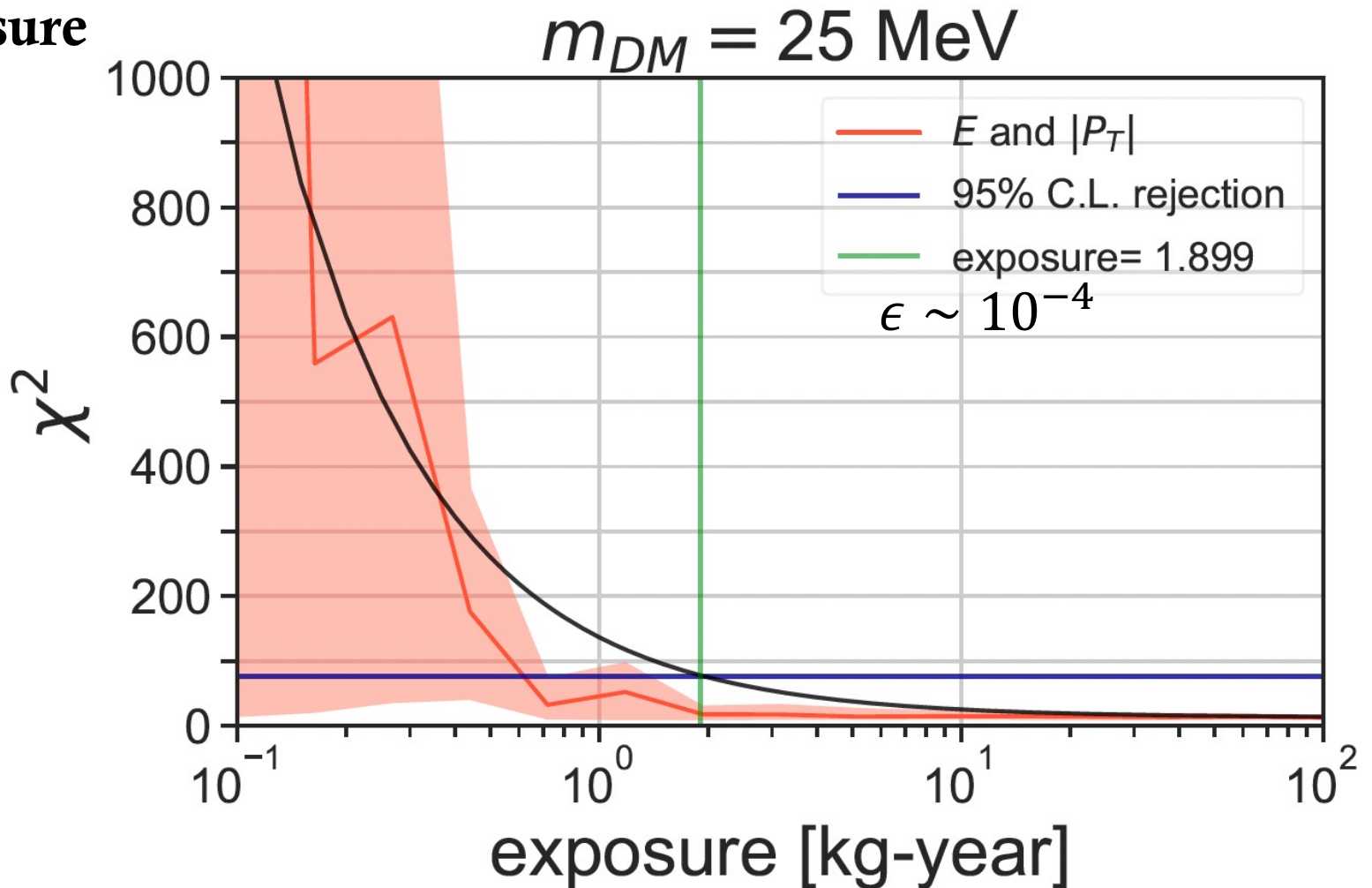
Required Direct Detection Exposure

$m_{DM} = 4 \text{ MeV} \rightarrow 0.0013 \text{ kg-yr}$

$m_{DM} = 25 \text{ MeV} \rightarrow 1.9 \text{ kg-yr}$

Experimental Status

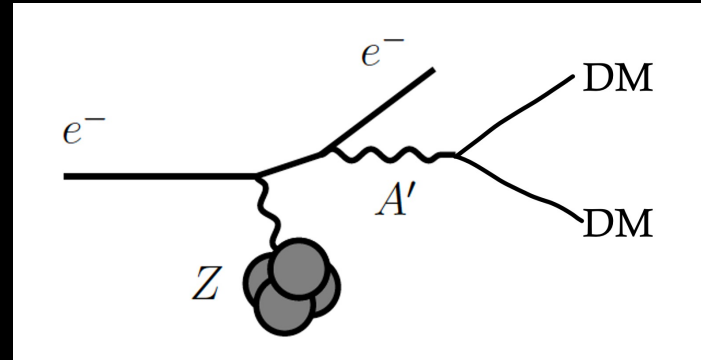
- **OSCURA**
future $\sim 30 \text{ kg-year}$
- **DAMIC-M**
current exposure of 0.0036 kg-yr
future $\sim \text{kg-yr}$
- **SENSEI**
current exposure of 0.00028 kg-yr



Summary

It's an exciting time for Light DM searches!

- upcoming experiments (such as **LDMX**)
- extending the sensitivity frontier for a rich array of **light DM theories**
- However, we need to:
 - Know how to deal with a **potential signal**
 - Properly model the signal + background
 - Nuclear structure can play a role!



Thank you for listening 😊

Additional Material



Light Dark Matter Status

Indirect Detection

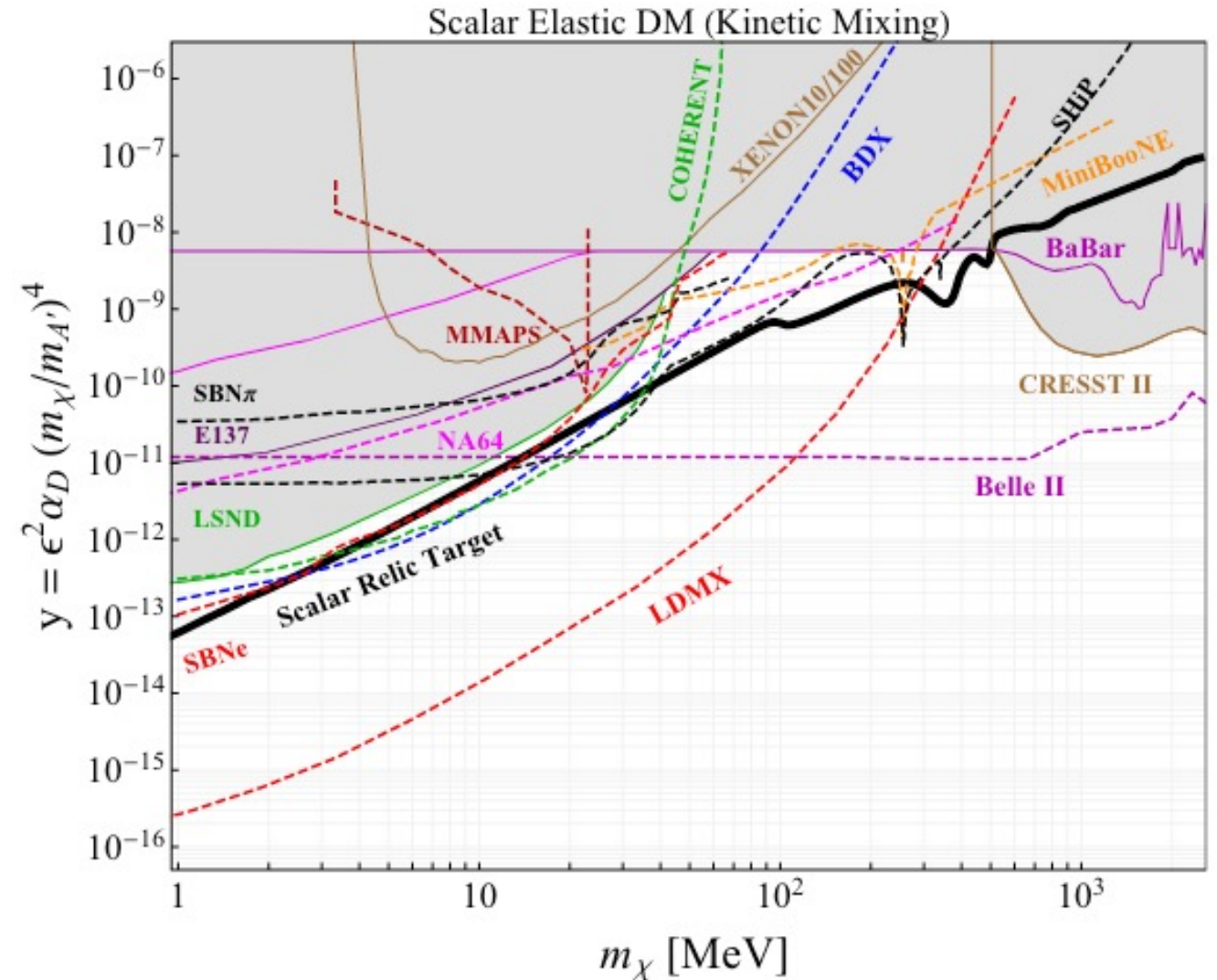
- CMB energy injection

Direct Detection

- Electron recoils
- Phonon excitations

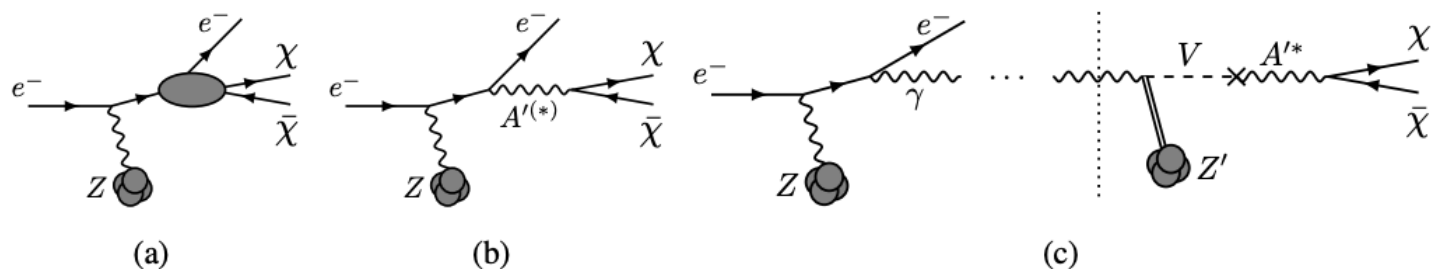
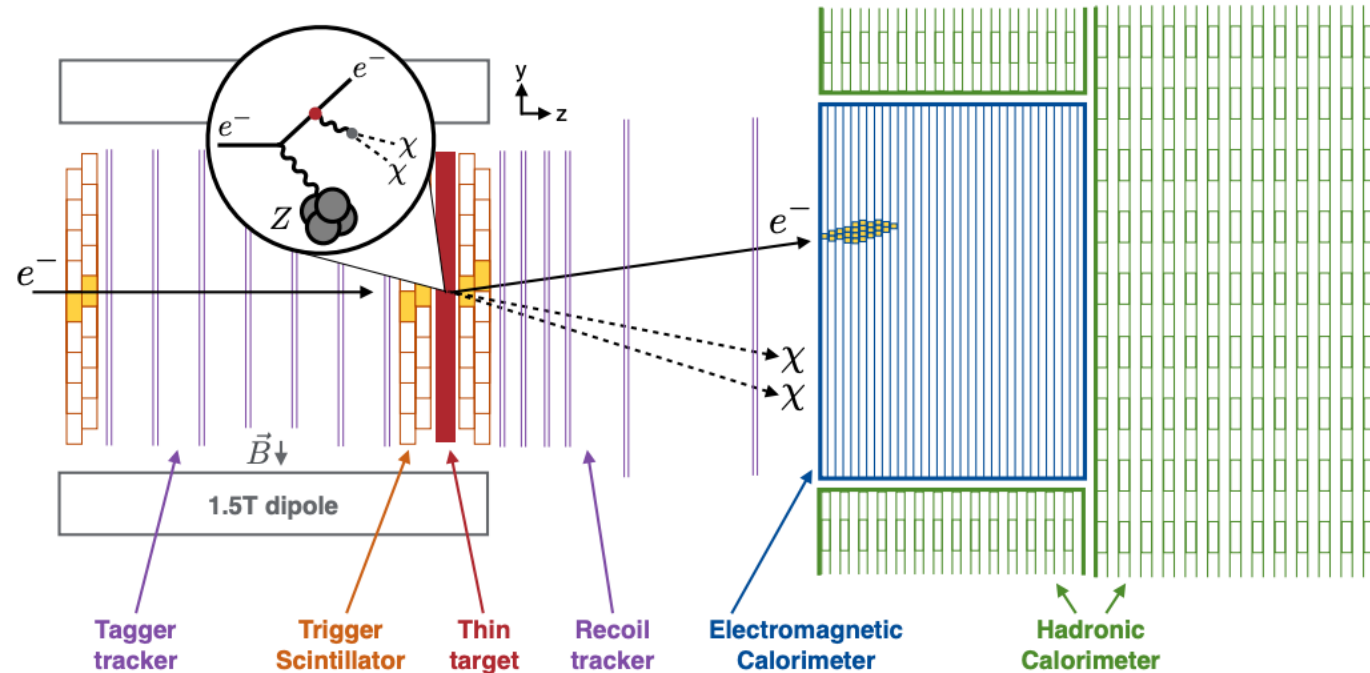
Collider

- Mono- γ
- Beam dumps
- **Fixed target**



Light Dark Matter eXperiment (LDMX)

- Future fixed target missing momentum exp
 - 2025: LESA delivers beam to LDMX allowing 4×10^{14} EOT
 - 2027: 10^{16} EOT
- e^- incident on a thin tungsten target
- Charged particle tracker and calorimeters to measure DM signature
 - Recoil electron pT accompanied by absence of other particle activity

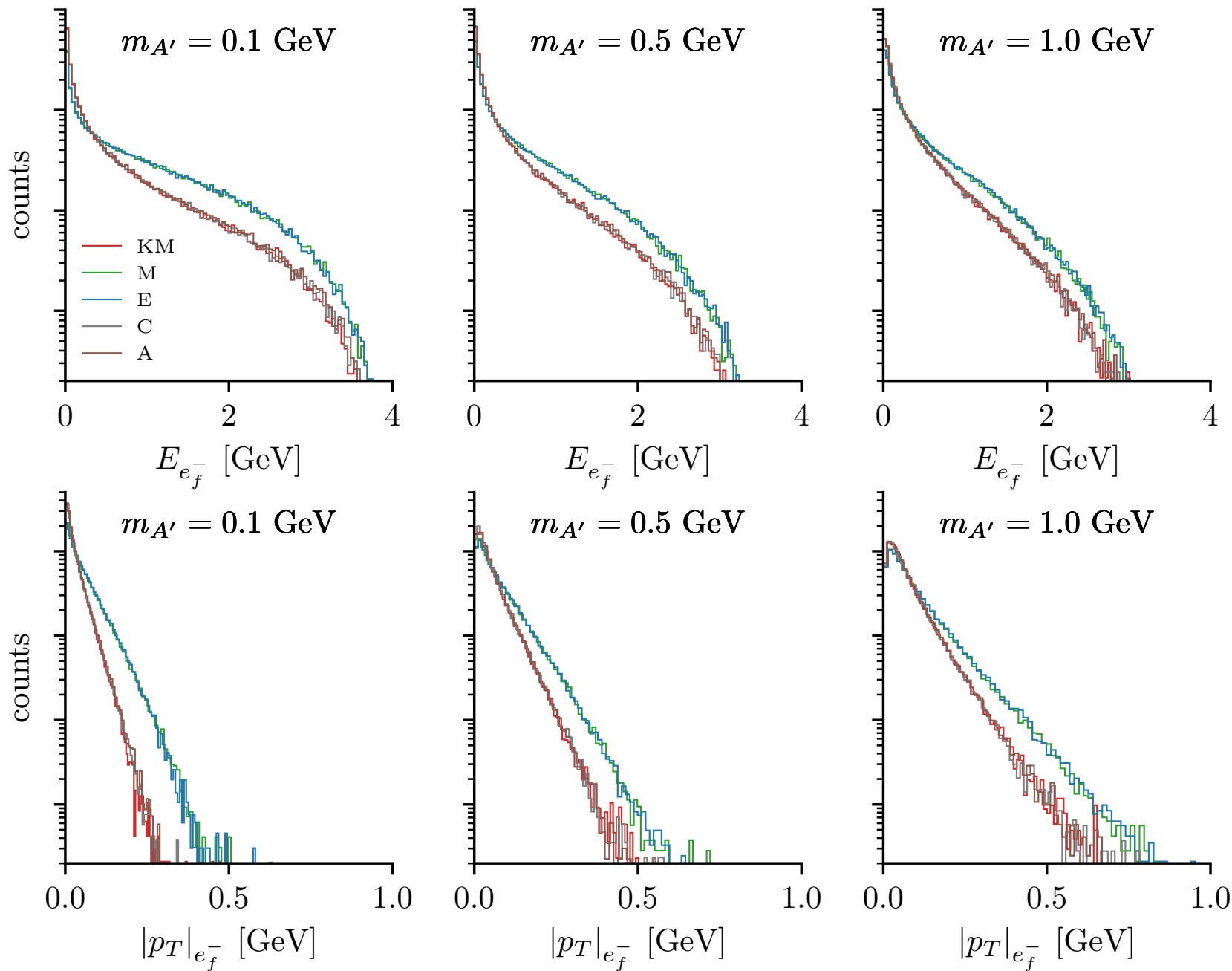


STATUS

- Completed tests of upstream beam (results available soon)
- LDMX beam into experimental area and the vertical slice test (This year)
- Measurements with HCal prototype in the experimental hall (early 2026)
- Begin construction in late 2026 (~3 years)

Kinematic Distributions at LDMX

- 3 groups
- Characterized by **overall cross section size** and **momentum dependence in Lorentz structures** of models



LDMX phase II projected exclusion bounds

