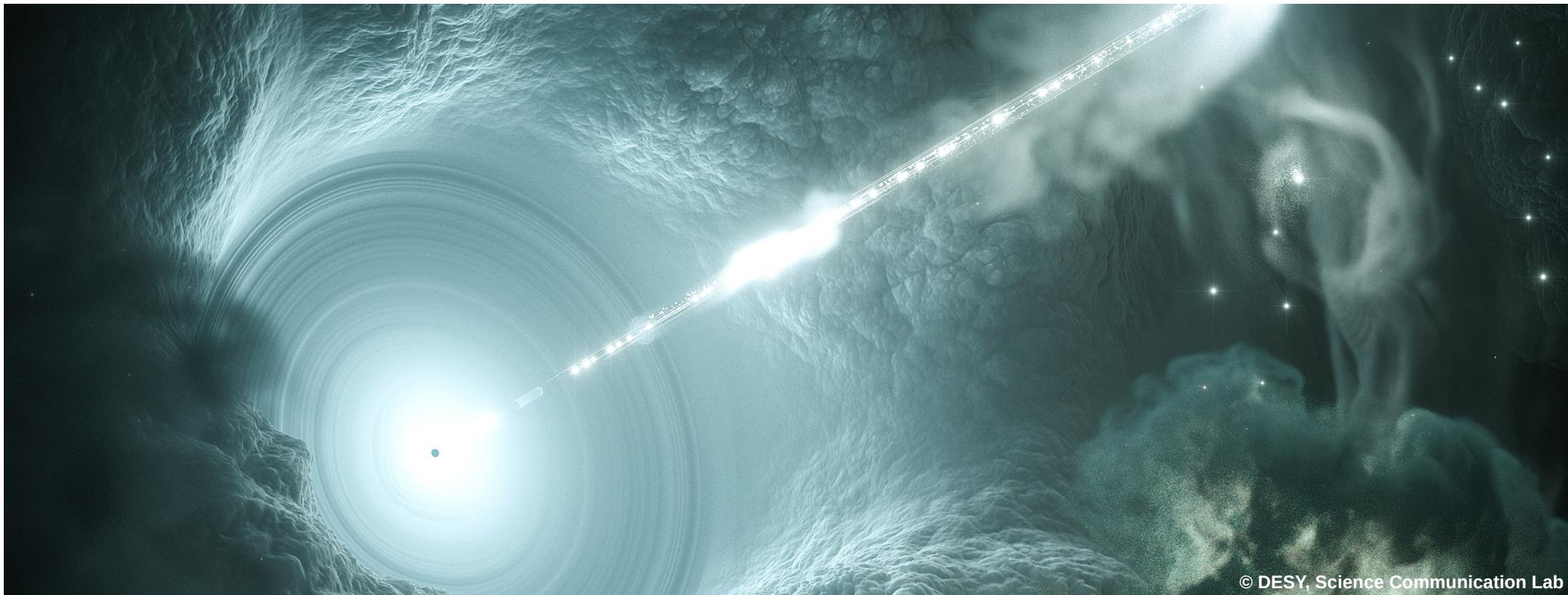




Facilitating multi-messenger modelling using Gammapy

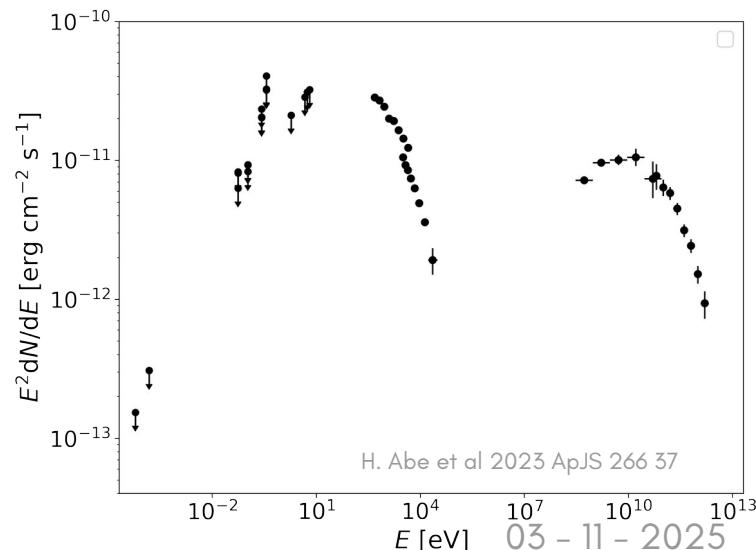
Lea Heckmann, Matteo Cerruti, Bruno Khélifi, Justin Albinet



Current AGN modelling

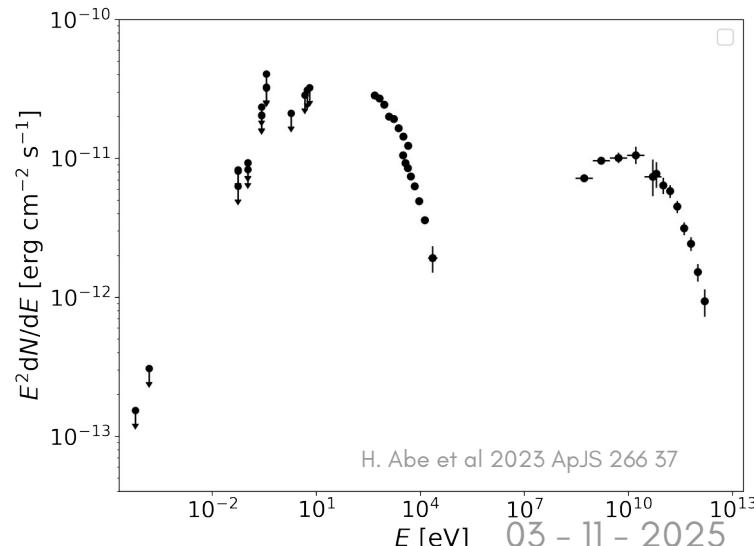
Current AGN modelling

- Combining MWL data on flux point level:



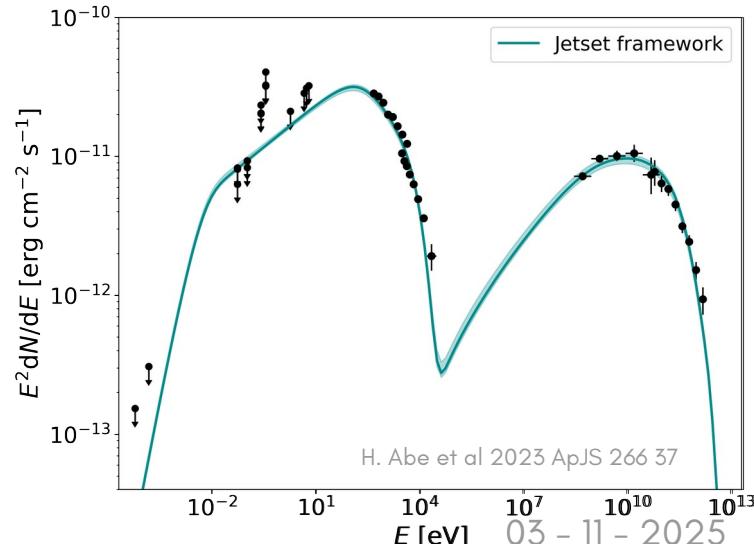
Current AGN modelling

- Combining MWL data on flux point level:
 - We construct flux points for single instruments separately
 - using different frameworks for different instruments
 - assuming simplistic models (e.g. power-law) - single instruments cannot constrain physical models (e.g. SSC)



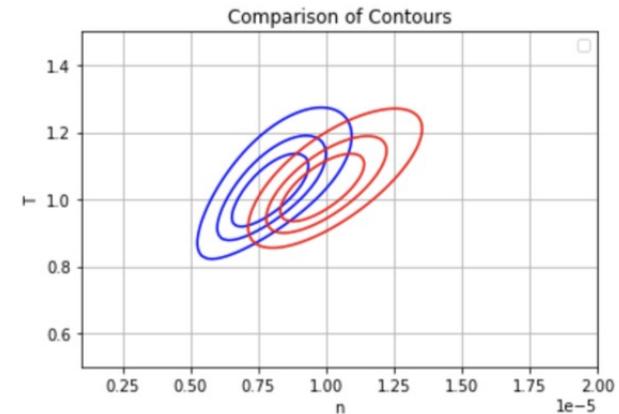
Current AGN modelling

- Combining MWL data on flux point level:
 - We construct flux points for single instruments separately
 - using different frameworks for different instruments
 - assuming simplistic models (e.g. power-law) - single instruments cannot constrain physical models (e.g. SSC)
 - Selection of the physical model based on flux points



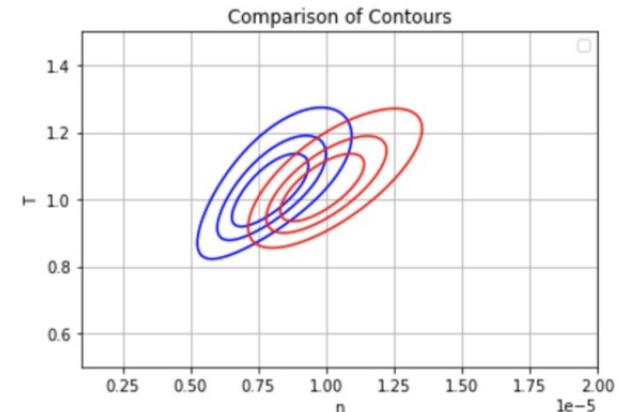
Current AGN modelling

- Combining MWL data on flux point level:
 - We construct flux points for single instruments separately
 - using different frameworks for different instruments
 - assuming simplistic models (e.g. power-law) - single instruments cannot constrain physical models (e.g. SSC)
 - Selection of the physical model based on flux points
- Fitting on flux points adds a bias:
 - Simple test with a blackbody spectrum using X-rays (XSPEC (1999ascl.soft10005A)) shows a bias of 1 sigma



Current AGN modelling

- Combining MWL data on flux point level:
 - We construct flux points for single instruments separately
 - using different frameworks for different instruments
 - assuming simplistic models (e.g. power-law) - single instruments cannot constrain physical models (e.g. SSC)
 - Selection of the physical model based on flux points
- Fitting on flux points adds a bias:
 - Simple test with a blackbody spectrum using X-rays (XSPEC ([1999ascl.soft10005A](https://www.mpe.mpg.de/~krolik/xspec.html))) shows a bias of 1 sigma
- **Long-term goal: Enable event-based multi-messenger fits in Gammapy**
 - We start with blazars and JetSet as an example here



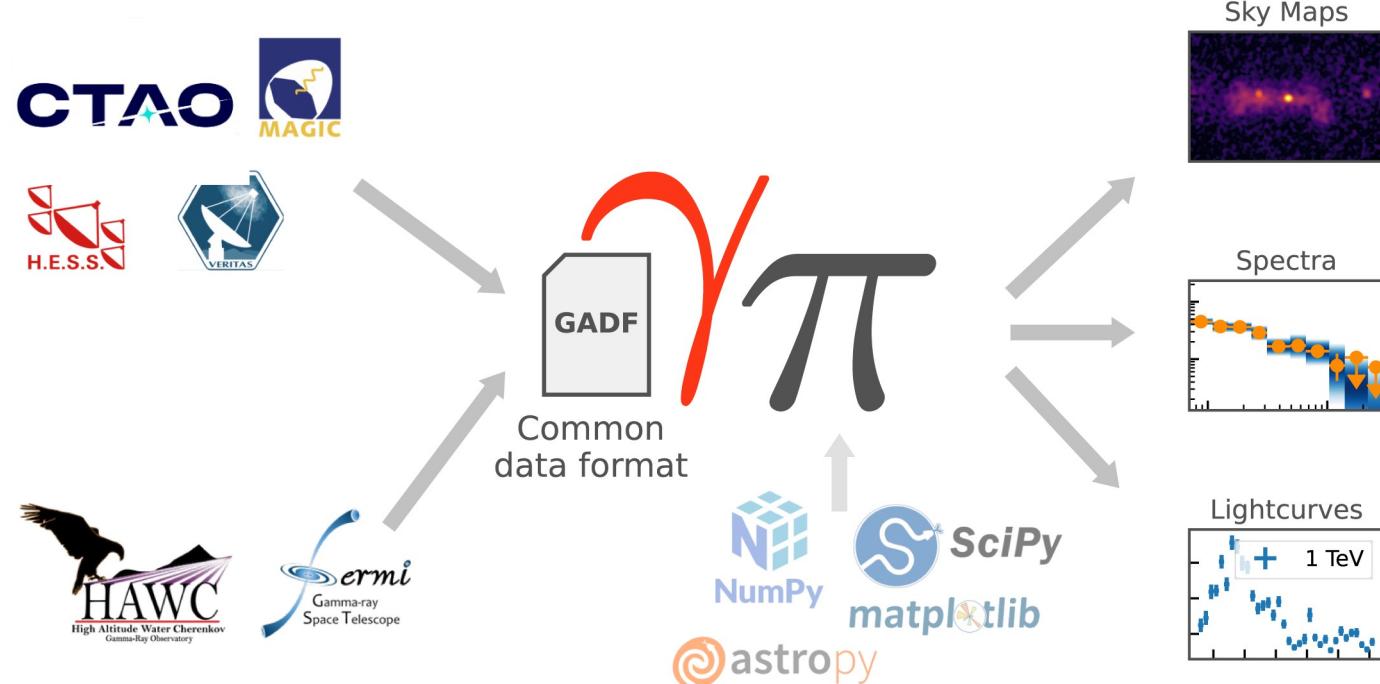
A new approach

- Joined MWL fit on event basis using Gammapy



A new approach

- Joined MWL fit on event basis using Gammapy
- Gammapy:
 - Python package for gamma-ray astronomy



Gammapy

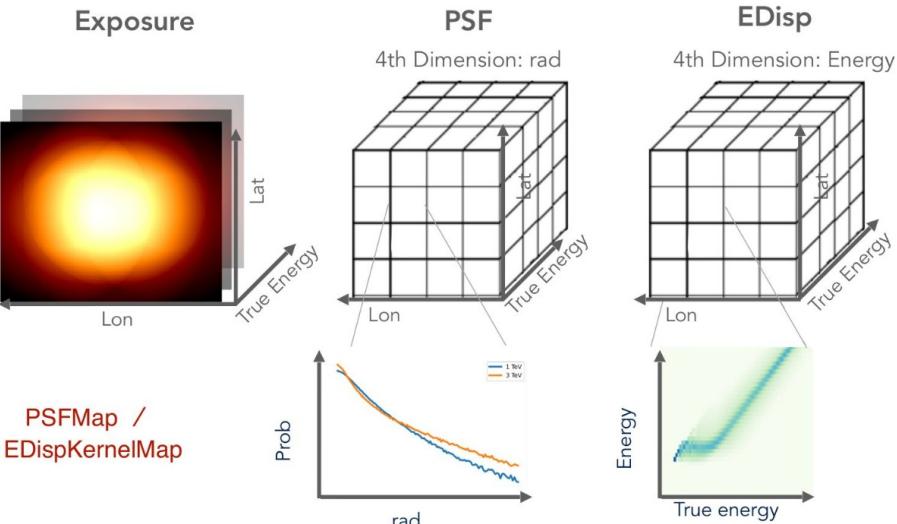
γ -like
events

DL3

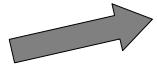
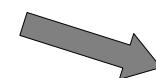
Time	RA	DEC	Energy	...
...				
...				

+

instruments
response
functions
(IRFs)



Gammapy

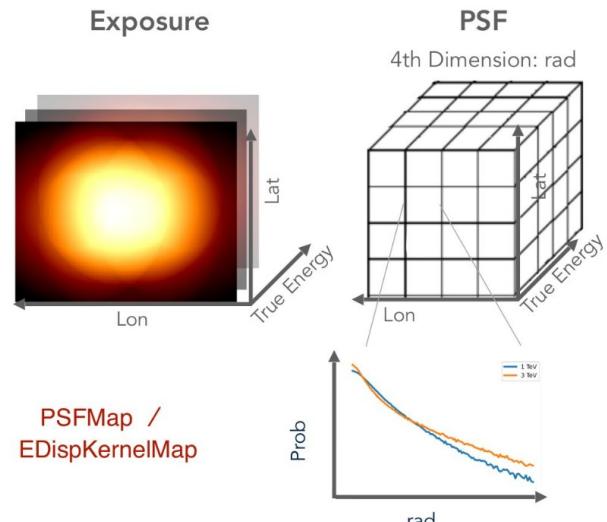
 $\gamma\pi$ 

γ -like
events

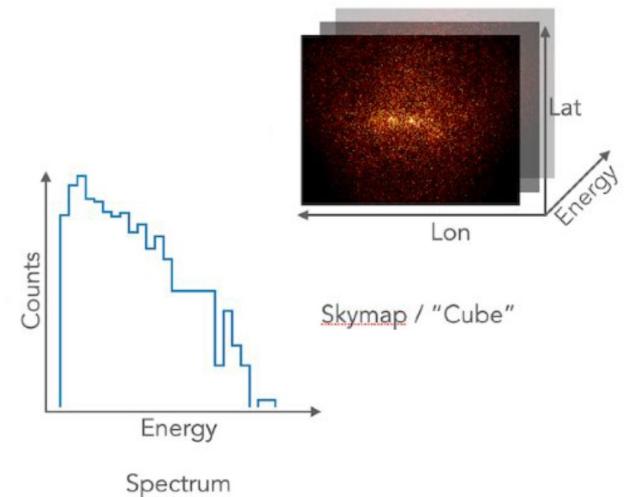
Time	RA	DEC	Energy	...
...				
...				

+

instruments
response
functions
(IRFs)

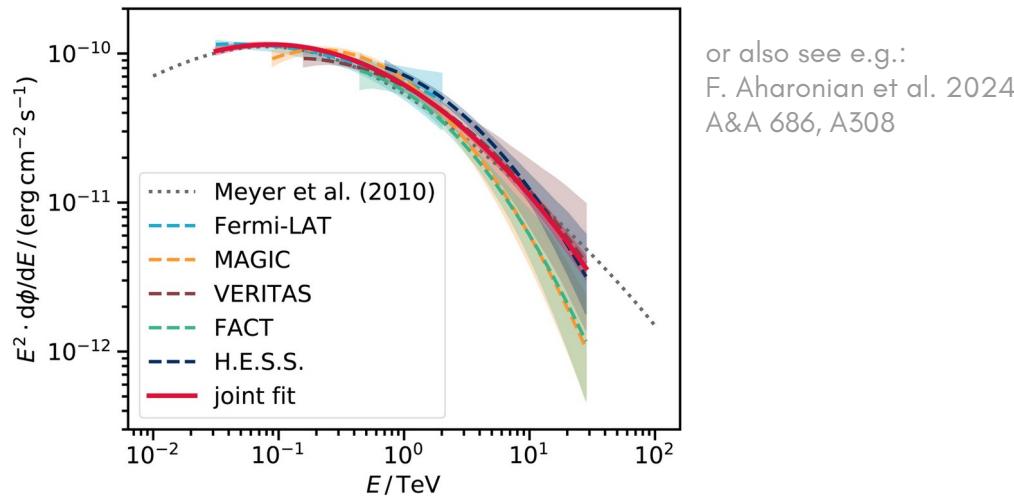


Binned data
 $\gamma\pi$ datasets:



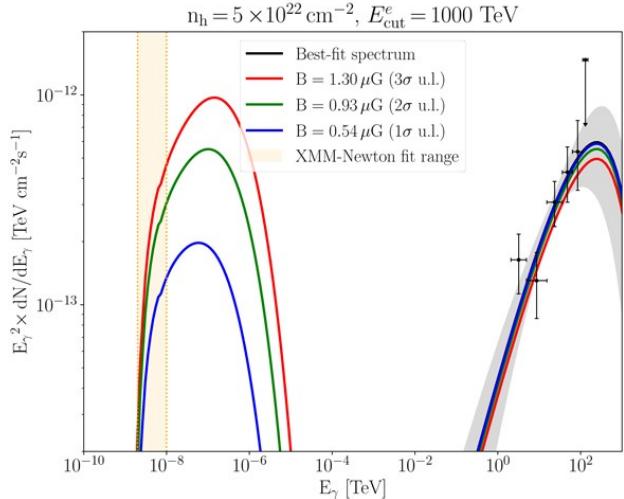
Gammapy & MWL data

- Works nicely for gamma-ray data and joint instrument fits:

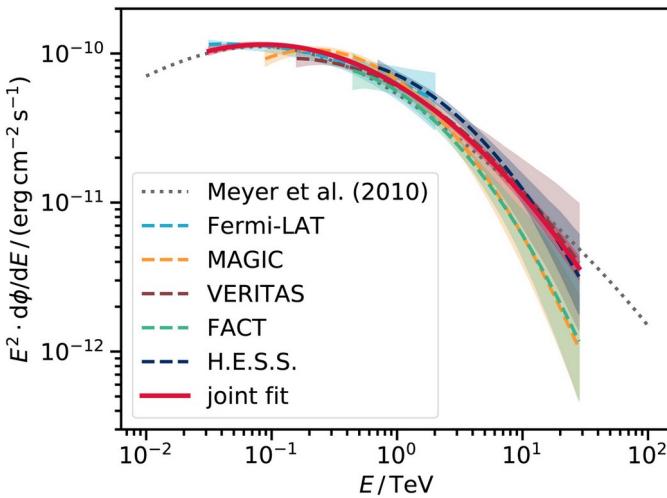


Gammapy & MWL data

- Works nicely for gamma-ray data and joint instrument fits:
- Extension to X-rays:
<https://github.com/registerrier/gammapy-ogip-spectra>



L. Giunti et al. 2022, A&A 667, A130

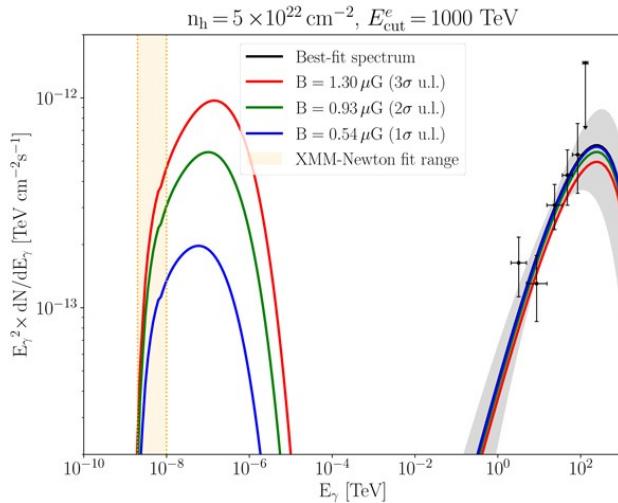


or also see e.g.:
F. Aharonian et al. 2024
A&A 686, A308

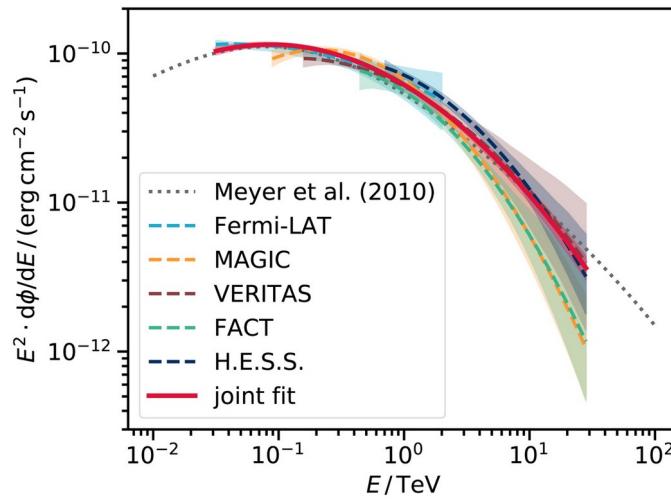
Gammapy & MWL data

- Works nicely for gamma-ray data and joint instrument fits:
- Extension to X-rays:

<https://github.com/registerrier/gammapy-ogip-spectra>

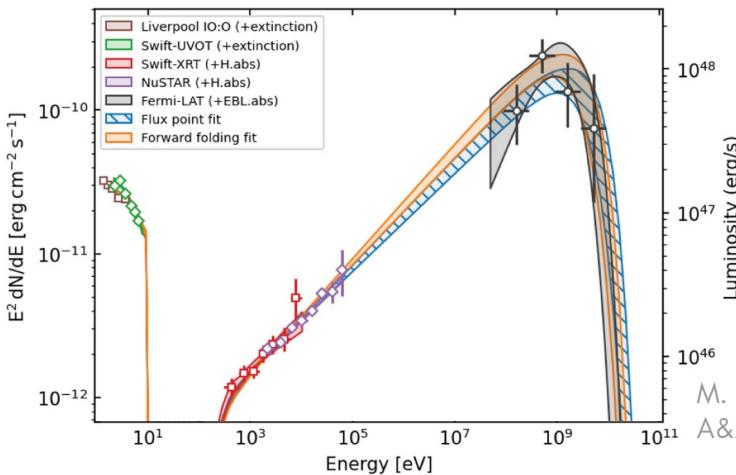


L. Giunti et al. 2022, A&A 667, A130



- Extension to UV/ optical:

https://github.com/mireianievas/gammapy_mwl_workflow



M. Nievas Rosillo et al. 2025,
A&A, 693, A287

or also see e.g.:
F. Aharonian et al. 2024
A&A 686, A308

New AGN modelling approach with Gammapy

- Building on this [gammapy_mwl_workflow](#) and public data

New AGN modelling approach with Gammapy

- Building on this [gammapy_mwl_workflow](#) and public data
- IRFs from UV to gamma-rays within Gammapy
 - Real data from Swift: UVOT + XRT, Fermi-LAT (UV, X-rays and gamma-rays)
 - Simulations from CTAO (Very-high-energy gamma-rays)

New AGN modelling approach with Gammapy

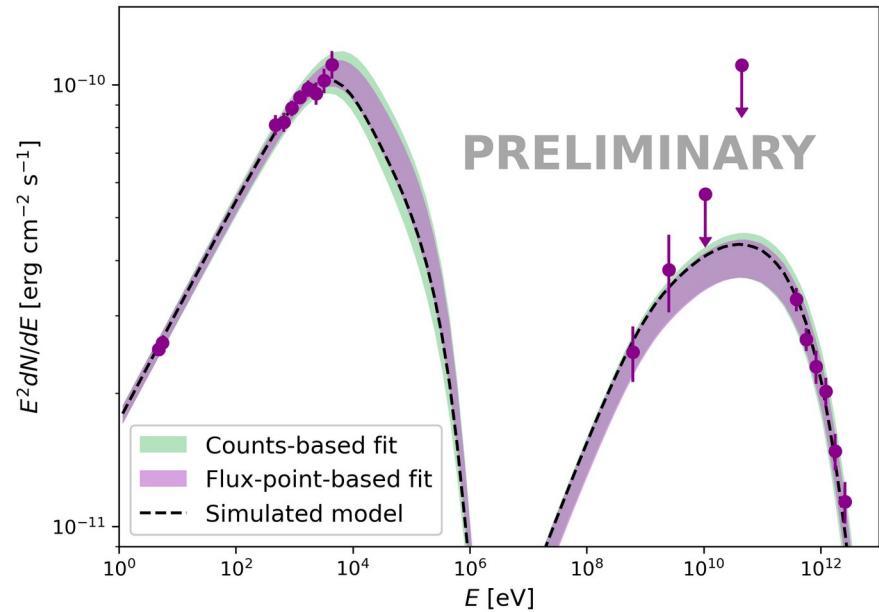
- Building on this [gammapy_mwl_workflow](#) and public data
- IRFs from UV to gamma-rays within Gammapy
 - Real data from Swift: UVOT + XRT, Fermi-LAT (UV, X-rays and gamma-rays)
 - Simulations from CTAO (Very-high-energy gamma-rays)
- Simulation of MWL data based on these IRFs using Gammapy:
 - Physical model:
 - Leptonic Self-Synchrotron (SSC) model with EBL absorption
 - JetSeT ([2020ascl.soft09001T](#)) with its Gammapy plugin
 - + Systematics for different bands
 - + Absorption processes via the `sherpa` ([10.5281/zenodo.593753.](#)) module

Gammapy & SSC

- We produce 2 MWL Gammapy datasets:
 - 1) Event data (DL3) set with event lists and IRFs
 - 2) Flux points data set:
 - Produced for each waveband
 - Assuming simple power laws
 - Mimics common approaches

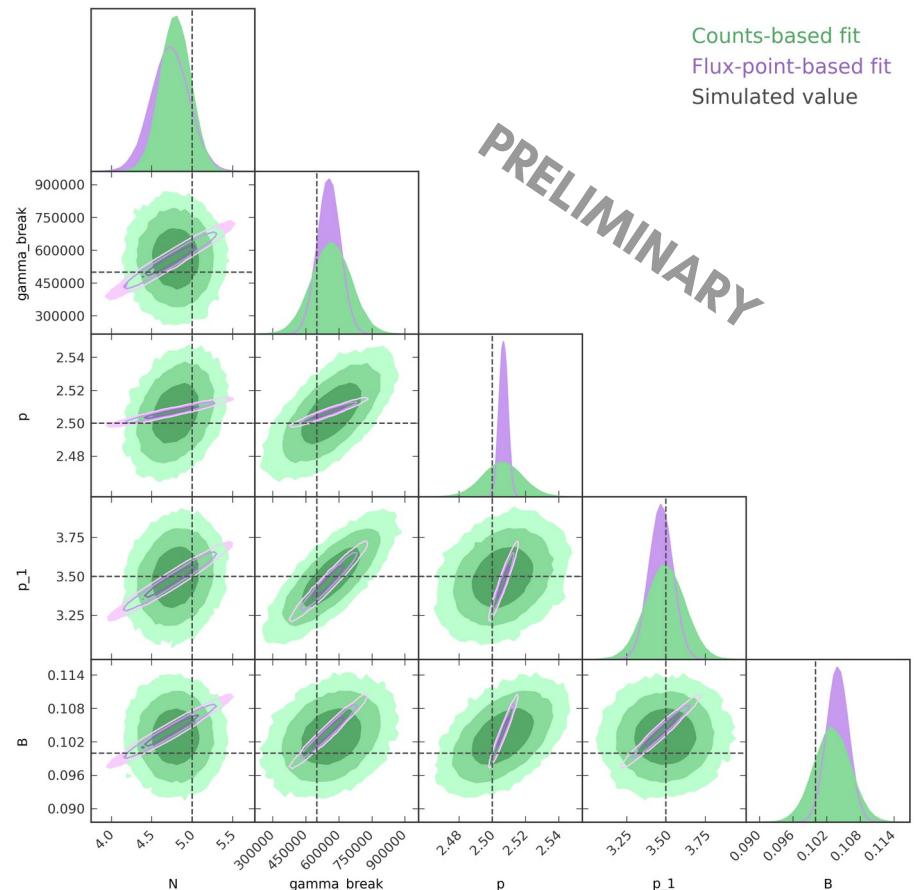
Gammapy & SSC

- We produce 2 MWL Gammapy datasets:
 - 1) Event data (DL3) set with event lists and IRFs
 - 2) Flux points data set:
 - Produced for each waveband
 - Assuming simple power laws
 - Mimics common approaches
- We fit both datasets with a SSC model using JetSeT



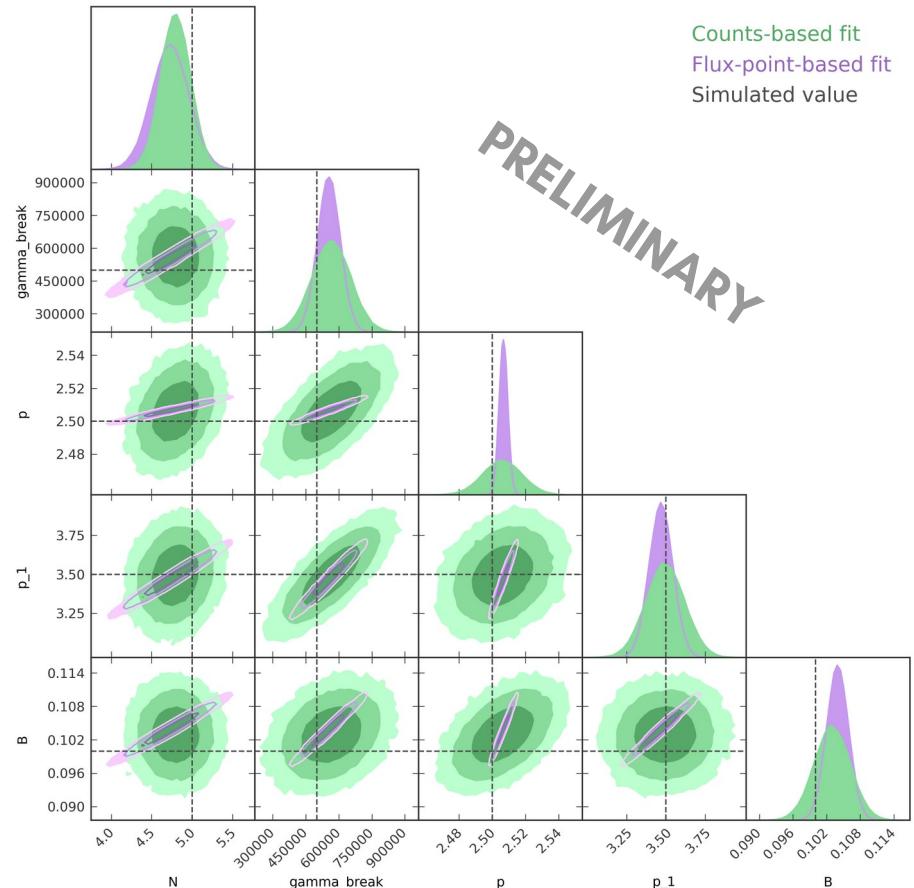
Facilitating multi-messenger modelling using Gammapy

- Works quite well and events fit shows
 - less bias/better error estimation
 - better convergence – up to ~ 5 free parameters



Facilitating multi-messenger modelling using Gammapy

- Works quite well and events fit shows
 - less bias/better error estimation
 - better convergence – up to ~ 5 free parameters
- But events fit takes longer (5-10x),
e.g. 35 vs 10 min for 3,
190 vs. 50 min for 5 free parameters
 - JetSet timing inside Gammapy at one E: 0.02 sec
 - More complex models (e.g. hadronic ones) > 1 min
- cannot be run on the fly



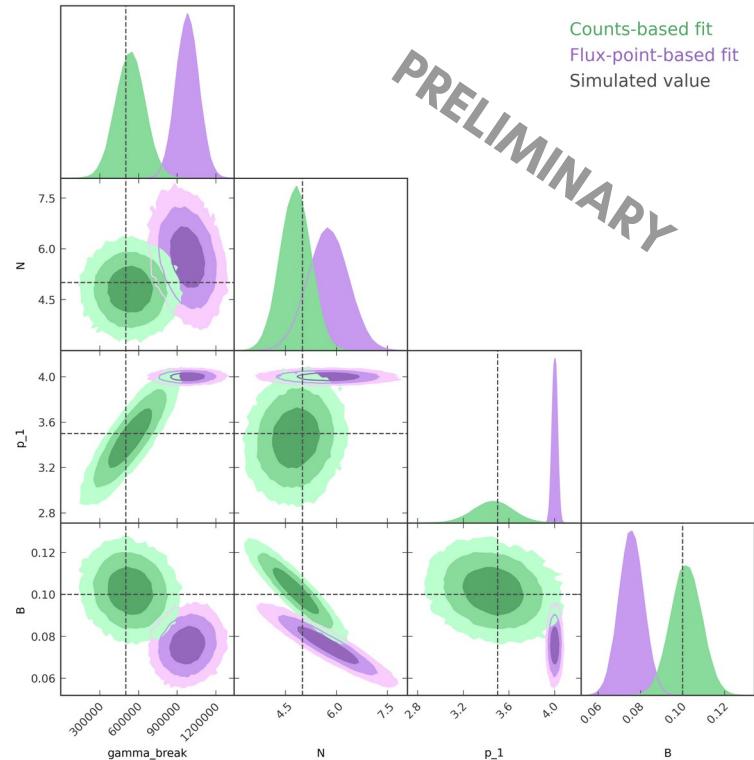
Facilitating multi-messenger modelling using Gammapy

- Work-around methods:
 - Grid of precomputed models
 - 10 grid points per parameters
 - 200 grid points in energy
 - Use grid interpolation during the fit

Facilitating multi-messenger modelling using Gammapy

- Work-around methods:

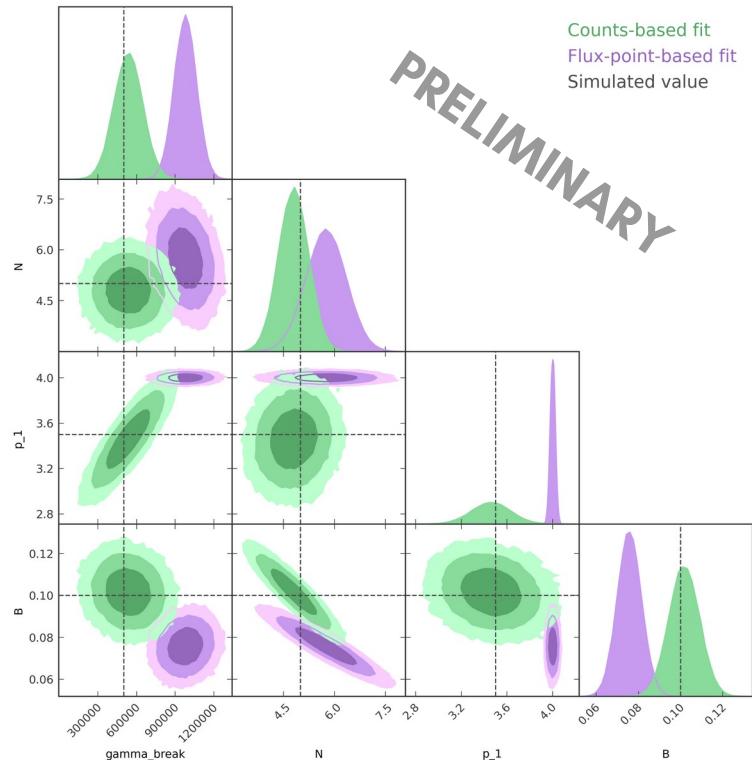
- Grid of precomputed models
 - 10 grid points per parameters
 - 200 grid points in energy
- Use grid interpolation during the fit
- Seems to work well



Facilitating multi-messenger modelling using Gammapy

- Work-around methods:

- Grid of precomputed models
 - 10 grid points per parameters
 - 200 grid points in energy
- Use grid interpolation during the fit
- Seems to work well
- BUT e.g. overloads memory for a high number of parameters



Lessons learnt so far

- **Gammapy** enables us to **fit physical models directly on event basis**
 - Reducing biases and assumptions
 - Taking into account absorption, systematics or e.g. also ULs naturally
 - But we are still limited by the large amount of parameters

Lessons learnt so far

- **Gammapy** enables us to **fit physical models directly on event basis**
 - Reducing biases and assumptions
 - Taking into account absorption, systematics or e.g. also ULs naturally
 - But we are still limited by the large amount of parameters
- For **fast models (e.g. leptonic ones)**
 - Fitting can be run by Gammapy **on the fly**
 - We also investigate various fitting methods and Bayesian nested sampling, which is limited by the Fermi analysis → minuit preferred currently

Lessons learnt so far

- **Gammapy** enables us to **fit physical models directly on event basis**
 - Reducing biases and assumptions
 - Taking into account absorption, systematics or e.g. also ULs naturally
 - But we are still limited by the large amount of parameters
- For **fast models (e.g. leptonic ones)**
 - Fitting can be run by Gammapy **on the fly**
 - We also investigate various fitting methods and Bayesian nested sampling, which is limited by the Fermi analysis → minuit preferred currently
- For **slow models (e.g. hadronic ones) - multimessenger fitting**
 - We are currently investigating **grid models + interpolation** methods
 - **Neural networks** and other machine learning methods might be a promising alternative

**Thank you
for your attention!**

