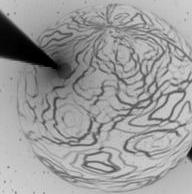


The MAGIC of Crab

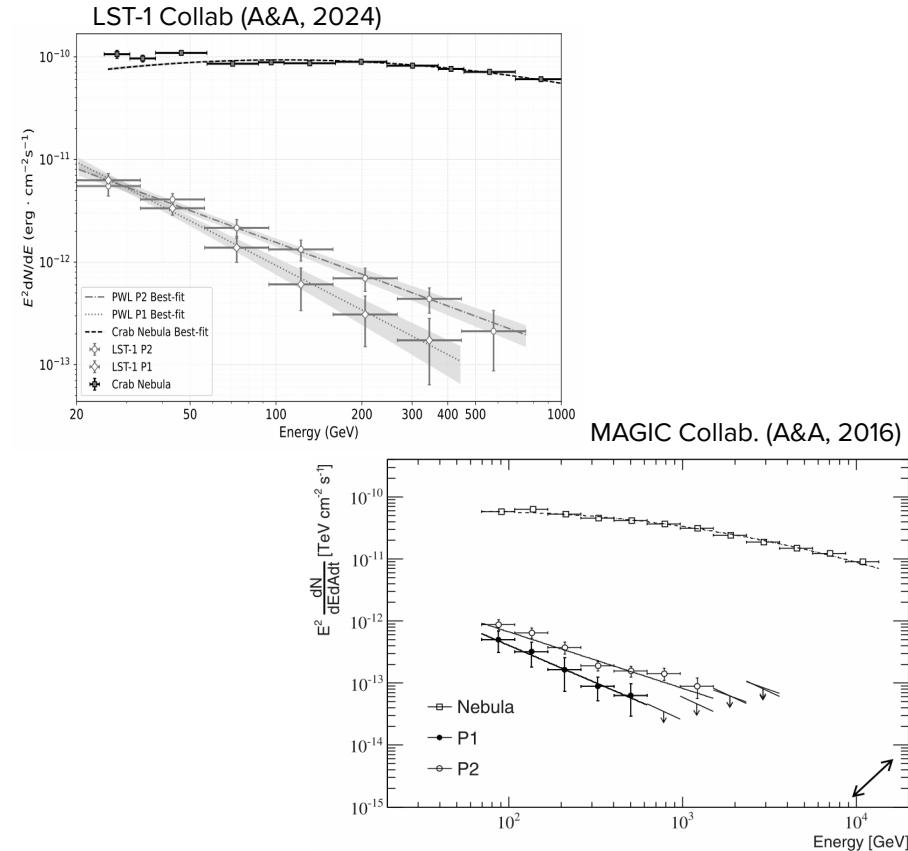


G. Ceribella, G. D'Amico, R. Mirzoyan
for the **MAGIC Collaboration**

04/11/2025

CRAB PULSAR (PSR J0534+2200)

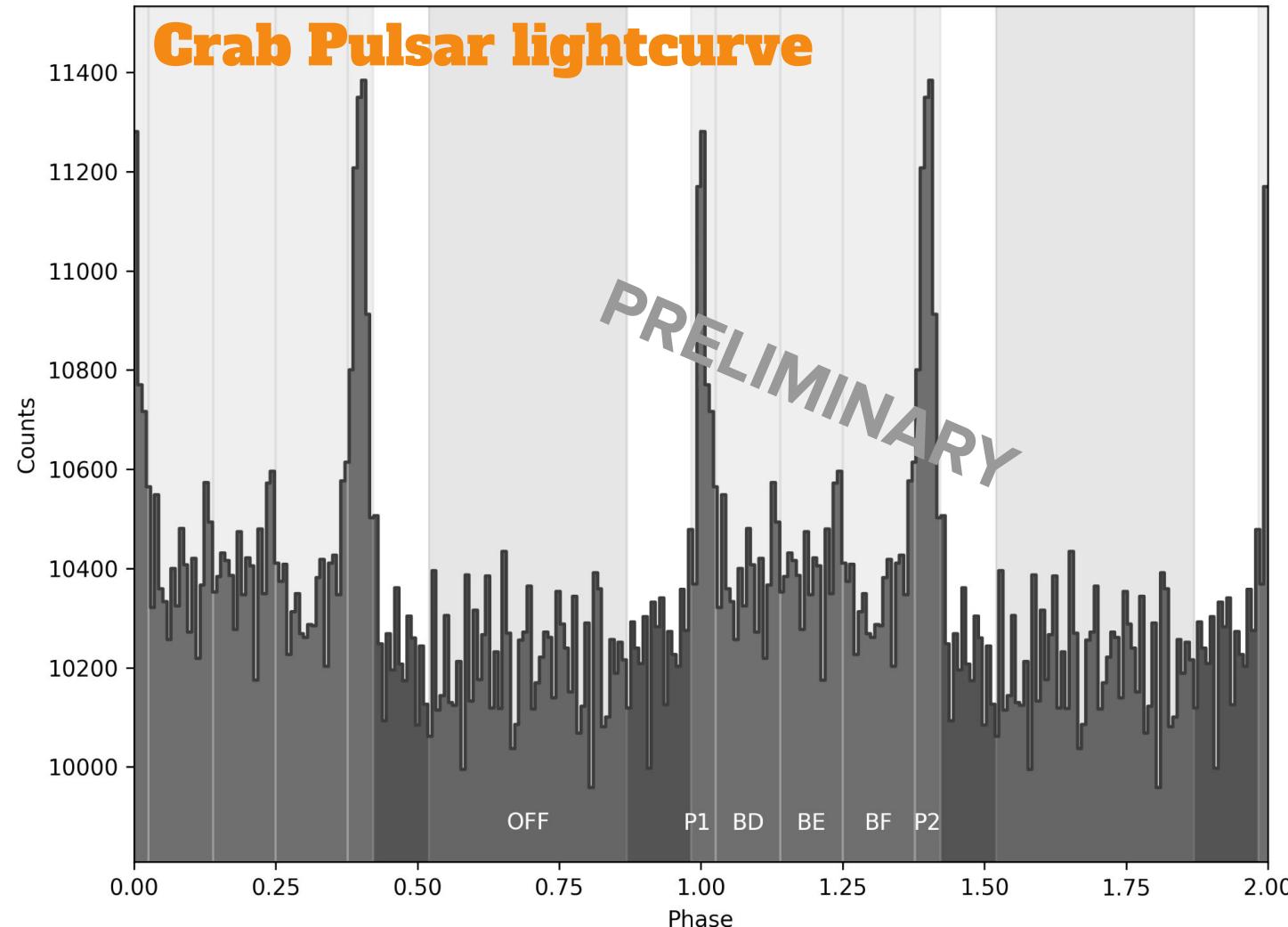
- Spin-down luminosity 4×10^{31} W (equivalent $10^5 L_\odot$)
- First known pulsar beyond **few GeV** (2008) and **1 TeV** (2016)
- Unique **brightness** beyond 100 GeV:
 - **Bridge emission** persisting at VHE
- Currently interpreted as **superposition** of multiple **curvature-radiation** and **inverse-Compton** components



MAGIC

- Two **IACTs** in La Palma operating in stereo since 2009.
- Absolute time resolution: ~ 200 ns
- Diameter of 17 m , stereo **effective area** at 50 GeV: 2×10^4 m².
- Energy range: **20 GeV** - 100 TeV
- Data used in this work:
 - **110 h** commissioning of **low-energy sum-trigger-II system**
 - Collected between 2015 and 2020
 - **Low-zenith** angle observations: $\cos(\text{zenith}) > 0.9$ (~ 26 deg)
 - Energy **threshold** (pulsar spectrum): ~ 20 GeV

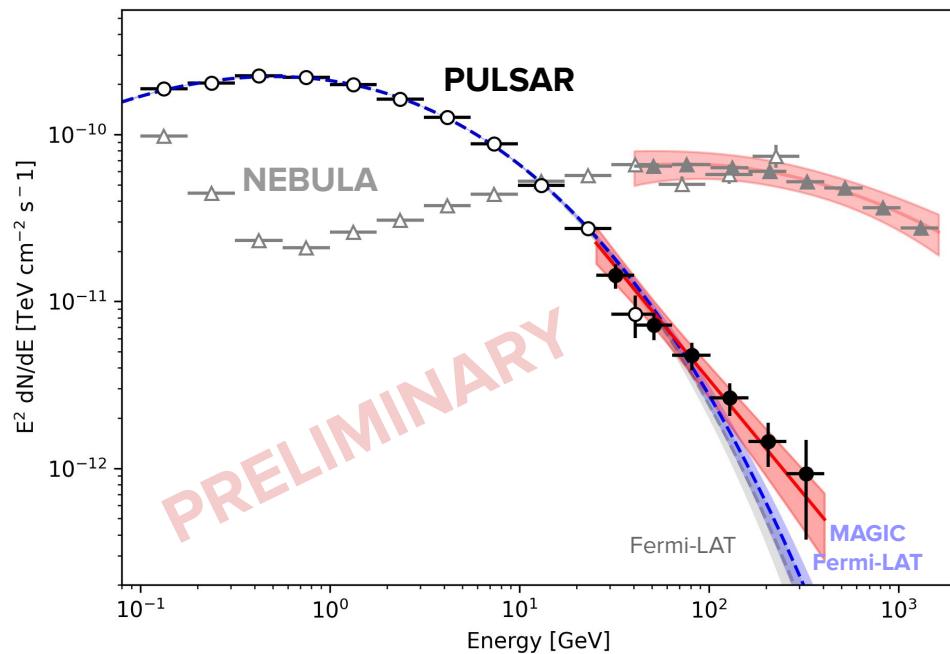
Crab Pulsar lightcurve



- ★ High-significant detection down to **20 GeV**
- ★ Cumulative significance $(P1+P2): \sim 22\sigma$
- ★ Sound statistics allows study of **bridge** emission in **sub-intervals**
 - 5σ in each of them
- ★ Background: **Crab Nebula**

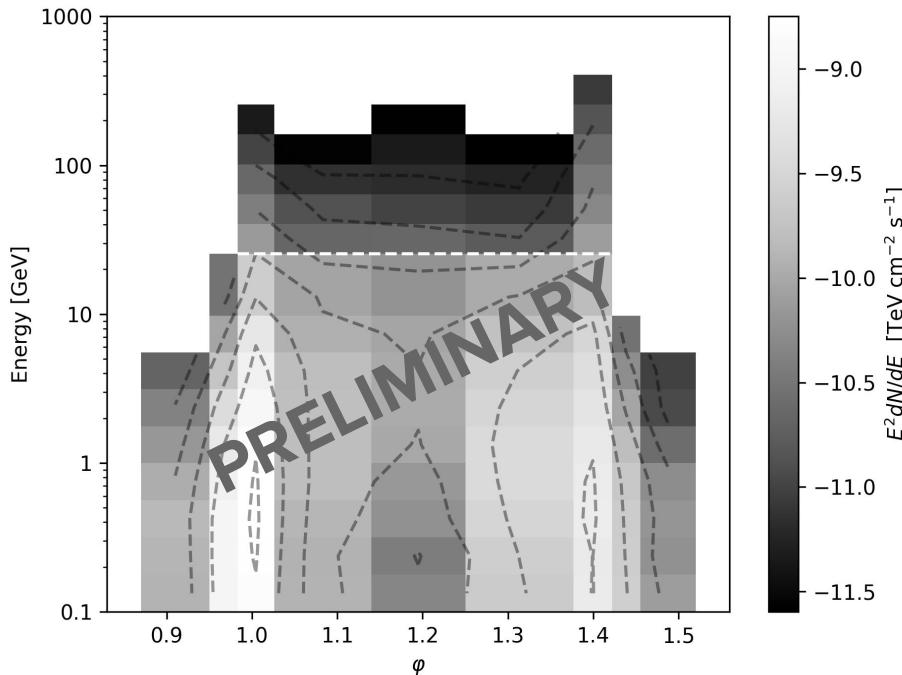
PHASE-AVERAGED SPECTRUM

- **Power-law** extension of the Fermi-LAT spectra:
 - Range: 25 GeV - 400 GeV
- **Nebula** gated in off phase interval (0.52,0.87)
- **Joint-fit** with LAT data:
 - MAGIC data “**drowned**” in LAT statistics
 - Compatible with cutoff from **curvature radiation?**

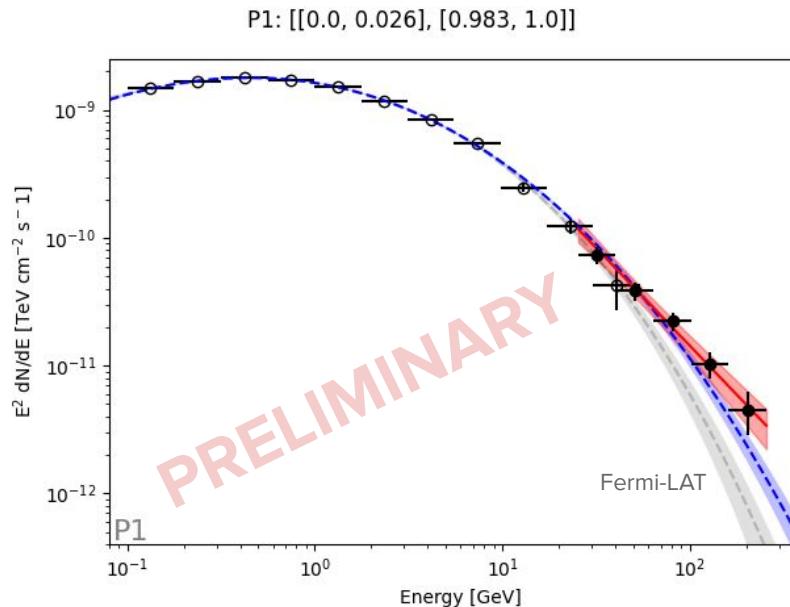


ADDING A DIMENSION: PULSED SPECTRA

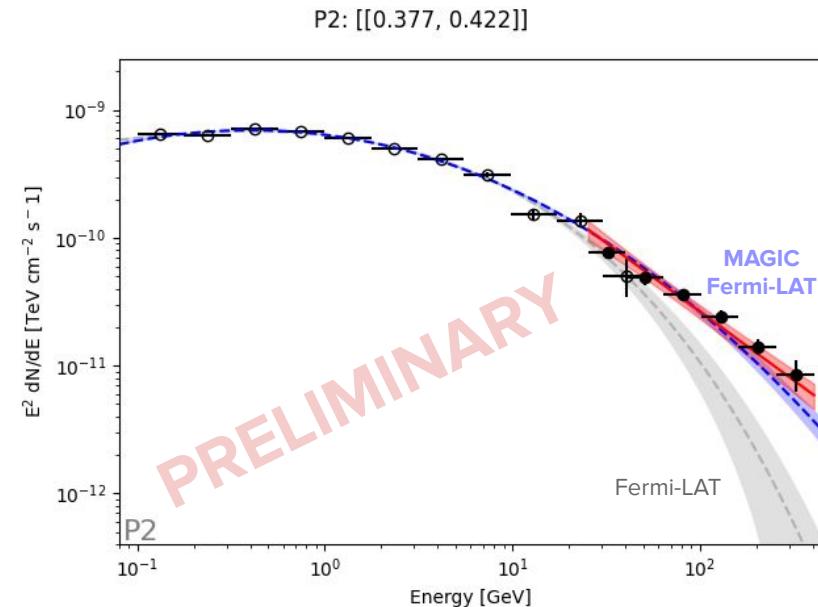
- Broad spectral changes within the **phase interval**
- Not confined to the two main **pulses** P1 and P2
- Bridge **substructure** studied for the **first time**
- **Power-law** extension tail much **more prominent** than in averaged spectrum
- **Joint Fermi-LAT + MAGIC** fits in all **phase intervals** with new **PLSECv4** model.



ADDING A DIMENSION: PULSED SPECTRA



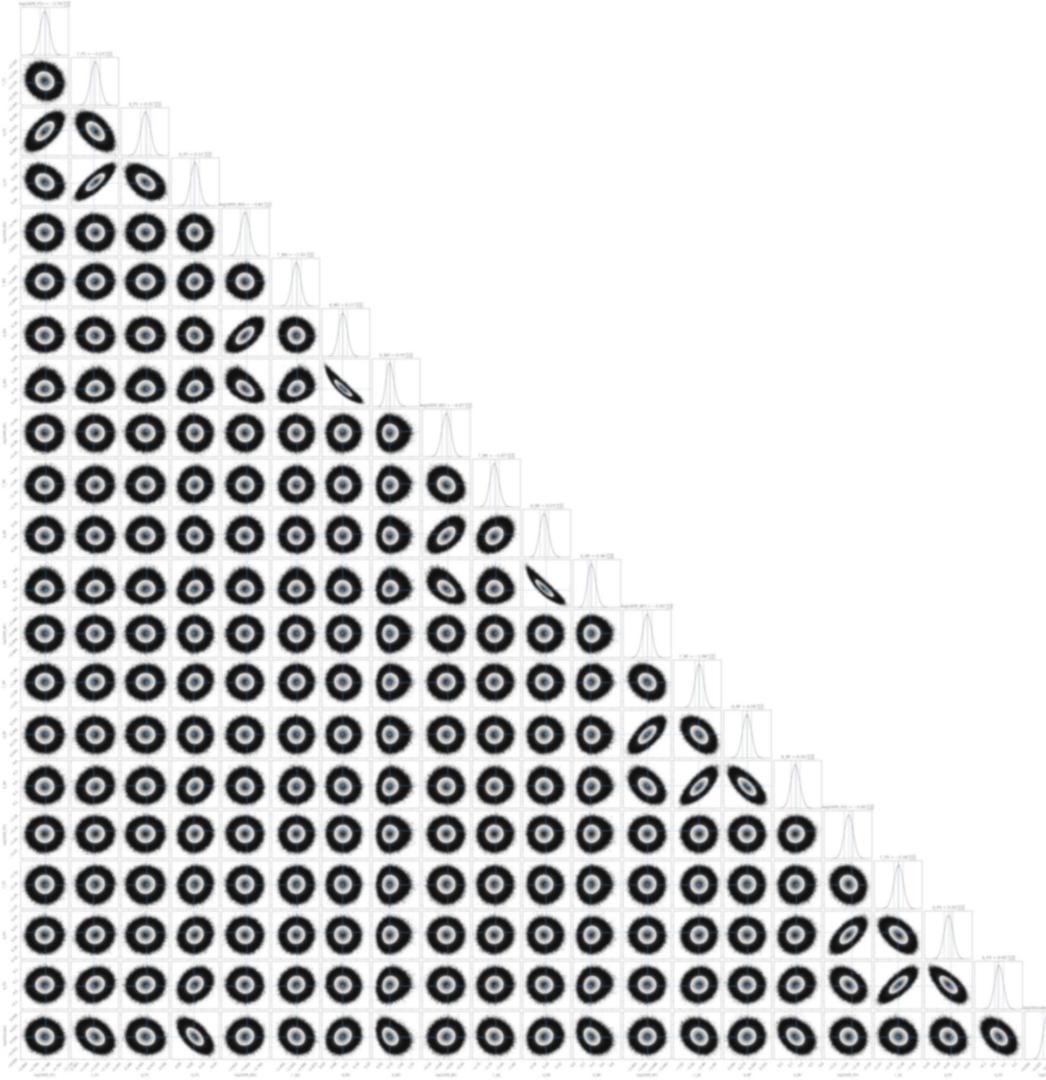
PWL Spectral Index (P1): **~3.5** (LST-1 ~3.4)
 $b(P1): 0.11 \pm 0.03$

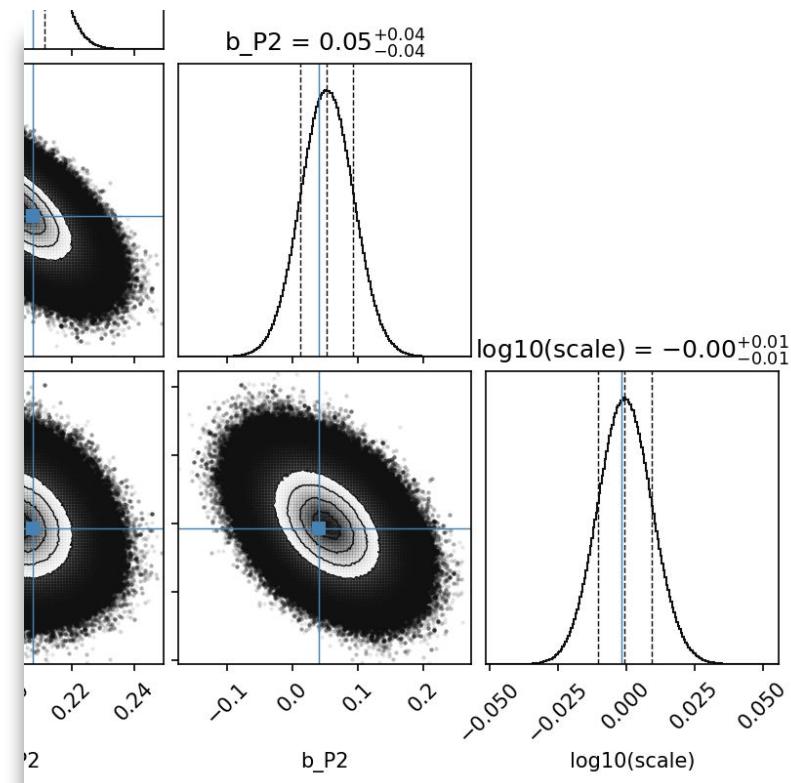
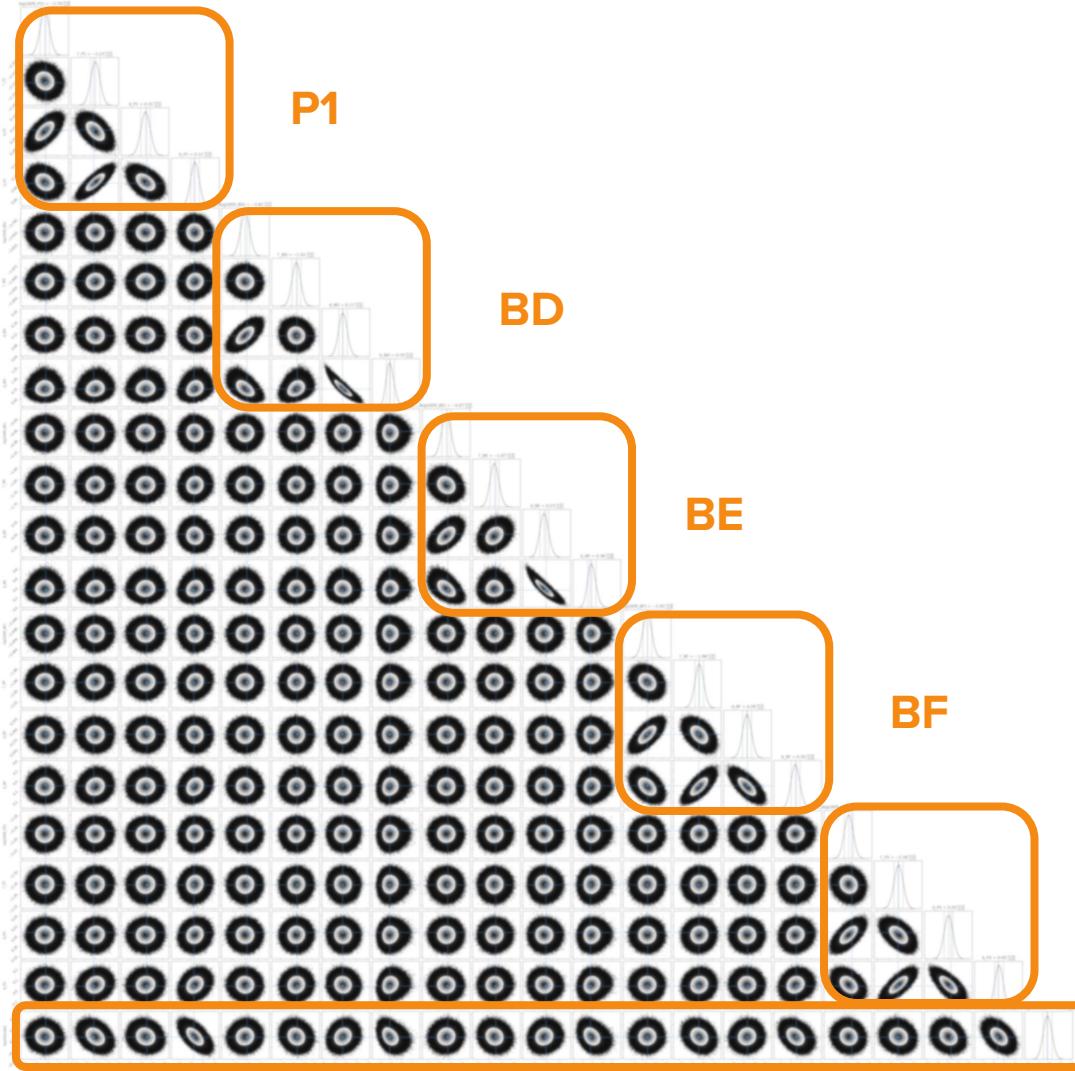


PWL Spectral Index (P2): **~3.1** (LST-1 ~3.0)
 $b(P2): 0.05 \pm 0.04$

SYSTEMATIC EFFECTS STUDY

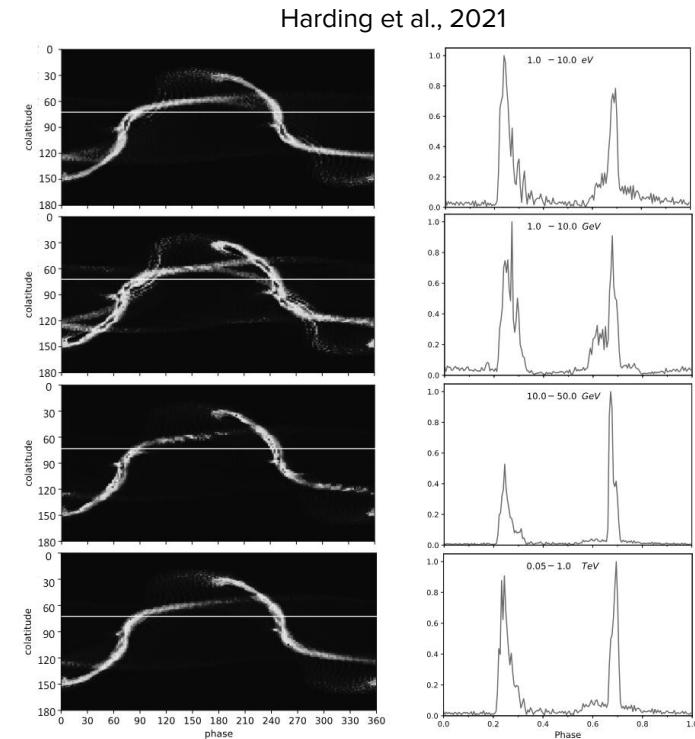
- Event selection
 - Gamma-Hadron cuts
- Pulsar **ephemeris**
- Phase gating suppressing **background-estimation systematics**
- Joint Fermi-LAT and MAGIC fits:
 - **Energy-scale** mismatch (equal for all phase regions)
 - MAGIC **flux-energy scaling law** determined from MC simulations
 - Energy scale modeled as **nuisance parameter**:
 - Single **simultaneous fit** of Fermi and MAGIC data in **all phase regions**
 - **Correlation** among phases through the **energy scaling** parameter
 - Results in an **estimate** of Fermi-LAT/MAGIC **energy mismatch**: **±2%**
- Sum of **systematic** contributions **comparable** to **statistical** uncertainties





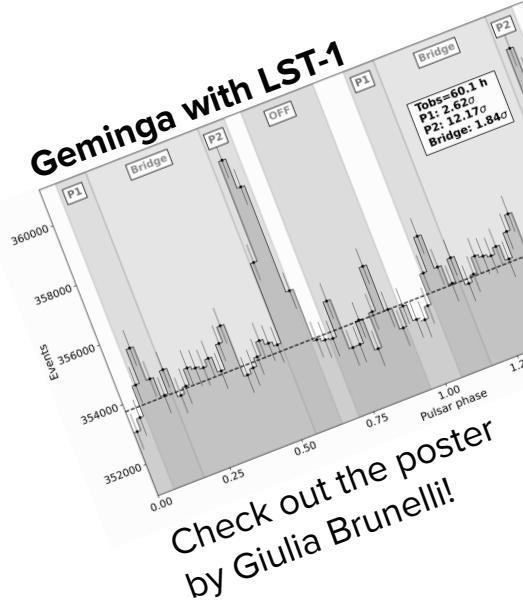
BATTLING MODELS

- Long march **towards** and beyond the **light cylinder**
- Two classes of models:
 - **Linear parallel E field** at the **Y point** (Harding et al., etc)
 - **Magnetic reconnection** in the **striped wind**
(Cerutti et al, etc)
- **Model degeneracy** impacted by **phase-dependent** spectra
- Need to account phase dependency in models.



CONCLUSIONS

- The **Crab pulsar** still offers a **rich playground for VHE pulsar model testing** after almost 20 years from its first discovery
- **Deep exposure** above **20 GeV** with **MAGIC** (sum-trigger-II) **yields unprecedented dataset** for joint **phase-spectral** studies (publication soon...)
- Novel **multi-instrument joint fit strategy** **incorporates systematic** effects and **provides energy scale** mismatch **estimate**
- **More pulsars** are out waiting for the **CTAO** era...



GRACIAS!