The design of a Master Thesis

Andrea Donini (curso organizado y supervisado con Pilar Hernández)

Scientific writing is not taught at school, nor during the undergraduate learning

WE, professional scientists, LEARN how to write by TRIALS AND ERRORS, after many years of scientific papers publication and refereeing

Many of us NEVER LEARN! You can be a very good scientist and a poor writer....

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Volume 50B, number 3

PHYSICS LETTERS

10 June 1974

DETAILED PREDICTIONS FOR THE p-n STRUCTURE FUNCTIONS IN THEORIES WITH COMPUTABLE LARGE MOMENTA BEHAVIOUR

G. PARISI*

Columbia University, New York, N.Y., USA

If strong interactions are described by non Abelian gauge theories, detailed predictions are obtained for the q2 nce at fixed ω of the difference of the structure functions of the proton and neutron.

It has been recently recognized that in a particular class or renormalizable theories the large momenta behaviour is computable from perturbation theory [1], at least for sufficient small coupling constant. Some of the non Abelian gauge theories belong to this class [2-4]. In these theories the Bjorken scaling law for deep inelastic scattering is satisfied [5] apart from logaritmical corrections which can be computed.

We consider a model with 3 × 3 colored quarks: the interaction is carried by an octet of colored gluons and it is symmetric under SU(3) × SU(3)'. In this model the following relation holds [6]:

$$\int_{1}^{\infty} F_{\mathrm{T}}^{2}(\omega, q^{2}) \, \omega^{-N} \, \mathrm{d}\omega \xrightarrow{q^{2} \to \infty} \bigoplus_{i=1}^{4} C_{\mathrm{T}}^{i,N} (\log q^{2})^{-A^{i,N}} \tag{1}$$

where T stands for the target name and $A^{i,N}$ are computed numbers which are independent from the target. Different powers behaviour comes from the contributions of different operators. However in the I=1t-channel only one operator is present; a simple relation is satisfied by the difference between the proton and

$$\int_{1}^{\infty} F_{p-n}^{2} (\omega, q^{2}) \omega^{-N} d\omega \xrightarrow{q^{2} \to \infty} C_{p-n}^{N} (\log q^{2})^{-A^{N}}$$

$$A^{N} = \frac{2}{27} \{1 - 2/N(N+1) + 4 \left[\psi(N+1) + \gamma^{-1}\right] \}$$
(2)

where ψ is the logaritmical derivative of the Euler Γ function and γ is the Euler-Mascheroni constant. A very nice test of the validity of the theoretical scheme would be provided by a comparison of (2) with the

The aim of this note is to point out that it is possible to derive from (2) simple consequences on the q^2 be

$$\begin{split} D(\omega, q^2) &\equiv q^2 \log q^2 \frac{\partial}{\partial q^2} F_{p-n}^2(\omega, q^2) \\ &= \frac{2}{27} \left\{ \left[3 + 4 \log (1 - 1/\omega) \right] F_{p-n}^2(\omega, q^2) + \int_1^{\omega} \frac{\mathrm{d}\mu}{\mu^2} \left[\left(2 - \frac{2}{\mu} + \frac{4}{\mu - 1} \right) F_{p-n}^2 \left(\frac{\omega}{\mu}, q^2 \right) - \frac{4\mu}{\mu - 1} F_{p-n}^2 \left(\omega, q^2 \right) \right] \right\}. \end{split} \tag{3}$$

Eq. (2) and eq. (3) are mathematically equivalent, however we believe that the second equation is easier to test than the first one. A salient feature of eq. (3) is that the integral in its r.h.s. is done from 1 up to ω : informations on the high ω behaviour of the structure functions are not needed.

* On leave of absence from Frascati (INFN)

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Nuclear Physics B126 (1977) 298-318
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ASYMPTOTIC FREEDOM IN PARTON LANGUAGE

G. ALTARELLI

Laboratoire de Physique Théorique de l'Ecole Normale Supérieure **, Paris, France

G. PARISI ***

Institut des Hautes Etudes Scientifiques, Bures-sur-Yvette, France

Received 12 April 1977

A novel derivation of the Q^2 dependence of quark and gluon densities (of given helicity) as predicted by quantum chromodynamics is presented. The main body of predictions of the theory for deep-inleastic scattering on either unpolarized targets is re-obtained by a method which only makes use of the simplest tree diagrams and is entirely phrased in parton language with no reference to the conventional operator formalism.

1. Introduction

The quark parton model [1] provides us with a very useful and simple description of the physics of deep inelastic phenomena [2]. The theoretical framework which justifies the parton model is given by the asymptotically free gauge theory of strong interactions based on the color degrees of freedom [3] (quantum chromodynamics, QCD). Although scaling is predicted to be broken by logarithms (a fact which appears to be well consistent with present experiments), the deviations from scaling can be and have been computed for deep inelastic structure functions for either unpolarized [4,5] or polarized targets [6,7]. In the leading logarithmic approximation, the results can again be phrased in the parton language by assigning a well determined Q^2 dependence to the parton densities. In spite of the relative simplicity of the final results, their derivation, although theoretically rigorous, is somewhat abstract and formal, being formulated in the language of renormalization group equations for the coefficient functions of the local operators which appear in the light cone expansion for the product of two currents.

- * On leave of absence from the Istituto di Fisica dell'Università di Roma
- ** Laboratoire propre du CNRS associé à l'Ecole Normale Supérieure et à l'Université de Paris-Sud. Postal address: 24, rue L'homond, 75231 Paris Cedex 05, France.
- *** On leave of absence from the Laboratori Nazionali di Frascati.

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* On leave

** Laborate

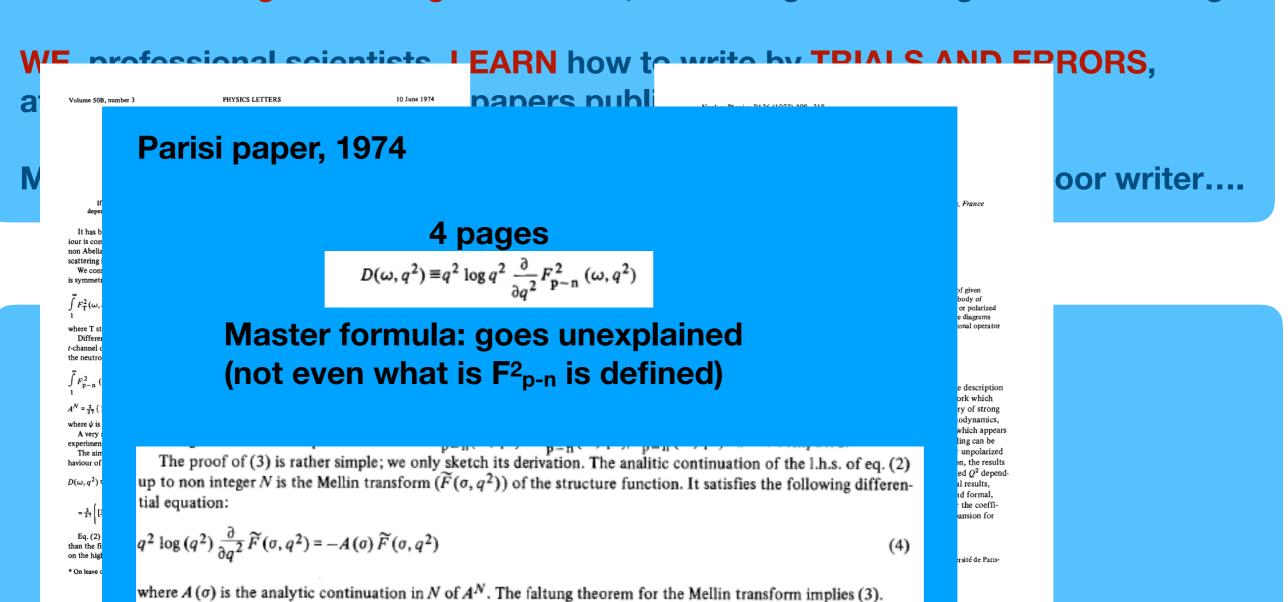
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8576 citations

Altarelli-Parisi equations

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Altarelli-Parisi paper, 1977: 20 pages

$$\frac{\mathrm{d}q^{\mathrm{NS}}(x,\,t)}{\mathrm{d}t} = \frac{\alpha(t)}{2\pi} \int_{x}^{1} \frac{\mathrm{d}y}{y} \, q^{\mathrm{NS}}(y,\,t) \, P\left(\frac{x}{y}\right),\,$$

or writer....

The paper is half devoted to explain what the master formula is, half to computations

The meaning of this equation is clear. Given a quark with momentum y there is a chance that is radiates a gluon, thus reducing its energy from y to x. The quantity

$$\mathcal{P}_{qq} + d\mathcal{P}_{qq} = \delta(z - 1) + \frac{\alpha}{2\pi} P(z) dt$$
 (21)

is the probability density of finding, inside a quark, another quark with fraction z of the parent momentum. The change with t of this probability produces the variation of the quark distribution function. Thus $P(z) \alpha/2\pi$ is the variation per unit t at order α of the probability density of finding inside a quark another quark with fraction z of the parent momentum.

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Many of us NEVER LEARN! You can be a very good scientist and a poor writer....

However,

we may learn from more experienced professional scientists

and

we may learn from more experienced professional writers

A Pulitzer Prize's suggestion

The 2006 Pulitzer Prize's winner Cormac McCarthy wrote a basic set of rules (suggestions...) on how to write a scientific article (in english)

Nature **574**, 441-442 (2019)

Some of them:

On the choice of the subject:

Decide on your paper's theme and two or three points you want every reader to remember. This theme and these points form the single thread that runs through your piece.

On simplicity:

If something isn't needed to help the reader to understand the main theme omit it.

Keep sentences short, simply constructed and direct.

Concise, clear sentences work well for scientific explanations.

On the style:

Inject questions and less-formal language to break up tone and maintain a friendly feeling. [...] Similarly, use a personal tone because it can help to engage a reader. If you must talk about arbitrary colours of an abstract sphere, it's more gripping to speak of this sphere as a red balloon or a blue billiard ball.

Maybe the most important:

Just enjoy writing.

Not everybody enjoys writing!

Accelerated learning

Our main objective is that you may learn the basic rules to write your Master's Thesis by this summer...

Our final goal is to have a uniform (minimum) level in the Theses, so as to simplify reading both for referees and normal readers (this may be useful for us to read your theses, and for you to sell your thesis outside the academy)

Starting from a "standard recipe", you may then evolve into your own scientific style

General view

- 1. Two classes on basic rules to write the Master's thesis
- 2. Practical exercise: organising the Thesis (email submission)
- 3. Correction of the practical exercise (either by mail or skype?)
- 4. Presentation to the class of one or two exercises
- 5. An introduction to the usage of AI for scientific writing?
- 6. One class on basic rules to present your Master's thesis

1. STARTING

- 1. Choosing a language
- 2. Choosing an editing language
- 3. Choosing an editor

2. ORGANIZATION OF THE THESIS

- 1. Understanding your subject
- 2. What is "old"
- 3. Appendices
- 4. What is "new"
- 5. Introduction and conclusions

3. GRAPHICS AND EQUATIONS

- 1. Drawing and inserting plots
- 2. Explaining plots (captions!)
- 3. Equations
- 4. REFERENCING

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- 1. Drawing and inserting plots
- 2. Explaining plots (captions!)
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- 4. REFERENCING

What this class it is not

We are not teaching you to write

We are not teaching you the language you use to write:
scientists write in a non-uniform language
that approximately resembles English;
however, I feel that "times they are a-changin' "...

We are not teaching you physics

1. Choose a language to write

Possible choices:

ENGLISH

CASTELLANO

VALENCIÁ

1. Choose a language to write

Possible choices:

ENGLISH

CASTELLANO

VALENCIÁ

I strongly suggest ENGLISH:

- 1. It is the (current) language of science
- 2. You read books and articles in english
- 3. It may be your first (or last) chance to write in english
- 4. Your thesis will be easier to show outside (to do a PhD abroad, to go to a private company, ...)

2. Choose a editing language

Possible choices:

NONE

2. Choose a editing language

There is (still) only one choice possible:

Possible choices:

NONE

LATEX

MAC: TEXSHOP https://pages.uoregon.edu/koch/texshop/

WINDOWS: MIKTEX https://miktex.org/download

LINUX: TEXLive (depends on which Linux)

https://www.tug.org/texlive/

3. Choose an editor

This is up to you (however, most Latex releases have their own editor)

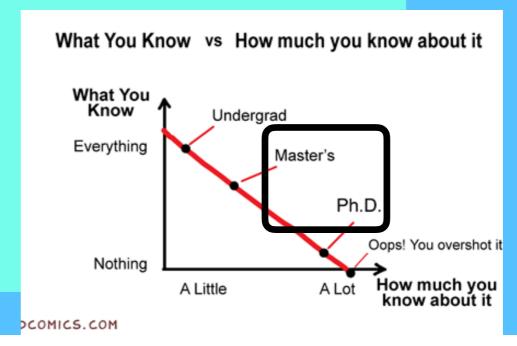
1. Understanding your subject

Usually, you and your advisor are the two persons that understand most of your thesis

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Sometimes, only your advisor understand (most of) your thesis



1. Understanding your subject

Usually, you and your advisor are the two persons that understand most of your thesis

The rest of the physicists that read your thesis divide into:

those that know the subject and those that do not know the subject

1. Understanding your subject

Usually, you and your advisor are the two persons that understand (most of) your thesis

The rest of the physicists that read your thesis divide into:

those that know the subject and those that do not know the subject

The jury that will examine you will be composed by both

WRITE FOR THE LATTER GROUP!

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and those that do not know the subject

The jury that will examine you will be composed by both

1. Understanding your subject

You should imagine that those that read

DO NOT KNOW DEEPLY THE SUBJECT OF YOUR THESIS

(give an overview of the subject)

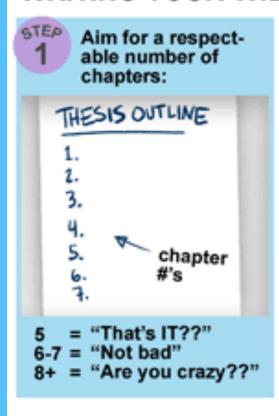
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THEY KNOW AND UNDERSTAND THE BASICS (do not start with undergraduate stuff)

1. Understanding your subject

WRITING YOUR THESIS OUTLINE

NOTHING SAYS "I'M ALMOST DONE" TO YOUR ADVISOR/ SPOUSE/PARENTS LIKE PRETENDING YOU HAVE A PLAN

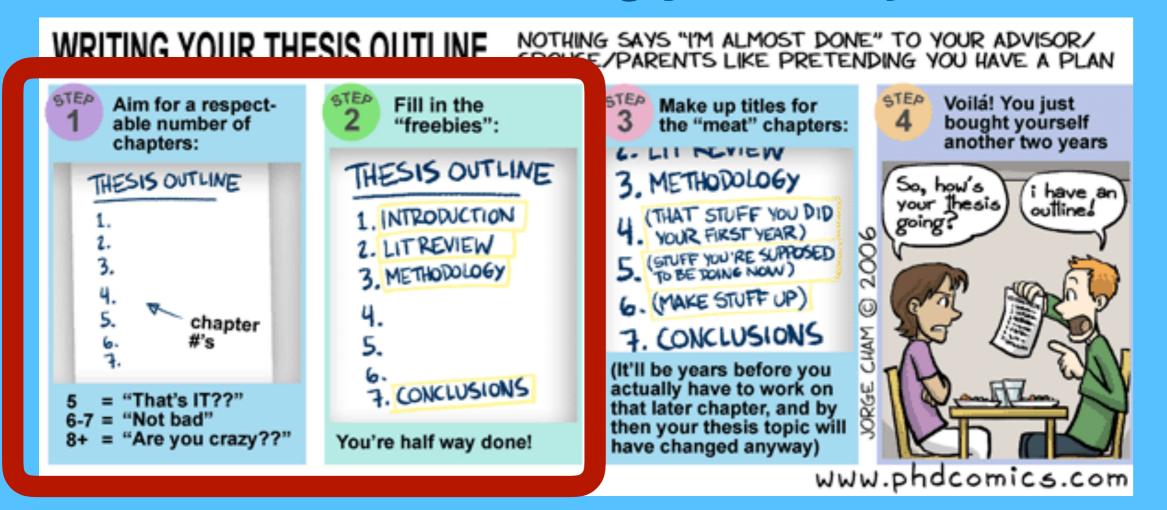








1. Understanding your subject



An example from old theses

Thesis on SUSY

1	M Problems
	.1 Observable Problems
	1.1.1 Neutrino Masses
	1.1.2 Dark Matter
	1.1.3 Baryon Asymmetry (Baryogenesis)
	1.1.4 Inflation Mechanism
	.2 Hierarchy Problem
	.3 Thoretical and Phenomenological Problems
2 '	The SUSY Way
	.1 LSP - Dark Matter
	2.1.1 R-parity
	.2 Stability in the Hierarchy Problem
	.3 Grand Unification
	.4 Untification of Couplings
	.5 Locality and Supergravity
3	ntroducing SUSY
	.1 SUSY Algebra
	3.1.1 Massless Supermultiplet
	3.1.2 Massive Supermultiplet
	.2 SUSY Lagrangians
	3.2.1 Non-interacting Wess-Zumino model
	3.2.2 Interacting chiral supermultiplet
	3.2.3 Gauge supermultiplet lagrangians
	3.2.4 Gauge and Chiral interactions
	.3 Necessity of Soft SUSY Breaking
4	MSSM
	.1 Electroweak symmetry breaking in the MSSM
	.2 MSSM mass spectrum
	4.2.1 Gluinos
	4.2.2 Neutralinos

An example from old theses

Introduction (motivation)

Thesis on SUSY

Cont	ents	
1 SM 1.1	Problems 1 Observable Problems 3 1.1.1 Neutrino Masses 3 1.1.2 Dark Matter 4 1.1.3 Baryon A symmetry (Baryogenesis) 5 1.1.4 Inflation Mechanism 5 Hierarchy Problem 5 Thoretical and Phenomenological Problems 8	
2.1 2.2 2.3 2.4 2.5	SUSY Way 9	
3.2	3.1.1 Massless Supermultiplet 18 3.1.2 Massive Supermultiplet 19 SUSY Lagrangians 20 3.2.1 Non-interacting Wess-Zumino model 20 3.2.2 Interacting chiral supermultiplet 22 3.2.3 Gauge supermultiplet lagrangians 24 3.2.4 Gauge and Chiral interactions 26 Necessity of Soft SUSY Breaking 27	
4 MS 4.1 4.2	SM 28 Electroweak symmetry breaking in the MSSM 29 MSSM mass spectrum 30 4.2.1 Gluinos 30 4.2.2 Neutralinos 31 4.2.3 Charginos 31 4.2.4 Chiral supermultiplet mixing 32	

An example from old theses

Introduction/ Motivation

Thesis on SUSY

" Lil' " review

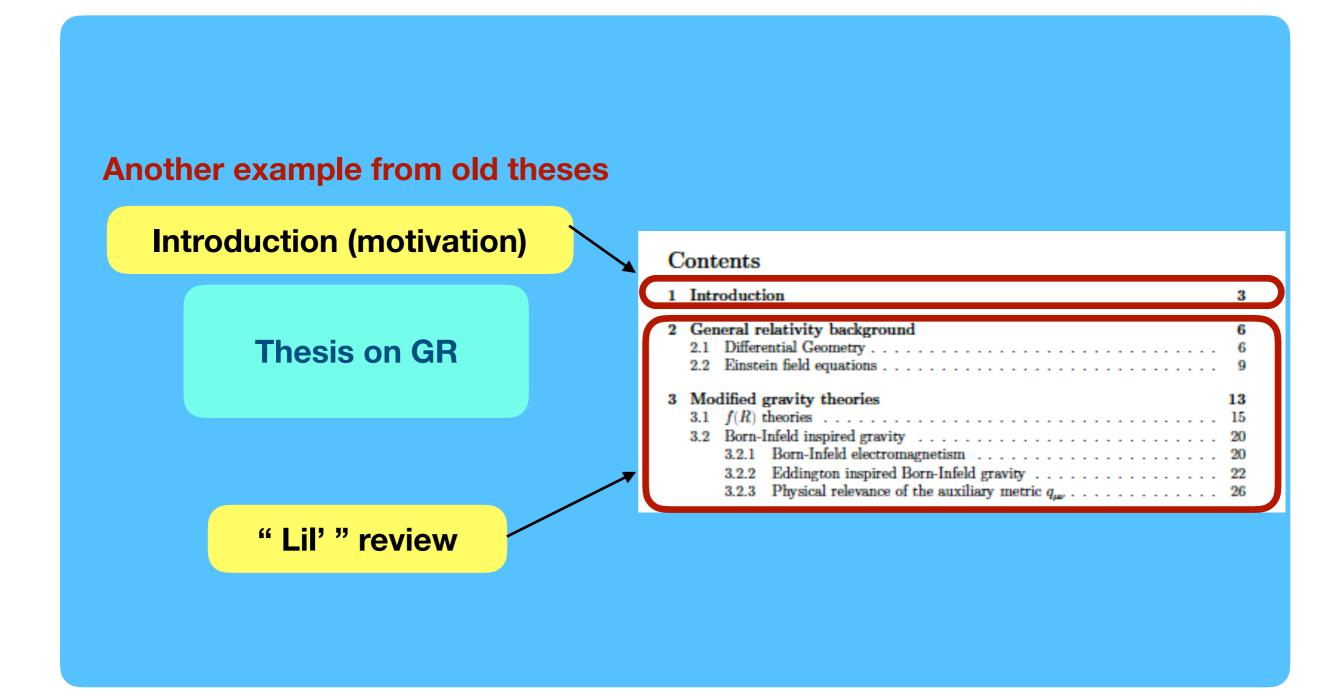
Contents
1 SM Problems 1 1.1 Observable Problems 3 1.1.1 Neutrino Masses 3 1.1.2 Dark Matter 4 1.1.3 Baryon Asymmetry (Baryogenesis) 5 1.1.4 Inflation Mechanism 5 1.2 Hierarchy Problem 5 1.3 Thoretical and Phenomenological Problems 8
2 The SUSY Way 9 2.1 LSP - Dark Matter 11 2.1.1 R-parity 12 2.2 Stability in the Hierarchy Problem 13 2.3 Grand Unification 15 2.4 Unification of Couplings 15 2.5 Locality and Supergravity 16
3 Introducing SUSY 17 3.1 SUSY Algebra 17 3.1.1 Massless Supermultiplet 18 3.1.2 Massive Supermultiplet 19 3.2 SUSY Lagrangians 20 3.2.1 Non-interacting Wess-Zumino model 20 3.2.2 Interacting chiral supermultiplet 22 3.2.3 Gauge supermultiplet lagrangians 24 3.2.4 Gauge and Chiral interactions 26 3.3 Necessity of Soft SUSY Breaking 27
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Another example from old theses

Thesis on GR

Contents

1	Introduction	3
2	General relativity background	6
	2.1 Differential Geometry	6
	2.2 Einstein field equations	
3	Modified gravity theories	13
	3.1 f(R) theories	15
	3.2 Born-Infeld inspired gravity	20
	3.2.1 Born-Infeld electromagnetism	20
	3.2.2 Eddington inspired Born-Infeld gravity	22
	3.2.3 Physical relevance of the auxiliary metric q	



Introduction/Motivation: 3-6 pages

"Old" Literature: 10 to 20 pages (from textbooks and articles)

START FROM HERE! START NOW!

Your stuff: 10 to 20 pages

Part II: ORGANIZATION

Introduction/Motivation: 3-6 pages

"Old" Literature: 10 to 20 pages (from textbooks and articles)

START FROM HERE! START NOW!

Your stuff: 10 to 20 pages

SECOND PART TO WRITE... (third if you have appendices)

Part II: ORGANIZATION

Introduction/Motivation: 3-6 pages

MOST IMPORTANT PART OF THE THESIS!
LAST TO WRITE!

"Old" Literature: 10 to 20 pages (from textbooks and articles)

START FROM HERE! START NOW!

Your stuff: 10 to 20 pages

SECOND PART TO WRITE... (third if you have appendices)

You should determine which amount of basic information is needed to introduce the subject

Who is your reader: a physicist that understand what you are doing, but do not know the basic literature on the subject (except for textbooks)

Example: Thesis on some topic in General Relativity

Do not explain what is General Relativity from scratch

If you need some non-standard formalism,

introduce it

2	General relativity background 2.1 Differential Geometry	
3	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	20 20 22

Example: Thesis on some topic in General Relativity

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STANDARD: 3 pages

2	2.1	eneral relativity background Differential Geometry											
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		3.2.2 Eddington inspired Born-Infeld gravity	22										
		3.2.3 Physical relevance of the auxiliary metric $q_{\mu\nu}$											

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3.2 Einstein field equations

3 Modified gravity theories
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3.2 Born-Infeld inspired gravity
3.2 Eddington inspired Born-Infeld gravity

This part of the thesis is the place where it is most probable to take a short-cut:

Cut&Pasting from holy books and articles

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A reasonable excuse: how may I write something about THIS better than THEM?

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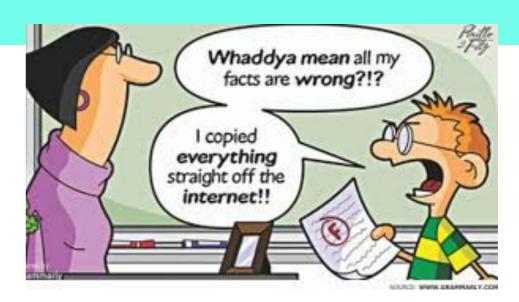
This is NOT ALLOWED!

Better to read the holy sources, try to understand and TRY TO WRITE BY YOURSELF

(see "Use of Al" class, though)

https://top-papers.com/how-to-avoid-plagiarism-in-research-papers/

- 1. Paraphrasing: your own words, same concept
- 2. Citing: (small) text from other paper [citenumber]
- 3. Quoting: "(complete) text from other paper" [citenumber]



Second: Appendices

They are the place where you must move TECHNICAL STUFF that you used but is:

- A) boring to read in the main text
- B) something that may be skipped

Thesis on Quantum Systems

Appendices										
A	Weak coupling ME calculations									
	A.1 $\text{Tr}_E[V_tH_I(t), \hat{\rho}(0)] = 0$ for an arbitrary state of the environment									
	A.2 Computation of $\text{Tr}_E[V_t^0 H_I(t), [V_\tau^0 H_I(\tau), \hat{\rho}(0)]]$									
	A.3 Switching to the Schrödinger picture									
	A.5 Switching to the Schrodinger picture									
В	Commutators of weak coupling ME for different OQS operators L									
	B.1 $L = \sigma (L^{\dagger} = \sigma_+)$									
	D.1 $L = 0 - \{L = 0 + \}$									
	R2 $L = L^{\dagger} = \sigma_{-}$									
	B.2 $L = L^{\dagger} = \sigma_z$ B.3 $[H_S, \rho_S(t)]$									

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As many as you need....

No more than a couple of pages each

Appendices													
A Weak coupling ME calculations A.1 $\operatorname{Tr}_E[V_tH_I(t),\hat{\rho}(0)]=0$ for an arbitrary state of the environment. A.2 Computation of $\operatorname{Tr}_E[V_t^0H_I(t),[V_\tau^0H_I(\tau),\hat{\rho}(0)]]$ A.3 Switching to the Schrödinger picture													
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Thesis on Quantum Systems

Write these SECOND! TOMORROW!

ΑĮ	Weak coupling ME calculations A.1 $\text{Tr}_E [V_t H_I(t), \hat{\rho}(0)] = 0$ for an arbitrary state of the environment																					
	A 1	Tre	V_*H_*G	$\hat{a}(0)$	11 =	0 fc	or ar	n ar	bitra /pH ctur	ary I _I (τ	sta), <i>į</i>	ate 6(0)	of t	he	envi	ron	me	nt.	:	 	- 1	3 2 2 2
)\	W	!		$=\sigma$. (: :	-		: :	: :						:		-	. 4

It must be very clear WHAT is NEW, and WHAT is NOT NEW

A typical question at the defense:

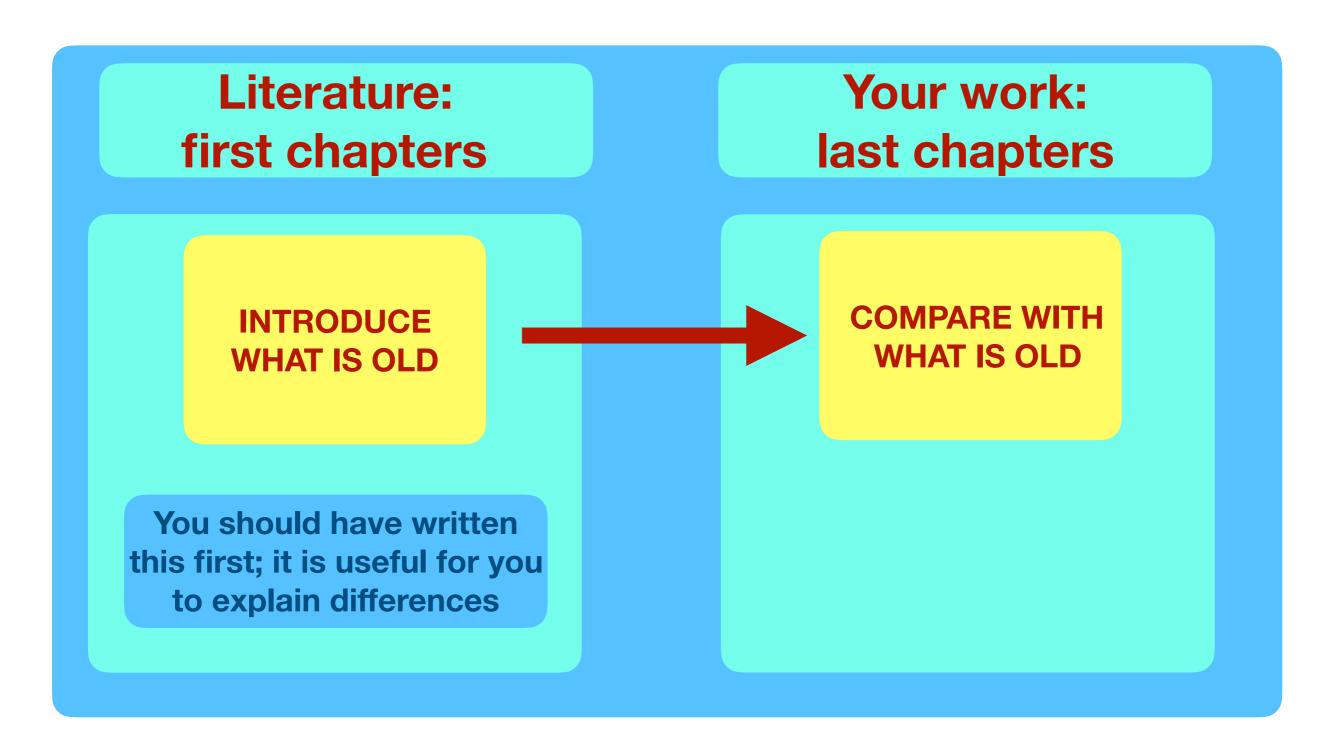
"Ok, very nice. But....
Can you explain me PRECISELY which is the difference between what YOU have done and the LITERATURE?"

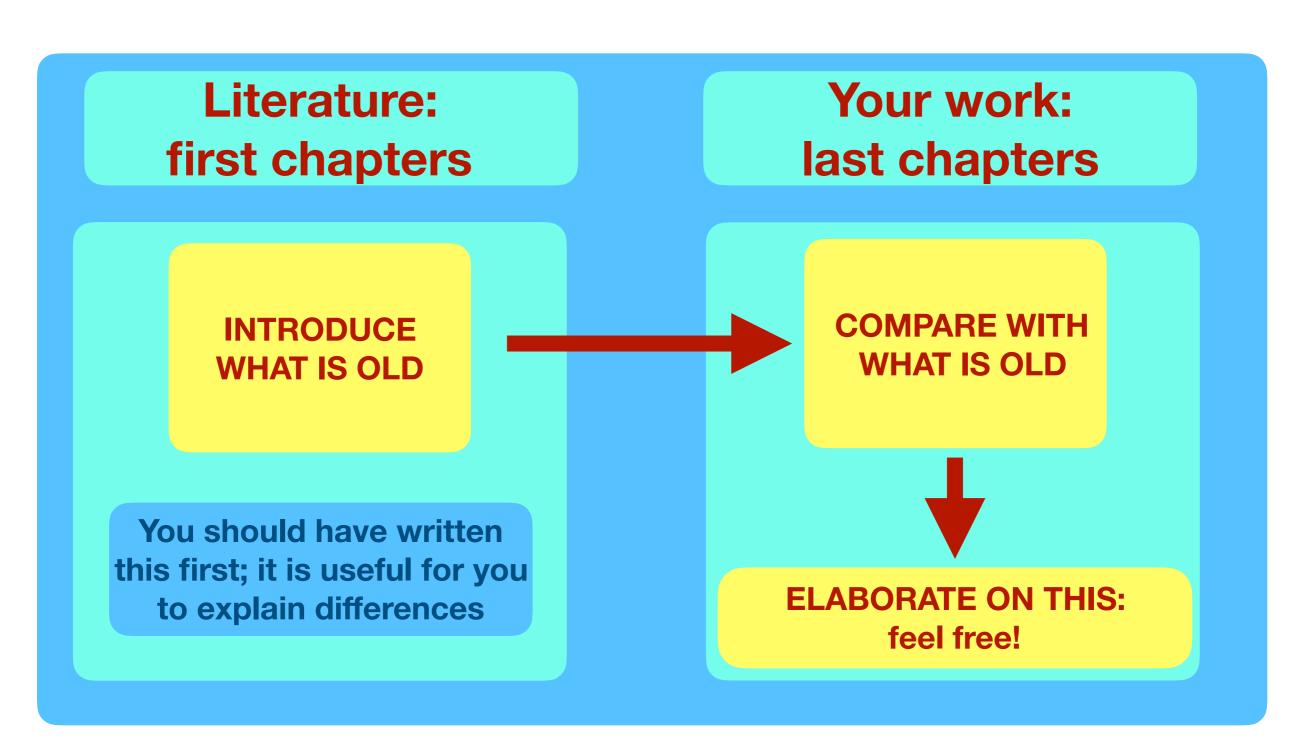
Literature: first chapters

Your work: last chapters

INTRODUCE WHAT IS OLD

You should have written this first; it is useful for you to explain differences





YOUR THESIS COMMITTEE

Also known as: an impossibly difficult group to get together in one room but who nevertheless hold your future in their hands depending on their ability to reach a civilized consensus.



Your Professor

Simultaneously your biggest ally and your worst enemy. Will be the first to suggest you do more work.



The Guru

Only here for the free cookies.
Don't forget to bring cookies.



Adversary The:Accircle

Has bitter rivalry with your Professor and will argue the exact opposite view. Work this to your advantage.



The Strawman/woman

Nice guy. No opinions.



The Assistant Professor

Still doesn't believe just a few months ago they were on the other side just like you. Pretends to be an adult.

NONE OF THEM WILL ACTUALLY READ YOUR ENTIRE THESIS.

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as a joke,
but
many times
overlaps
one-to-one
with reality
(true both for
PhD and Master)

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It is a complete summary of the thesis: one should understand what is written in the main body just from reading this. If you succeed, the reader will feel happy!

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Some text framing your area of research

Examples:

"Standard Model"
"General Relativity"
"Complex systems"

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Motivation for your specific research

Most important part of the Introduction: show that you know WHY you have done what you have done!

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Organization of the thesis: "In Chapt. 1 I review; in Chapt. 2...."

Five: Conclusions

It is a short summary of the thesis: remind shortly the motivation and the results with respect to the literature

This is the place to insert your work as just one single brick in a big construction to be carried over many years

Use a few words to explain possible future development, "beyond the scope of this thesis."