

# Recollections from more than 20 years of collaboration and friendship

(Contribution to HirschFest)

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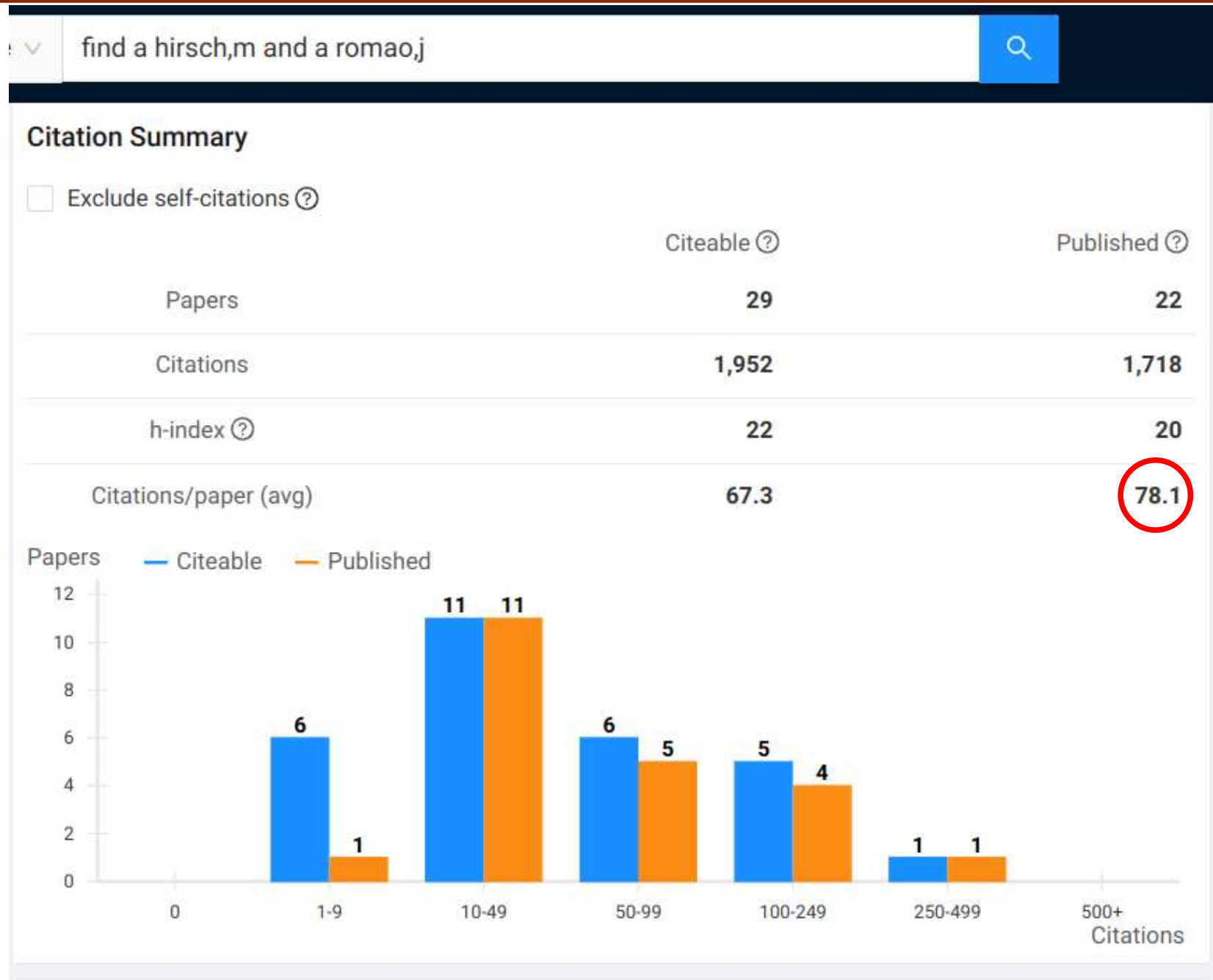
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- ❑ R-Parity Violation and neutrino masses masses
  - ◆ **Bilinear R-Parity Violation**
  - ◆ Spontaneous R-Parity Violation
- ❑ Testing R-Parity at the colliders
  - ◆ SUSY decays
  - ◆ Higgs decays
- ❑ **Lepton Flavour Violation (LFV) and Dark Matter**
  - ◆ Type I, II and III seesaw
  - ◆ Left-Right (LR) model
- ❑ Unification in GUT's
  - ◆ LR models with sliding scale

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- R-Parity

- BRpV

- Testing @ LHC

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$$R_P = (-1)^{2J+3B+L}$$

- ❑ BRpV: **Explicit R-Parity Violation**
  - ◆ Bilinear R-Parity Violation
  - ◆ Same particle content as the MSSM
- ❑ SRpV: **Spontaneously R-Parity Violation**
  - ◆ More Complicated Higgs Structure
  - ◆ Majoron  $J$
  - ◆ Invisible Higgs Decay  $h \rightarrow JJ$

Rich Phenomenology

Hierarchical Spectrum: No visible  $\beta\beta_{0\nu}$

J.W.F. Valle, M. Diaz, M. Hirsch, W. Porod, A. Villanova del Moral, JCR, ...

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- The superpotential of the BRpV model is

$$W^{\text{BRpV}} = W^{\text{MSSM}} + \varepsilon_{ab} \epsilon_i \hat{L}_i^a \hat{H}_2^b$$

- The superfield content is as in the MSSM.
- The last term in is the only  $R$ -parity violating term. The  $\epsilon_i$  are parameters with units of mass
- The electroweak symmetry is broken when the two Higgs doublets  $H_d$  and  $H_u$ , and the neutral component of the slepton doublets  $\tilde{L}_i^0$  acquire vacuum expectation values.
- The scalar potential is

$$V_{\text{total}}^0 = \sum_i \left| \frac{\partial W}{\partial z_i} \right|^2 + V_D + V_{\text{SB}}^{\text{MSSM}} + V_{\text{SB}}^{\text{BRpV}}$$

where  $z_i$  is any one of the scalar fields in the superpotential

- The  $7 \times 7$  neutralino/neutrino mass matrix  $\mathbf{M}_N$  has two zero eigenvalues. Therefore two of the  $\nu$ 's remain massless at tree level.
- The values of the parameters  $\epsilon_i$  needed to account for the neutrino masses and mixings are small compared with the electroweak scale.
- Using this we can approximately bring  $\mathbf{M}_N$  to block diagonal form,

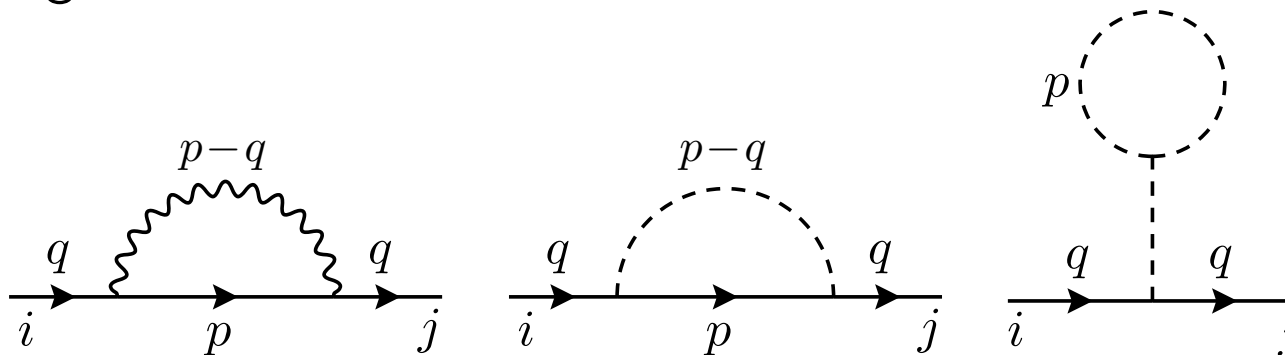
$$\mathbf{M}_N \simeq \begin{bmatrix} m_{\text{eff}} & 0 \\ 0 & \mathcal{M}_{\chi^0} \end{bmatrix}, m_{\text{eff}} = \frac{M_1 g^2 + M_2 g'^2}{4 \det(\mathcal{M}_{\chi^0})} \begin{bmatrix} \Lambda_e^2 & \Lambda_e \Lambda_\mu & \Lambda_e \Lambda_\tau \\ \Lambda_e \Lambda_\mu & \Lambda_\mu^2 & \Lambda_\mu \Lambda_\tau \\ \Lambda_e \Lambda_\tau & \Lambda_\mu \Lambda_\tau & \Lambda_\tau^2 \end{bmatrix}$$

where  $\mathcal{M}_{\chi^0}$  is the neutralino mass matrix and  $\Lambda_i \equiv \mu v_i + v_d \epsilon_i$ .

- It is clear from the projective nature of  $m_{\text{eff}}$  that two of the  $\nu$ 's remain massless at tree level. The non zero neutrino mass is

$$m_\nu = \text{Tr}(m_{\text{eff}}) = \frac{M_1 g^2 + M_2 g'^2}{4 \det(\mathcal{M}_{\chi^0})} |\vec{\Lambda}|^2$$

- The diagrams contributing to the self-energy include the exchange of gauge bosons as well of scalars.



- These diagrams can be calculated in a straightforward way. For instance the  $W$  diagram in the  $\xi = 1$  gauge gives

$$\Sigma_{ij}^V = -\frac{1}{16\pi^2} \sum_{k=1}^5 2 (O_{Ljk}^{\text{ncw}} O_{Lki}^{\text{cnw}} + O_{Rjk}^{\text{ncw}} O_{Rki}^{\text{cnw}}) B_1(p^2, m_k^2, m_W^2)$$

$$\Pi_{ij}^V = -\frac{1}{16\pi^2} \sum_{k=1}^5 (-4) (O_{Ljk}^{\text{ncw}} O_{Rki}^{\text{cnw}} + O_{Rjk}^{\text{ncw}} O_{Lki}^{\text{cnw}}) m_k B_0(p^2, m_k^2, m_W^2)$$

where  $B_0$  and  $B_1$  are the Passarino-Veltman functions, and  $O^{\text{cnw}}$ ,  $O^{\text{ncw}}$  are coupling matrices

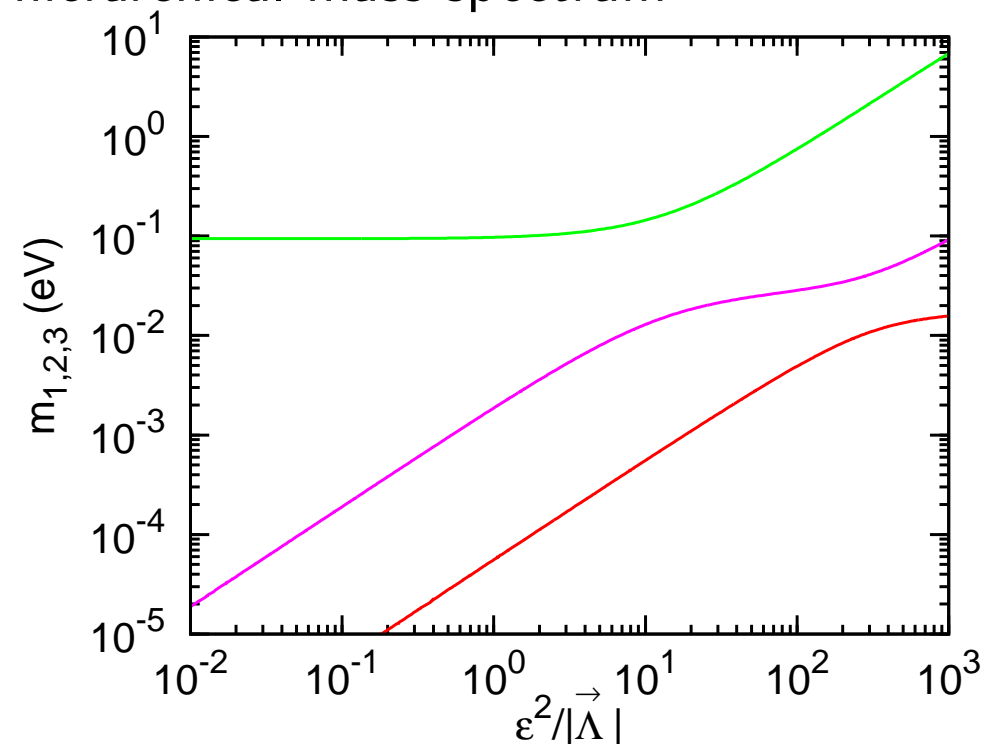
□ The BRpV model produces a hierarchical mass spectrum

□ The largest mass scale can be estimated by the tree level value

□ The solar mass scale can be obtained at one-loop

□ The correct mixing angles can be obtained by an appropriate choice of  $|\vec{\Lambda}|$  and  $\Lambda_i$  such that

$$\Lambda_e \ll \Lambda_\mu, \Lambda_\tau$$



M. Hirsch, M. Diaz, W. Porod, JCR, J.W.F. Valle  
Phys.Rev.D62:113008,2000

**At the time BRpV was proposed the neutrino data was not very well known. It is interesting that with the present high precision data on neutrino physics, the model can still accommodate it.**



# The mixings $\theta_{13}$ , $\theta_{23}$ and $\theta_{12}$ could be accommodated

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• R-Parity

• **BRpV**

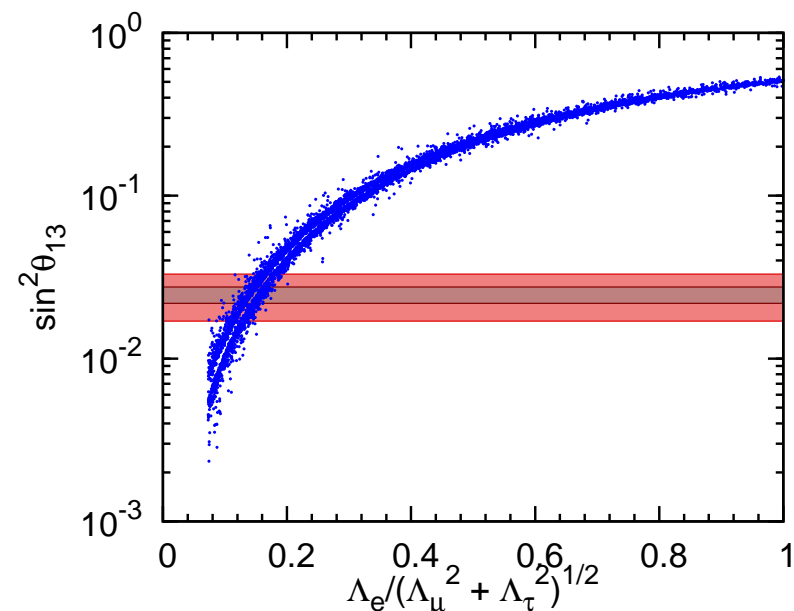
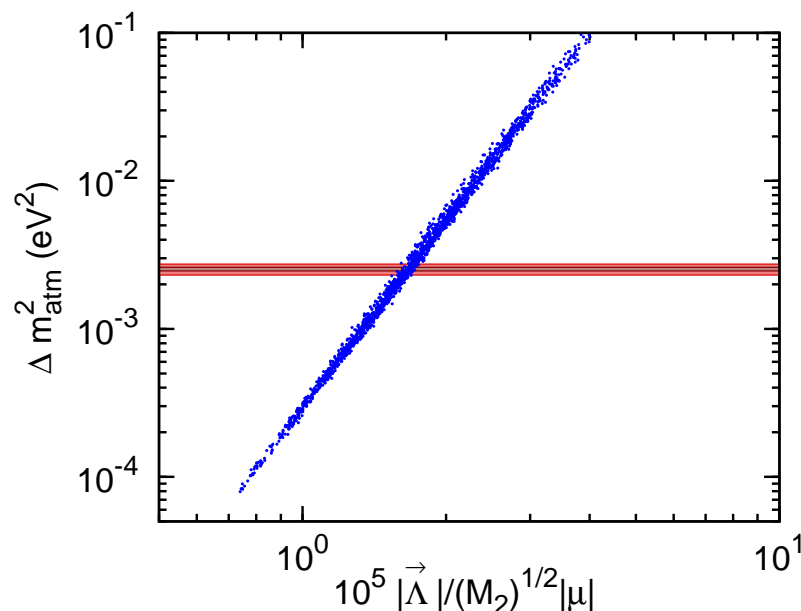
• Testing @ LHC

[Seesaw Models](#)

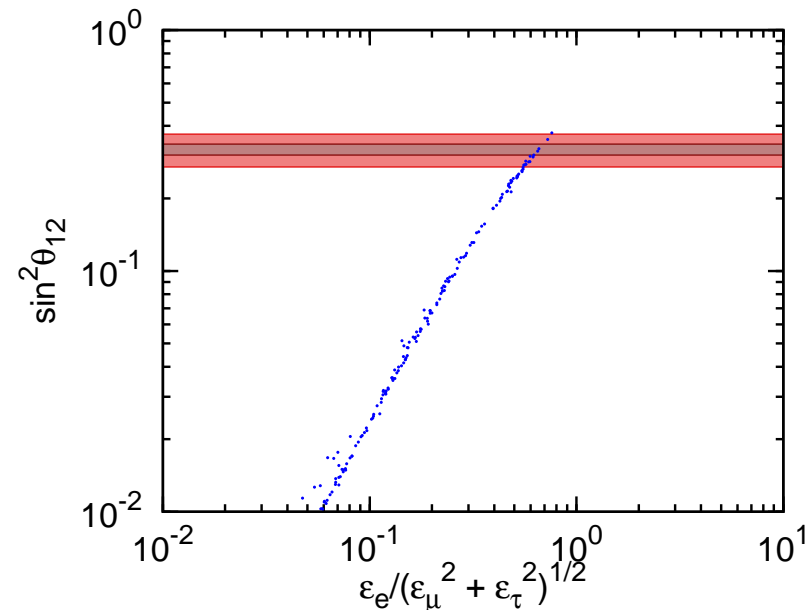
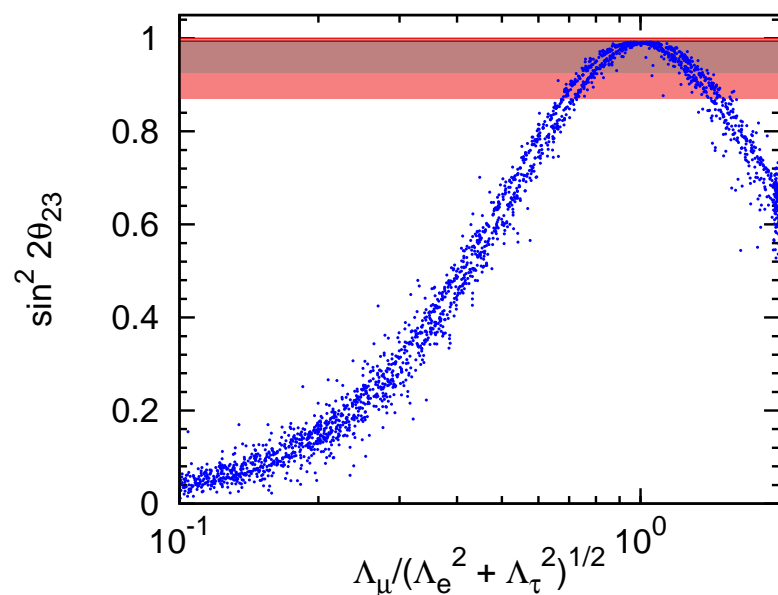
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J.W.F. Valle and JCR, Neutrinos in high energy and astroparticle physics (Wiley, 2015)



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## □ LSP Decays: (mSUGRA)

The fact that, in these models, the LSP decays through R-parity violating processes allows it to be either neutral or charged.

- ◆ In most cases the LSP is the lightest neutralino , like in the MSSM.
- ◆ For some regions of the parameter space the LSP can also be the scalar tau .
- ◆ In both cases we have shown that despite the smallness of  $m_\nu$  the LSP decays inside the detector.

## □ LSP Decays: (non mSUGRA)

If we depart from mSUGRA then the LSP can be almost any particle. This gives complementary information .

# Example: Probing the Atmospheric Angle (old slide)

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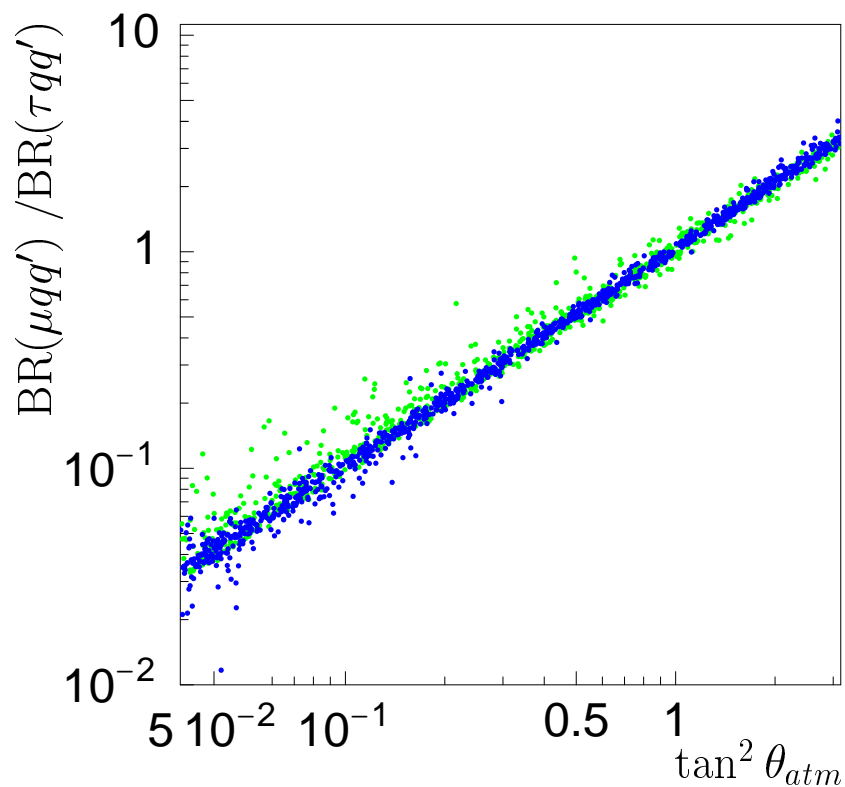
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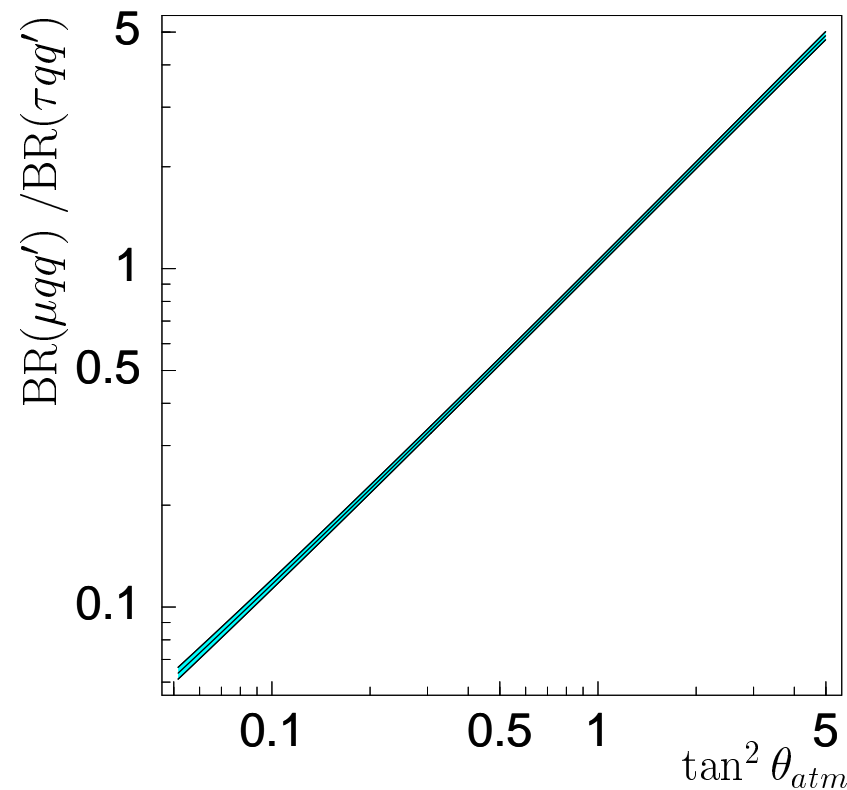
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Before SUSY



W. Porod, M. Hirsch, JCR, J.W.F. Valle  
Phys.Rev.D63:115004,2001

After SUSY



$M_2 = 120 \text{ GeV}, \mu = 500 \text{ GeV}, \tan \beta = 5$   
 $m_0 = 500 \text{ GeV}, A = -500 \text{ GeV}.$

# Seesaw models for neutrino masses: Type-I-II-III

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Seesaw Models

• Type-I-II-III

• LR

• Effect on Spectra

• LFV

• DM & LFV

•  $\mathcal{A}(\mu^+ \rightarrow e^+ \gamma)$

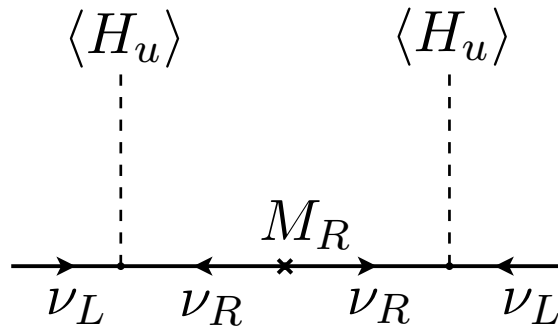
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$$\mathcal{L} = \dots + H_u \bar{\nu}_L Y_\nu^I \nu_R - \frac{1}{2} \nu_R^T C^{-1} M_R \nu_R$$

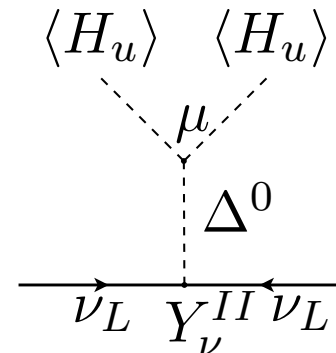
type-I



$$m_{\text{eff}}^I = -(v Y_\nu) M_R^{-1} (v Y_\nu)^T$$

$$\mathcal{L} = \dots - \frac{1}{2} Y_\nu^{\text{II}} \bar{\nu}_L^c i \tau_2 \Delta_L \nu_L - \mu H_u^T \Delta_L H_u - M_\Delta^2 \Delta_L^\dagger \Delta_L$$

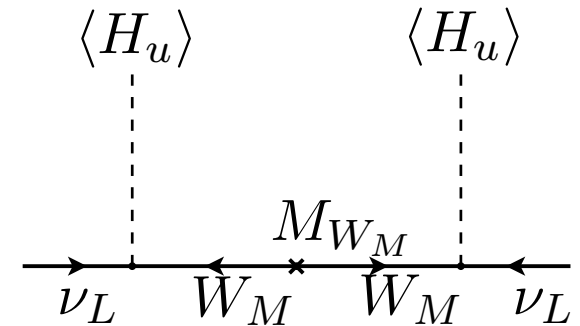
type-II



$$m_{\text{eff}}^{\text{II}} = \frac{v^2 \mu Y_\nu^{\text{II}}}{M_\Delta^2}$$

$$\mathcal{L} = \dots + H_u \bar{W}_M Y_\nu^{\text{III}} \nu_L - \frac{1}{2} W_M^T C^{-1} M_{W_M} W_M$$

type-III



$$m_{\text{eff}}^{\text{III}} = -(v Y_\nu^{\text{III}}) M_{W_M}^{-1} (v Y_\nu^{\text{III}})^T$$

- ❑ Exchanged particle: Type-I-III: neutral fermion. Type-II: neutral scalar
- ❑ Type-I: gauge singlet. Type-II-III; gauge triplet → Stronger running
- ❑  $m_\nu \sim 1\text{eV}$  and  $Y_\nu \sim \mathcal{O}(1)$  →  $M_{\text{Seesaw}} \sim \mathcal{O}(10^{12-14})\text{GeV}$  Not directly observable

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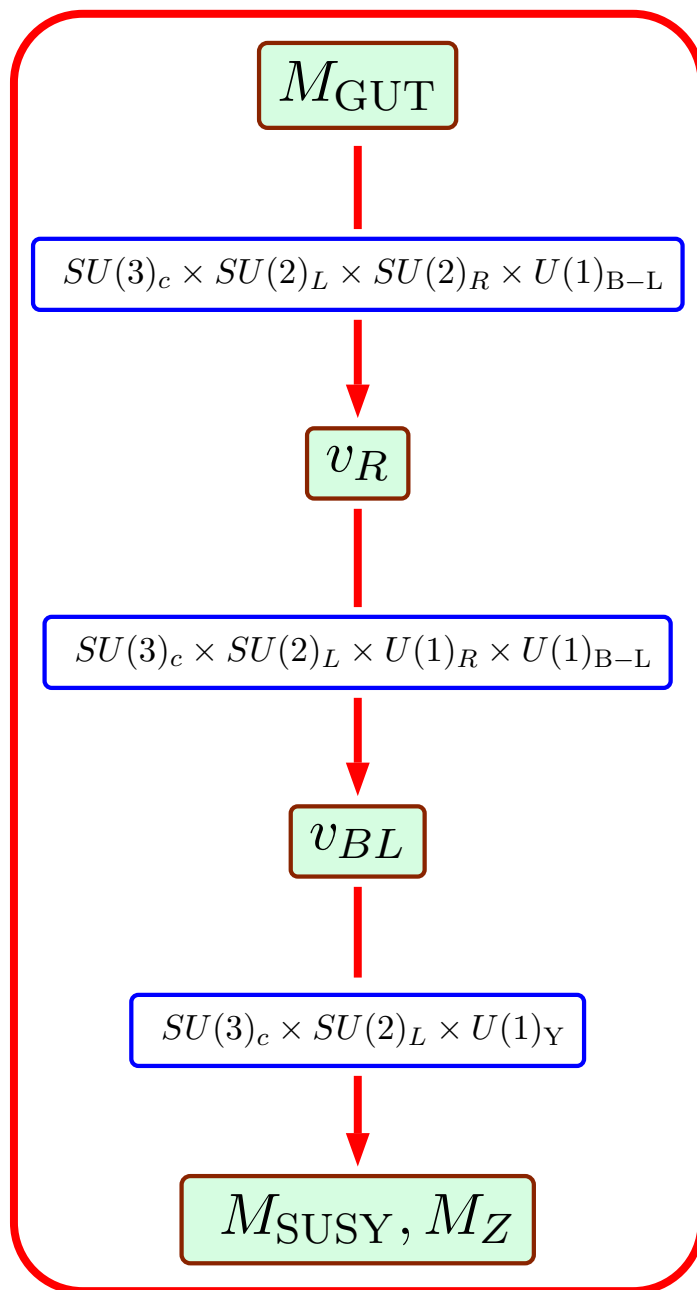
• DM & LFV

•  $\mathcal{A}(\mu^+ \rightarrow e^+ \gamma)$

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## □ Superpotential

$$\mathcal{W} = Y_L L \Phi L^c - f_c L^c \Delta^c L^c + \dots$$

◆  $Y_L$  and  $f_c$  complex  $3 \times 3$  matrices

## □ Lagrangian at $v_{BL} = \langle \Delta_c^0 \rangle$

$$\mathcal{L} = H_u \bar{\nu}_L Y_\nu^\dagger \nu_R - \frac{1}{2} \nu_R^T C^{-1} (f_c v_{BL}) \nu_R + \dots$$

## □ Effective neutrino mass matrix (type-I)

$$m_{\text{eff}}^{\text{LR}} = -(v Y_\nu) (f_c v_{BL})^{-1} (v Y_\nu)^T$$

◆  $Y_\nu$  fit  $\rightarrow f_c = \mathbb{1}$ ,  $Y_\nu$  arbitrary

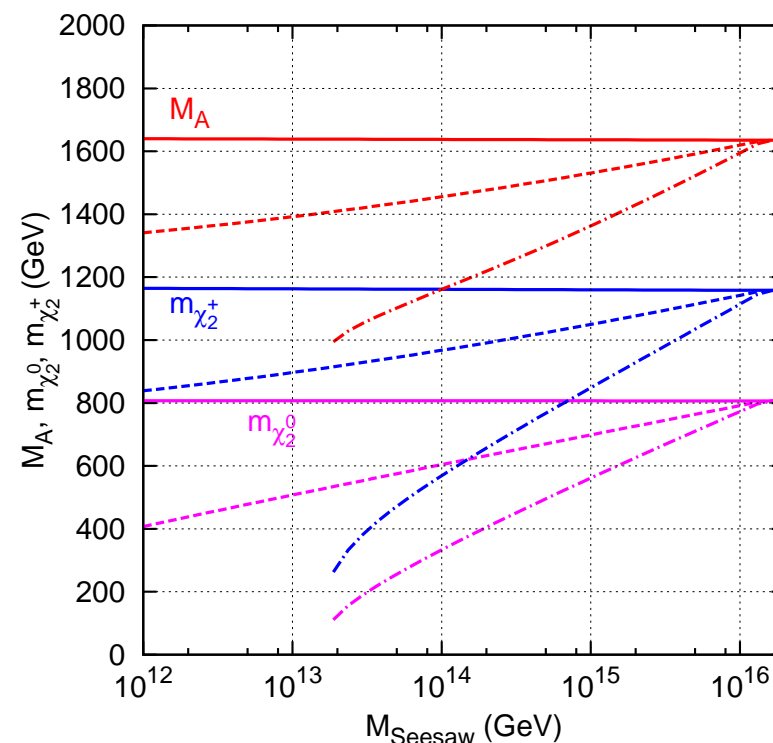
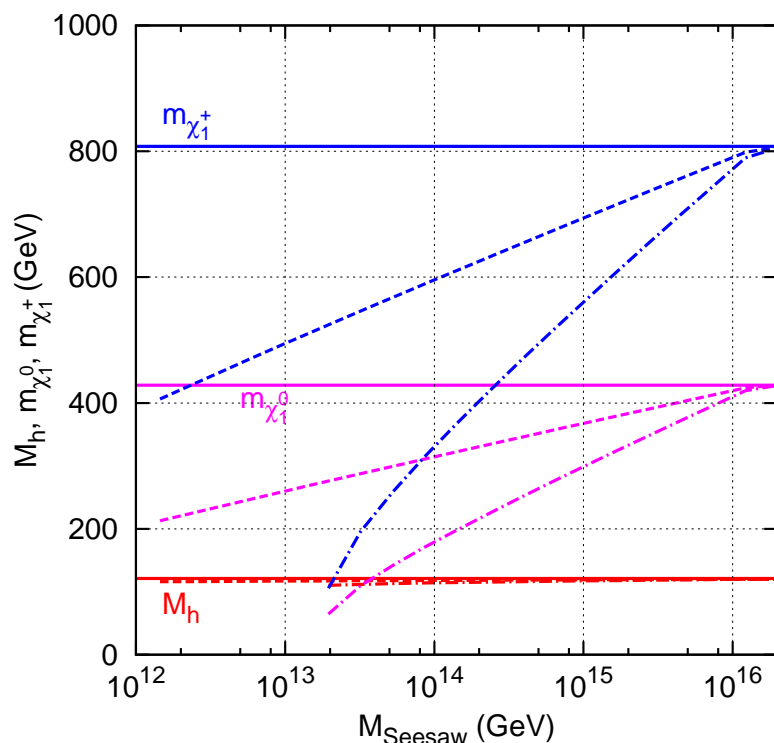
◆  $f$  fit  $\rightarrow Y_\nu = \mathbb{1}$ ,  $f_c$  arbitrary

◆ Different imprints on RGE running

# Effects of the heavy particles on the MSSM spectrum

- The appearance of particles with charges under the gauge group at scales between  $M_{\text{Seesaw}}$  and  $M_{\text{GUT}}$  leads to **changes in the beta functions of the gauge couplings**
- Example of spectra at  $Q = 1$  TeV versus the seesaw scale
- SUSY parameters  $m_0 = M_{1/2} = 1$  TeV,  $A_0 = 0$ ,  $\tan \beta = 10$  and  $\mu > 0$
- Type-I (full lines), type-II (dashed) and type-III (dash-dotted)

J.N. Esteves, JCR, M. Hirsch, F. Staub and W. Porod, Phys.Rev.D83,013003,2011



Starting with universal (mSUGRA) boundary conditions @  $M_{\text{GUT}}$

## □ Seesaw type-I-II-III

$$\Delta m_{L,ij}^2 \simeq -\frac{a_k}{8\pi^2} (3m_0^2 + A_0^2) \left( Y_N^{k,\dagger} L Y_N^k \right)_{ij}, \quad L = \ln\left(\frac{M_{\text{GUT}}}{M_N}\right)$$

$$\Delta m_{E,ij}^2 \simeq 0 \quad a_{\text{I}} = 1, \quad a_{\text{II}} = 6 \quad \text{and} \quad a_{\text{III}} = \frac{9}{5}$$

## □ Left-Right Model

 $M_{\text{GUT}}$ 

 $v_R$ 

$$\Delta m_L^2 \simeq -\frac{1}{4\pi^2} \left( 3f f^\dagger + Y_L^{(k)} Y_L^{(k)\dagger} \right) (3m_0^2 + A_0^2) \ln\left(\frac{M_{\text{GUT}}}{v_R}\right)$$

$$\Delta m_E^2 \simeq -\frac{1}{4\pi^2} \left( 3f^\dagger f + Y_L^{(k)\dagger} Y_L^{(k)} \right) (3m_0^2 + A_0^2) \ln\left(\frac{M_{\text{GUT}}}{v_R}\right)$$

 $v_R$ 

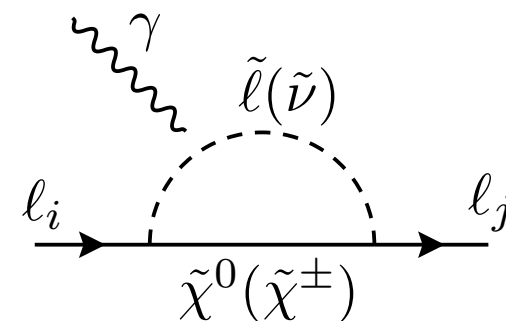
 $v_{BL}$ 

$$\Delta m_L^2 \simeq -\frac{1}{8\pi^2} Y_\nu Y_\nu^\dagger (m_L^2|_{v_R} + A_e^2|_{v_R}) \ln\left(\frac{v_R}{v_{BL}}\right)$$

$$\Delta m_E^2 \simeq 0$$

□ BR( $l_i \rightarrow l_j \gamma$ ): (**MEG**...)

$$\mathcal{L}_{eff} = e \frac{m_i}{2} \bar{l}_i \sigma_{\mu\nu} F^{\mu\nu} (A_L^{ij} P_L + A_R^{ij} P_R) l_j + h.c.$$



$$\text{BR}(l_i \rightarrow l_j \gamma) = \frac{48\pi^3 \alpha}{G_F^2} \left( |A_L^{ij}|^2 + |A_R^{ij}|^2 \right) \text{BR}(l_i \rightarrow l_j \nu_i \bar{\nu}_j)$$

□ For seesaw models:  $A_L^{ij} \sim \frac{(\Delta m_L^2)_{ij}}{m_{SUSY}^4}$ ,  $A_R^{ij} \sim \frac{(\Delta m_E^2)_{ij}}{m_{SUSY}^4}$

- ◆ type-I-II-III  $\rightarrow$  only  $A_L$
- ◆ Left-Right model: In principle **both**  $A_L$  and  $A_R$
- ◆ Distinguish models:  $\rightarrow$  Positron polarization asymmetry(**MEG**)

$$\mathcal{A}(\mu^+ \rightarrow e^+ \gamma) = \frac{|A_L|^2 - |A_R|^2}{|A_L|^2 + |A_R|^2} = \begin{cases} 1 & \text{type-I-II-III} \\ \neq 1 & \text{LR} \end{cases}$$

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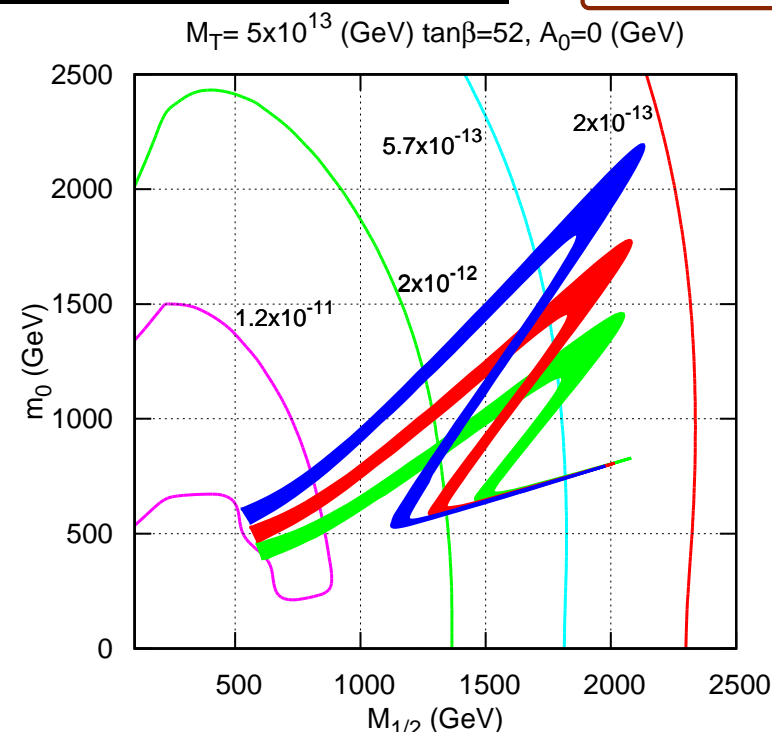
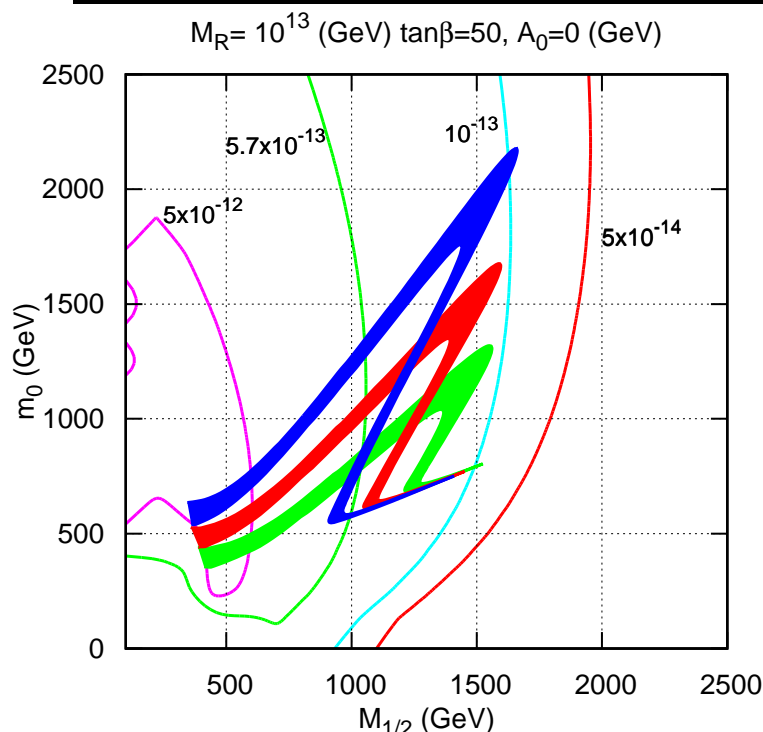
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- Parameters: SUSY:  $\{m_0, M_{1/2}, A_0 = 0, \tan \beta = 10, 52, \text{sign}(\mu) = +\}$
- Seesaw:  $M_R = 10^{13}$  GeV and  $M_T = 5 \times 10^{13}$  GeV
- Dark matter region: WMAP ( $3\sigma$ ),  $0.081 \leq \Omega h^2 \leq 0.129$
- SPheno(W.Porod), SARAH(F.Staub)

J.N. Esteves, S. Kaneko, JCR, M. Hirsch, and W. Porod, Phys.Rev.D80,095003,2009

Higgs funnel



- $m_{top} = 169.1$  GeV (blue),  $171.2$  GeV (red),  $173.3$  GeV (green)
- Superimposed the contour lines for the  $Br(\mu \rightarrow e\gamma)$ . With the present bound of  $4.2 \times 10^{-13}$  most of parameter space shown is excluded

# A low-energy observable for Left-Right model: $\mathcal{A}(\mu^+ \rightarrow e^+ \gamma)$

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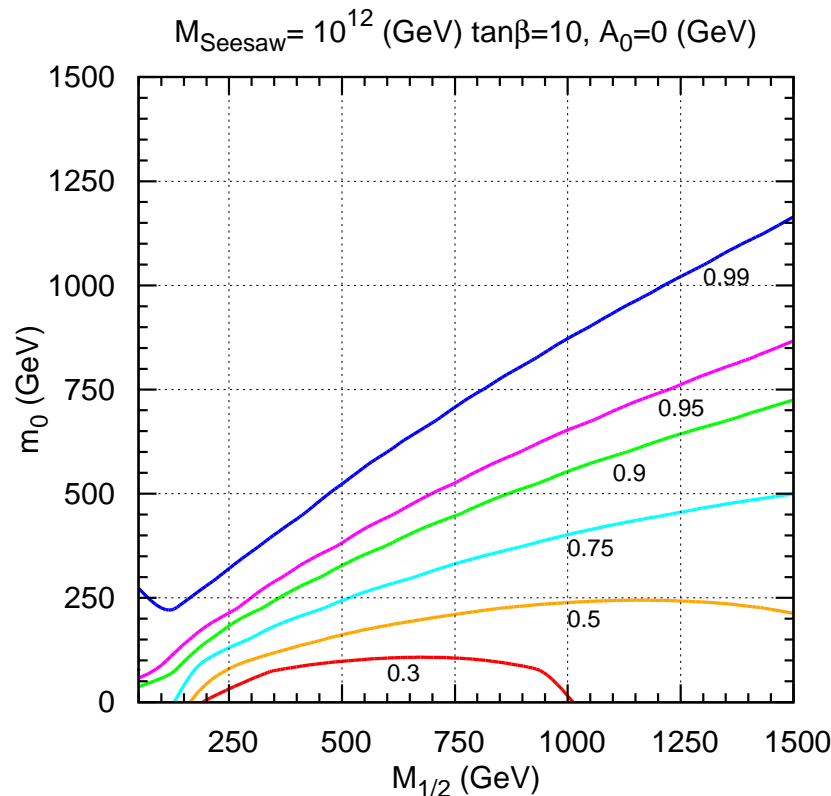
- Type-I-II-III
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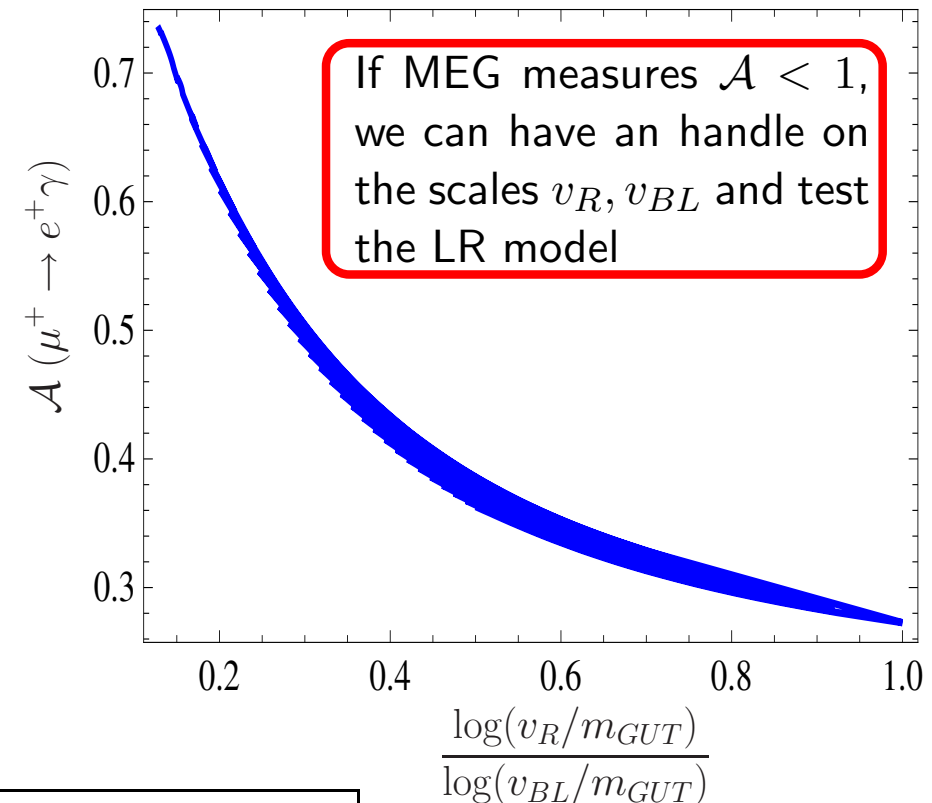
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- Positron polarization asymmetry:  $\mathcal{A}(\mu^+ \rightarrow e^+ \gamma) = \frac{|A_L|^2 - |A_R|^2}{|A_L|^2 + |A_R|^2}$
- In seesaw type-I-II-III:  $\mathcal{A}(\mu^+ \rightarrow e^+ \gamma) = 1$ , as  $A_R \simeq 0$
- Parameters:
  - ◆ SUSY: SPS3  $\{m_0 = 90, M_{1/2} = 400, A_0 = 0, \tan \beta = 10, \text{sign}(\mu) = +\}$
  - ◆ LR:  $v_{BL} = 10^{15}$  GeV,  $v_R \in [10^{14}, 10^{15}]$  GeV,  $Y_\nu$  fit



J.N. Esteves, JCR, M. Hirsch, A. Vicente and W. Porod, JHEP 12 (2010) 077



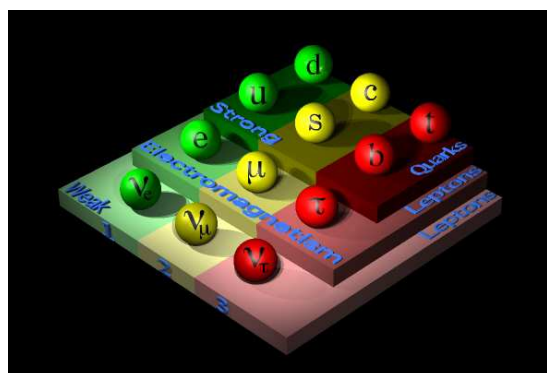
## □ Number of particles with Spin 1

Fixed by the choice of Symmetry Group  $SU(3) \times SU(2) \times U(1)$

Properties of the Interactions				
The strengths of the interactions (forces) are shown relative to the strength of the electromagnetic force for two u quarks separated by the specified distances.				
Property	Gravitational Interaction	Weak Interaction (Electroweak)	Electromagnetic Interaction	Strong Interaction
Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge
Particles experiencing:	All	Quarks, Leptons	Electrically Charged	Quarks, Gluons
Particles mediating:	Graviton (not yet observed)	$W^+$ $W^-$ $Z^0$	$\gamma$	Gluons
Strength at $\begin{cases} 10^{-18} \text{ m} \\ 3 \times 10^{-17} \text{ m} \end{cases}$	$10^{-41}$ $10^{-41}$	0.8 $10^{-4}$	1 1	25 60

## □ Number of Particles with Spin $\frac{1}{2}$

There is no principle. Fixed by experiment



## □ Number of particles with Spin 0

There is no principle.

Therefore should be fixed by experiment!

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- Theory
    - ◆ Perturbative Unitarity
    - ◆ Bounded From Below
    - ◆ Precision observables
    - ◆ CP violation and EDM constraints
    - ◆ Contributions to  $b \rightarrow s\gamma$
    - ◆ Many parameters: Symmetry constrained models
    - ◆ Dark Matter
    - ◆ Flavor changing neutral interactions (FCNI)
    - ◆ ...
  - Experiment
    - ◆ Confronting LHC data
    - ◆ Confronting the new limits on the EDM
    - ◆ Checking flavor constraints
    - ◆ ...
- Also lots of fun in the last 10 years ...

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These are the students and postdocs that I have shared with Martin:

☐ Albert Villanova del Moral

☐ Solveig Skadhauge

☐ Cyril Hugonie

☐ João N. Esteves

☐ Satoru Kaneko

☐ Thomas Kernreiter

☐ Avelino Vicente

☐ Carolina Arbelaez

☐ Renato M. Fonseca

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- ❑ It has been a lot of fun to work all these years with Martin
- ❑ We have established a very strong link between the Valencia and Lisbon groups, with the exchange of many students.
- ❑ I hope that the next years will be as rewarding as the ones that have elapsed.

**Thank you Martin**

# Multi-Higgs Workshops

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## Breaking R-parity

## Seesaw Models

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# Workshop on Multi-Higgs Models

## 3 - 6 Sept 2024

Lisbon - Portugal

This Workshop brings together those interested in the theory and phenomenology of Multi-Higgs models. The program is designed to include talks given by some of the leading experts in the field, and also ample time for discussions and collaboration between researchers. A particular emphasis will be placed on identifying those features of the models which are testable at the LHC and DM searches.

For registration and/or to propose a talk, send an email to:

2hdmwork@cftp.tecnico.ulisboa.pt

Web Page : <http://cftp.tecnico.ulisboa.pt/~2hdmwork/>

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