

# An event-type based analysis for LST-1

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TeV Particle Astrophysics  
**TeVPA**  
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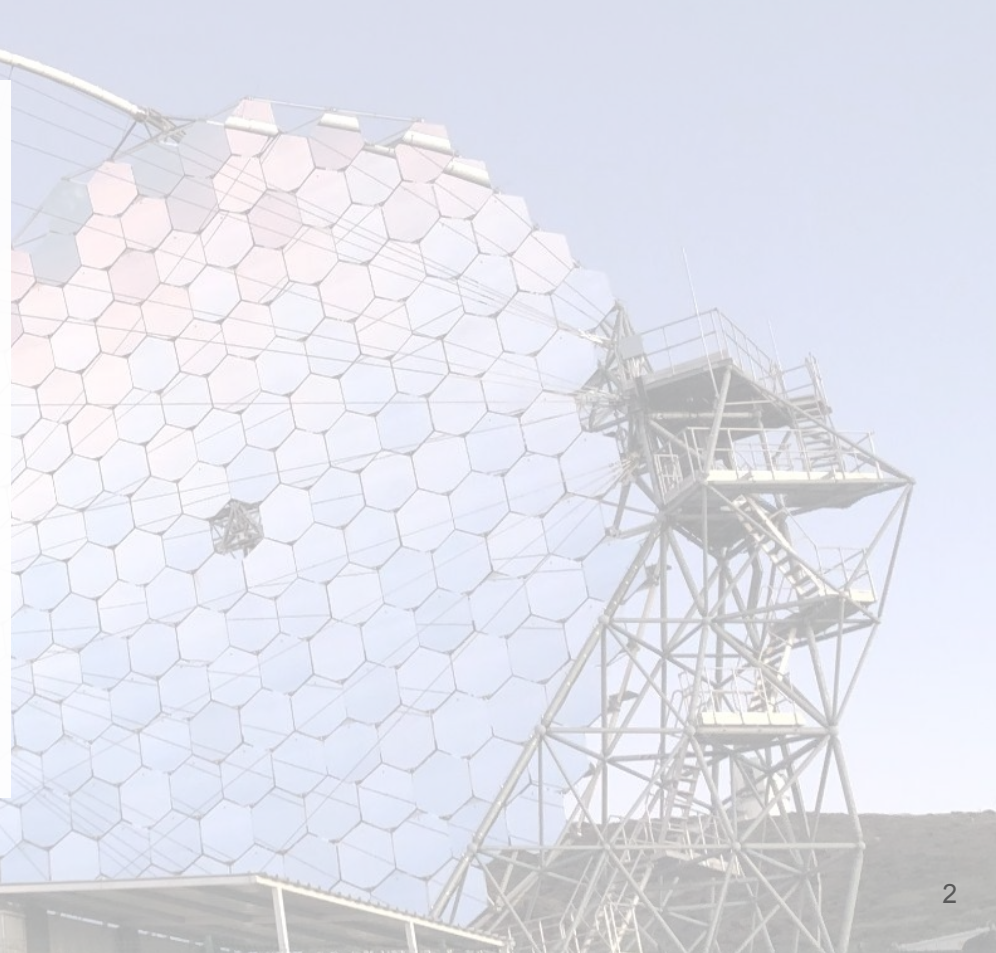
- The event-type based analysis

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- Real LST-1 data
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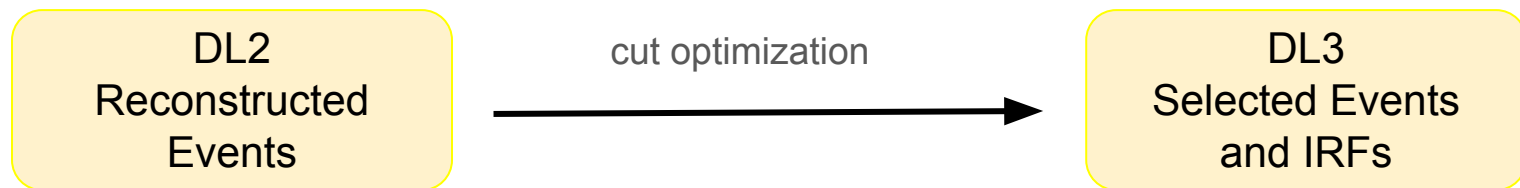
## High-level performance and analysis

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# The event-type based analysis

## Standard analysis

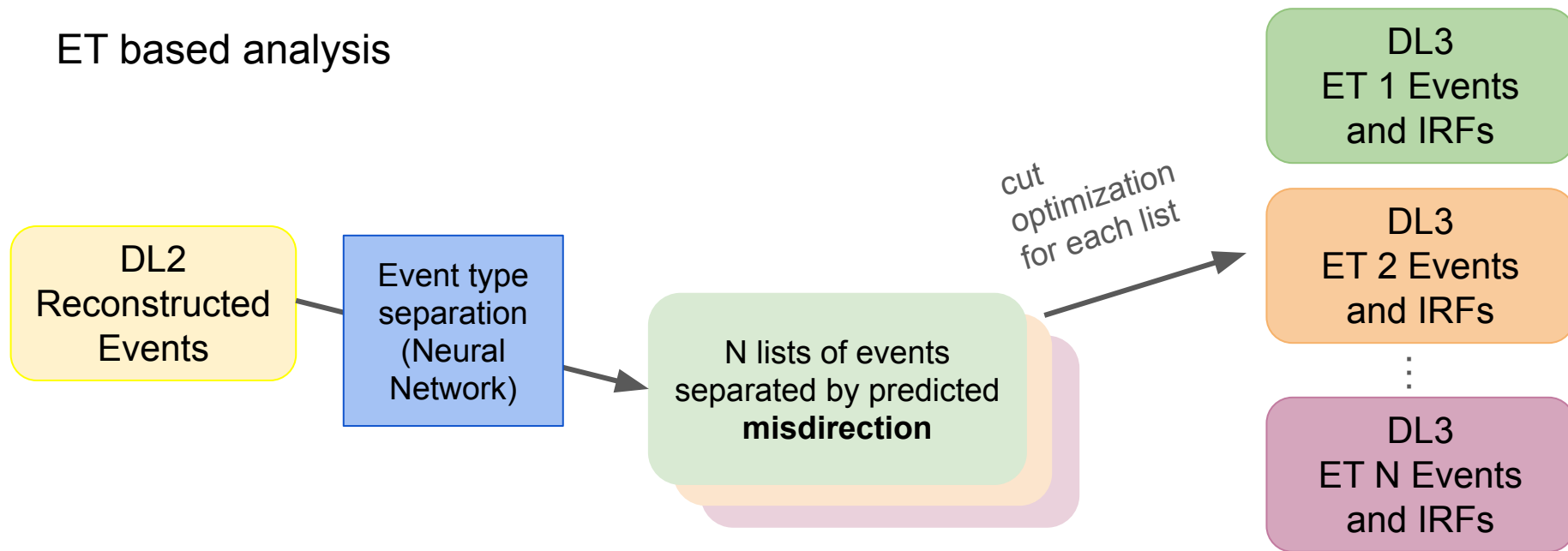


→ All selected events are treated equally as if they had the same quality

Is there an alternative? → the event types approach is already successfully used by Fermi-LAT

# The event-type based analysis

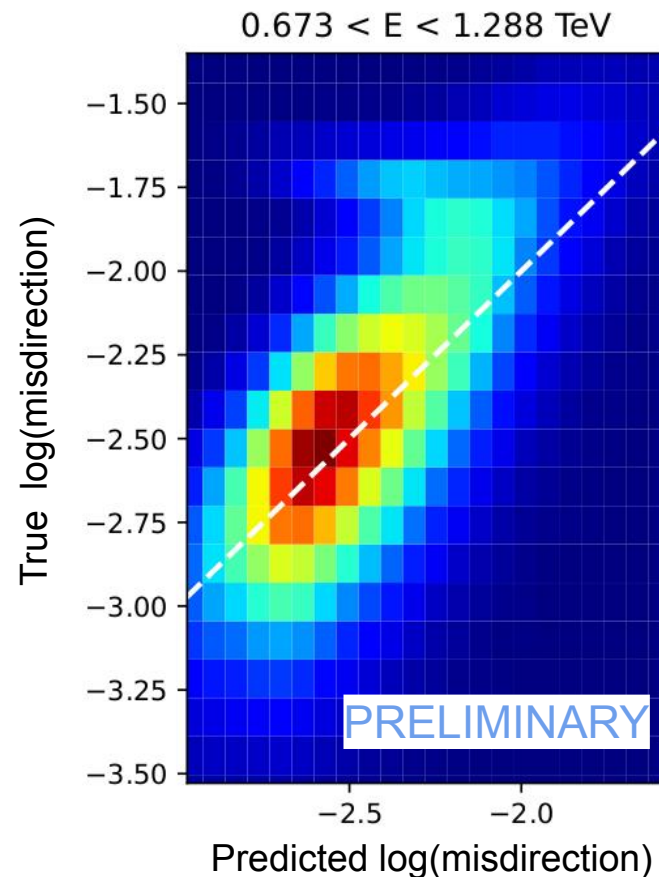
## ET based analysis



**Event-type analysis:** separate events in subsamples according to their expected reconstruction quality and generate event-type-wise IRFs

# The event-type based analysis

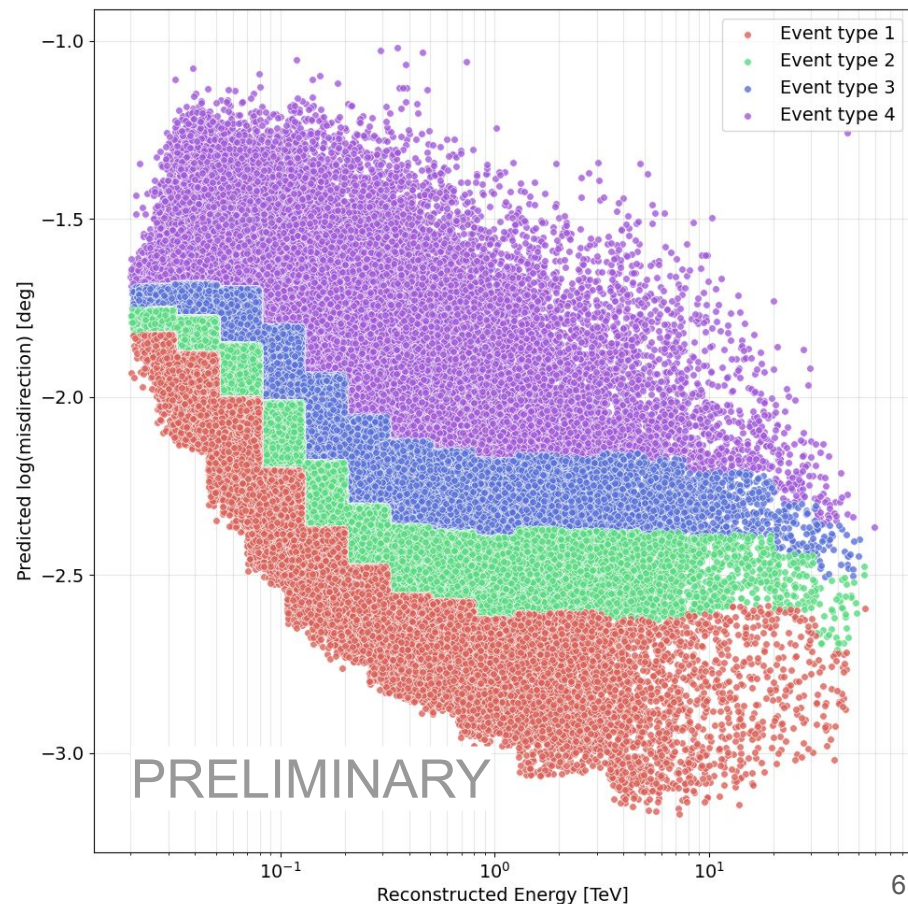
- ET separation is done predicting the “misdirection” of events
  - angular difference between simulated and reconstructed events
- Training/test statistics 25/75 %
- We use a regression algorithm
  - more flexibility and information about the quality is preserved





# The event-type based analysis using real LST-1 data

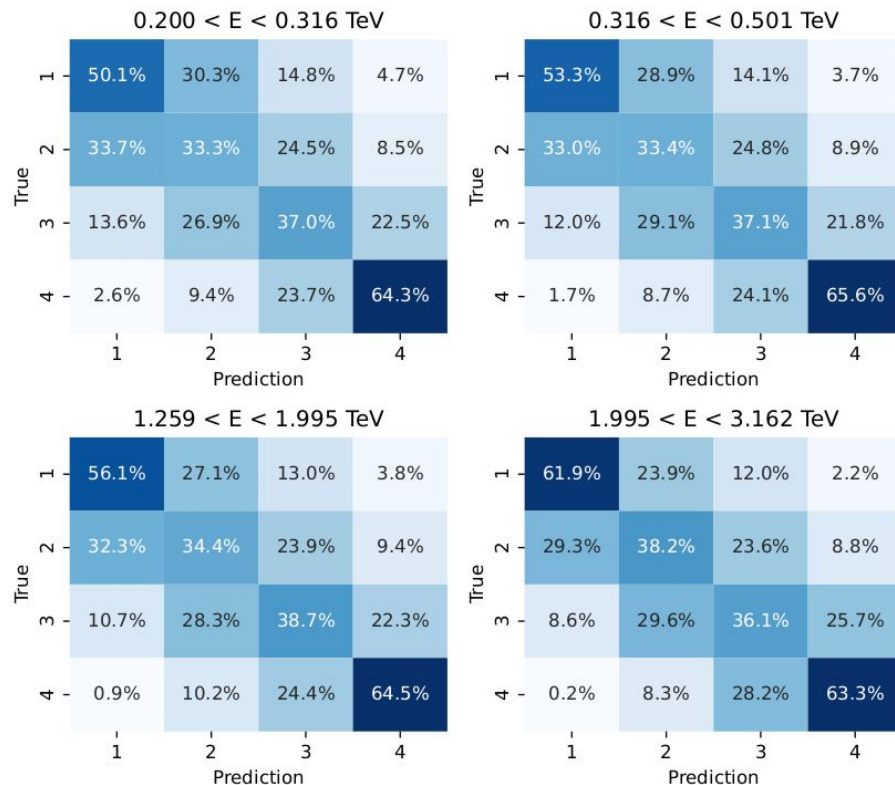
- Thresholds are defined in bins of reconstructed energy
- In this work, the events are partitioned into 15-15-30-40%\*
- The Diffuse MC used are the corresponding ones for each run based on its nsb tuning, declination line and theta-azimuth node



\*based on previous results by J. Bernete

# The event-type based analysis using real LST-1 data

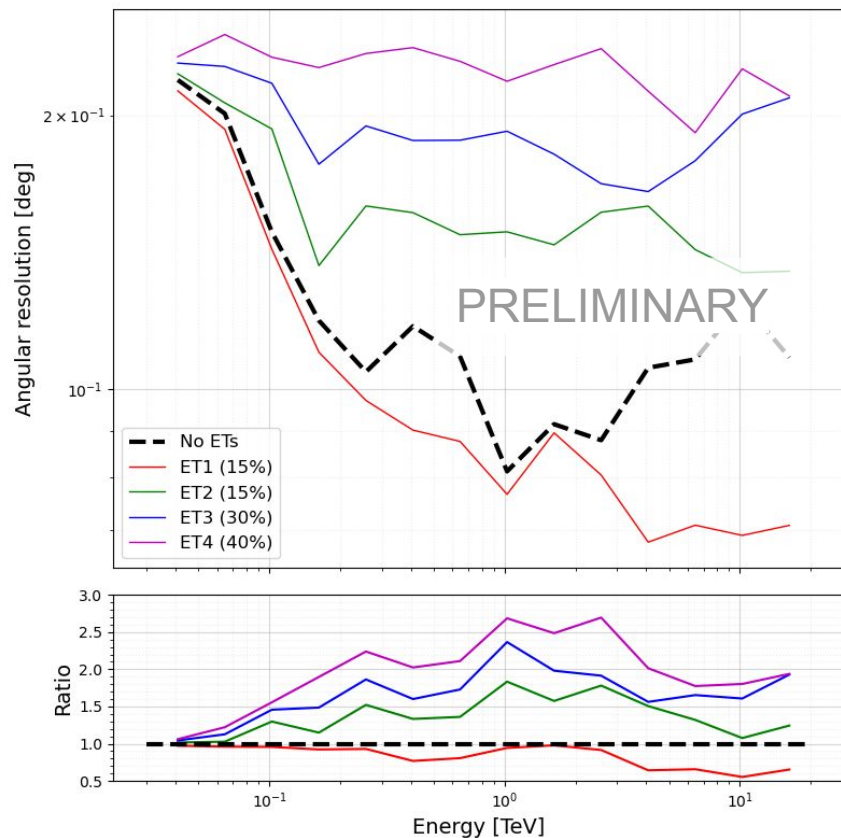
## PRELIMINARY



Example of the performance for the  
theta\_23.161\_az\_260.739 MC node

- Very few cases of confusion between the event type 1 (best 15% quality events) and event type 4 (worst 40%)

# Instrument Response Functions



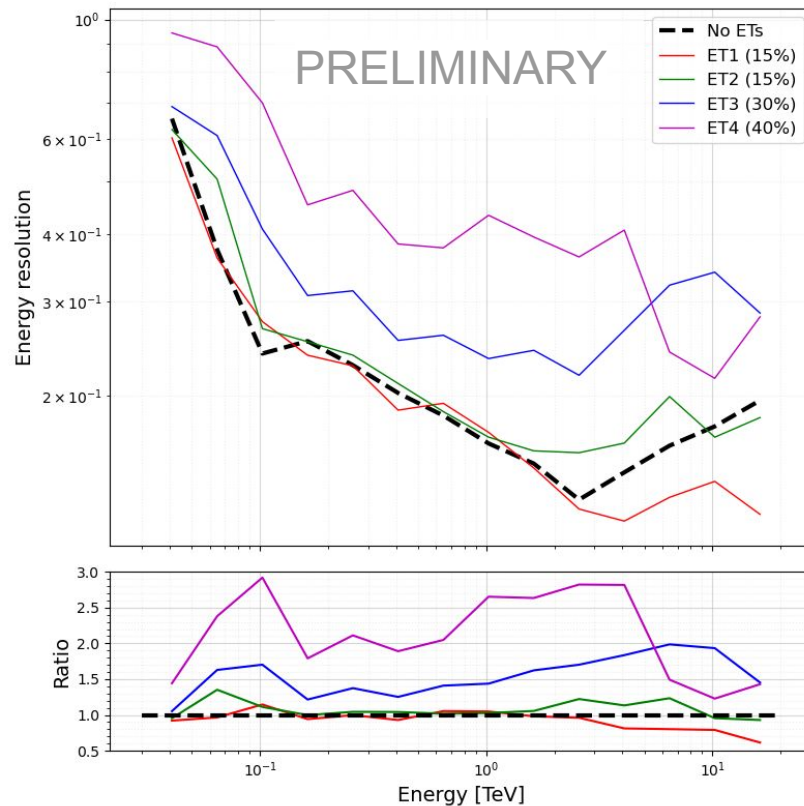
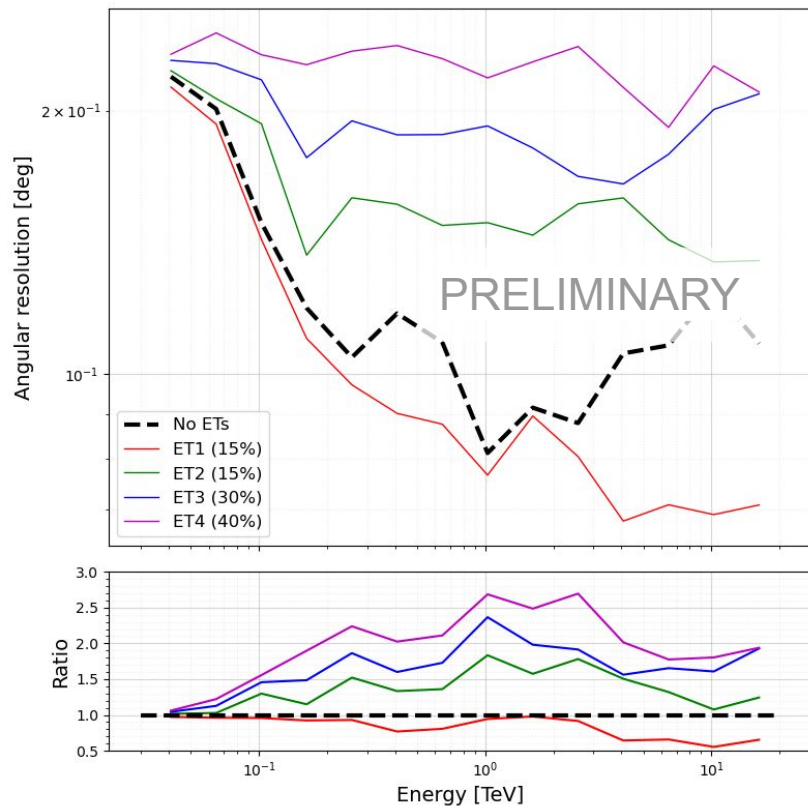
Example for the theta\_23.161\_az\_260.739 node IRFs

- Event-type-wise IRFs can be generated using optimized cuts
- Angular performance improves for the event types with better quality
- Events in the top 15% (ET 1) lead to a significant improvement with respect to the standard analysis



# Instrument Response Functions

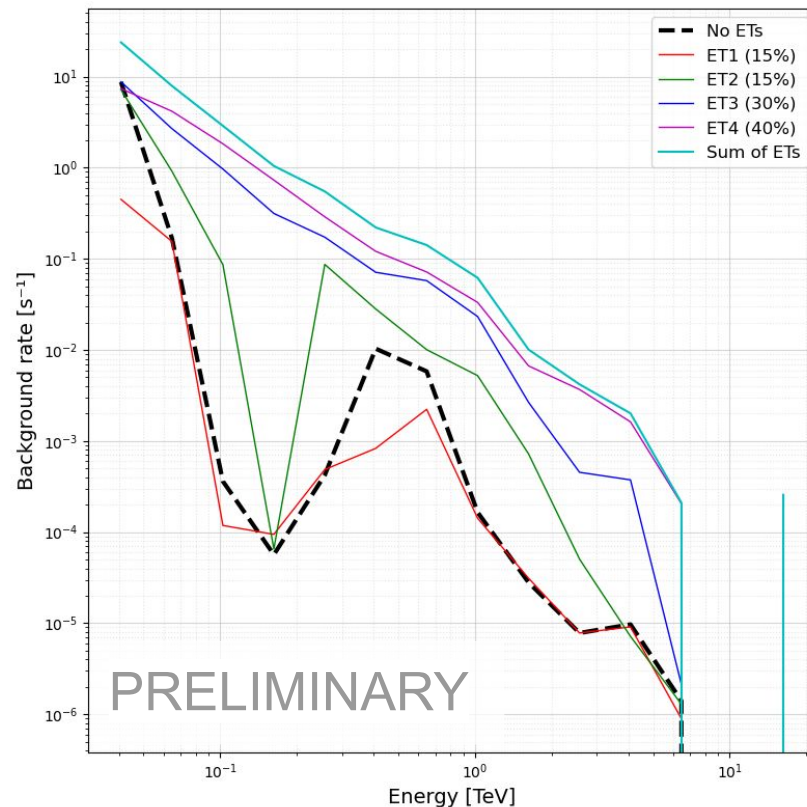
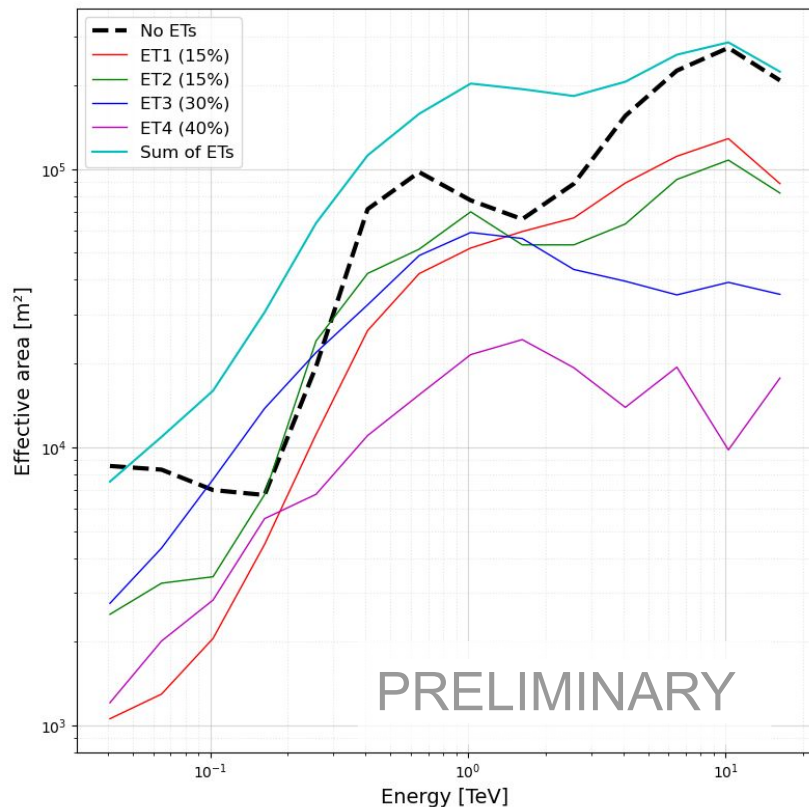
→ The 15% top events show better **angular and energy** resolution



Example for the theta\_23.161\_az\_260.739 node IRFs

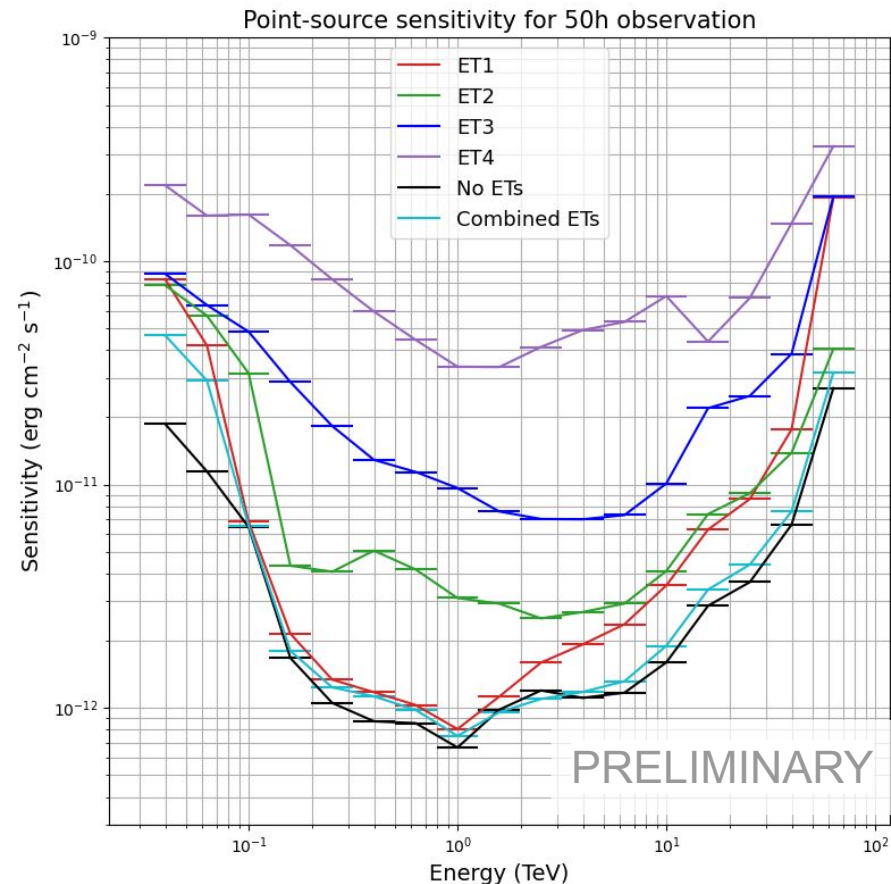
# Instrument Response Functions

- Need for custom proton MCs to produce accurate bkg IRF

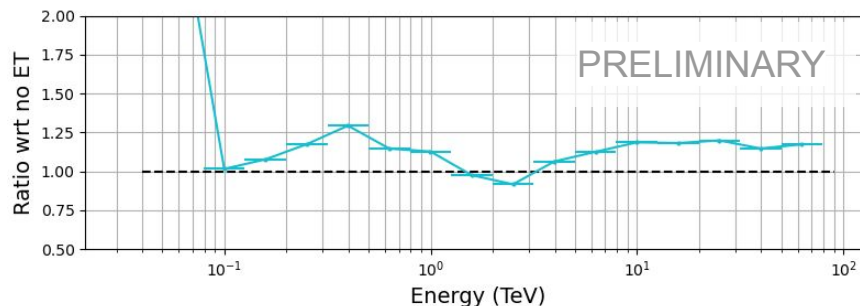


Example for the theta\_23.161\_az\_260.739 node IRFs

# Sensitivity Estimation



- Done using Gammapy's *FluxPointsEstimator* and simulating a 50h observation with a PWL spectrum with  $\Gamma=2.62$
- Combined sensitivity compatible to standard analysis
- Next steps: decide proper sensitivity estimation method and apply directly to real data

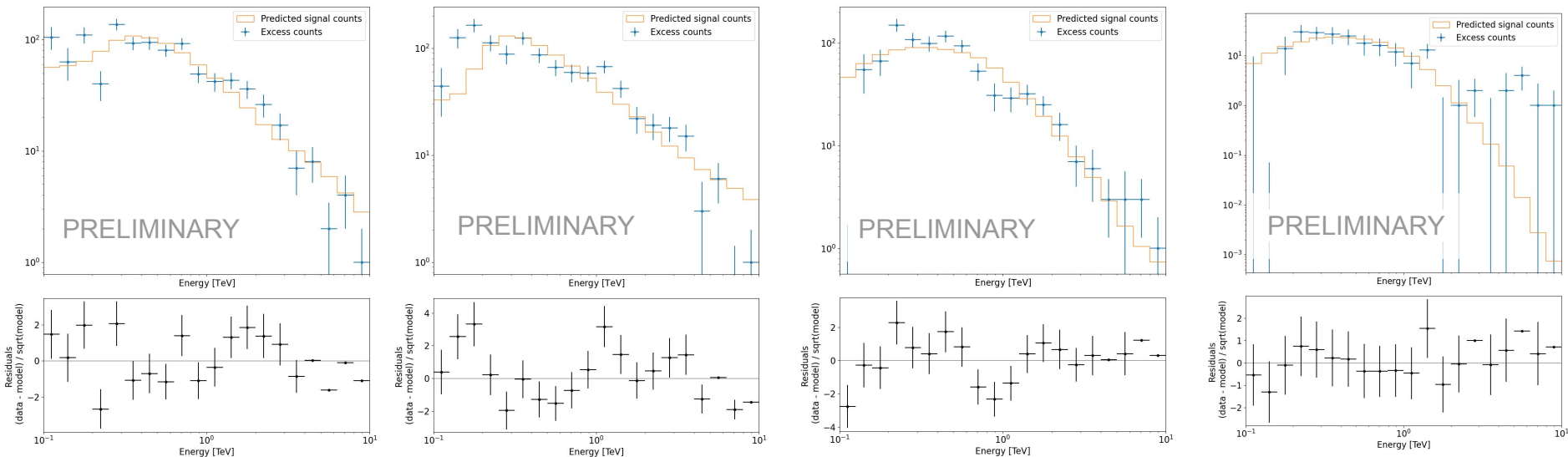


Example for the theta\_23.161\_az\_260.739 node IRFs

# The event-type based analysis using real LST-1 data

Crab Nebula dataset: 2.00h of observations taken on March 4th and 5th, 2022  
(standard DL3 are publicly available<sup>1</sup>)

➔ No clear trends when comparing the fit residuals of the 4 event types used in this work



<sup>1</sup>Morcuende, D. (2024). Crab Nebula DL3 example dataset from the LST-1 performance study [Data set]. Zenodo

# Conclusions

- Preliminary event-type based analysis ready for LST-1 data analysis
- The current implementation of the NN used works with real data:
  - ◆ Partition ranking shows IRFs improving in resolution
  - ◆ High correlation between well-characterized events in energy and direction
- Next steps:
  - ◆ Optimize cuts to improve combined sensitivity
  - ◆ Evaluate the impact of ET-based analysis on spectral reconstruction systematics

Thank you!



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