

Galactic Cosmic Ray spectral measurements with the DAMPE space mission



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on behalf of the DAMPE Collaboration



DAMPE science goals



High energy particle detection in space

- Study of the cosmic electron spectrum
- Study of cosmic ray protons and nuclei
- High energy gamma ray astronomy
- Search for dark matter signatures in e/γ spectra



Detection of

10 GeV - 10 TeV e/γ

50 GeV – 0.5 PeV protons and nuclei

**with excellent (e.m.) energy resolution , tracking precision
and particle identification capabilities**

- Exotica and “unexpected” , e.g. GW e.m. counterpart in the FoV

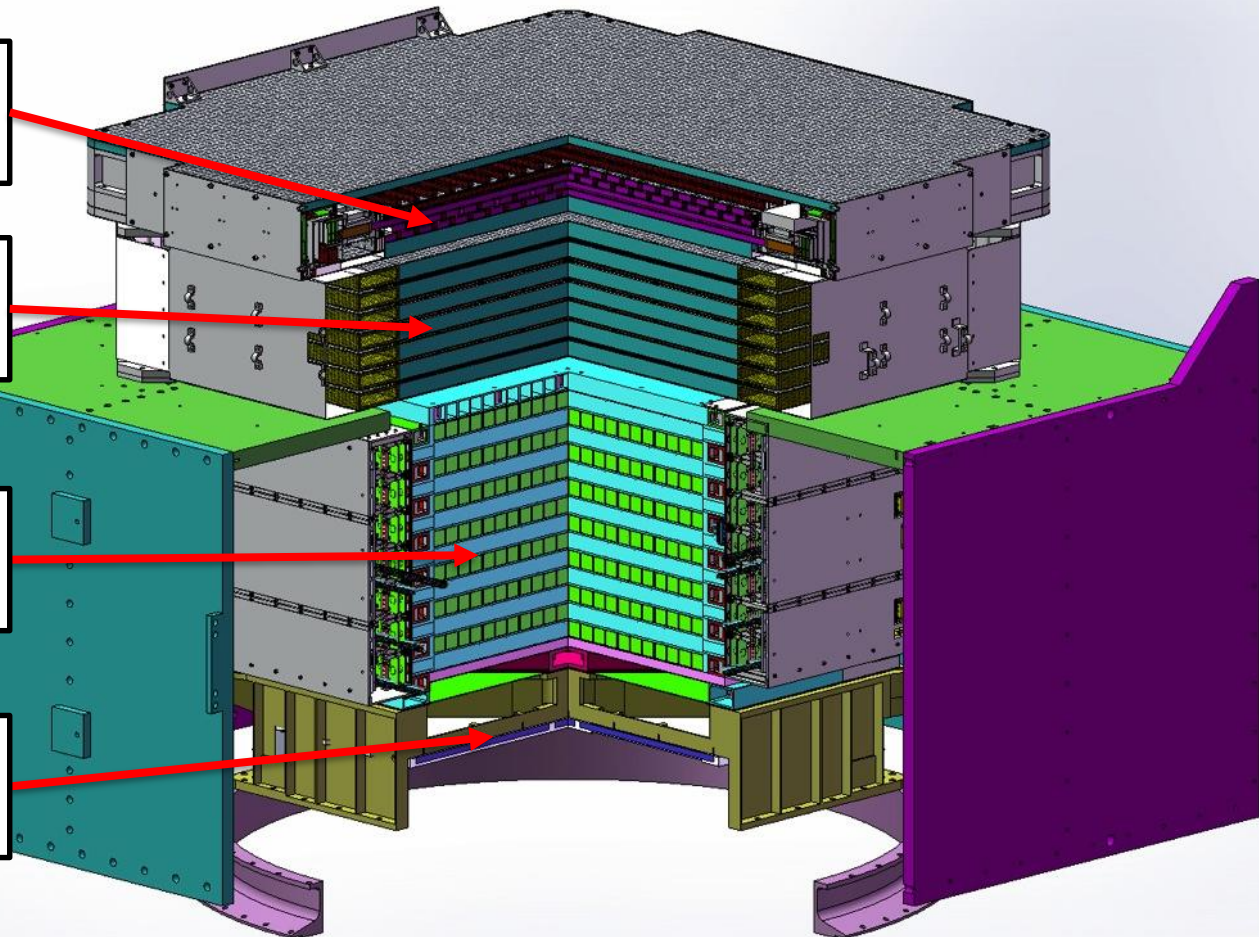
The detector

Plastic Scintillator Detector
(PSD) ($\sim 0.2e$)

Silicon-Tungsten Tracker
(STK) ($\sim 50\mu m$)

BGO Calorimeter
(CALO) ($31 x_0$)

Neutron Detector
(NUD)



- Charge measurement (dE/dx in PSD , STK and BGO)
- Tungsten converter (pair production)
- Precise tracking (silicon strips)
- Thick calorimeter (BGO bars)
- Hadron rejection (neutron detector)



high energy
 γ -ray, electron and cosmic ray
telescope

The launch: Dec 17th 2015, 0:12 UTC

Jiuquan Satellite Launch Center
Gobi desert

CZ-2D rocket

Mass: 1850 kg (scientific payload 1400 kg)

Power : 640 W (scientific payload 400 W)

Orbit: sun synchronous

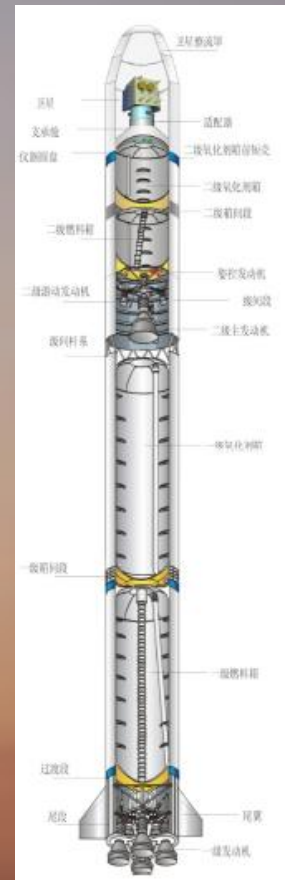
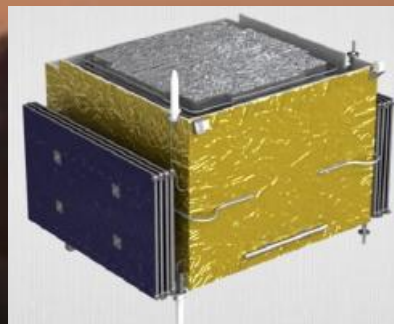
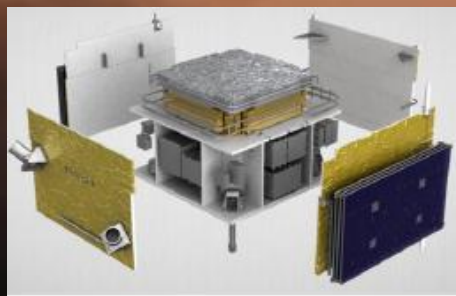
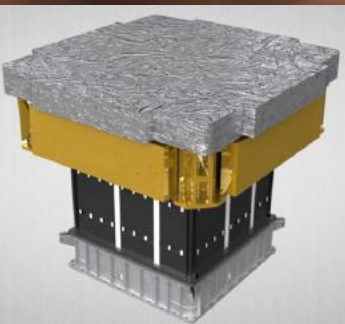
Altitude: 500km

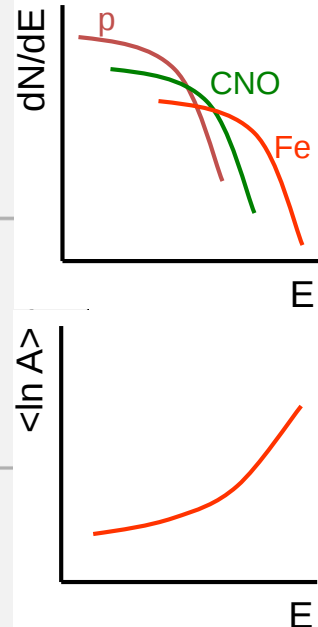
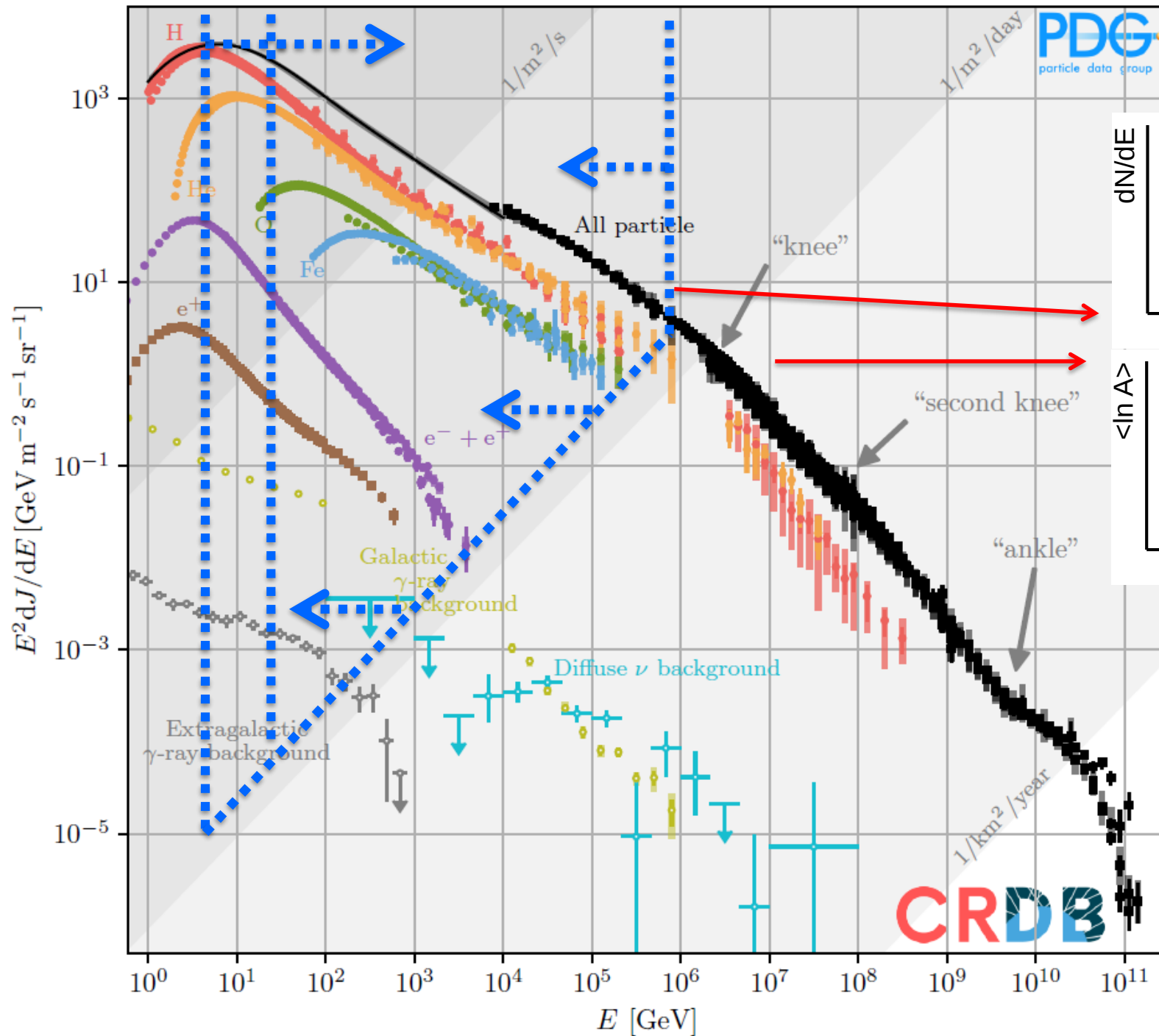
Inclination: 97.41°

Period: 95 minutes

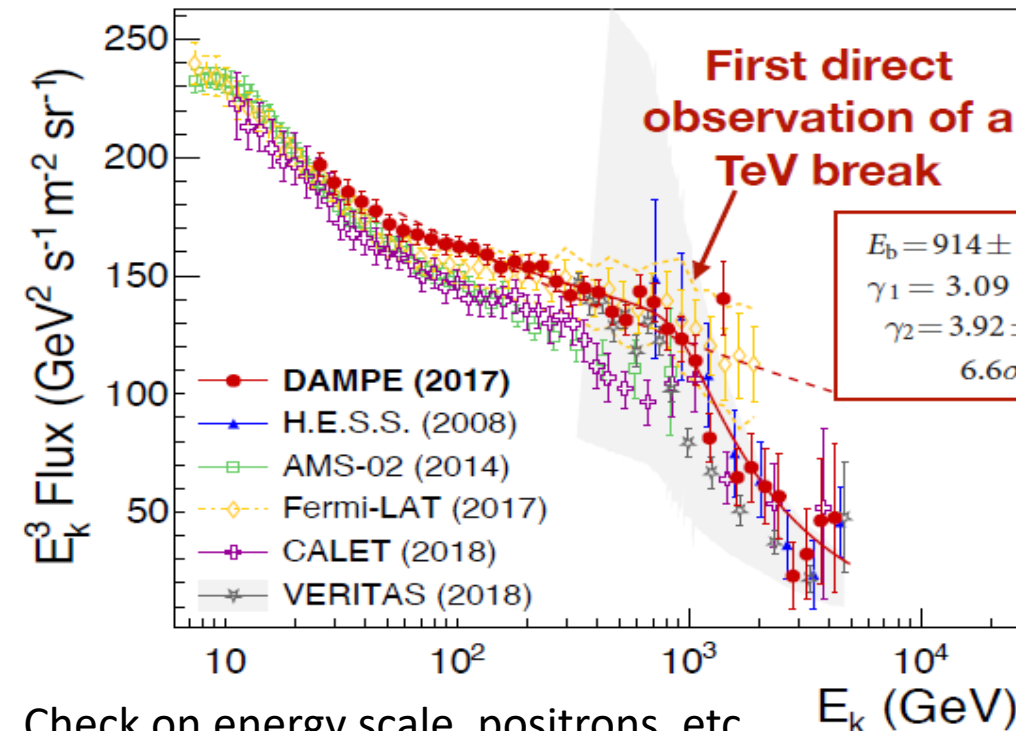
Downlink: 16 GB / day

Lifetime: > 3 years





The DAMPE ($e^+ + e^-$) spectrum



LETTER

nature

International weekly journal of science

doi:10.1038/nature24475

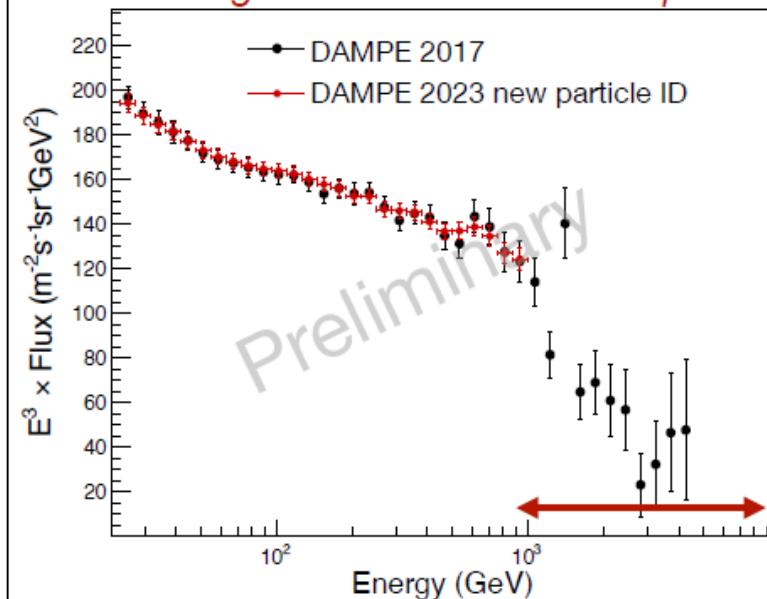
Direct detection of a break in the teraelectronvolt cosmic-ray spectrum of electrons and positrons

DAMPE Collaboration*

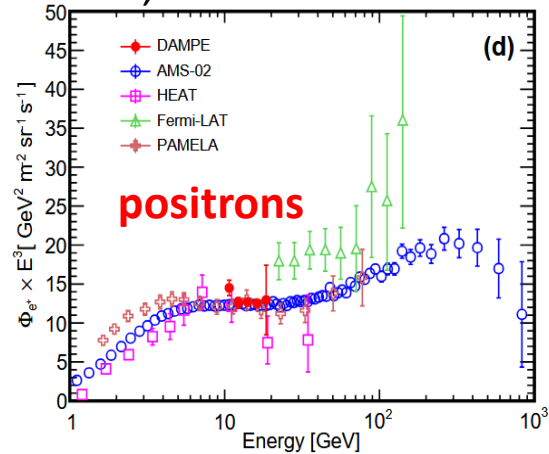
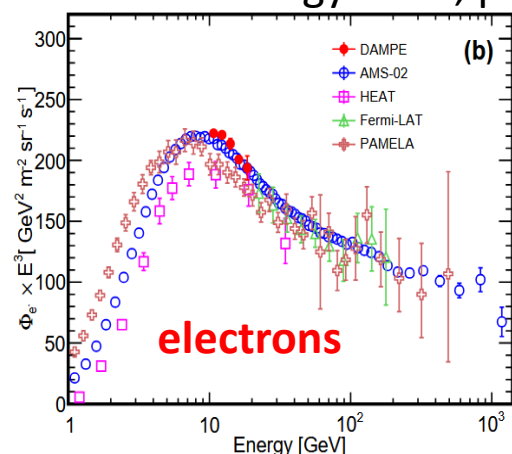
- 530 days
- 2.8 billions CR events
- 1.5 million CREs above 25 GeV

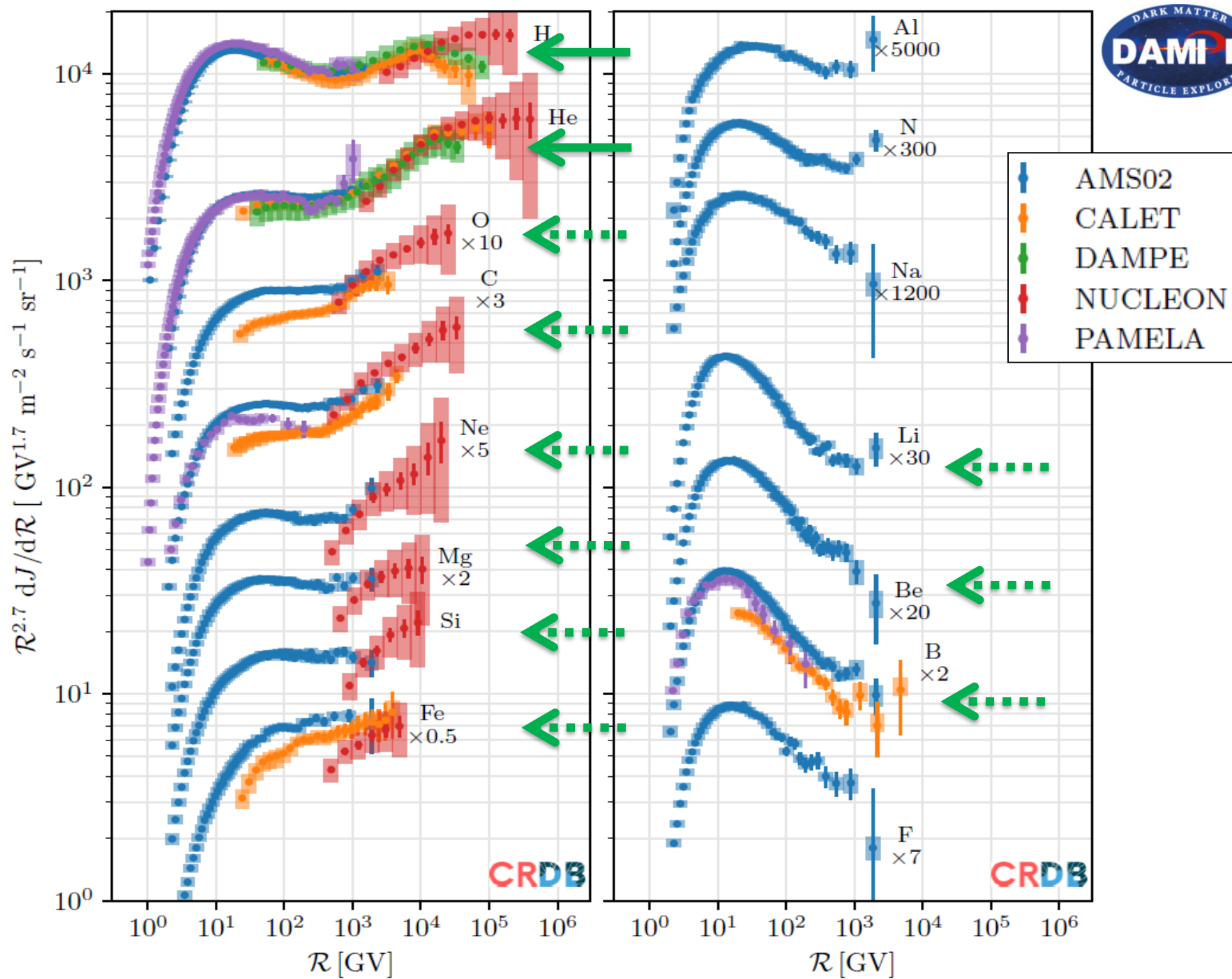
New analyses (NN, NL, ..) ongoing

Excellent agreement with standard particle ID

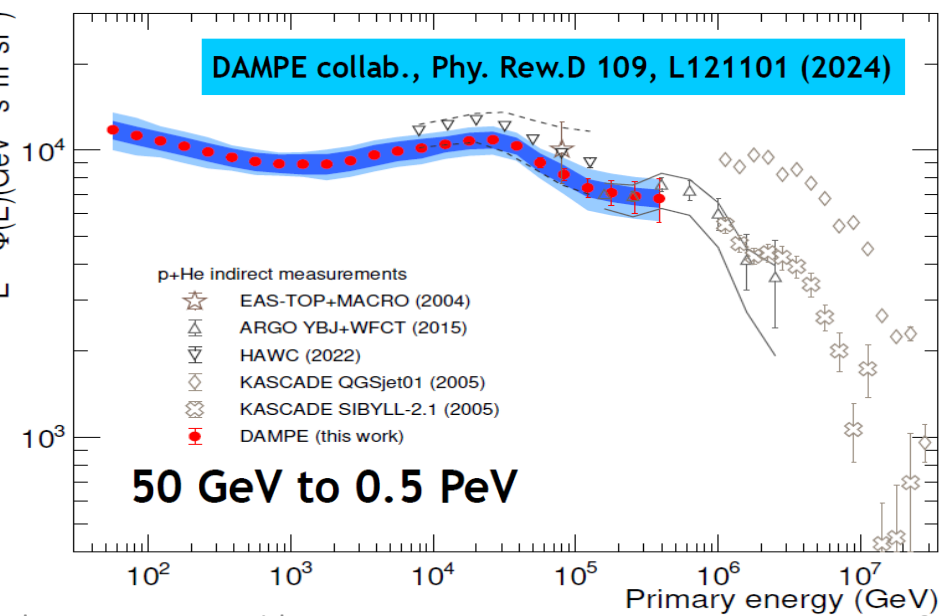
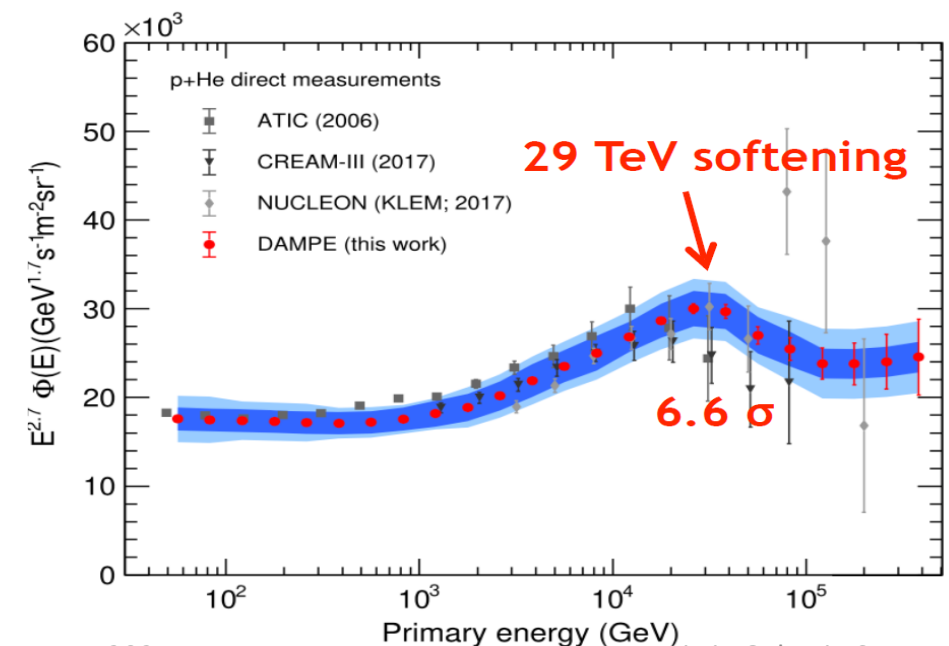
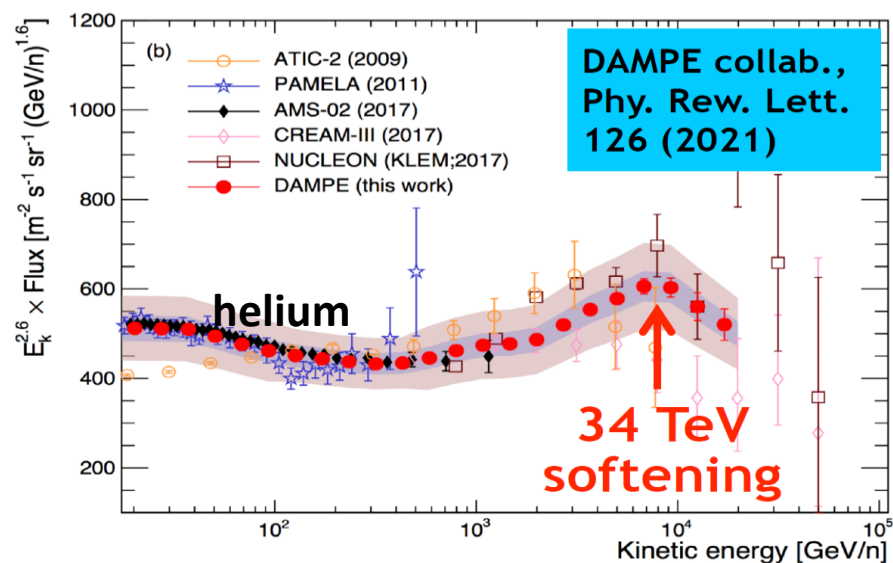
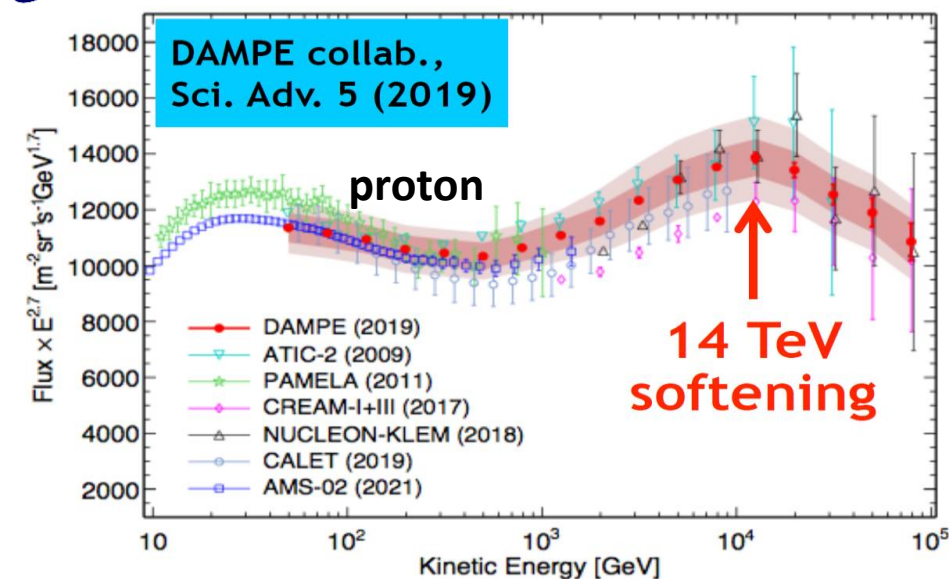


Check on energy scale, positrons, etc..

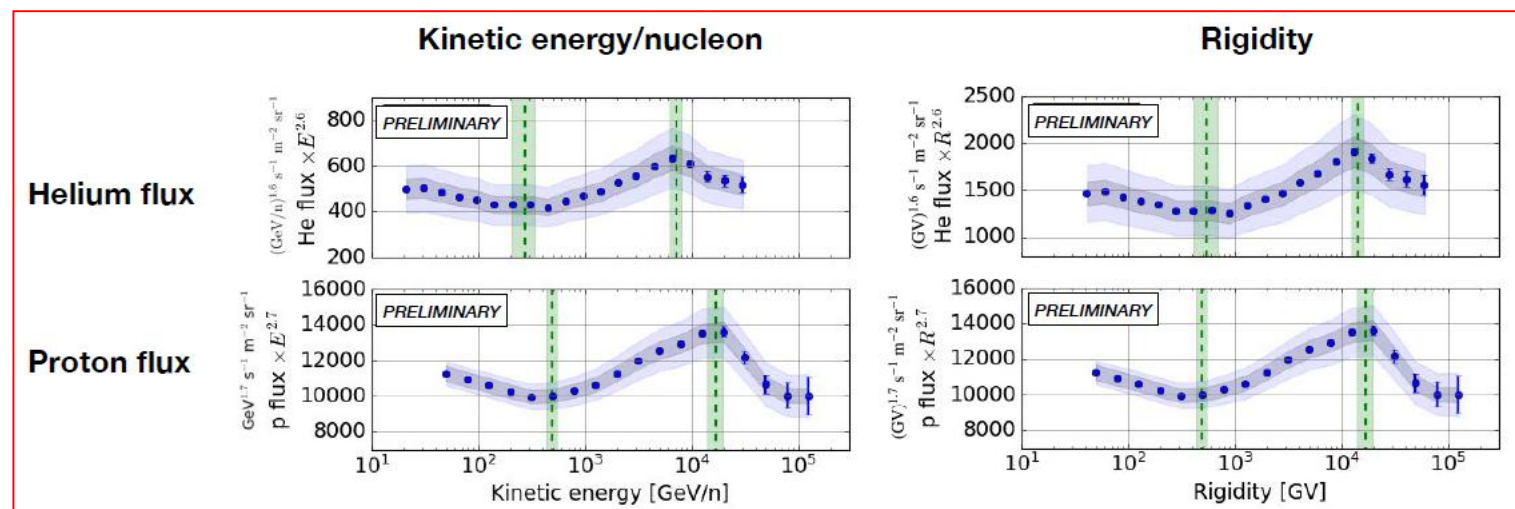
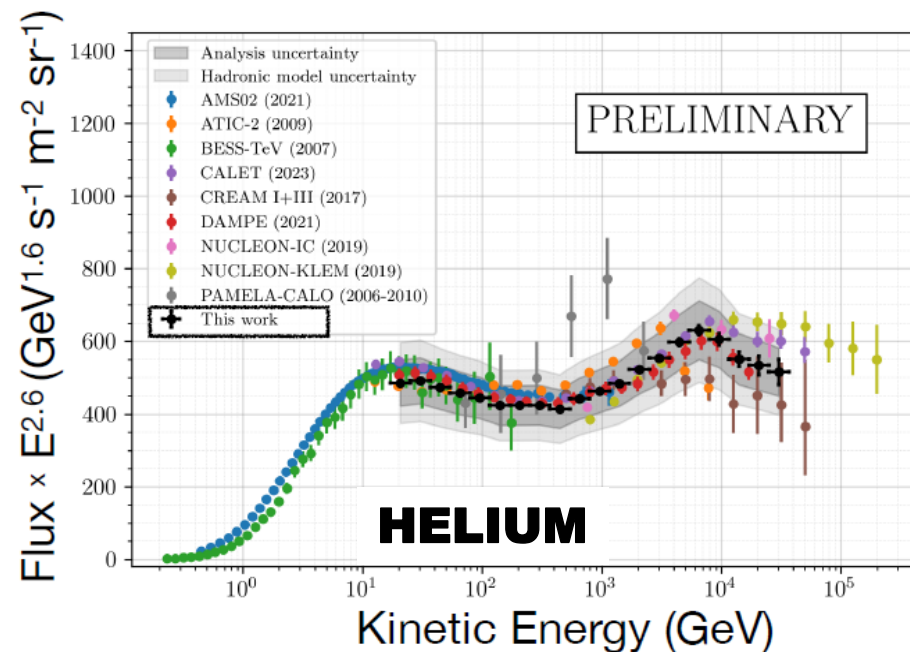
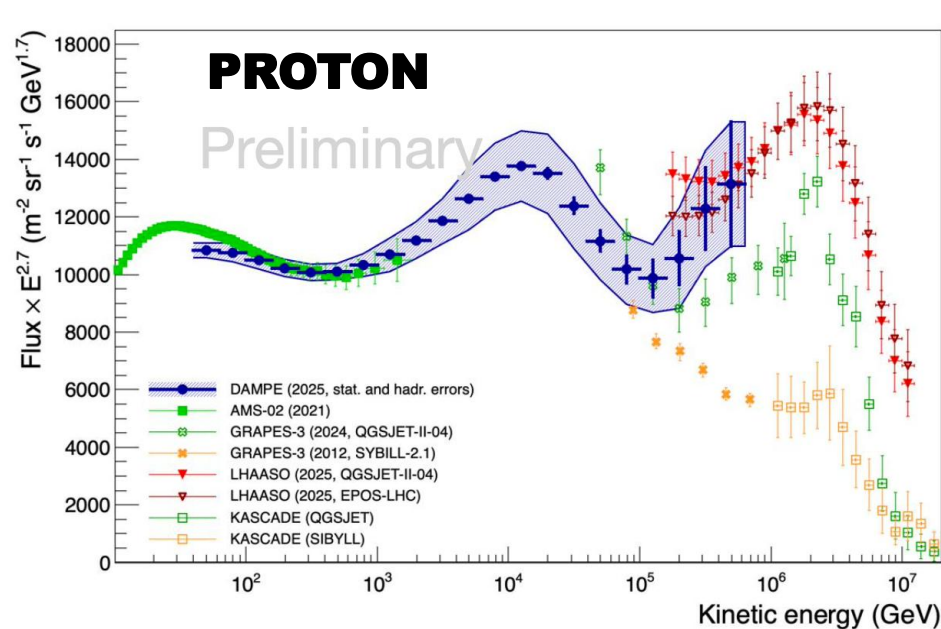




Proton, Helium, p+He



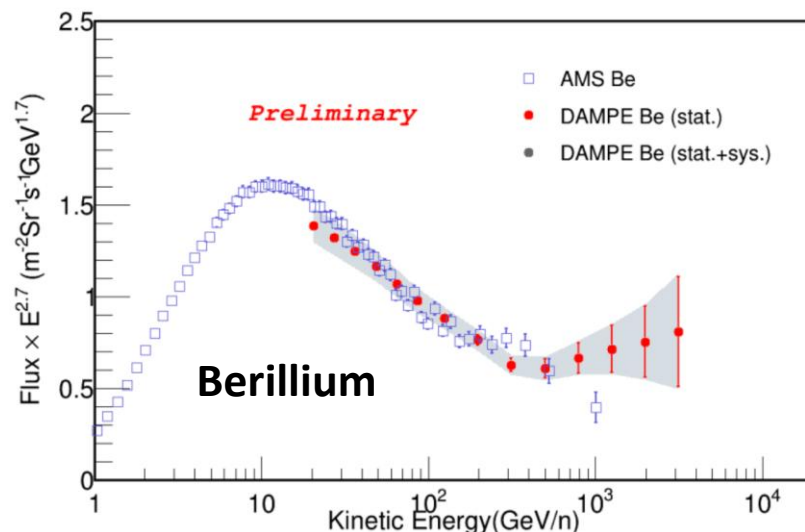
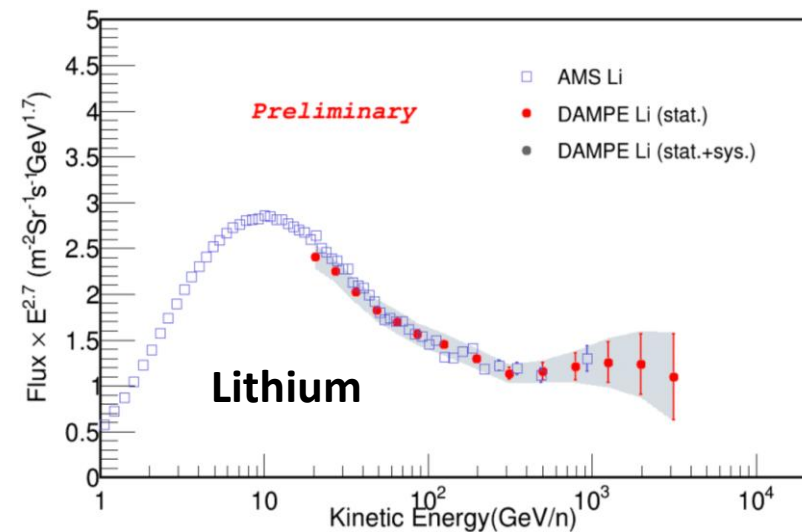
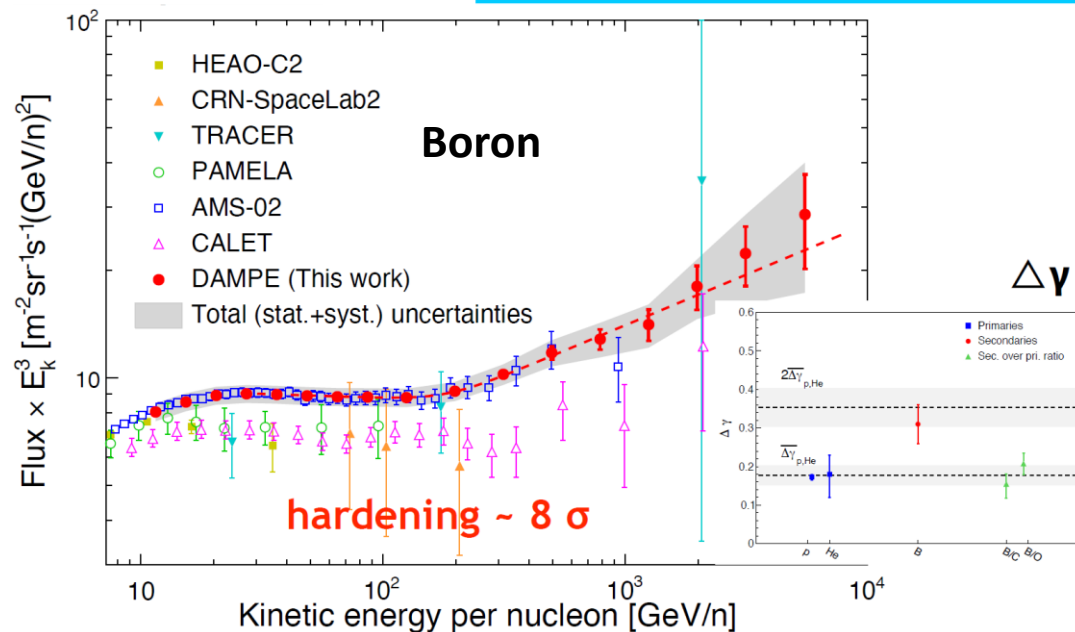
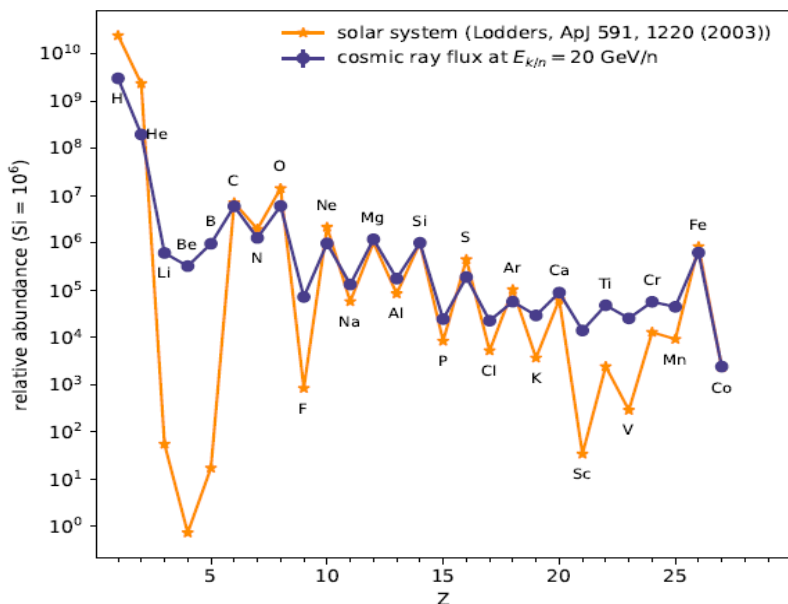
p and He spectra: updates



A rigidity dependence of both hardening and softening is favoured by data

Secondaries: Li, Be and B

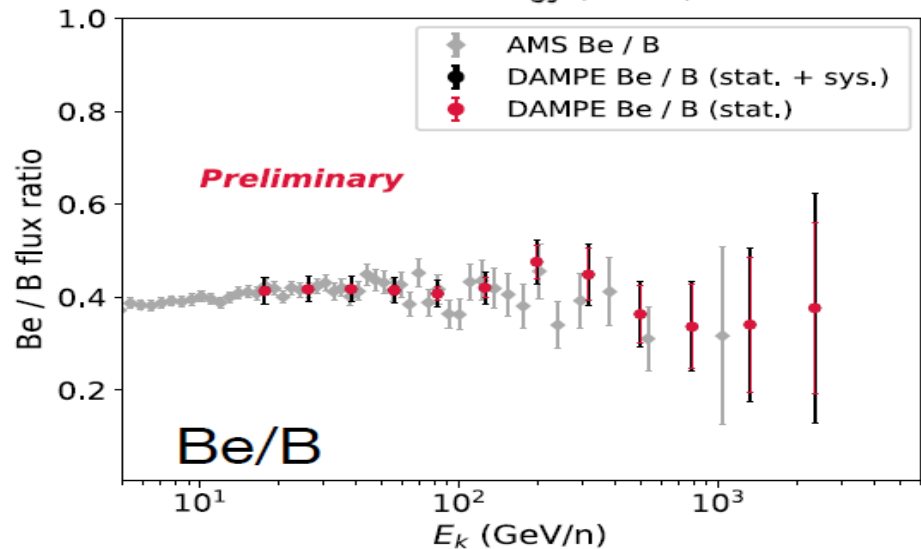
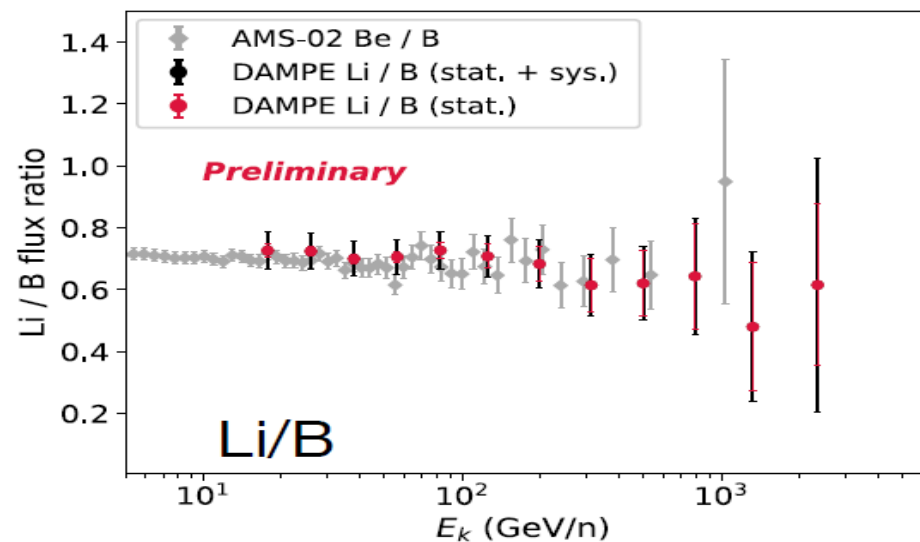
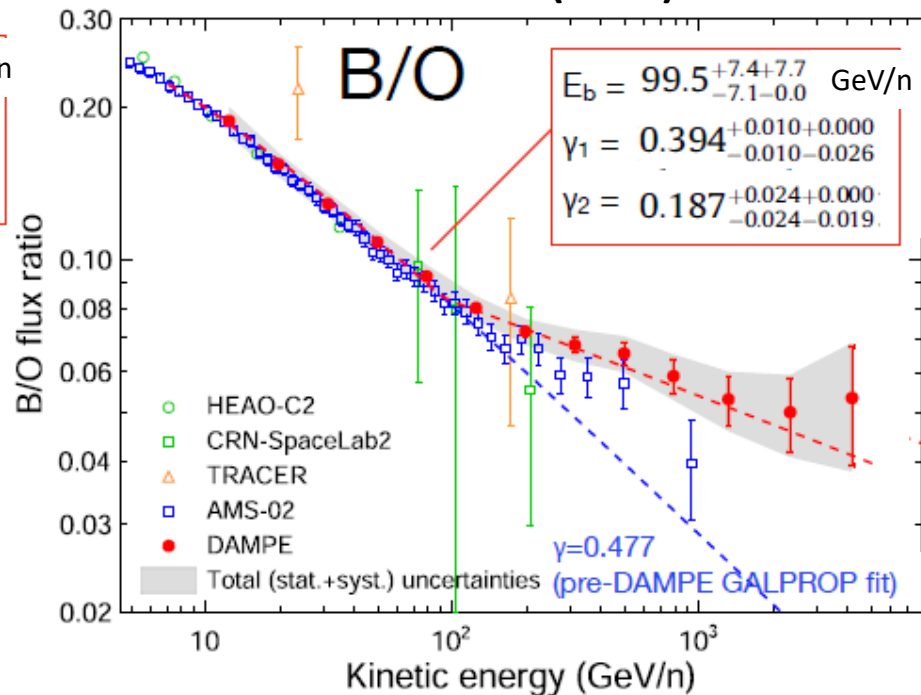
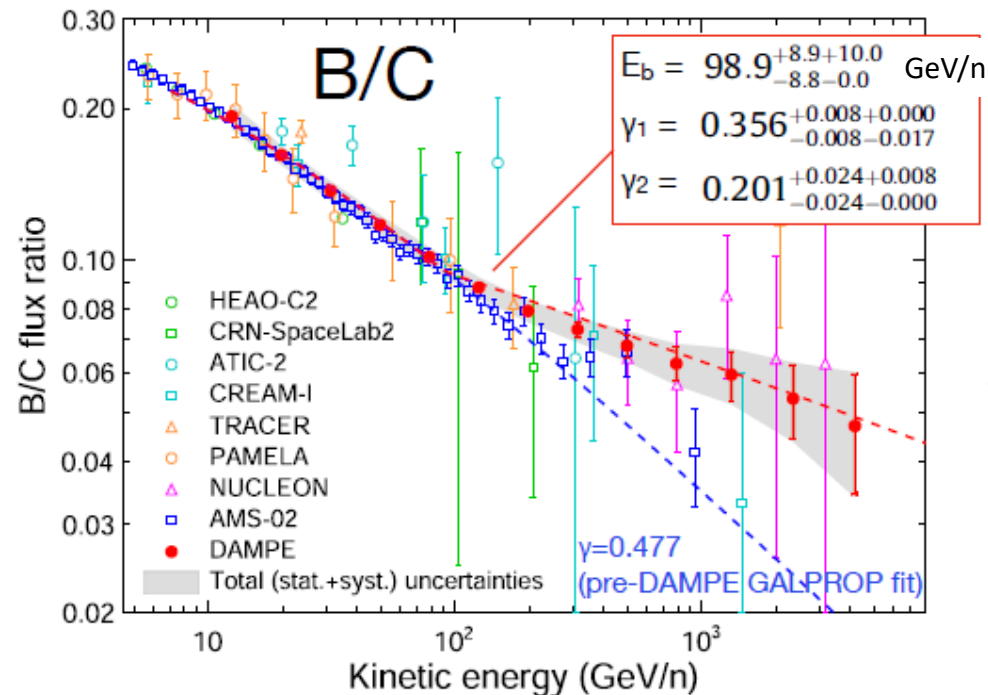
DAMPE collab., *Phys. Rev. Lett.* 134, 191001 (2025)



Flux ratios

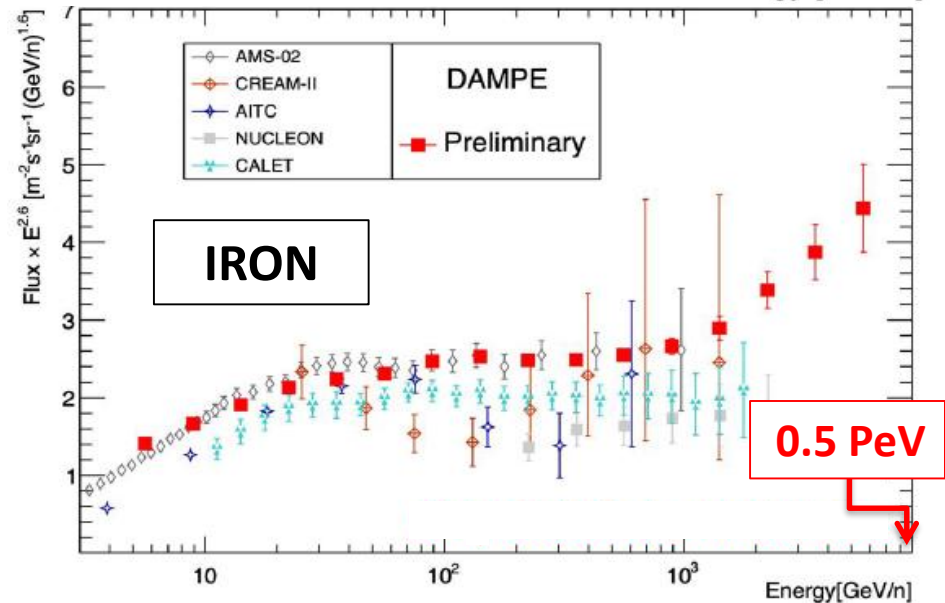
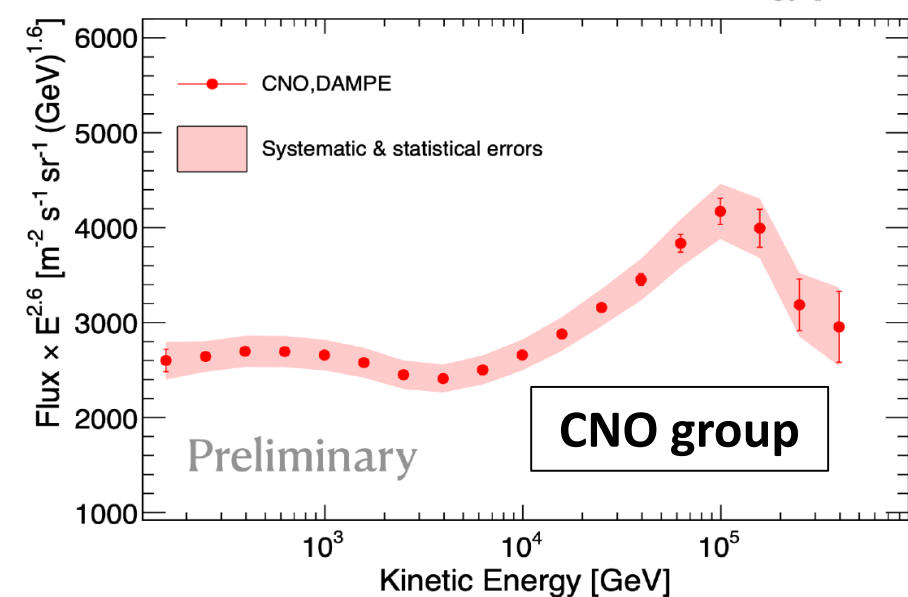
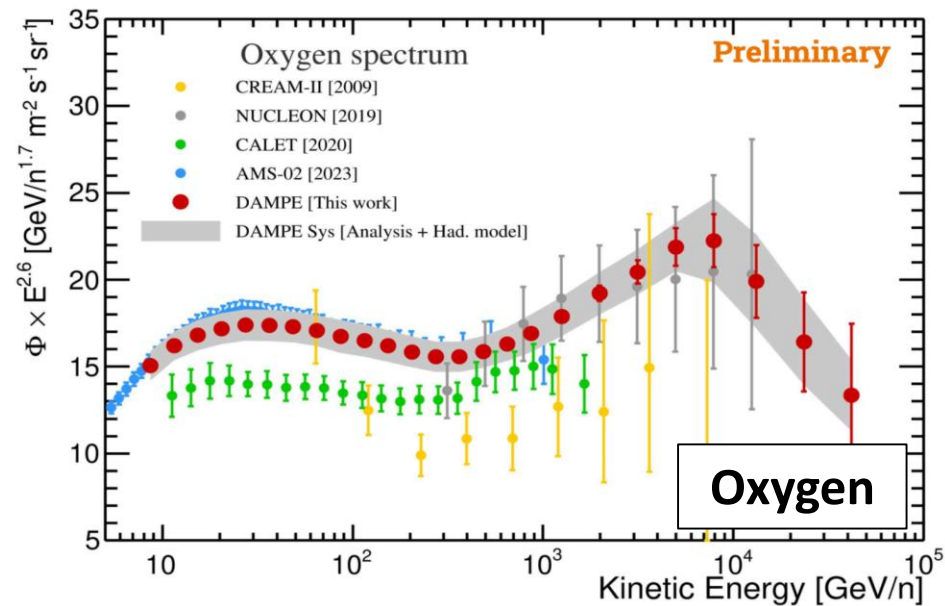
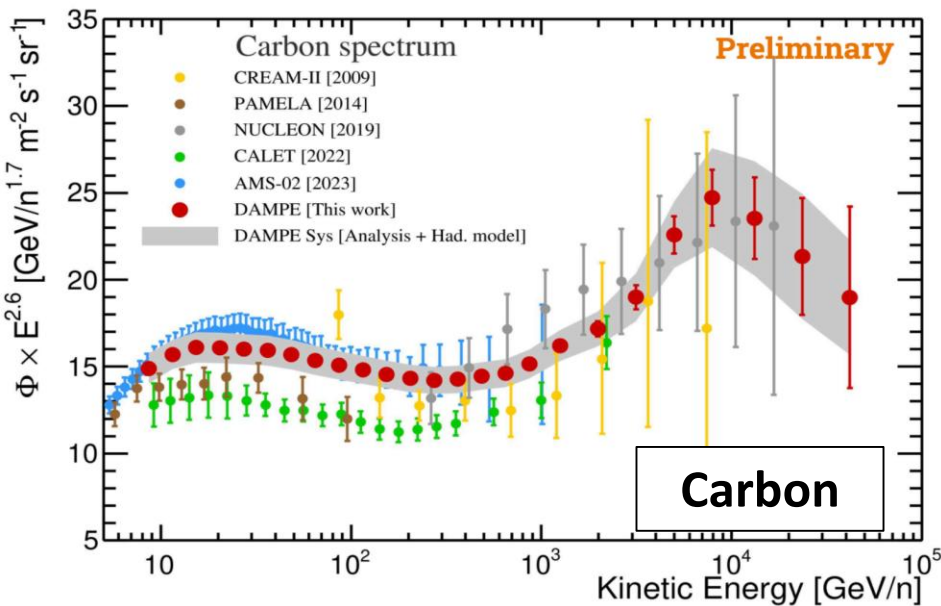
B/C and B/O

DAMPE Coll. Science Bull. 67 (2022) 21



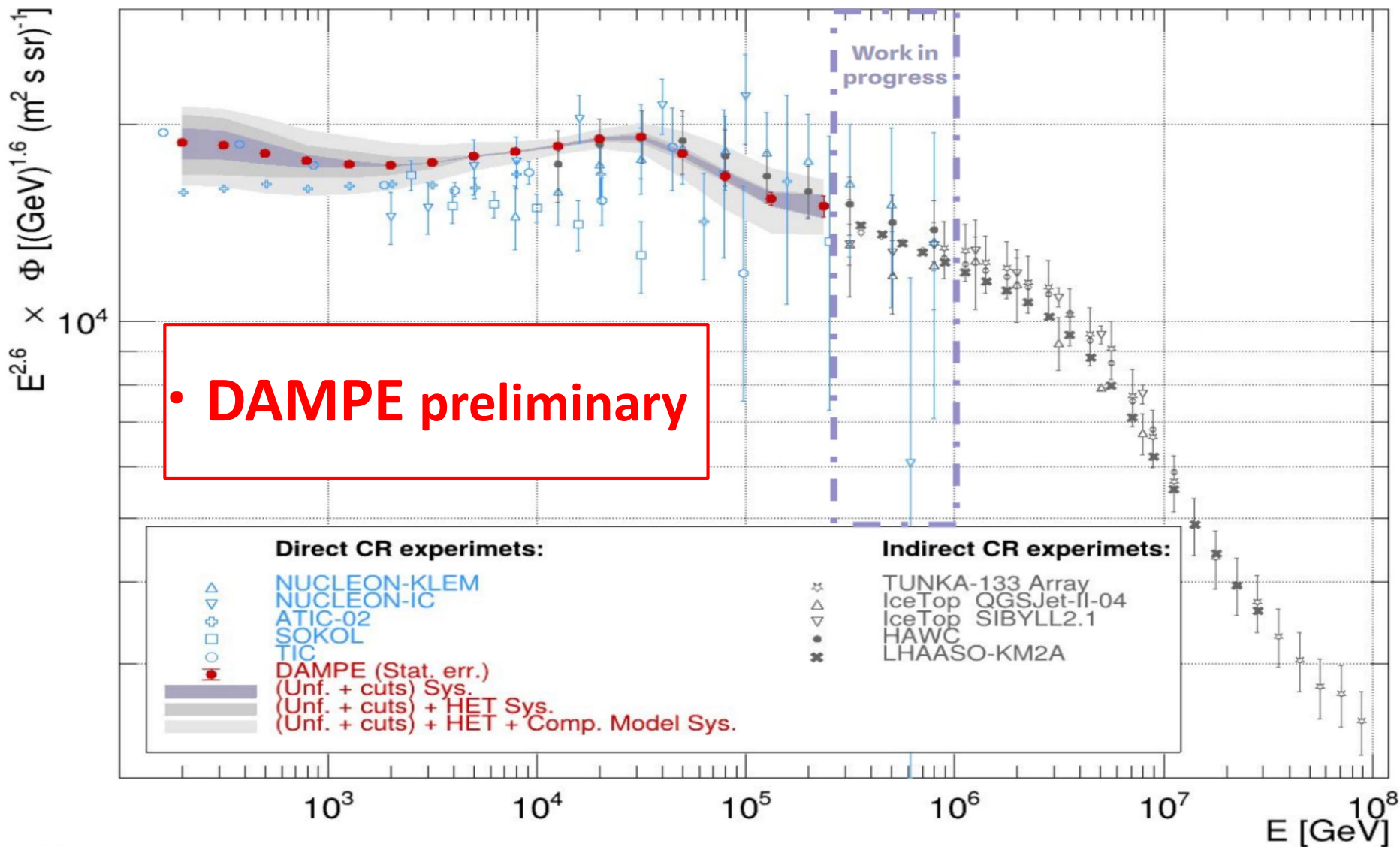
Heavier elements...

Several analyses are in progress towards the highest energies (~0.8 PeV for Fe)



The all-particle spectrum

A single measurement across almost 4 orders of magnitude



Summary

The detector

- Large geometric factor instrument ($0.3 \text{ m}^2 \text{ sr}$ for p and nuclei)
- Precision Si-W tracker ($50\mu\text{m}$, 0.2°)
- Thick calorimeter ($31 X_0$, σ_E/E better than 1% above 50 GeV for e/γ , $\sim 35\%$ for hadrons)
- “Multiple” charge measurements (0.2-0.3 e resolution)
- e/p rejection power $> 10^5$ (topology alone, plus neutron detector)

Launch and performances

- Succesfull launch on Dec 17th, 2015
- On orbit operation steady and with high efficiencies (50 Hz, more than 13 billion events)
- Absolute energy calibration by using the geomagnetic cut-off (+1.25% at 13 GeV)
- Absolute pointing cross check by use of the photon map ($\text{PSF} = 0.3^\circ$ for 10GeV photons)

Science:

- Evidence for a cutoff at $\sim 1 \text{ TeV}$ in the all electron spectrum
- Evidence for a softening in the proton spectrum at $\sim 14 \text{ TeV}$
- Evidence for a softening in the helium spectrum at $\sim 34 \text{ TeV}$ (suggest Z dependence)
- Measurement of p+He confirms the softening and suggest a hardening around 100TeV
- Break in secondary to primary ratios (B/C and B/O) at 100 GeV/n
- Undergoing spectral measurements of heavier nuclei (C, N, O, Ne, Mg, Si, Fe, Ni, ...) and light secondaries (Li, Be, B)
- Detected new features in Forbush decrease
- Cross section measurements, Upper limits for fractionally charged particles, ...
-

More Stuff

The collaboration

- **CHINA**

- Purple Mountain Observatory, CAS, Nanjing
- Institute of High Energy Physics, CAS, Beijing
- National Space Science Center, CAS, Beijing
- University of Science and Technology of China, Hefei
- Institute of Modern Physics, CAS, Lanzhou



- **ITALY**

- INFN Bari and University of Bari
- INFN Lecce and University of Salento
- INFN LNGS and Gran Sasso Science Institute
- INFN Perugia and University of Perugia

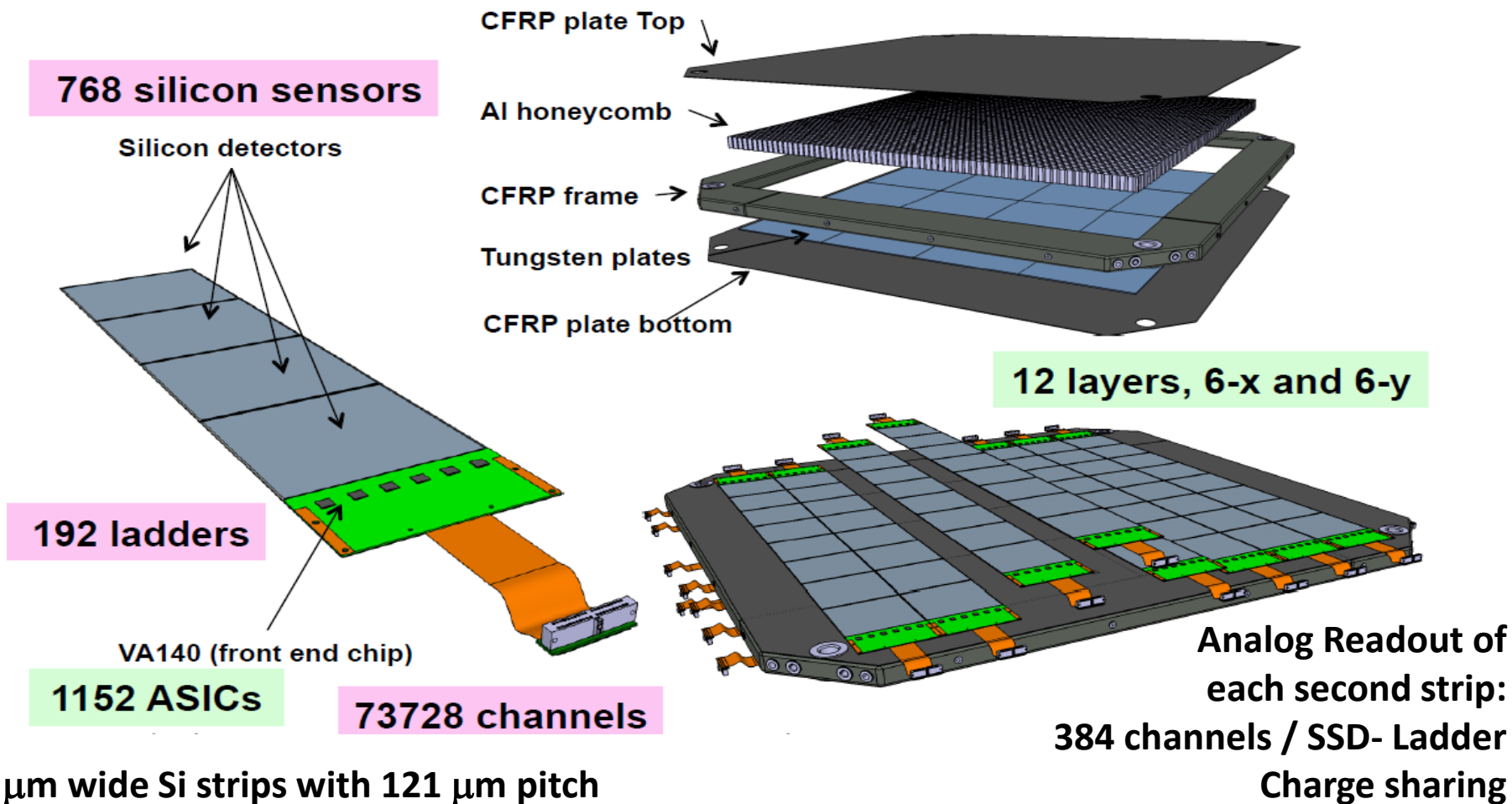


- **SWITZERLAND**

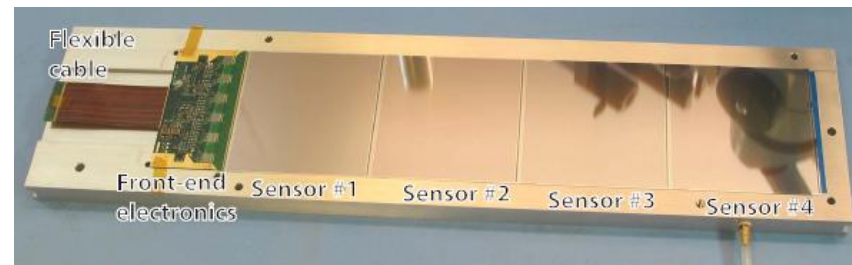
- University of Geneva



The Silicon Tracker (STK)

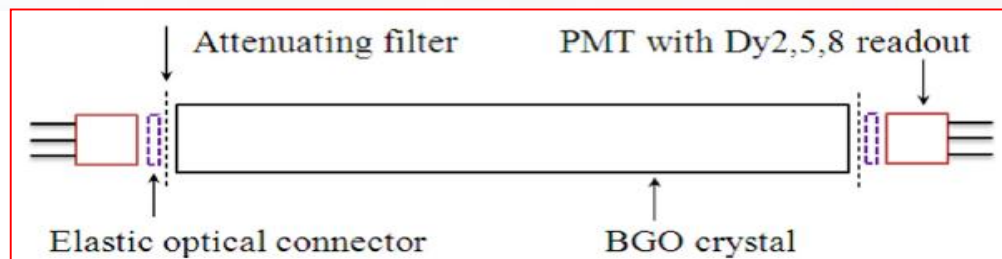
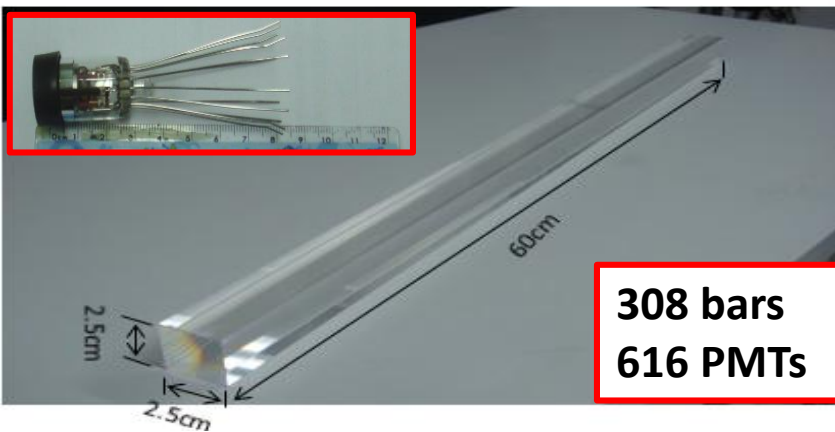
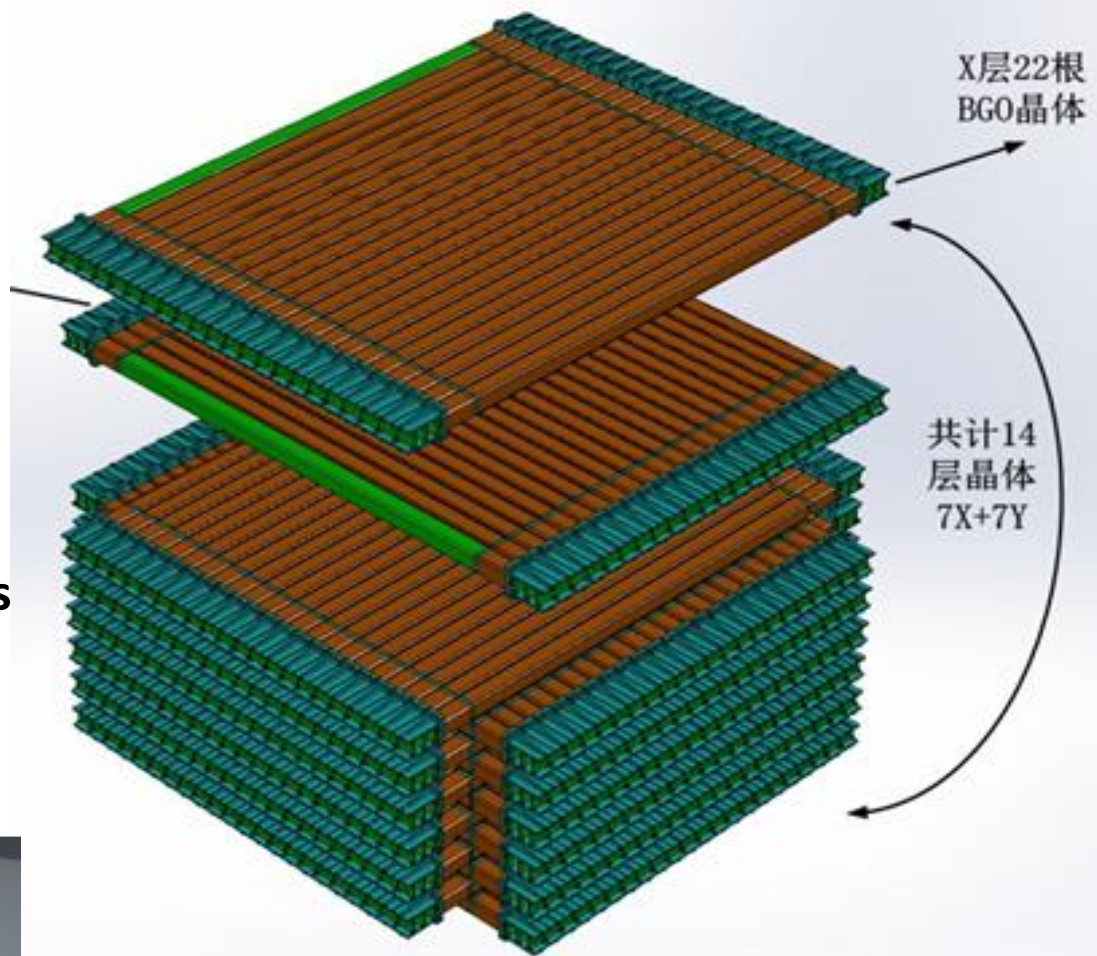


- 48 μm wide Si strips with 121 μm pitch
- (95 $\times$ 95 $\times$ 0.32 mm³) Silicon Strip Detector (SSD)
- 768 strips in each SSD
- One ladder composed by 4 (SSD)
- 16 Ladders per layer (76 cm $\times$ 76 cm)
- 12 layers (6x + 6y)



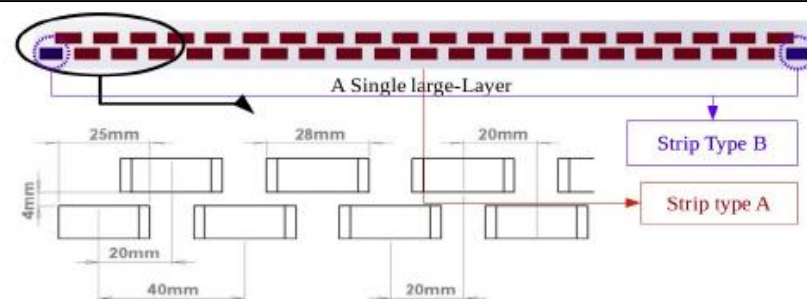
The CALOrimeter

- 14 layers of 22 BGO bars
 - $2.5 \times 2.5 \times 60 \text{ cm}^3$ bars
 - 14 hodoscopic stacking alternating orthogonal layers
 - depth $\sim 32X_0$
- Two PMTs coupled with each BGO crystal bar at the two ends
- Electronics boards attached to each side of module

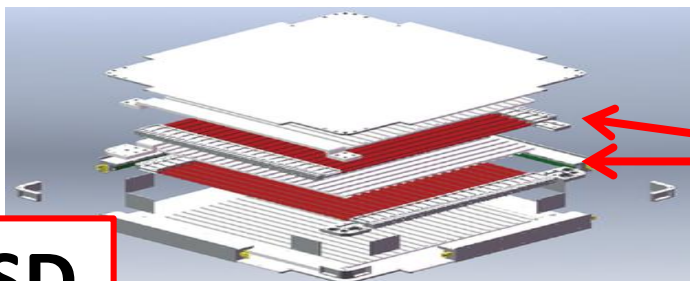


The Plastic Scintillator Detector and the NeUtron Detector

- 1.0 cm thick ,2.8cm wide and 82.0 cm long scintillator strips
- staggered by 0.8 cm in a layer
- 82 cm × 82 cm layers
- 2 layers (x and y)

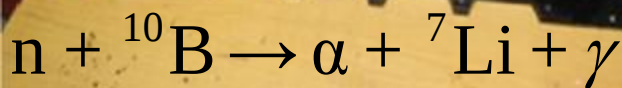
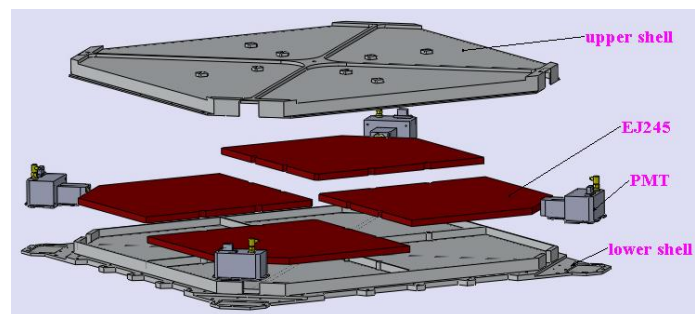


PSD



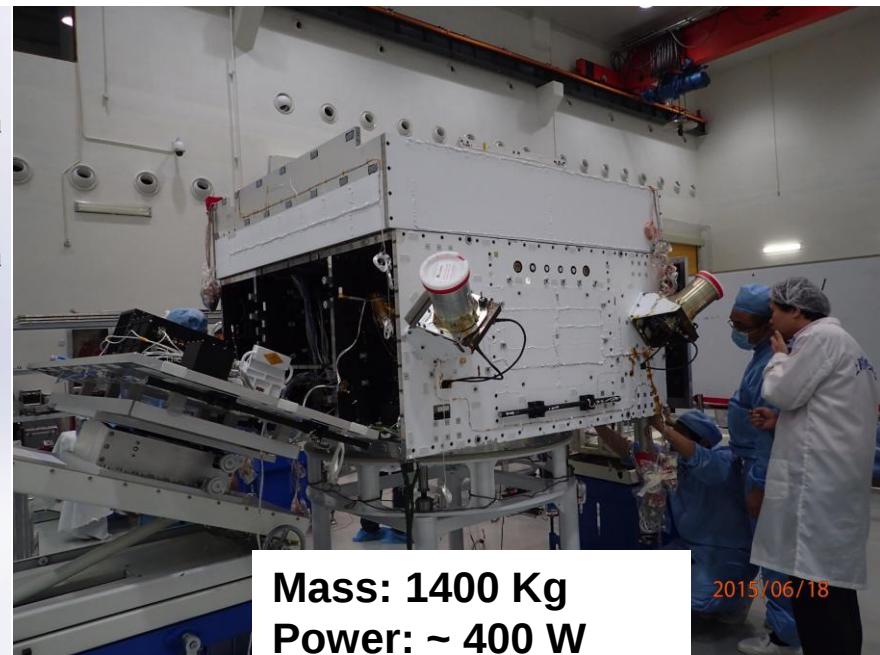
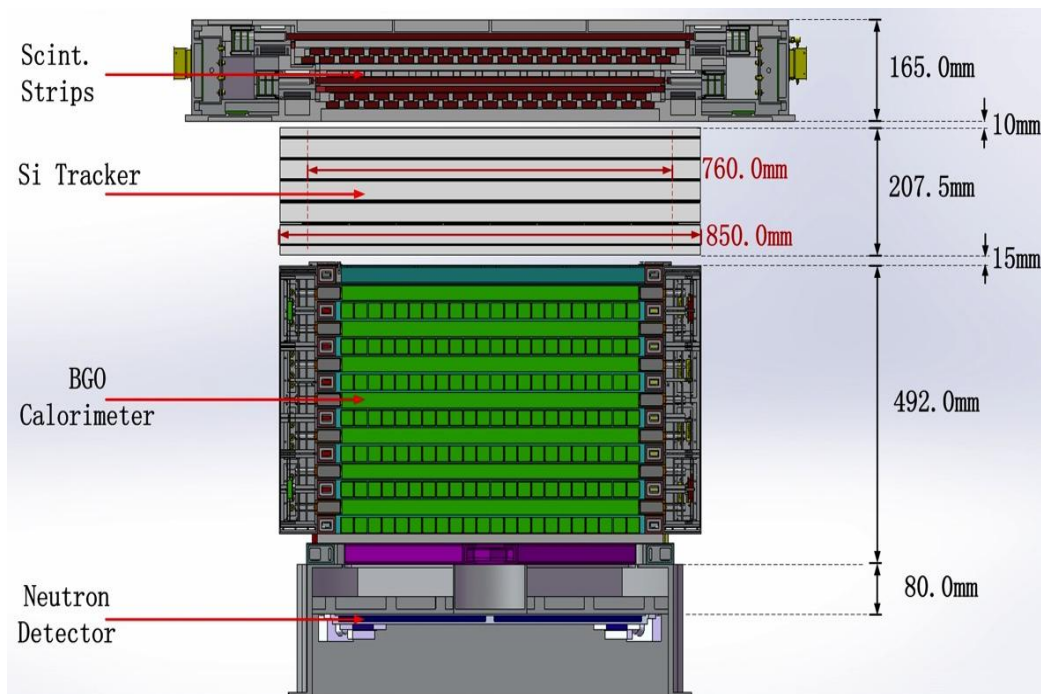
NUD

- 4 large area boron-doped plastic scintillators (30 cm × 30 cm × 1 cm)



Comparison with AMS-02 and FERMI

	DAMPE	AMS-02	Fermi LAT
e/ γ Energy res.@100 GeV (%)	1.2	3	10
e/ γ Angular res.@100 GeV (deg)	0.2	0.3	0.1
e/p discrimination	10^5-10^6	$10^5 - 10^6$	10^3
Calorimeter thickness (X_0)	32	17	8.6
Geometrical accep. (m^2sr)	0.3	0.09	1

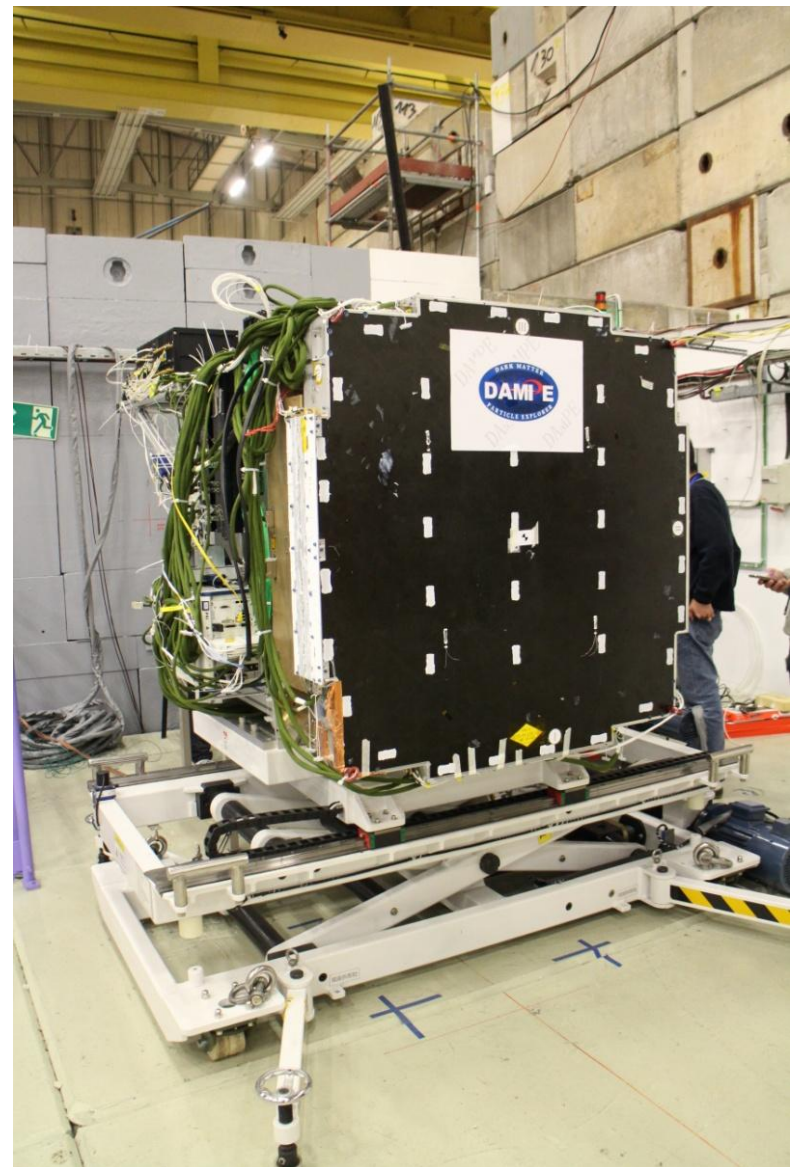


Mass: 1400 Kg
Power: ~ 400 W
Livetime: > 3 years

2015/06/18

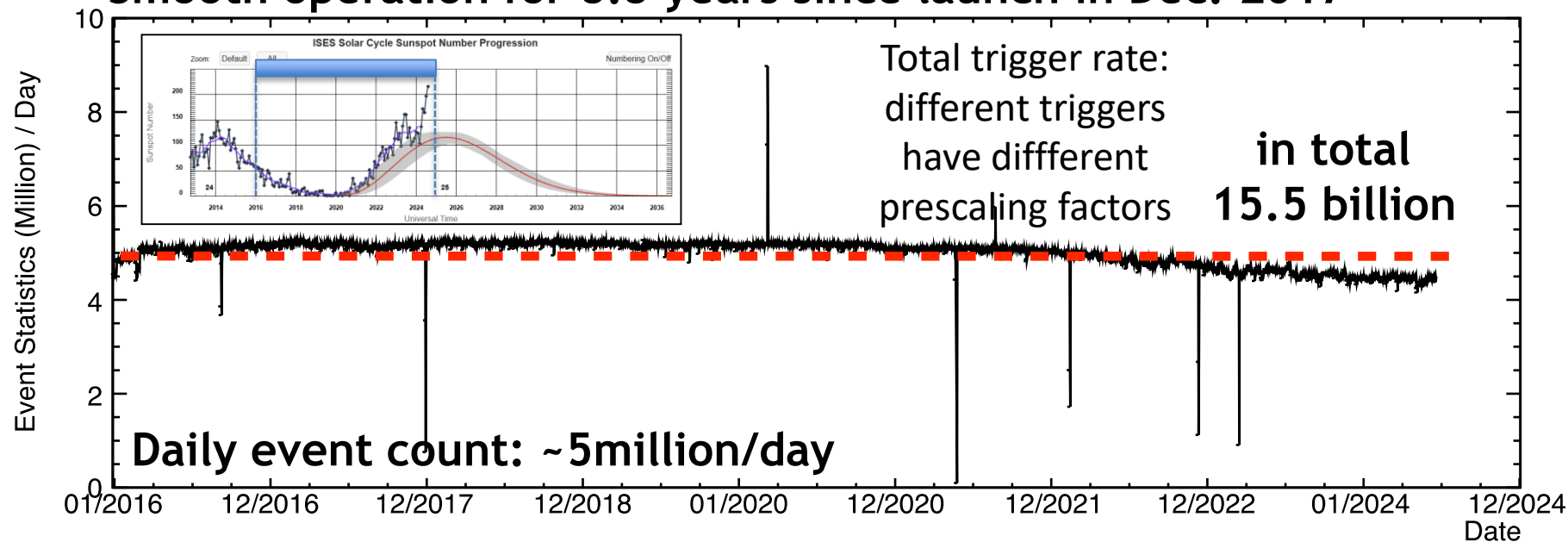
Test beam activity at CERN

- 14days@PS, 29/10-11/11 2014
 - e @ 0.5GeV/c, 1GeV/c, 2GeV/c, 3GeV/c, 4GeV/c, 5GeV/c
 - p @ 3.5GeV/c, 4GeV/c, 5GeV/c, 6GeV/c, 8GeV/c, 10GeV/c
 - π^- @ 3GeV/c, 10GeV/c
 - γ @ 0.5-3GeV/c
- 8days@SPS, 12/11-19/11 2014
 - e @ 5GeV/c, 10GeV/c, 20GeV/c, 50GeV/c, 100GeV/c, 150GeV/c, 200GeV/c, 250GeV/c
 - p @ 400GeV/c (SPS primary beam)
 - γ @ 3-20GeV/c
 - μ @ 150GeV/c,
- 17days@SPS, 16/3-1/4 2015
 - Fragments: 66.67-88.89-166.67GeV/c
 - Argon: 30A- 40A- 75AGeV/c
 - Proton: 30GeV/c, 40GeV/c
- 21days@SPS, 10/6-1/7 2015
 - Primary Proton: 400GeV/c
 - Electrons @ 20, 100, 150 GeV/c
 - γ @ 50, 75 , 150 GeV/c
 - μ @ 150 GeV /c
 - π^+ @10, 20, 50, 100 GeV/c
- 10days@SPS, 11/11-20/11 2015
 - Pb 30AGeV/c (and fragments) (HERD)
- 6days@SPS, 20/11-25/11 2015
 - Pb 030 AGeV/c (and fragments)

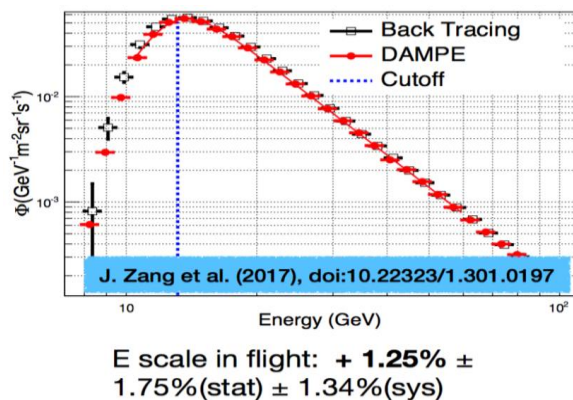


On orbit operation

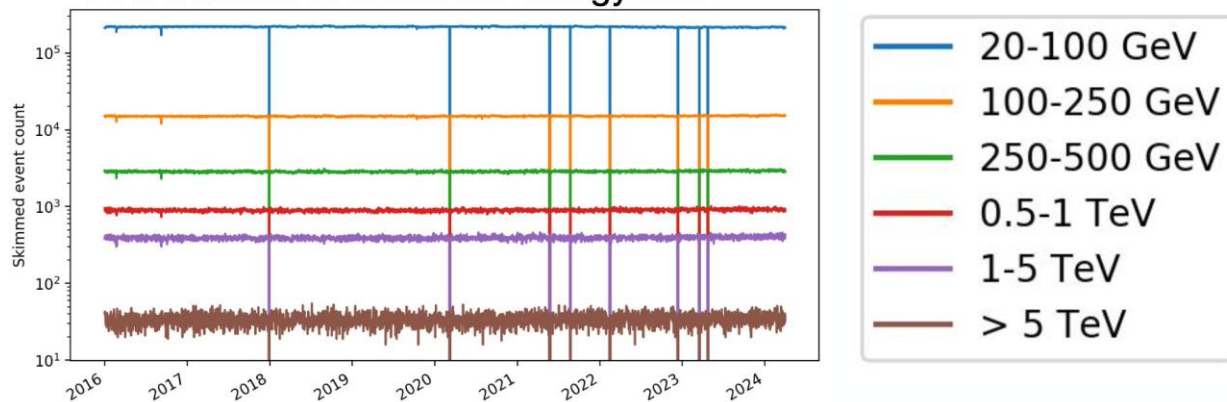
Smooth operation for 8.5 years since launch in Dec. 2017



Energy scale calibration by using geomagnetic cutoff on electrons



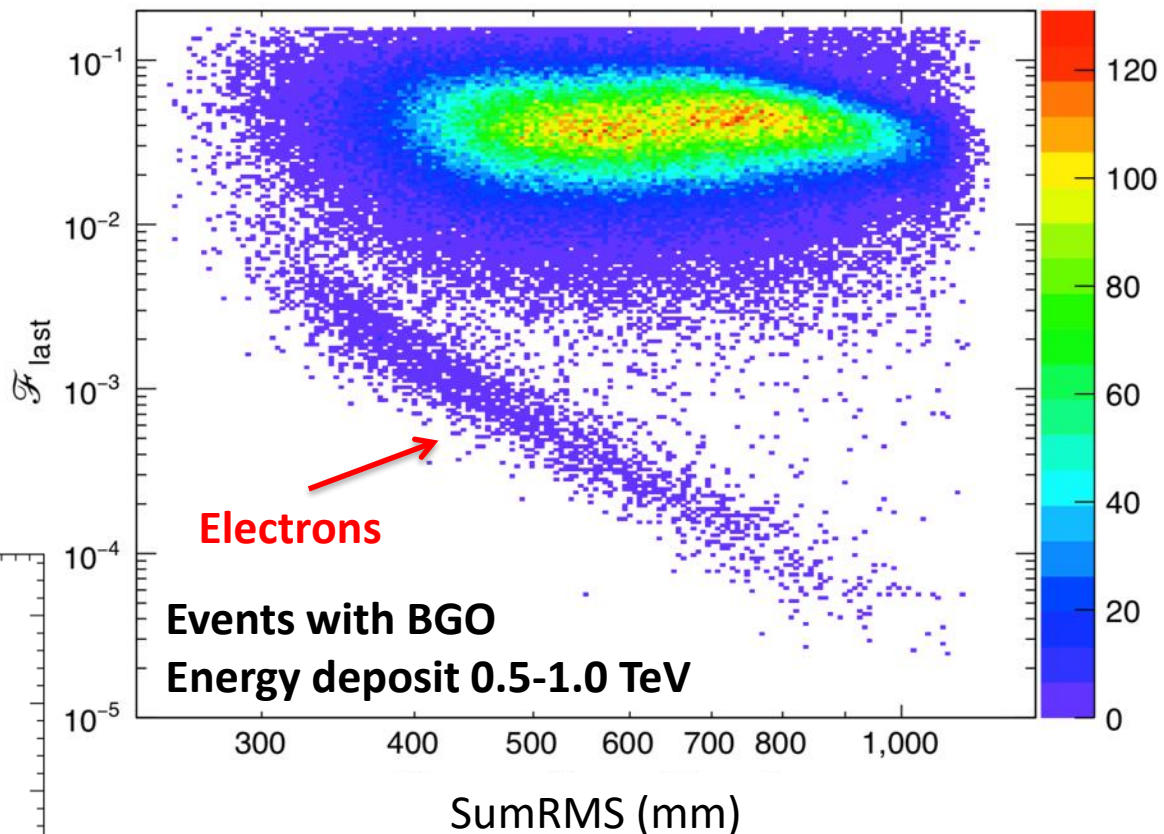
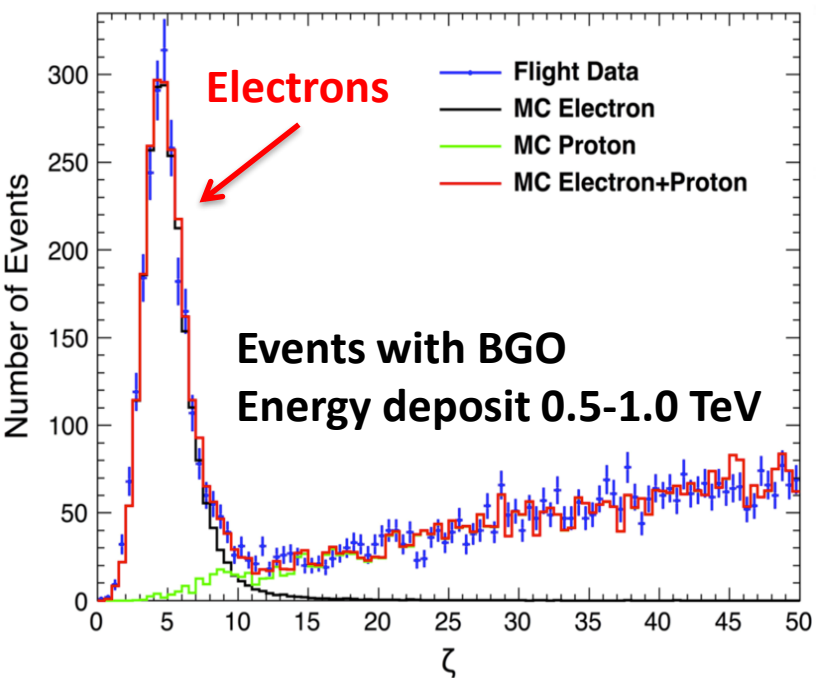
Event counts in difference energy bins



Electron IDentification

$\mathcal{F}_{\text{last}}$ = fraction of energy
 deposit in the last BGO
 layer with hits

$$RMS_i = \sqrt{\frac{\sum_j (x_{j,i} - x_{c,i})^2 E_{j,i}}{\sum_j E_{j,i}}}$$



SumRMS = Sum of single layer
 RMS values

$$\zeta = \mathcal{F}_{\text{last}} \times (\sum_i RMS_i / \text{mm})^4 / (8 \times 10^6)$$

Hadronic cross section measurements

