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DEGLI STUDI
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Gravitational waves from supermassive black holes at pulsar timing arrays

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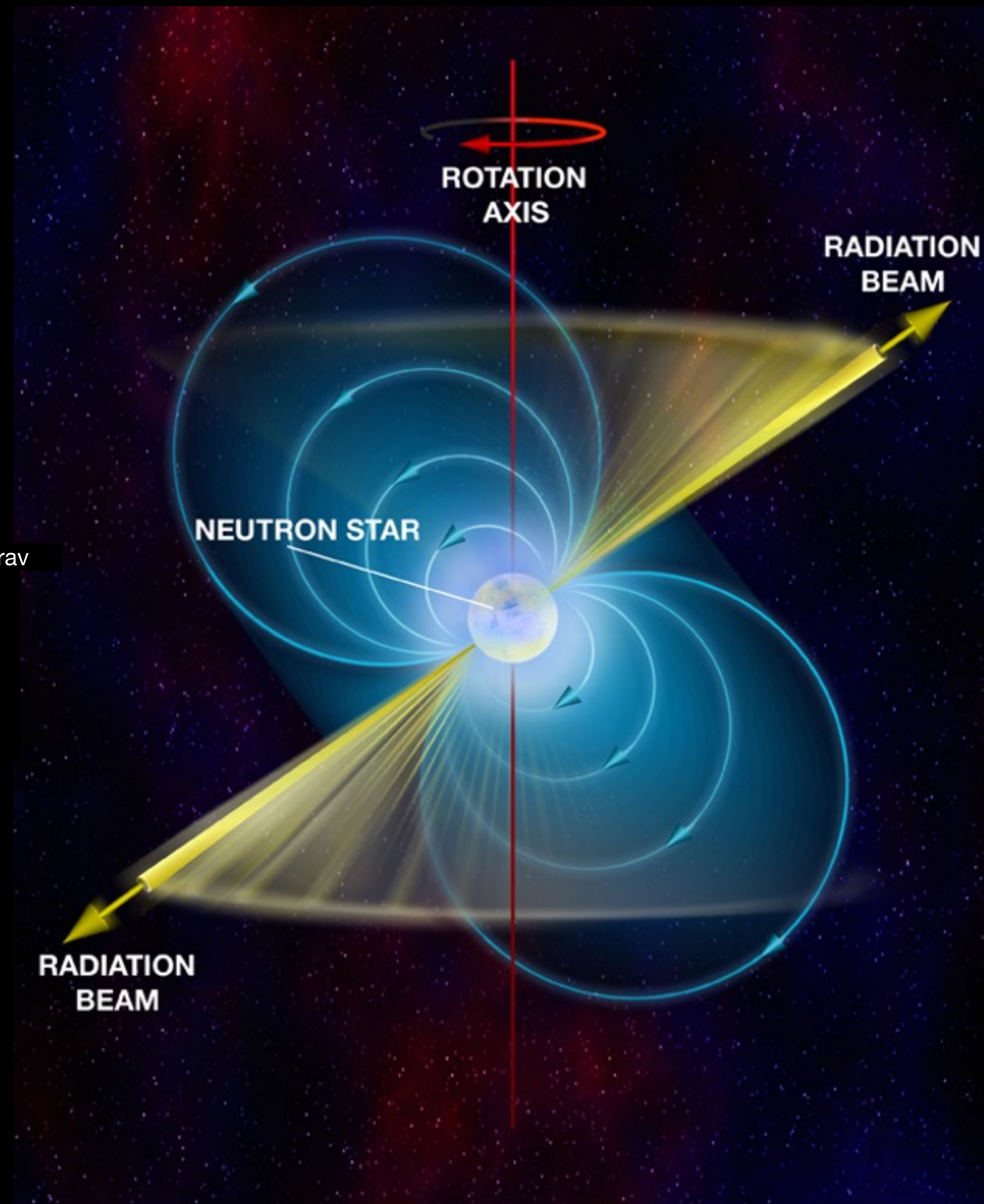
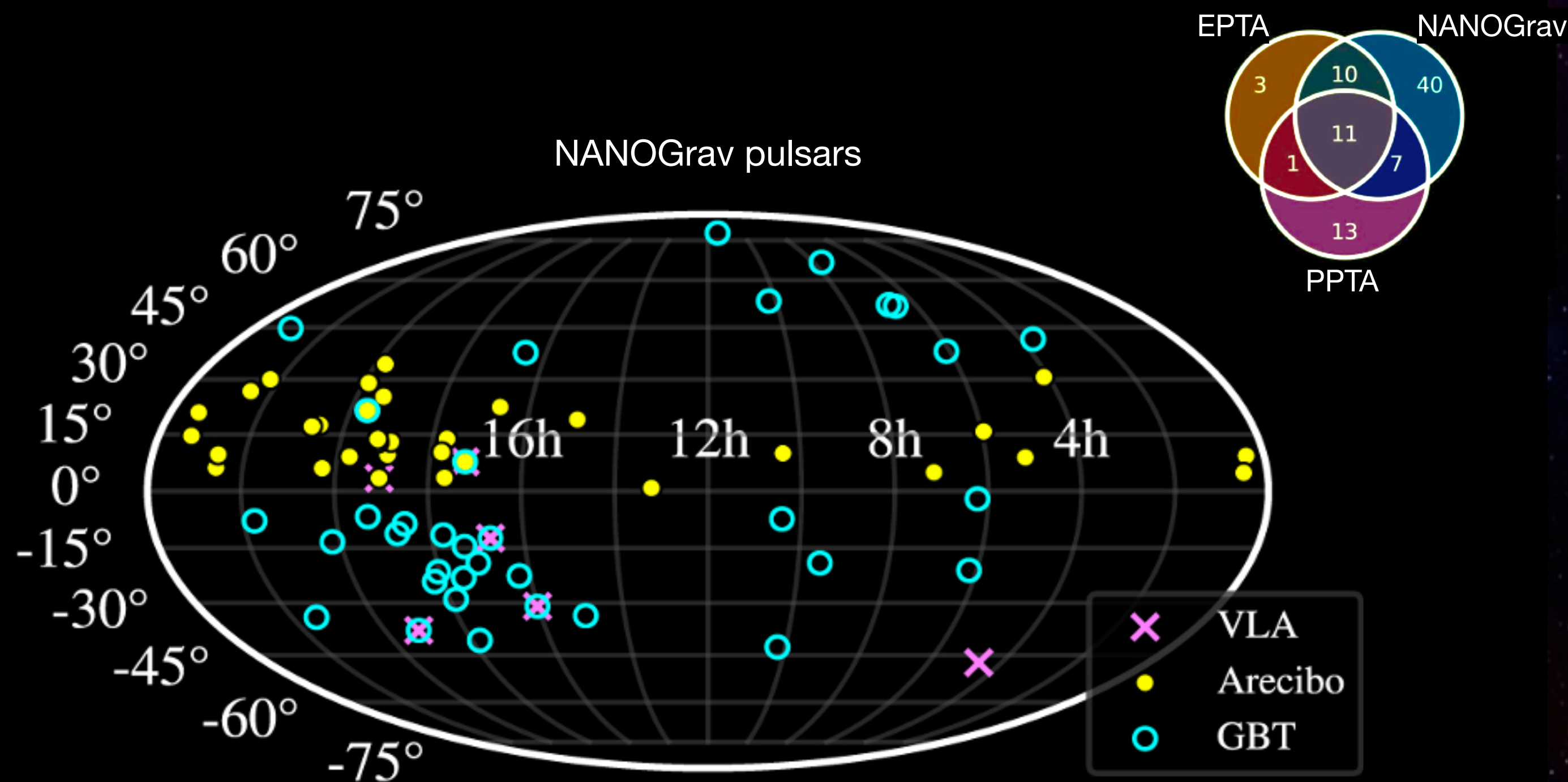


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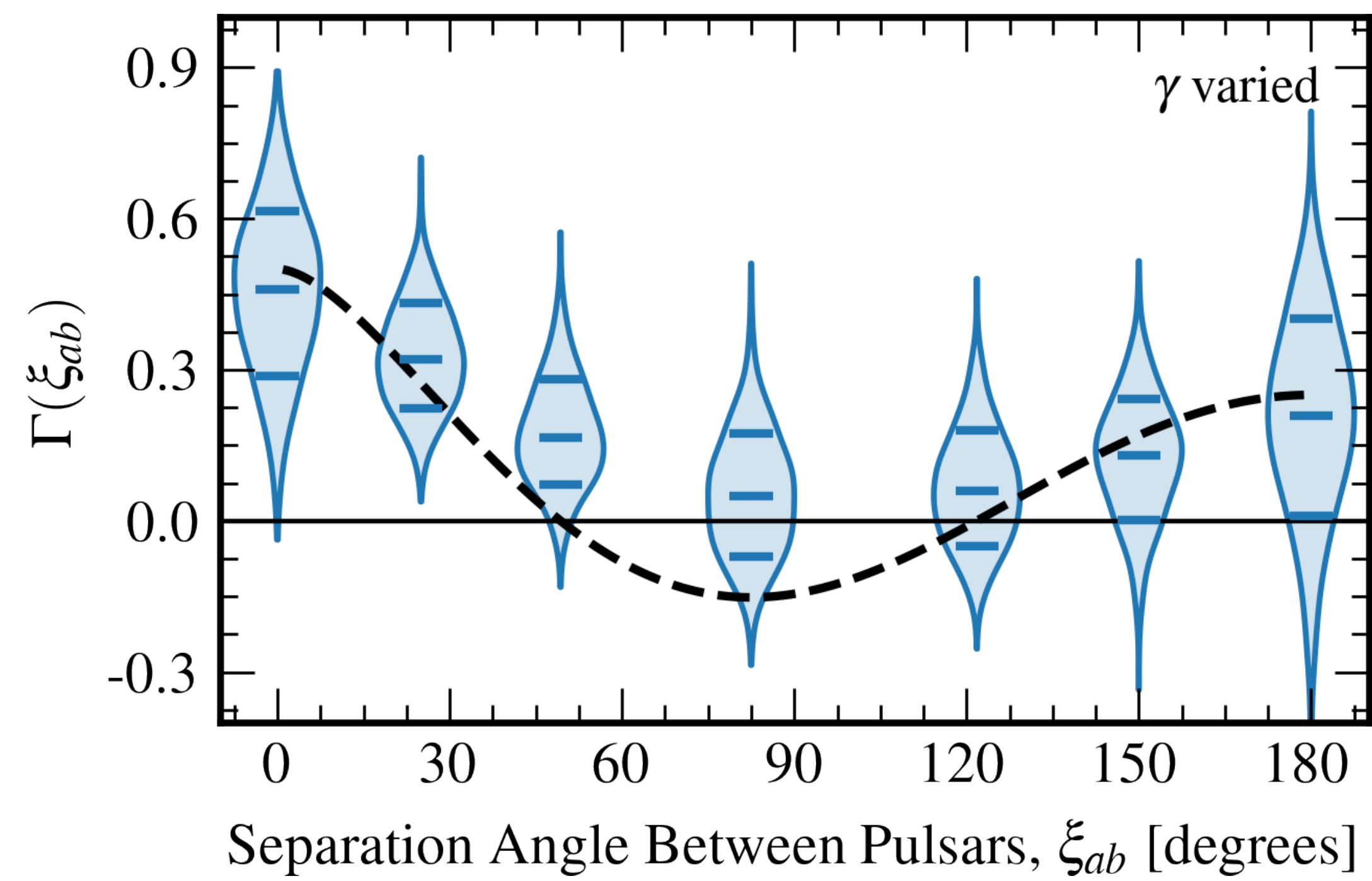
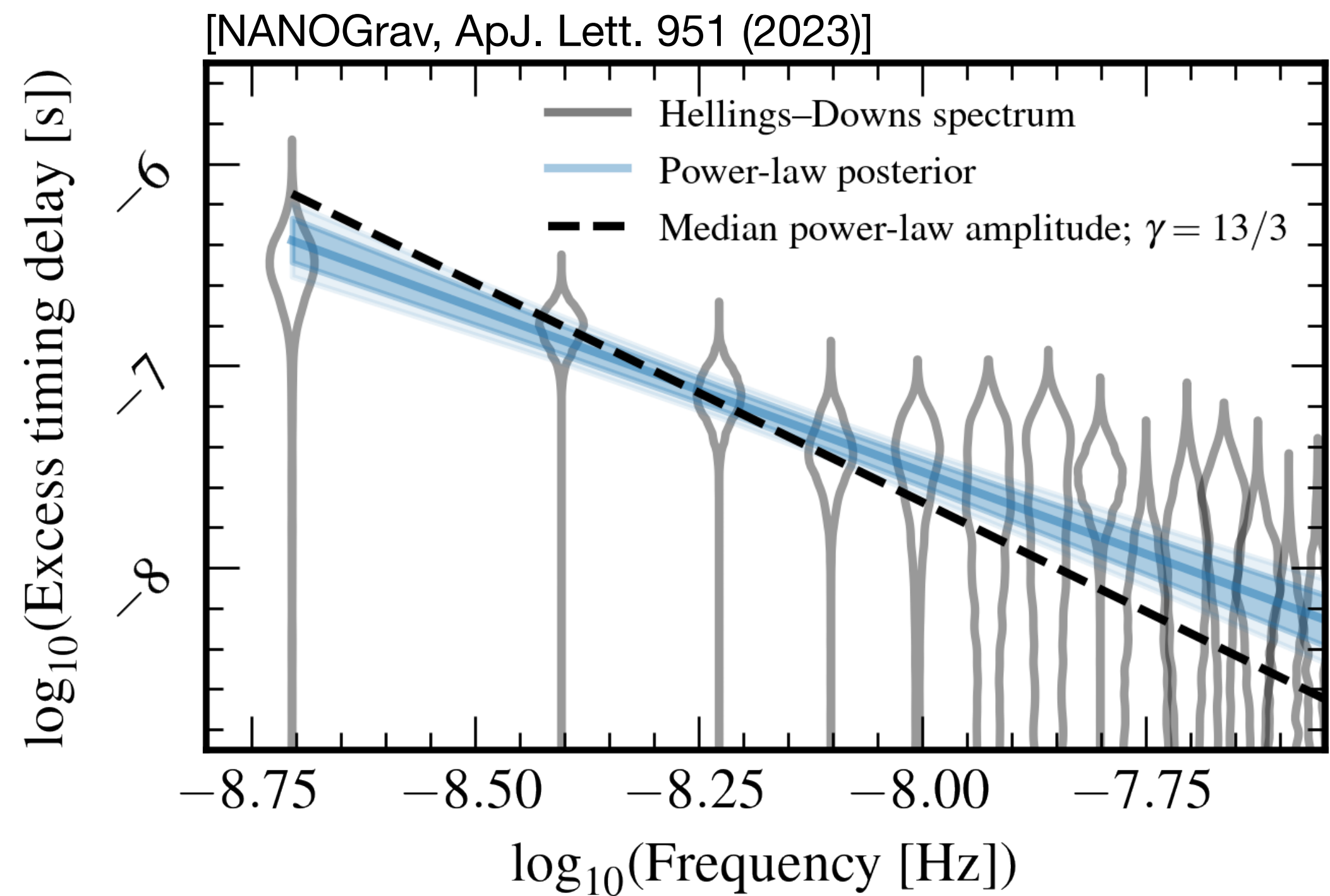
Valencia, 3 November 2025

Pulsar timing arrays

- Measure **timing series** from a set of millisecond pulsars.
- GWs induce noise in that is **correlated** between pulsars.



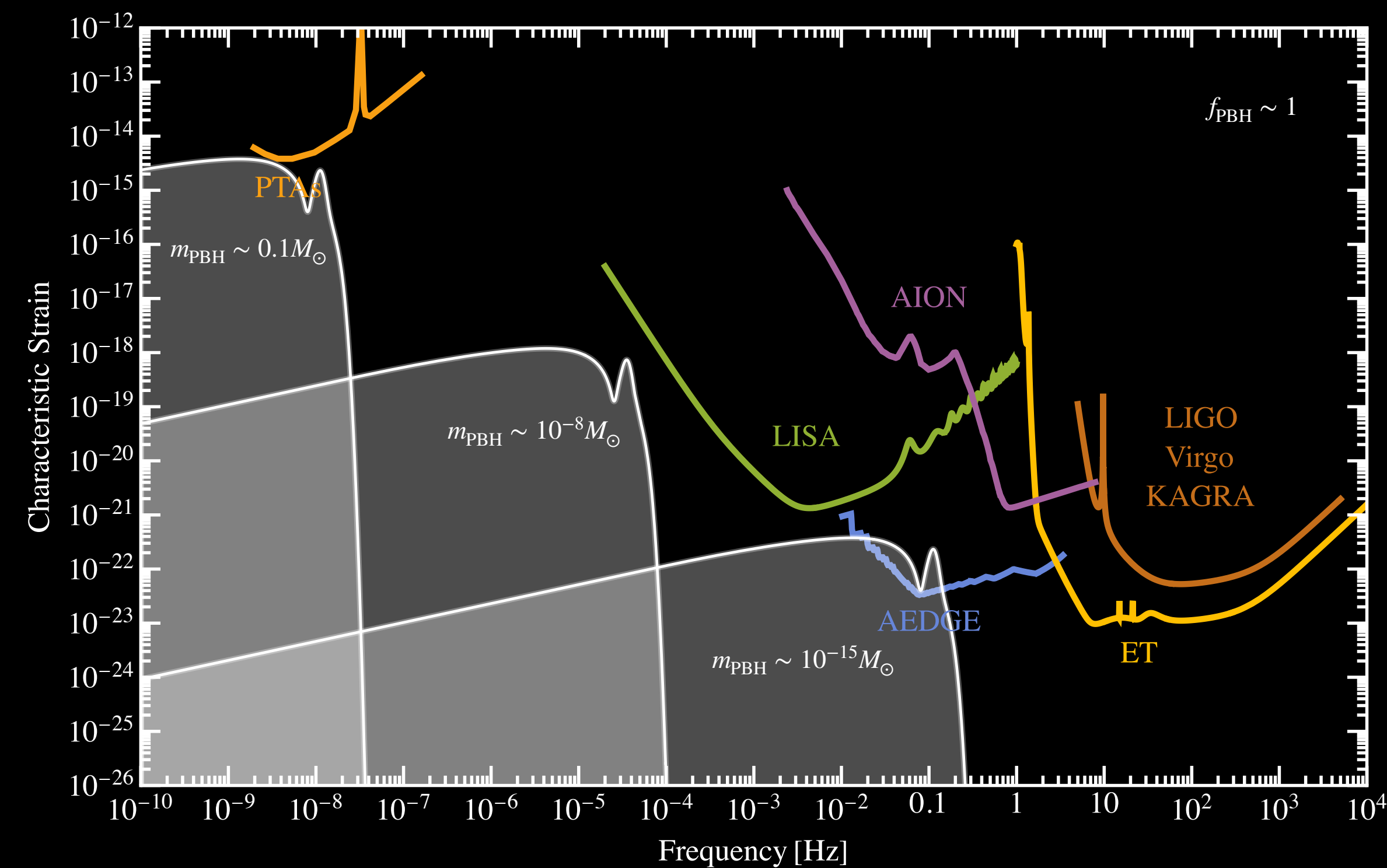
Strong evidence for a GW background at PTAs



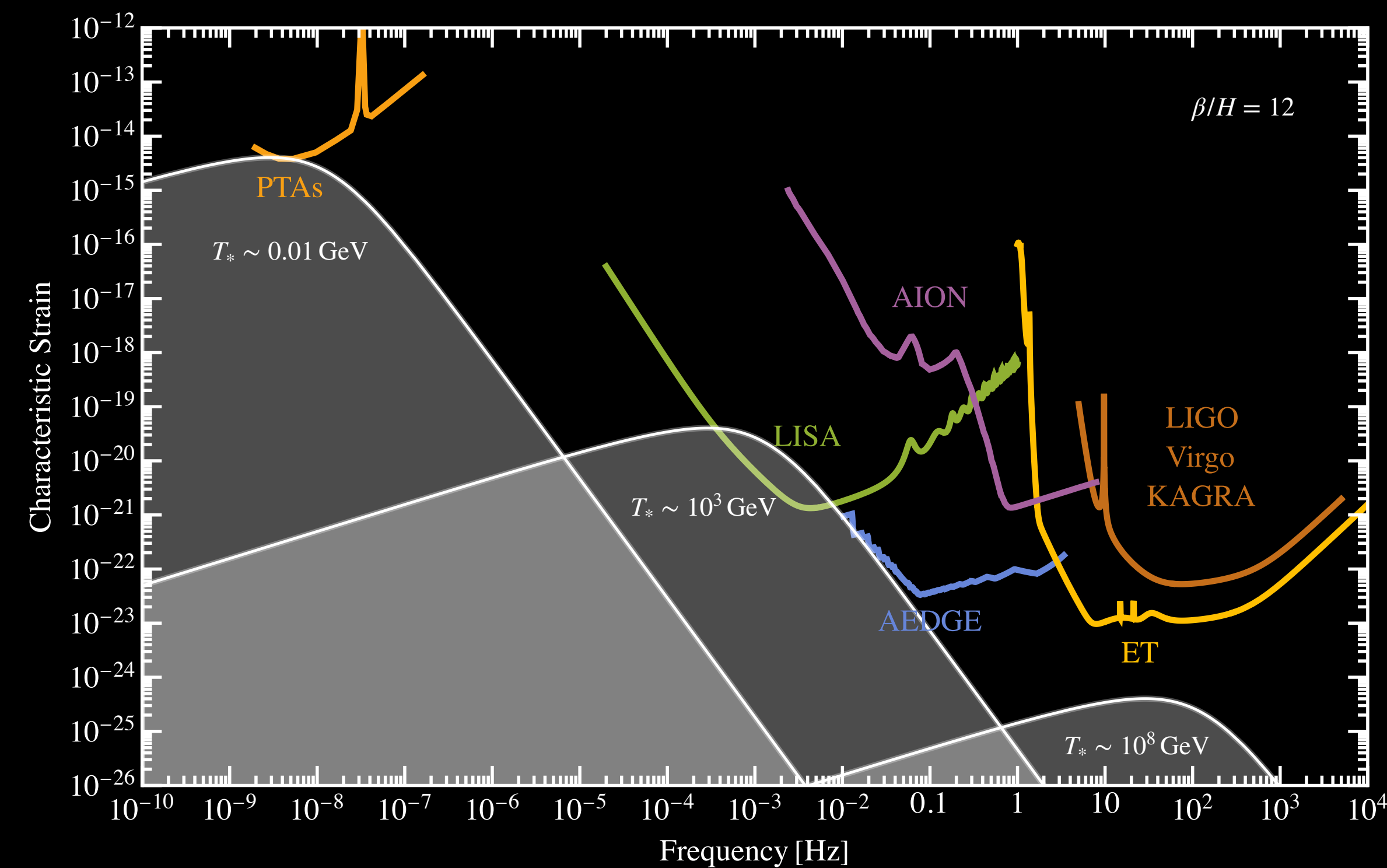
Common spectrum red noise + Hellings-Downs angular correlations

Cosmological sources

Scalar-induced GWs

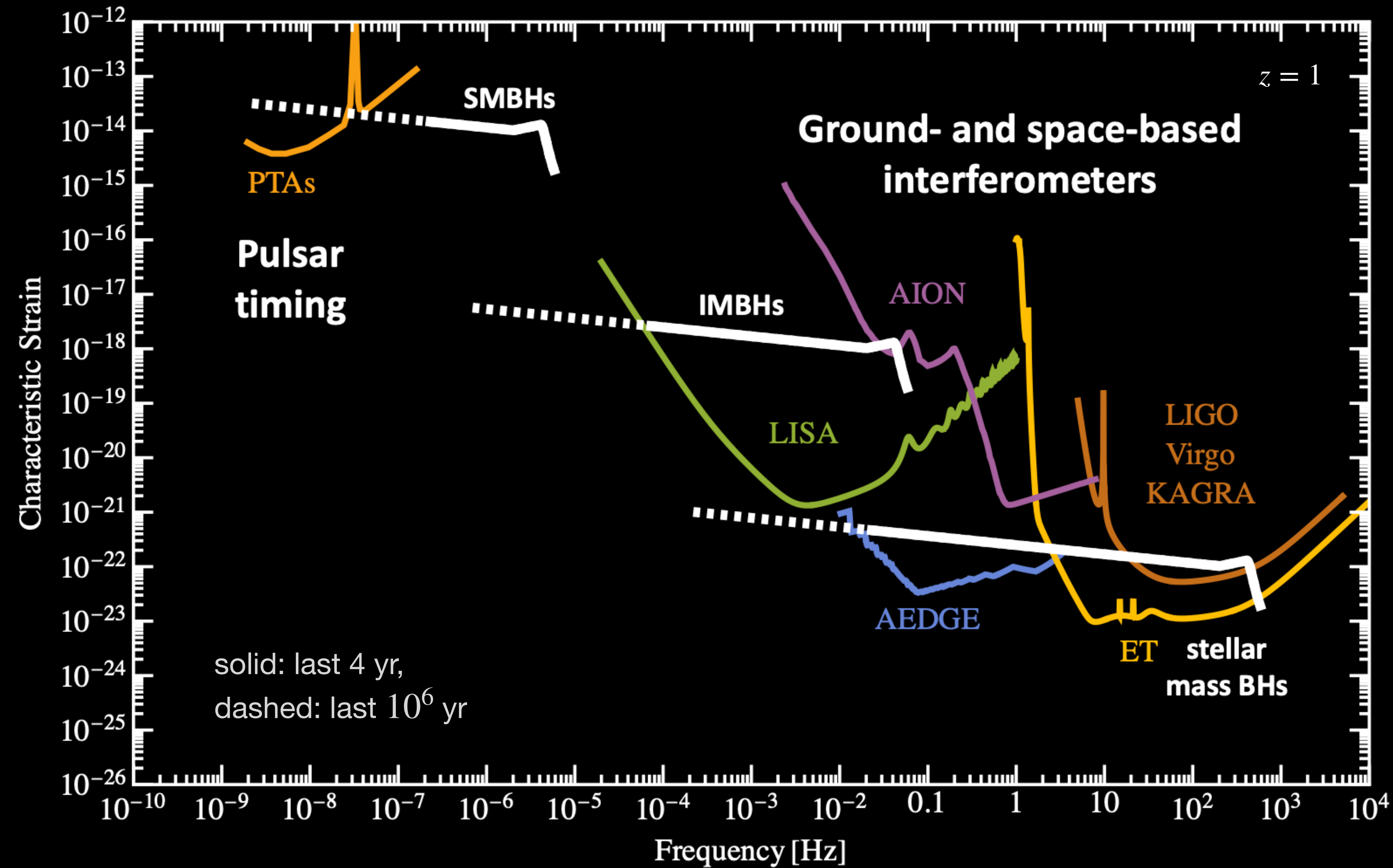


GWs from phase transitions



cosmic strings, domain walls etc.

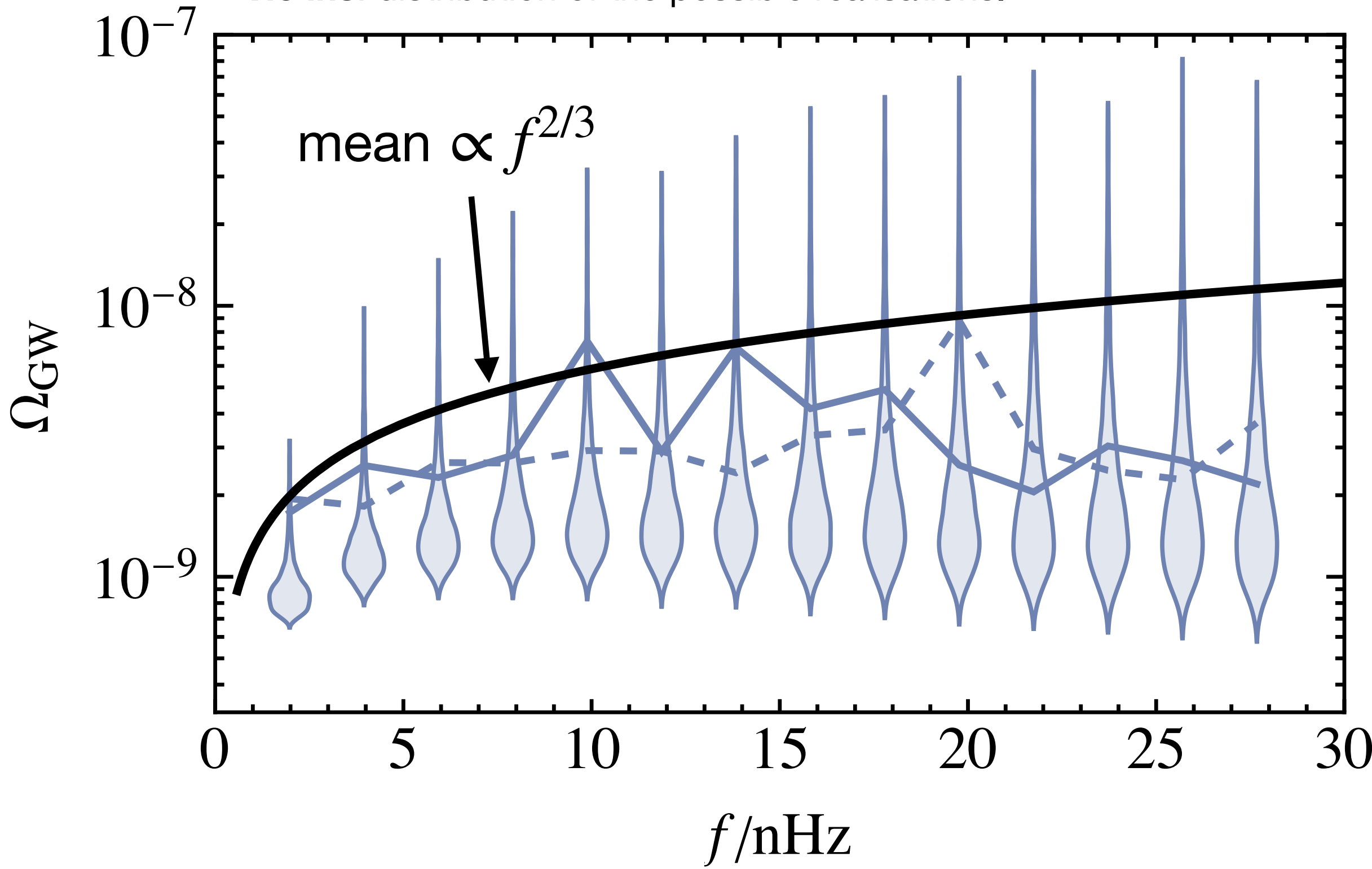
Black hole binaries



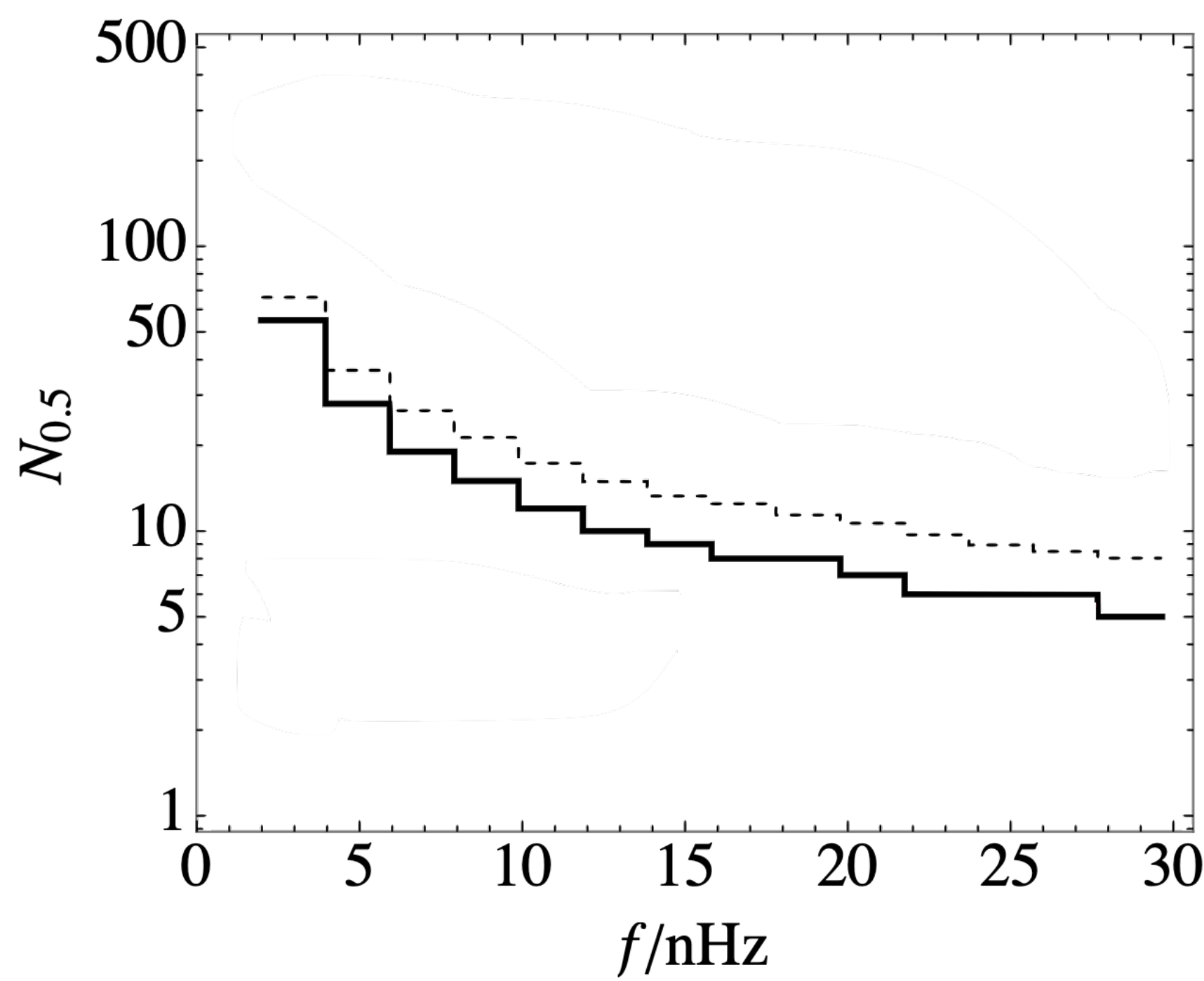
SMBH GW background

$$\Omega_{\text{tot}}(f_i) = \sum_j \Omega_{\text{GW},j}^{(1)} \text{ where } \Omega_{\text{GW}}^{(1)} \propto (\mathcal{M}f)^{10/3} D_L^{-2}$$

- **Solid and dashed curves:** two realizations of the background.
- **Violins:** distribution of the possible realisations.



Sources that make 50% of the signal:



SMBH GW background

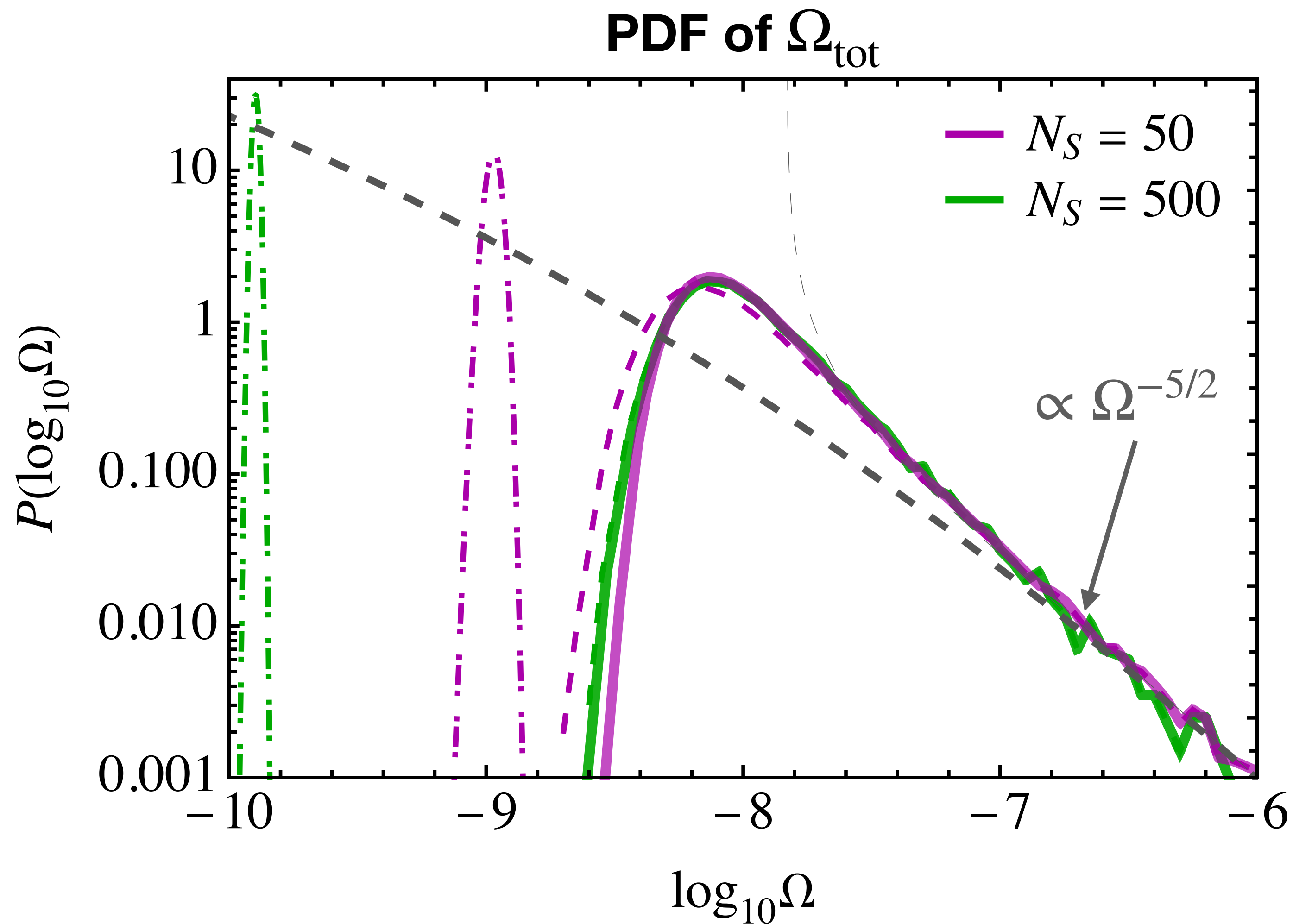
[Ellis et. al., *PRD* 109 (2023); *A&A* 676 (2023)]

$$P^{(1)}(\Omega) \propto \int d\lambda \frac{dt}{d \ln f_r} \delta(\Omega - \Omega^{(1)})$$

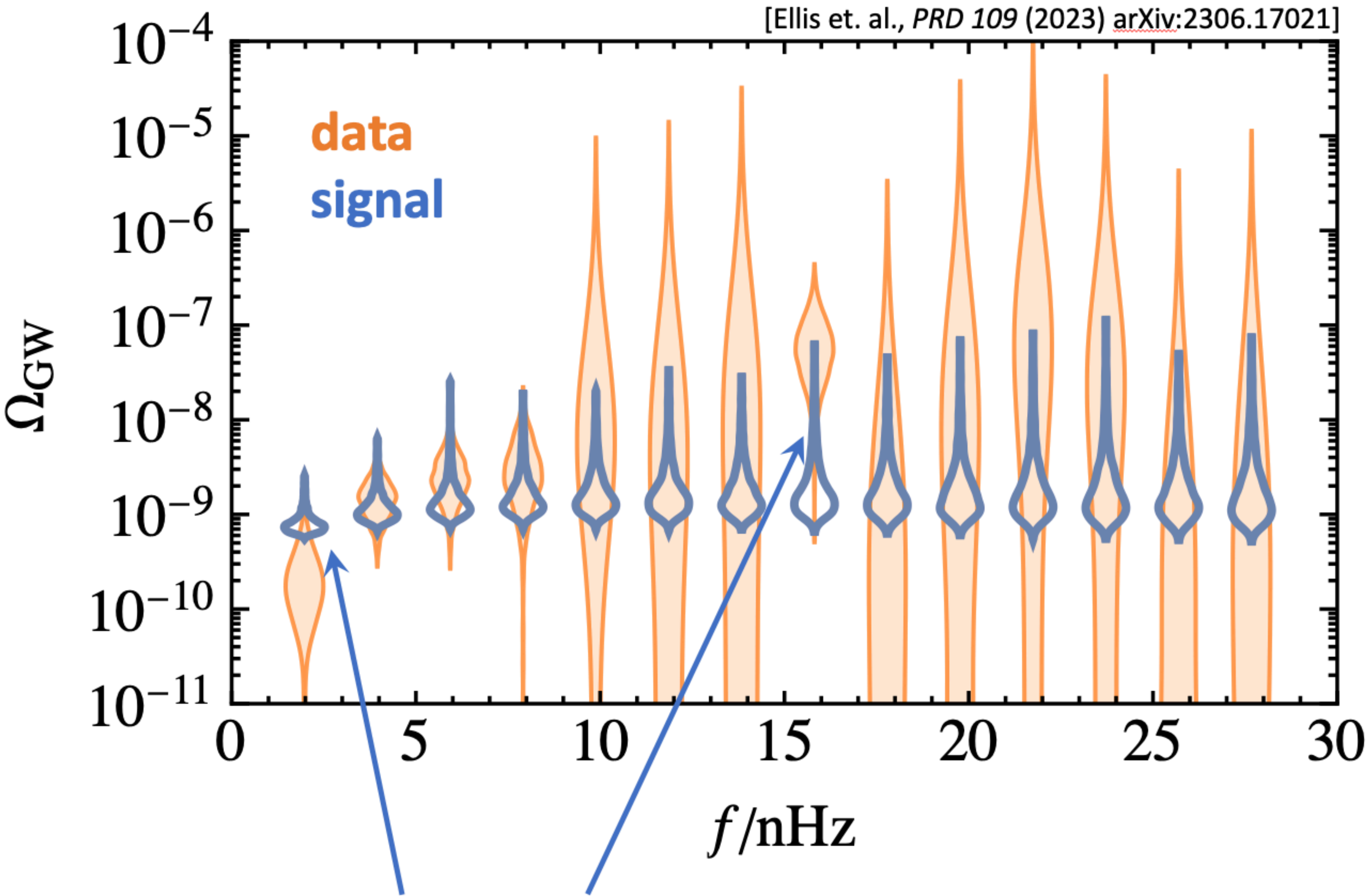
$P^{(1)}(\Omega)$ has a shallow power-law tail,
caused by nearby binaris,

central limit theorem does not hold,

PDF of Ω_{tot} is not Gaussian.



PTA fit: GW driven binaries



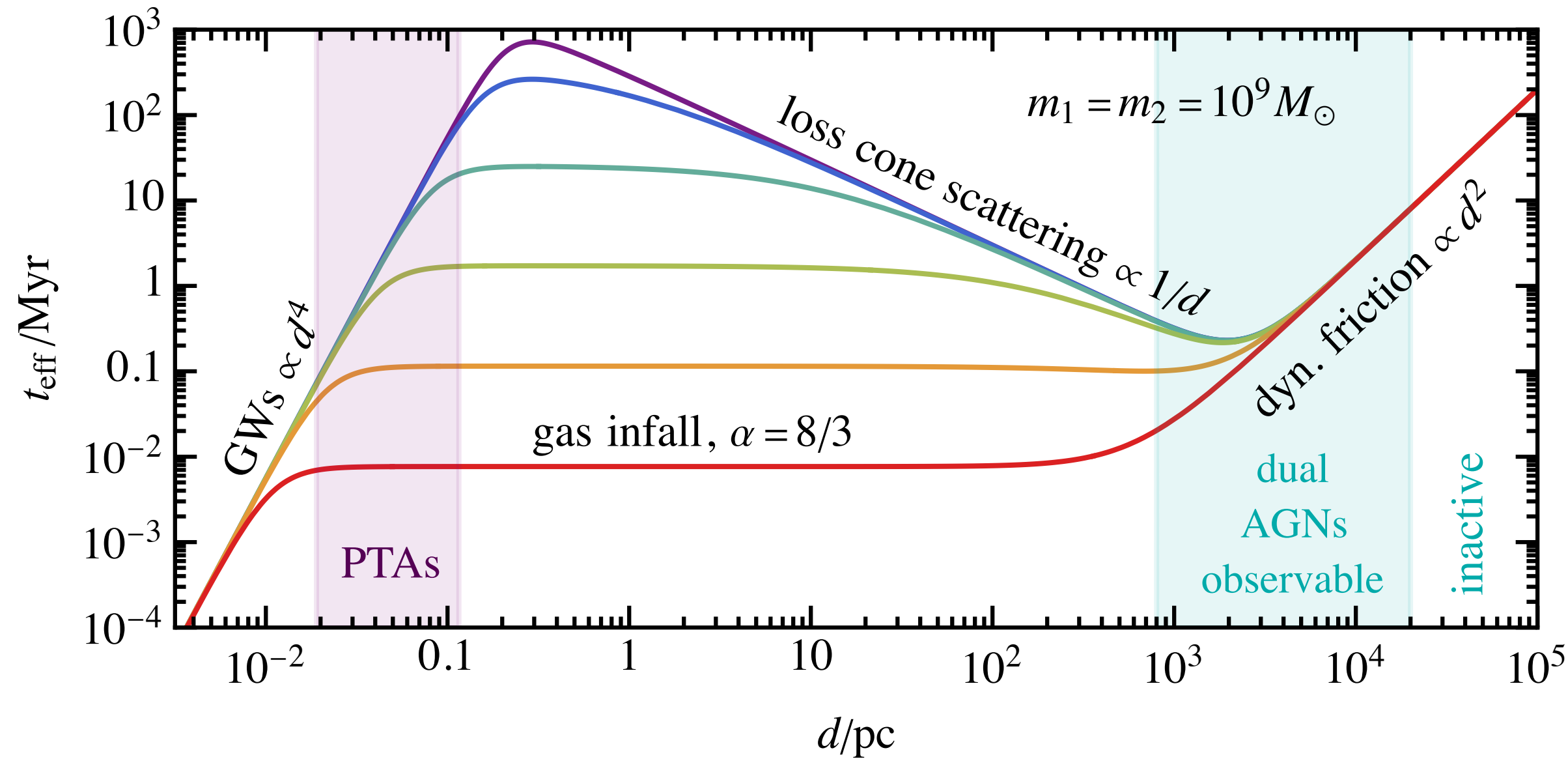
Best fit merger
efficiency:

$$p_{\text{BH}} \approx 0.1$$

GW driven evolution gives
too flat spectrum.

Low-frequency suppression

Environmental effects



- Shorter residence time:

$$\frac{dt}{d \ln f_b} = \frac{dt_{\text{GW}}}{d \ln f_b} \left[1 + \frac{t_{\text{GW}}}{t_{\text{env}}} \right]$$

- GW-driven evolution takes over at small separations.

Eccentricities

$$\frac{dE_n}{dt_r} = \frac{dE_{\text{circ}}}{dt_r} g_n(e), \quad \mathcal{F}(e) = \sum_{n=1}^{\infty} g_n(e) > 1$$

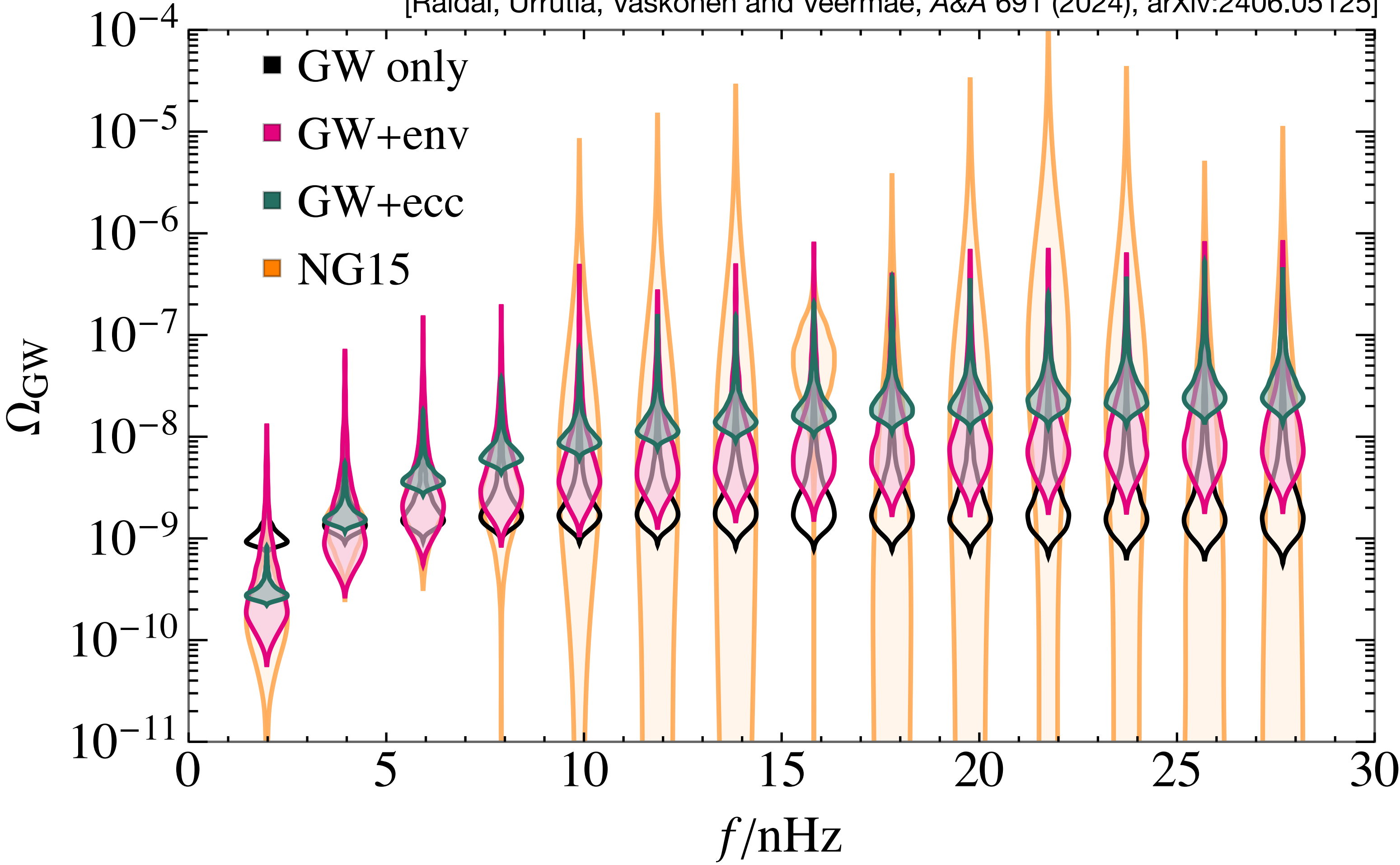
- Shorter residence time:

$$\frac{dt}{d \ln f_b} = \frac{dt_{\text{circ}}}{d \ln f_b} \frac{1}{\mathcal{F}(e)}$$

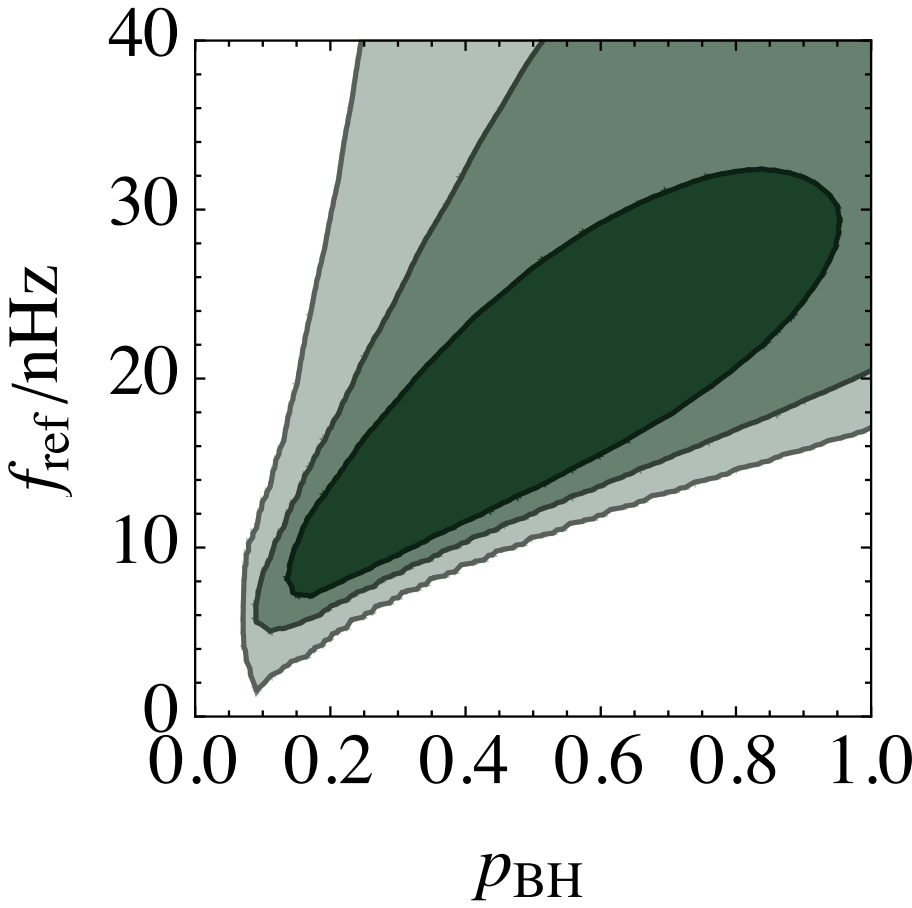
- Signal from single binary spreads over multiple frequencies.
- GW emission circularises binaries.

PTA fit: low-frequency suppression

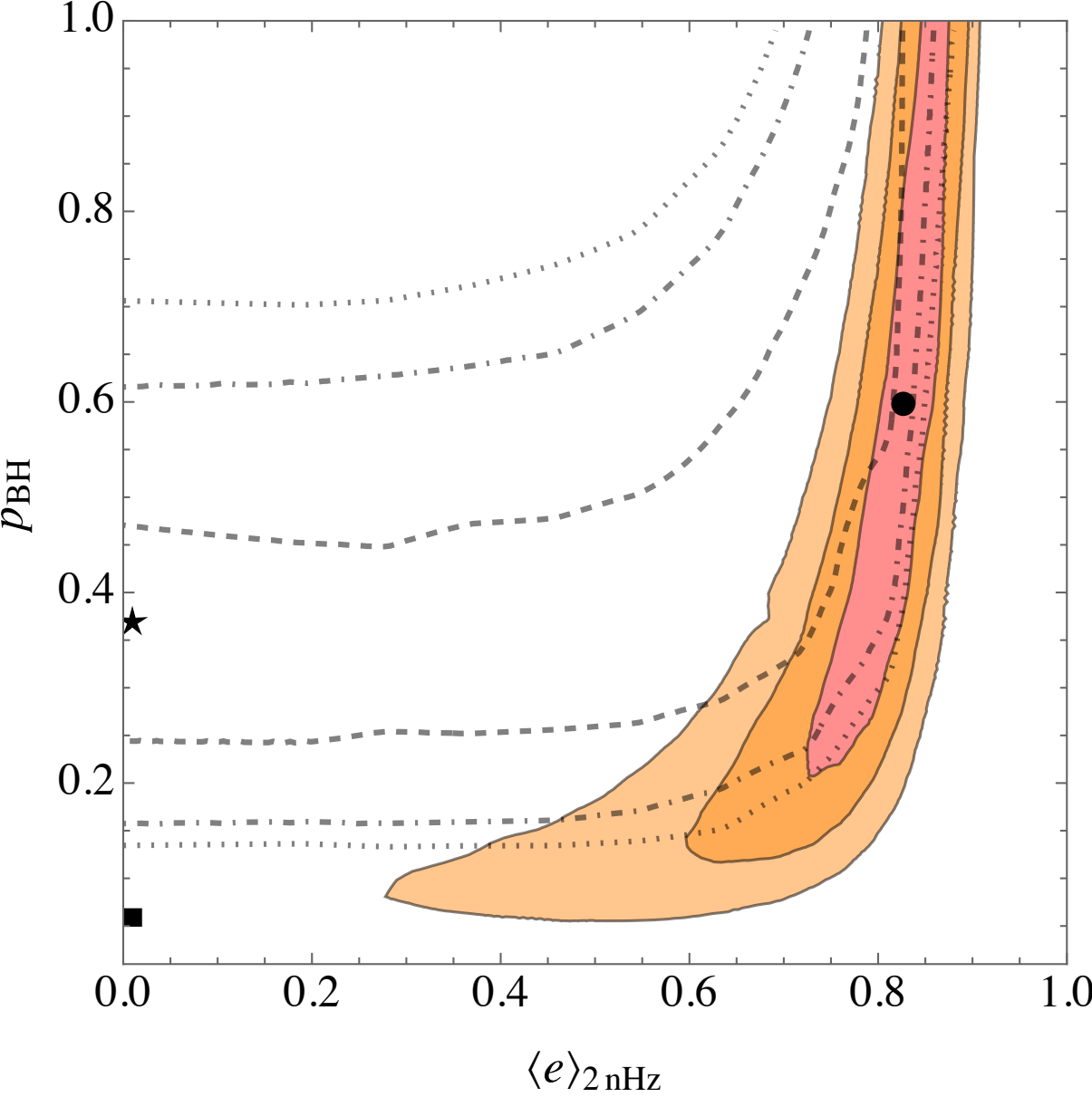
[Raidal, Urrutia, Vaskonen and Veermäe, A&A 691 (2024), arXiv:2406.05125]



env. effects at $f \gtrsim 5 \text{ nHz}$

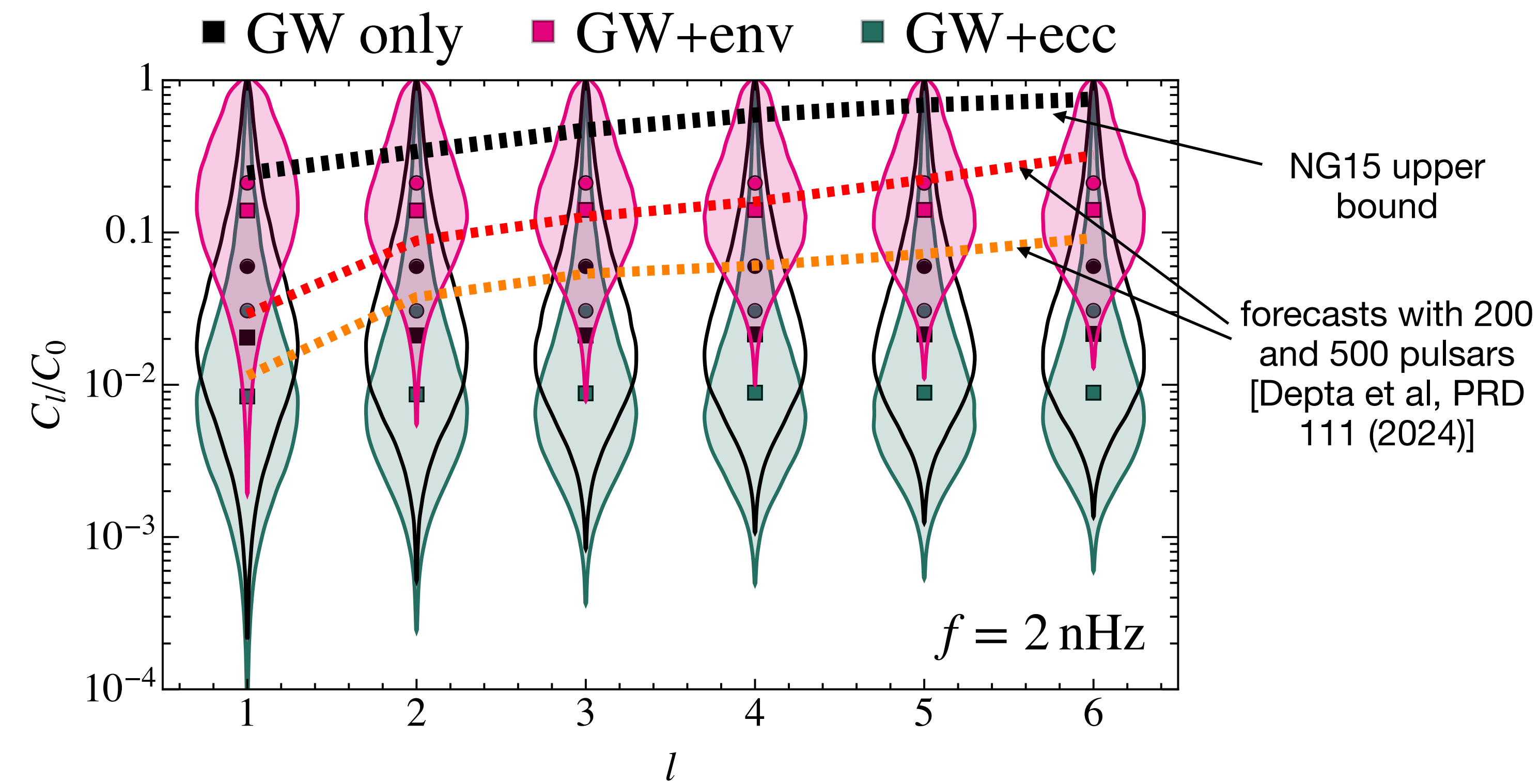


large eccentricities, $e \gtrsim 0.6$

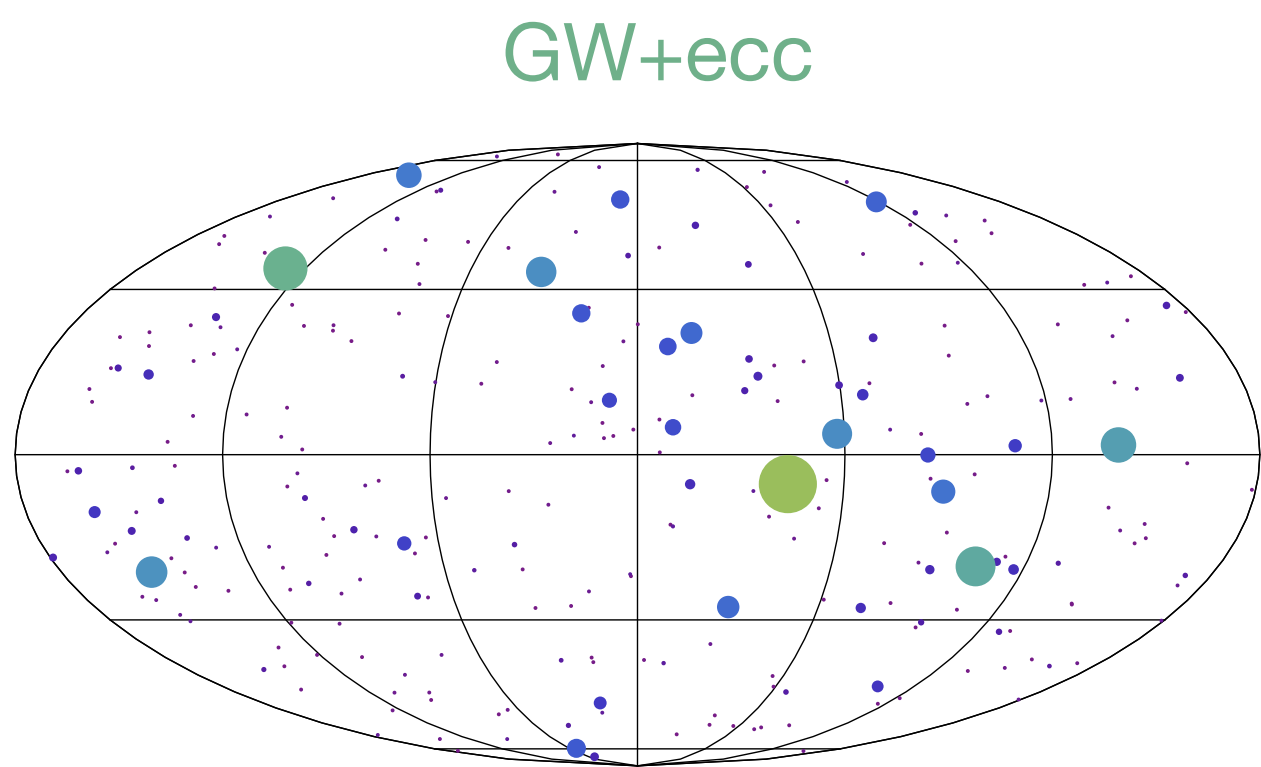
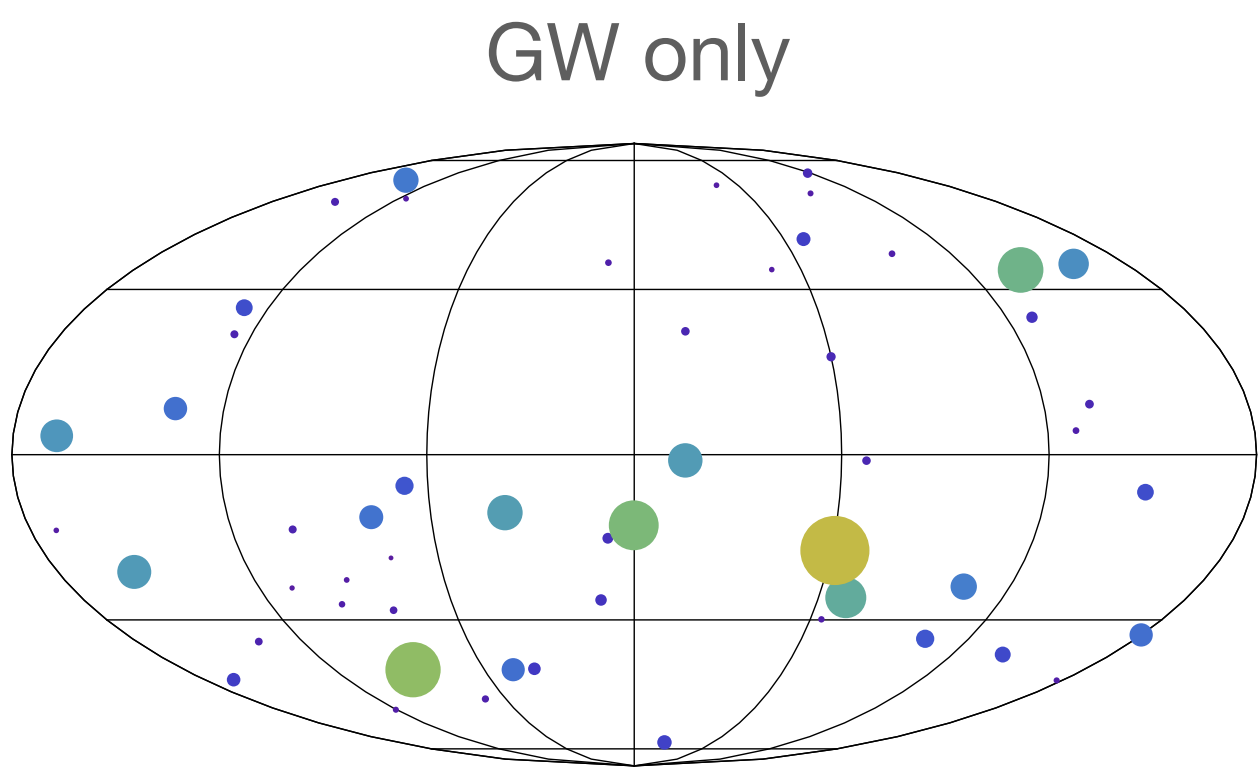
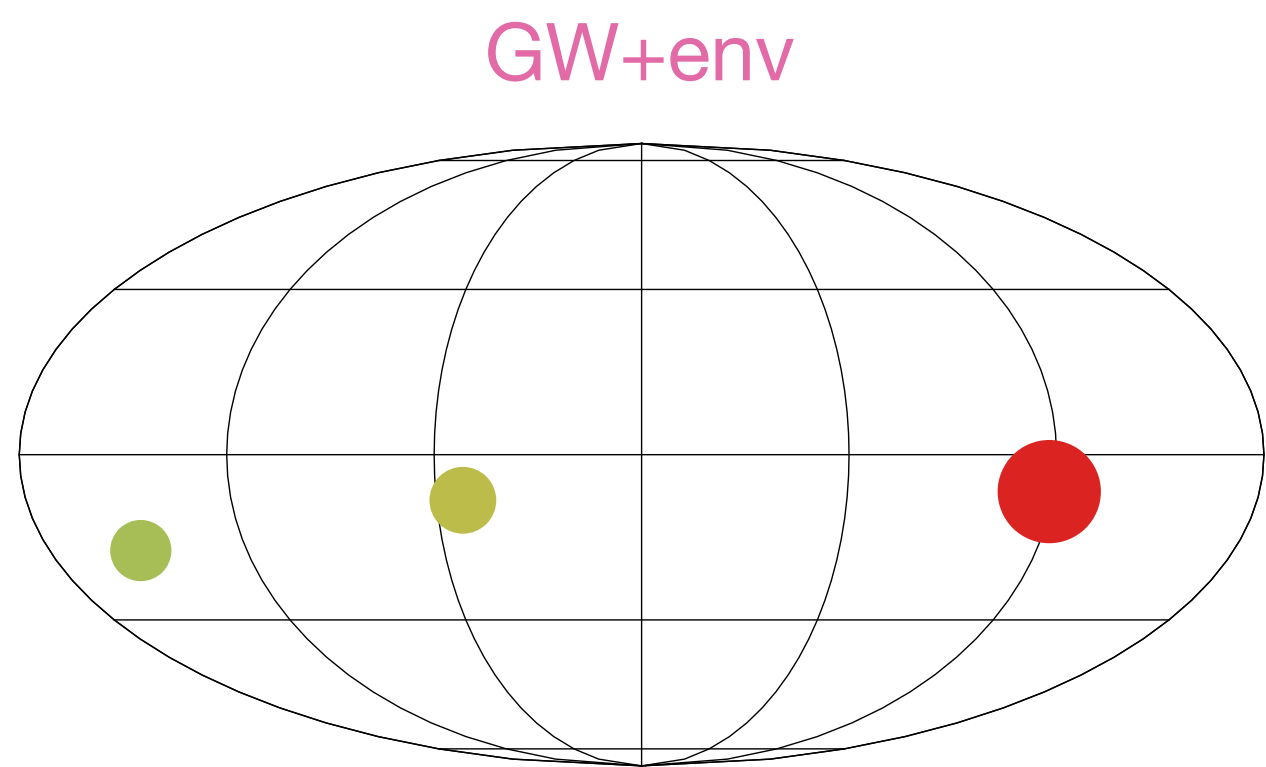
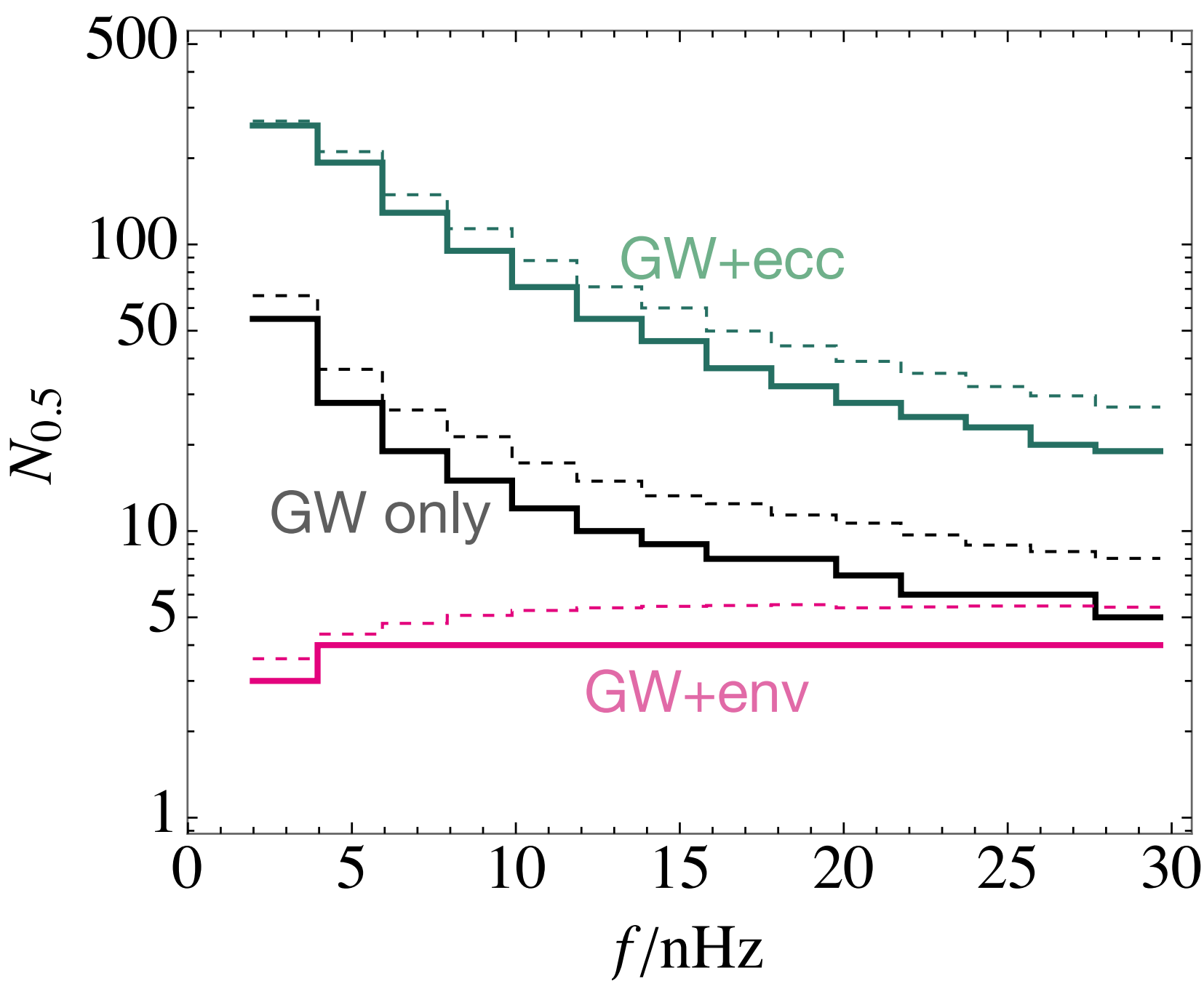


Anisotropies

[Raidal, Urrutia, Vaskonen and Veermäe, arXiv:2411.19692]



Sources that make 50% of the signal:



Summary

1. The PDF of SMBH GWB realizations has a power-law tail, $P(\Omega) \propto \Omega^{-5/2}$, arising from the possibility of having nearby sources.
2. The PTA GW spectrum shape indicates either strong environmental effects or highly eccentric binaries.
3. Spectral fluctuations and anisotropies provide a way to distinguish SMBH origin of the PTA signal from cosmological sources.