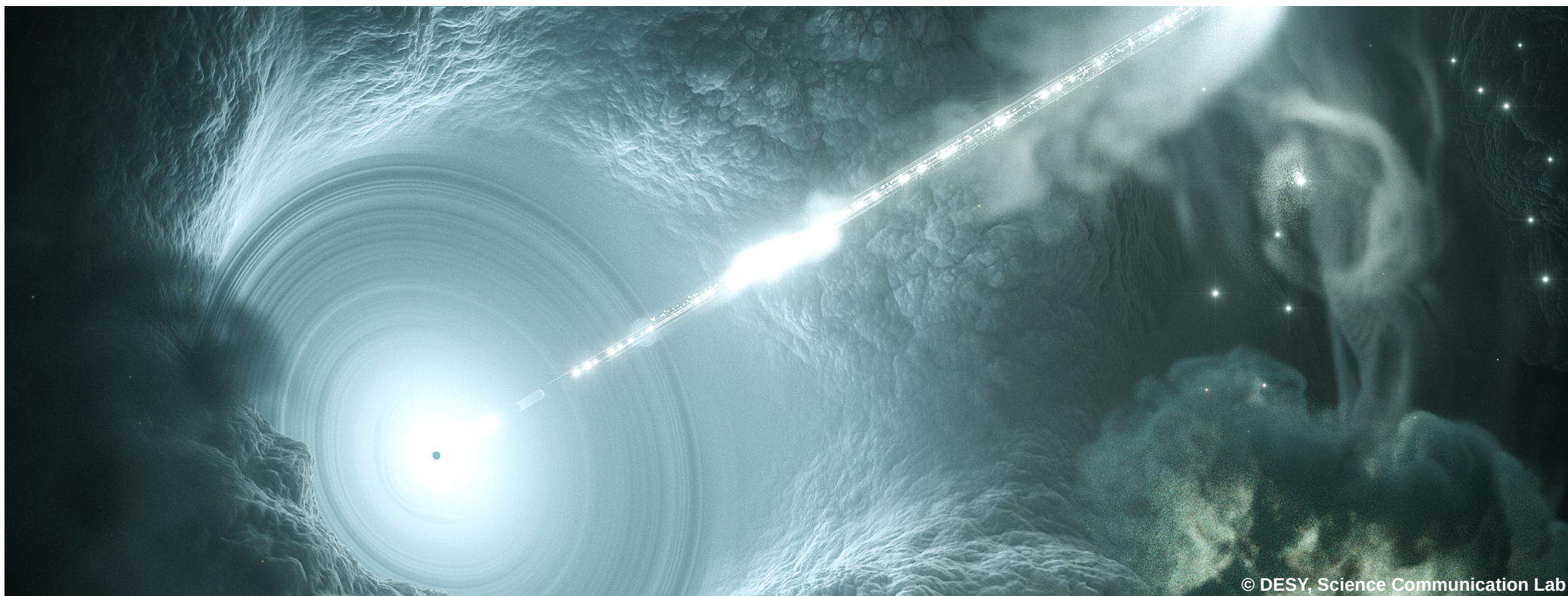




Facilitating multi-messenger modelling using Gammapy

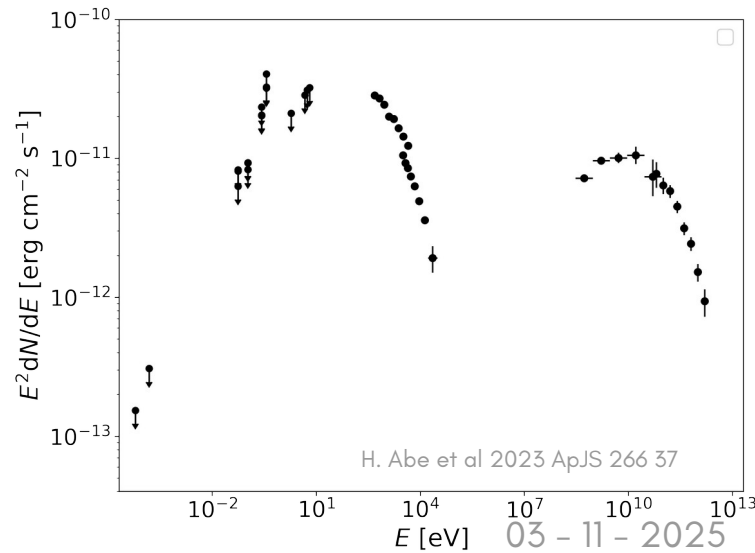
Lea Heckmann, Matteo Cerruti, Bruno Khélifi, Justin Albinet



Current AGN modelling

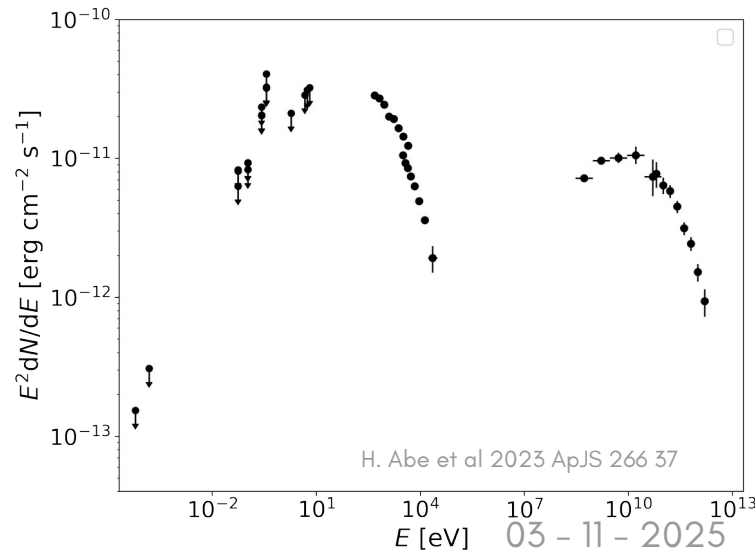
Current AGN modelling

- Combining MWL data on flux point level:



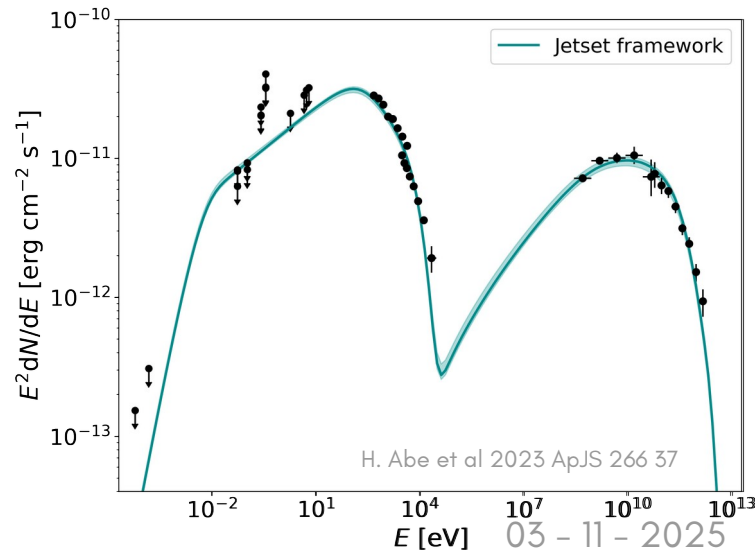
Current AGN modelling

- Combining MWL data on flux point level:
 - We construct flux points for single instruments separately
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 - assuming simplistic models (e.g. power-law) – single instruments cannot constrain physical models (e.g. SSC)



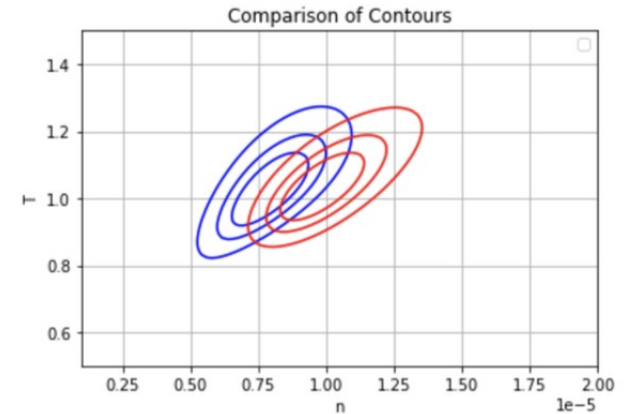
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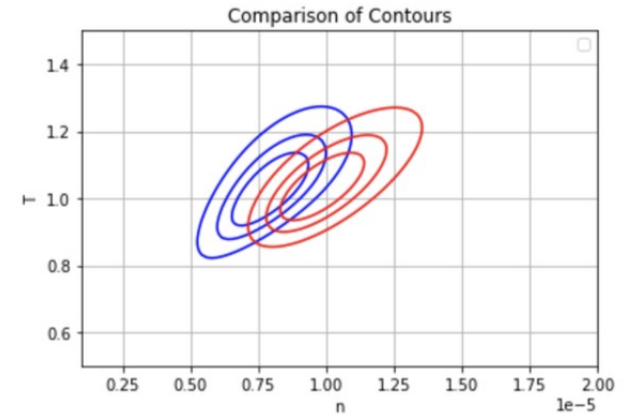
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- **Long-term goal: Enable event-based multi-messenger fits in Gammapy**
 - We start with blazars and JetSet as an example here



A new approach

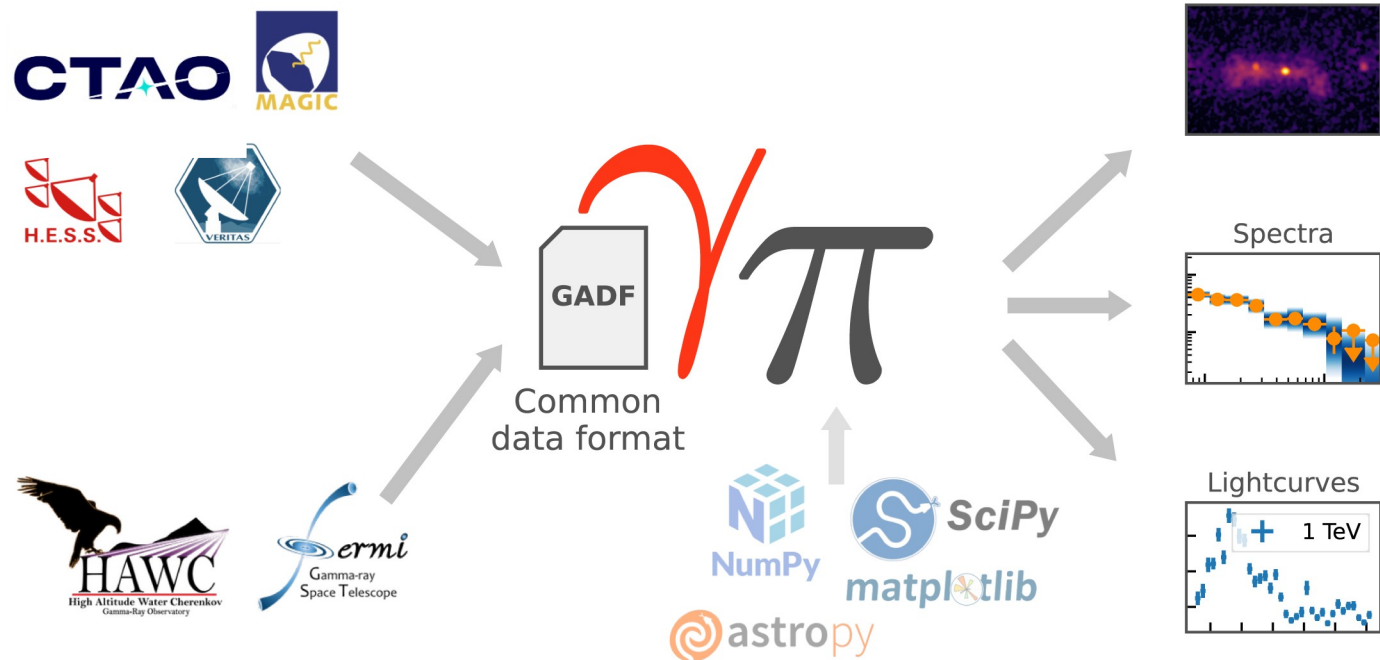
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A new approach

- Joined MWL fit on event basis using Gammapy
- Gammapy:
 - Python package for gamma-ray astronomy

 A **Python** package for **gamma-ray** astronomy



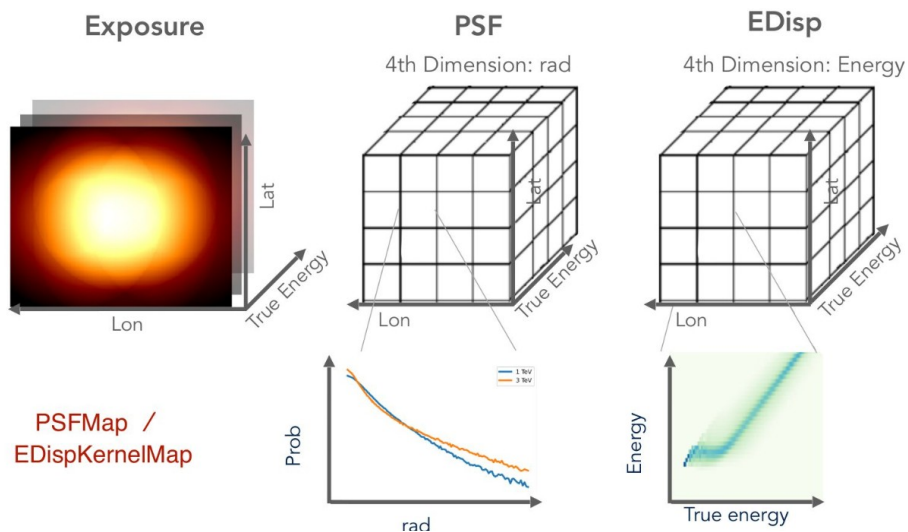
Gammapy

DL3

γ -like
events

Time	RA	DEC	Energy	...
...				
...				

+
instruments
response
functions
(IRFs)



Gammapy

DL3

$\gamma\pi$

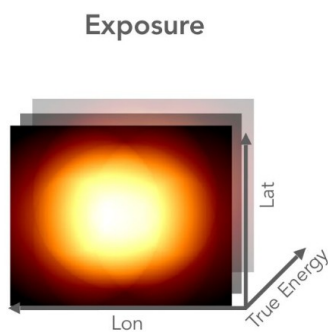
DL4

y-like
events

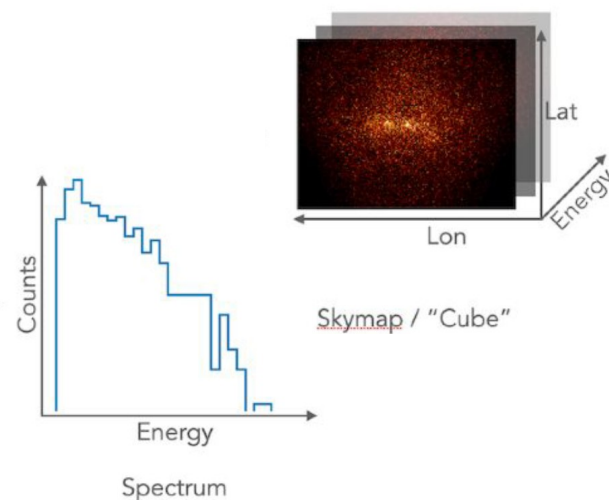
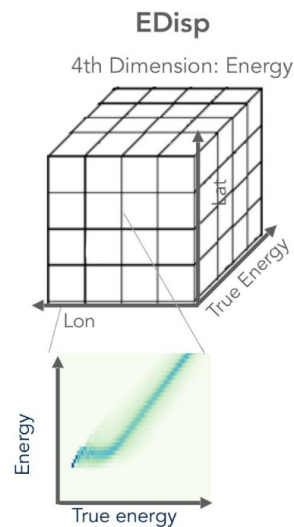
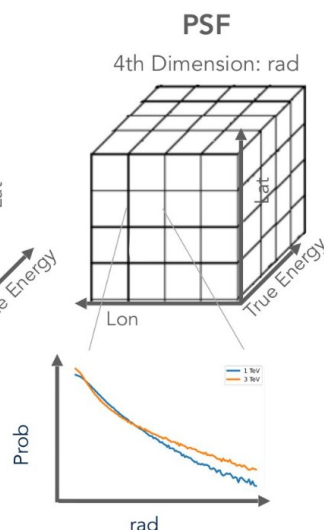
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...				
...				

Binned data
 $\gamma\pi$ datasets:

+
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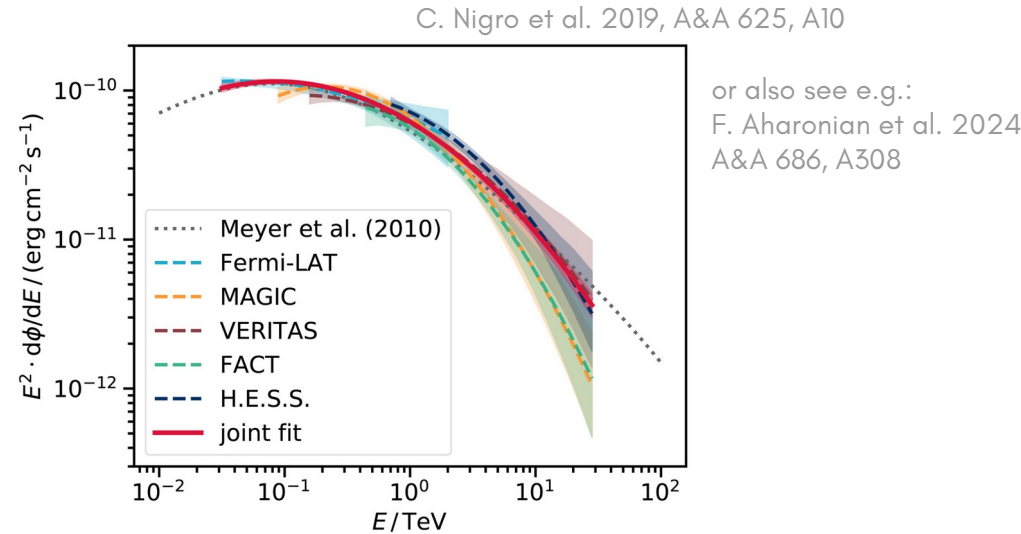


PSFMap /
EDispKernelMap



Gammapy & MWL data

- Works nicely for gamma-ray data and joint instrument fits:

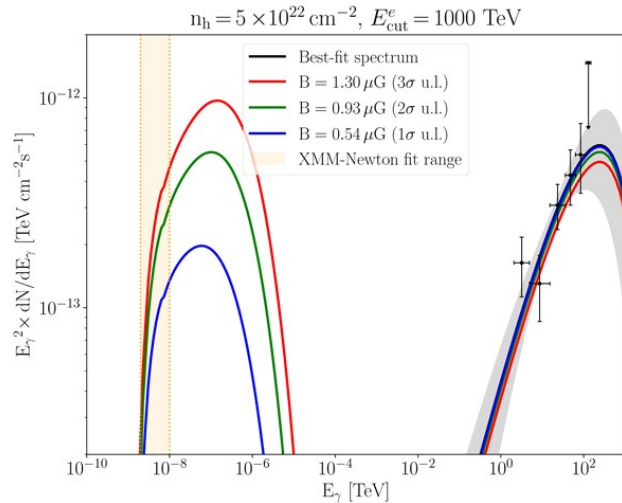


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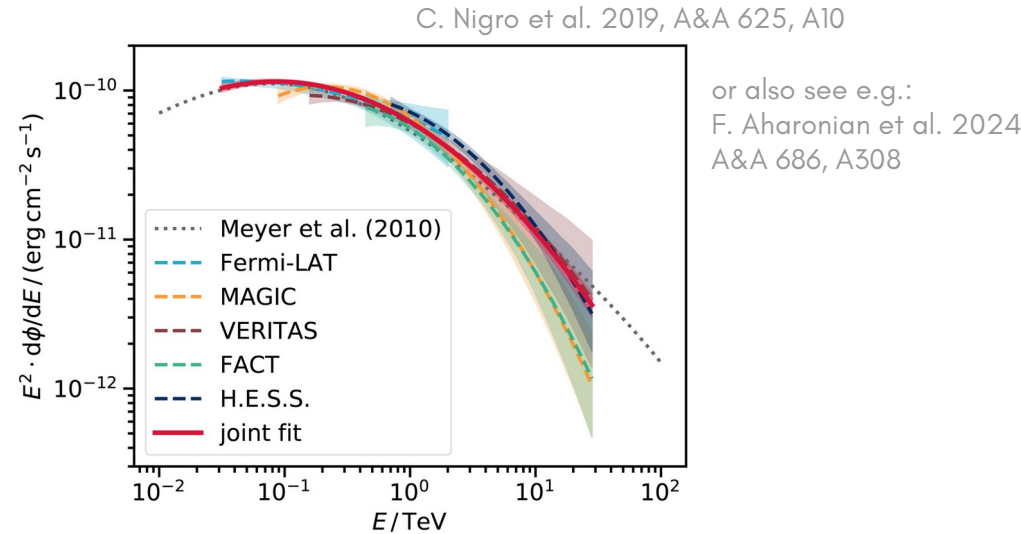
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<https://github.com/registerier/gammapy-ogip-spectra>



L. Giunti et al. 2022, A&A 667, A130

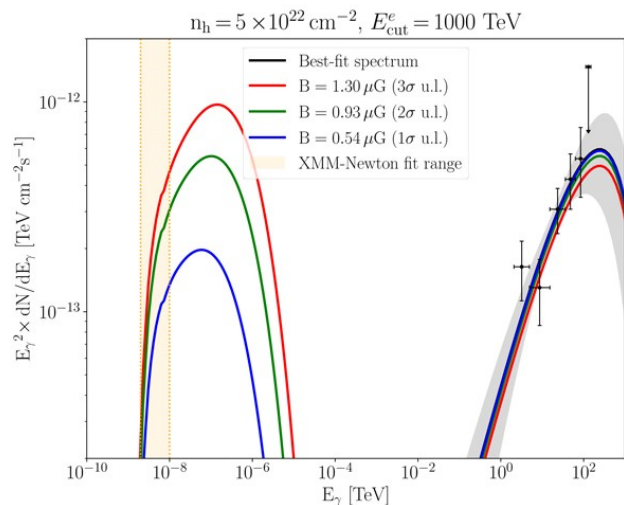


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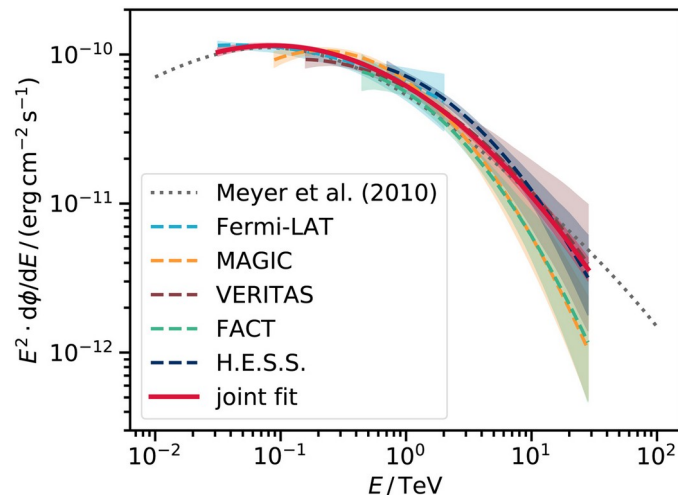
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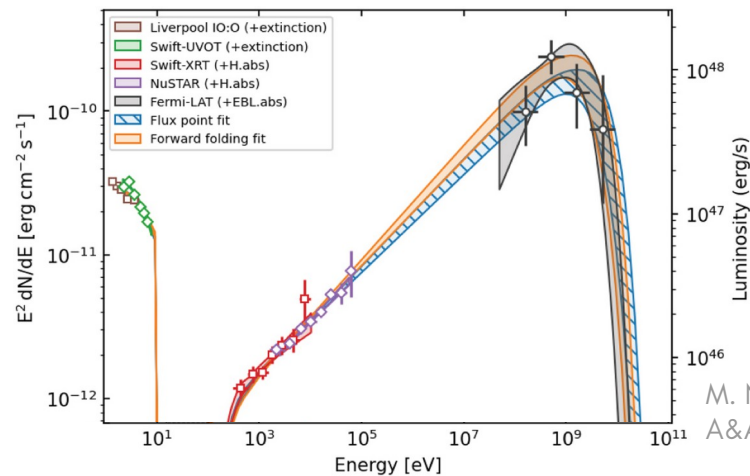
C. Nigro et al. 2019, A&A 625, A10



or also see e.g.:
F. Aharonian et al. 2024
A&A 686, A308

- Extension to UV/ optical:

https://github.com/mireianievas/gammapy_mwl_workflow



M. Nieves Rosillo et al. 2025,
A&A, 693, A287

New AGN modelling approach with Gammapy

- Building on this [gammapy_mwl_workflow](#) and public data

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New AGN modelling approach with Gammapy

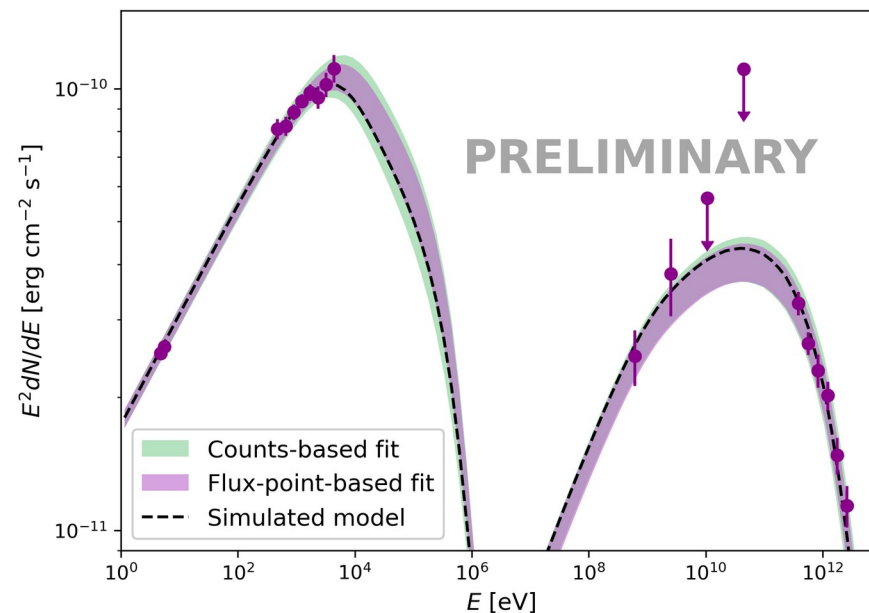
- Building on this [gammapy_mwl_workflow](#) and public data
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 - Real data from Swift: UVOT + XRT, Fermi-LAT (UV, X-rays and gamma-rays)
 - Simulations from CTAO (Very-high-energy gamma-rays)
- Simulation of MWL data based on these IRFs using Gammapy:
- Physical model:
 - Leptonic Self-Synchrotron (SSC) model with EBL absorption
 - JetSeT ([2020ascl.soft09001T](#)) with its Gammapy plugin
- + Systematics for different bands
- + Absorption processes via the sherpa ([10.5281/zenodo.593753.](#)) module

Gammapy & SSC

- We produce 2 MWL Gammapy datasets:
 - 1) Event data (DL3) set with event lists and IRFs
 - 2) Flux points data set:
 - Produced for each waveband
 - Assuming simple power laws
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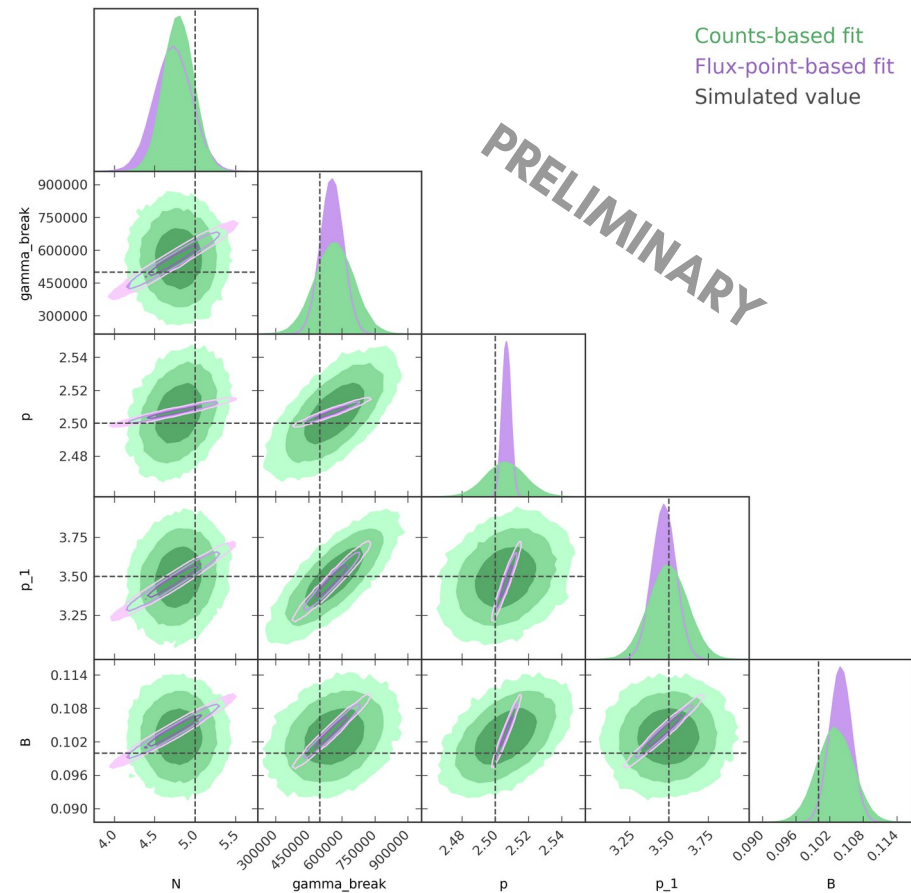
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- We fit both datasets with a SSC model using JetSeT



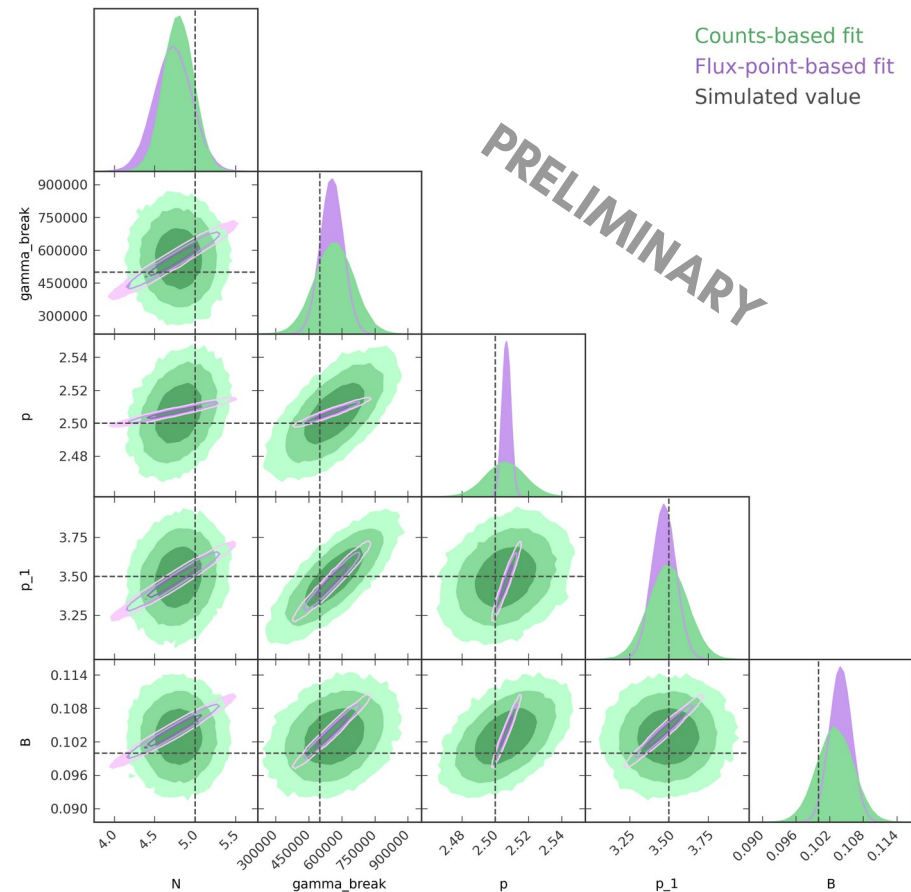
Facilitating multi-messenger modelling using Gammapy

- Works quite well and events fit shows
 - less bias/better error estimation
 - better convergence – up to ~ 5 free parameters



Facilitating multi-messenger modelling using Gammapy

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 - less bias/better error estimation
 - better convergence – up to ~ 5 free parameters
- But events fit takes longer (5-10x),
e.g. 35 vs 10 min for 3,
190 vs. 50 min for 5 free parameters
 - JetSet timing inside Gammapy
at one E: 0.02 sec
 - More complex models
(e.g. hadronic ones) > 1 min
→ cannot be run on the fly

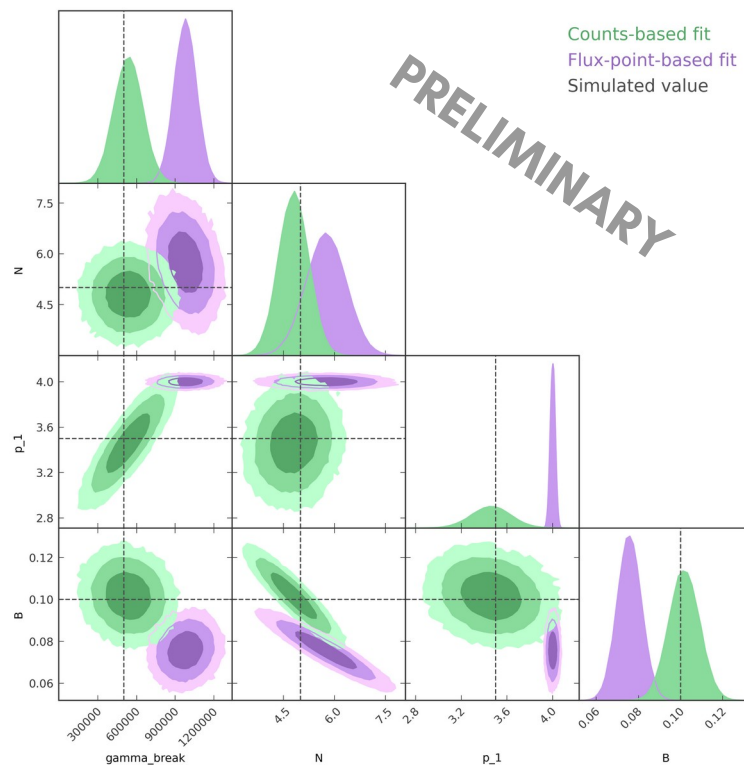


Facilitating multi-messenger modelling using Gammapy

- Work-around methods:
 - Grid of precomputed models
 - 10 grid points per parameters
 - 200 grid points in energy
 - Use grid interpolation during the fit

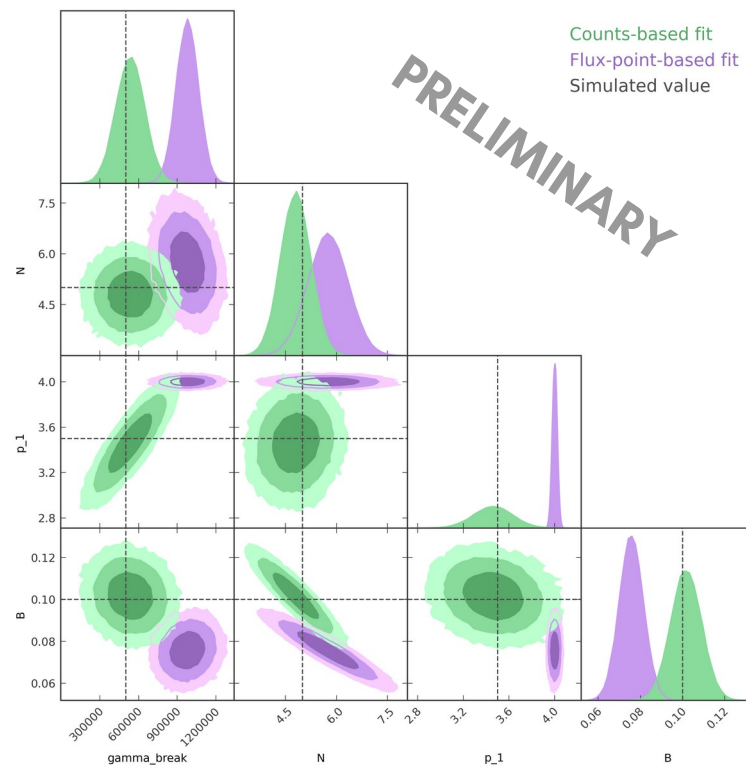
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 - Seems to work well
 - BUT e.g. overloads memory for a high number of parameters



Lessons learnt so far

- **Gammapy** enables us to **fit physical models directly on event basis**
 - Reducing biases and assumptions
 - Taking into account absorption, systematics or e.g. also ULs naturally
 - But we are still limited by the large amount of parameters

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- For **slow models (e.g. hadronic ones) – multimessenger fitting**
 - We are currently investigating **grid models + interpolation** methods
 - **Neural networks** and other machine learning methods might be a promising alternative

**Thank you
for your attention!**

