

L4: Baryogenesis, dark matter and cosmic messengers

(a brief survey of topics studied at IFIC)



New name: Astroparticles, dark matter and cosmology



December 15, 2023

Coordinators:

Sergio Palomares, Óscar Vives and Juande Zornoza

L4: Baryogenesis, dark matter and cosmic messengers

at IFIC...



Matter-antimatter asymmetry

(mainly via leptogenesis)

collider and indirect
dark sector searches

Dark matter
(experiment and theory/phenomenology)

model building,
phenomenology
(cosmo, astro, colliders, etc.)

(mainly neutrinos)

Cosmic messengers
(from astrophysical sources)

(based on slides provided by a number of people)

Dark matter faces

MODEL BUILDING AND PRODUCTION MECHANISMS

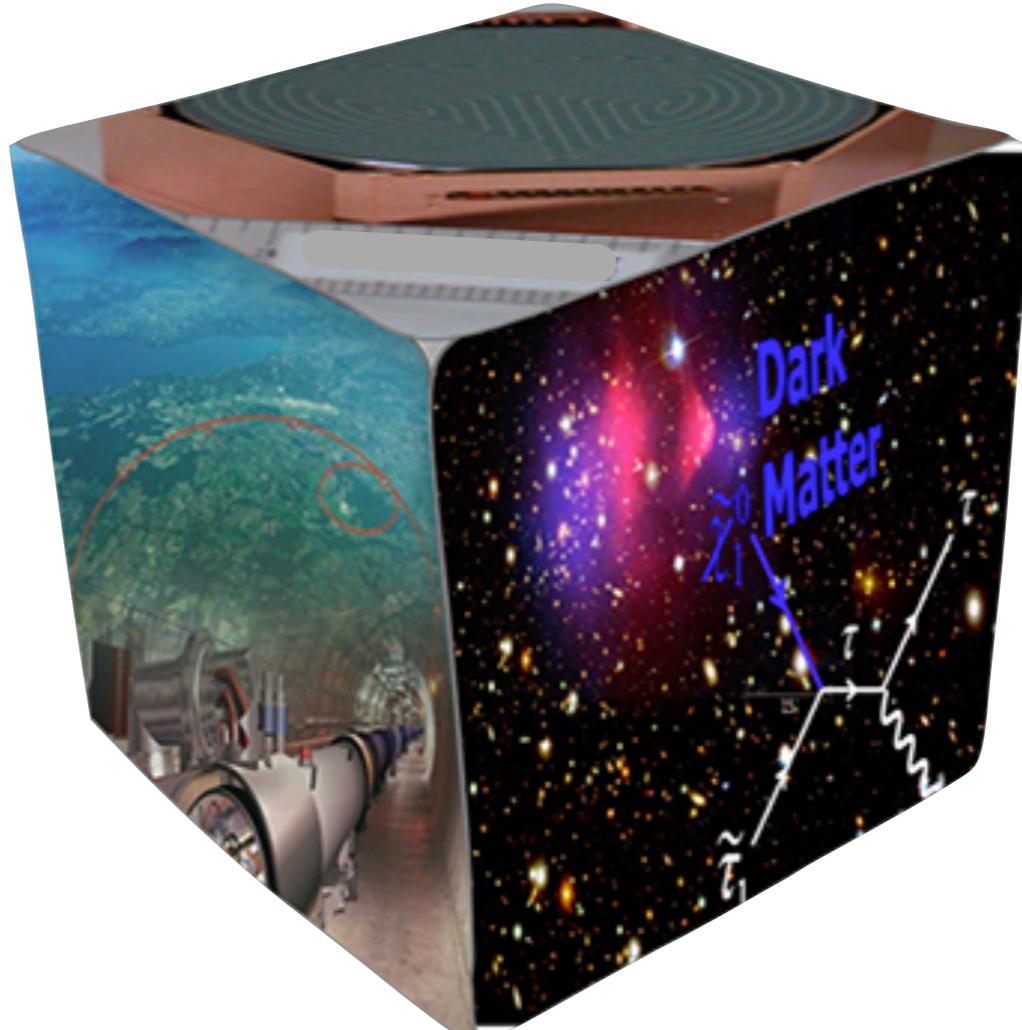


ASTRO/COSMO PROBES

BBN, CMB, 21cm, structure formation...

DIRECT DETECTION

Nuclear recoil produced by DM scattering



ACCELERATOR SEARCHES

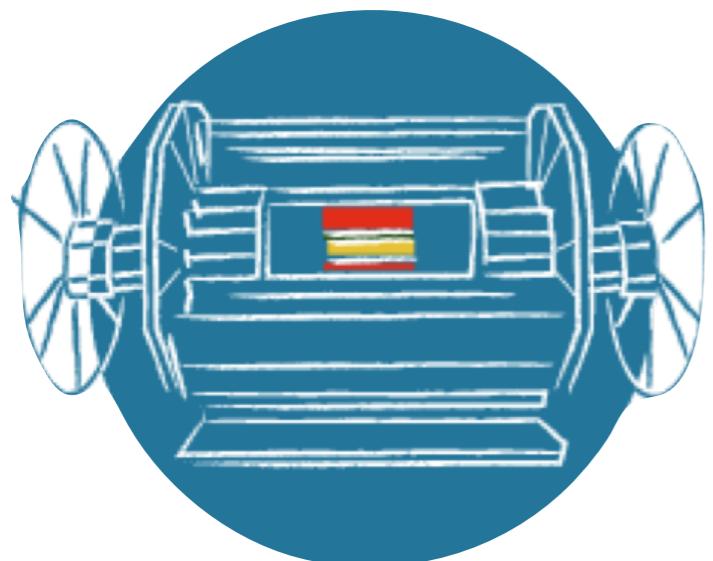
LLPs, Missing energy, mono-jets(bosons)

INDIRECT DETECTION

Products of DM annihilation/decay

Significant connections with other lines

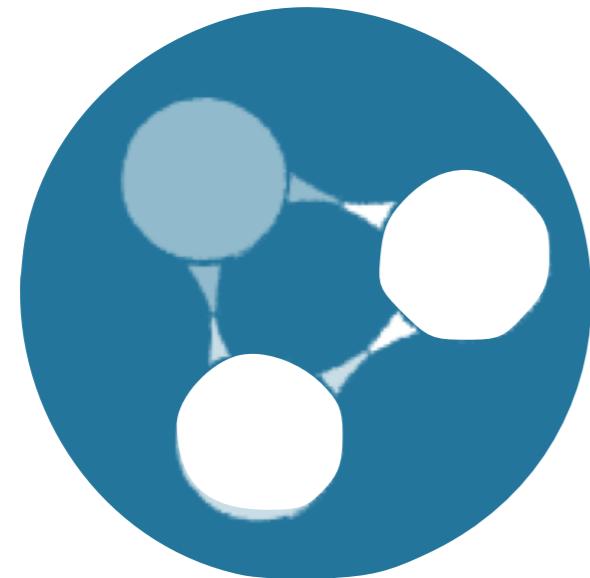
Line 1: The Higgs Force



Line 2: Neutrinos and Lepton Flavour



Line 3: Flavor and Quark Matter



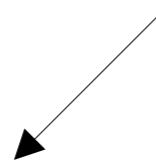
Line 5: Gravity and the Dark Universe:
Gravitational Waves and Black Holes



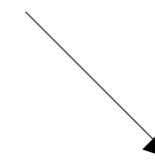
Dark matter model building

Model building boils down to (properly) choosing...

Symmetry + Representations



DM stability



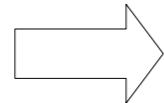
DM candidate

In the process of making a DM model one should consider...

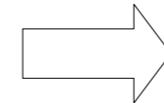
- Is the model compatible with the current constraints?
- Does the model address any other problem of the SM?
- Would a variation of the model lead to a different phenomenology?



MANY
(really many)
models



Recent
works just
at IFIC



- Scotogenic model (+ variants)
- DM from flavor or CP
- Supersymmetric DM
- DM from extra dimensions
- DM related to lepton number

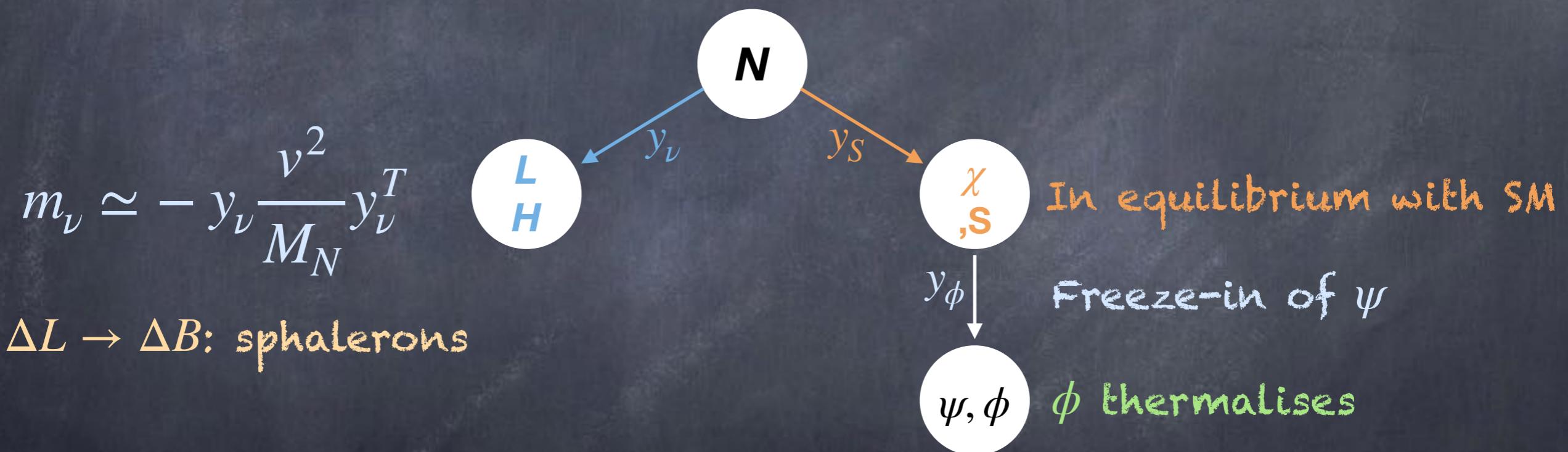
Dark matter in the light of no signals

Well-motivated DM models with suppressed signals are needed:
PNGB DM, annihilations into ν_R , multi-DM, asymmetric freeze-in...

extended co-ogenesis framework

[G. Landini, JHG, D. Vatsyayan, JHEP05 (2023) 049]

$$\mathcal{L}_{\text{int}} = -y_\nu^{\alpha i} \bar{L}^\alpha \tilde{H} N_R^i - y_\sigma^{ij} \sigma \bar{N}_R^{ic} N_R^j - y_S^i S \bar{N}_R^i \chi - y_\phi \phi \bar{\psi} \chi + \text{H.c.}$$



Idea: Dark asymmetry transferred via late decays
 $\chi \rightarrow \psi + \phi$, after χ symmetric population has been erased

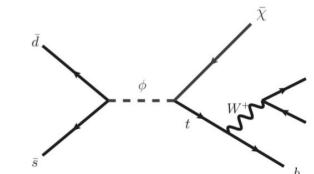
We consider $\eta_\chi \equiv \eta_D \simeq \eta_B$

Search for invisible particles produced in association with single top quarks

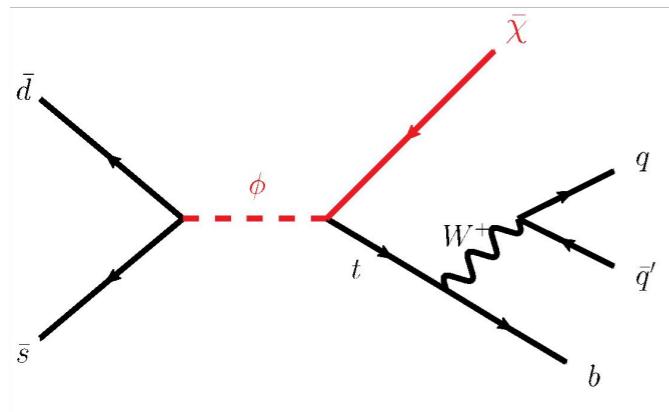
- Looking at signatures of a top + invisible with the full Run-2 dataset.
- All-hadronic channel considered.
- Focused on models:
 - Dark Matter (DM) Monotop model (scalar and vector DM mediators).
 - Vector-like top production.

Simplified models

- Signature-driven first-order description of new physics.
- Bridge gap between EFT and complete models.



Scalar DM mediator

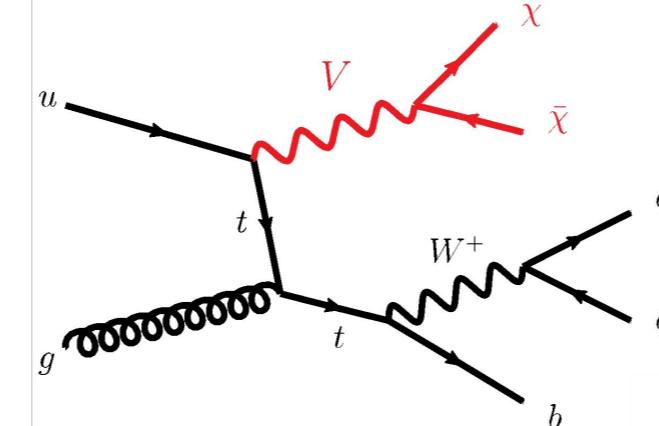


Colored-charged scalar decaying to a top quark and a DM particle.

Parameters:

- Couplings λ_q and y_χ .
- Masses of the mediator ϕ and the DM particle χ .

Vector DM Mediator

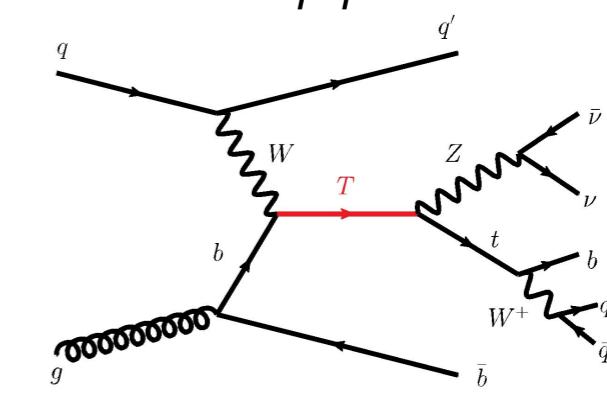


FCNC interaction between u and top quarks and a vector mediator, which further decays to a couple of DM particles.

Parameters:

- Couplings a and g_χ
- Masses of the mediator V and the DM particle χ .

Vector-like top production



Here focused on the single Vector-Like Top production, considering the invisible decay of the Z boson. Additional forward jet is expected.

Parameters:

- Overall coupling κ_T (related to W , Z and H couplings).
- Mass of the VLT.

And also...

Search for dark matter produced in association with single top quark and an energetic W-boson

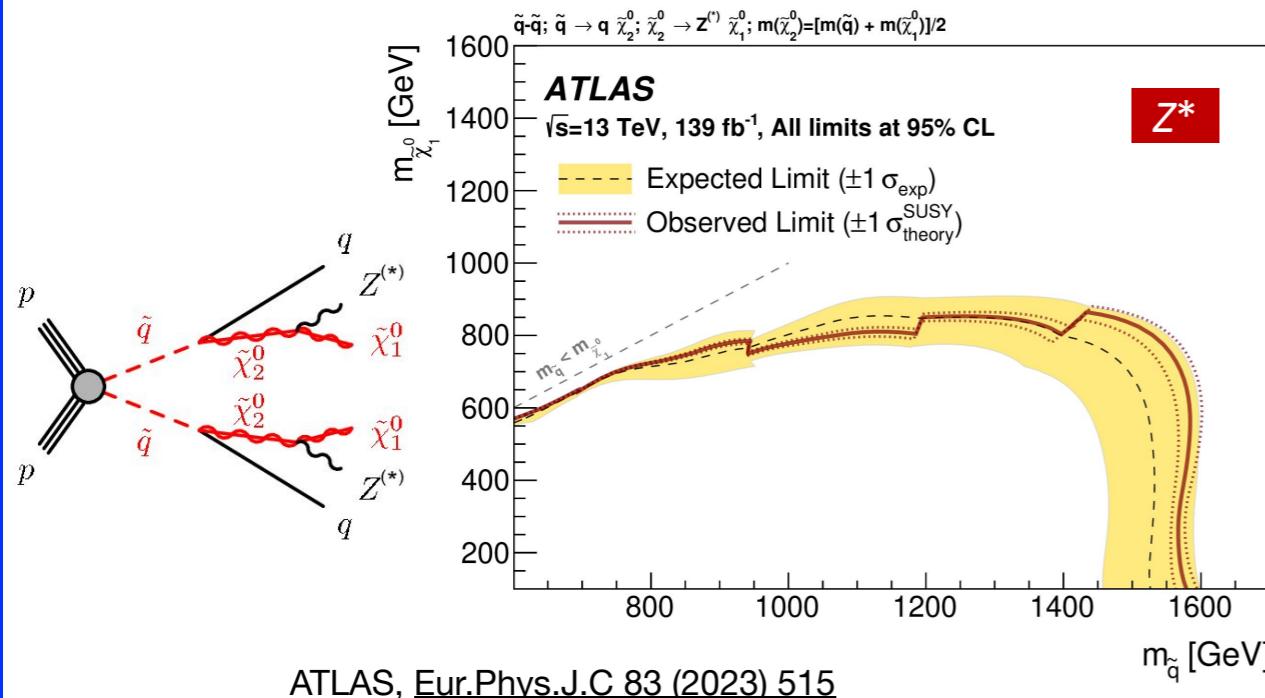
Search for neutral long-lived particles that decay into displaced hadronic jets in the ATLAS calorimeter



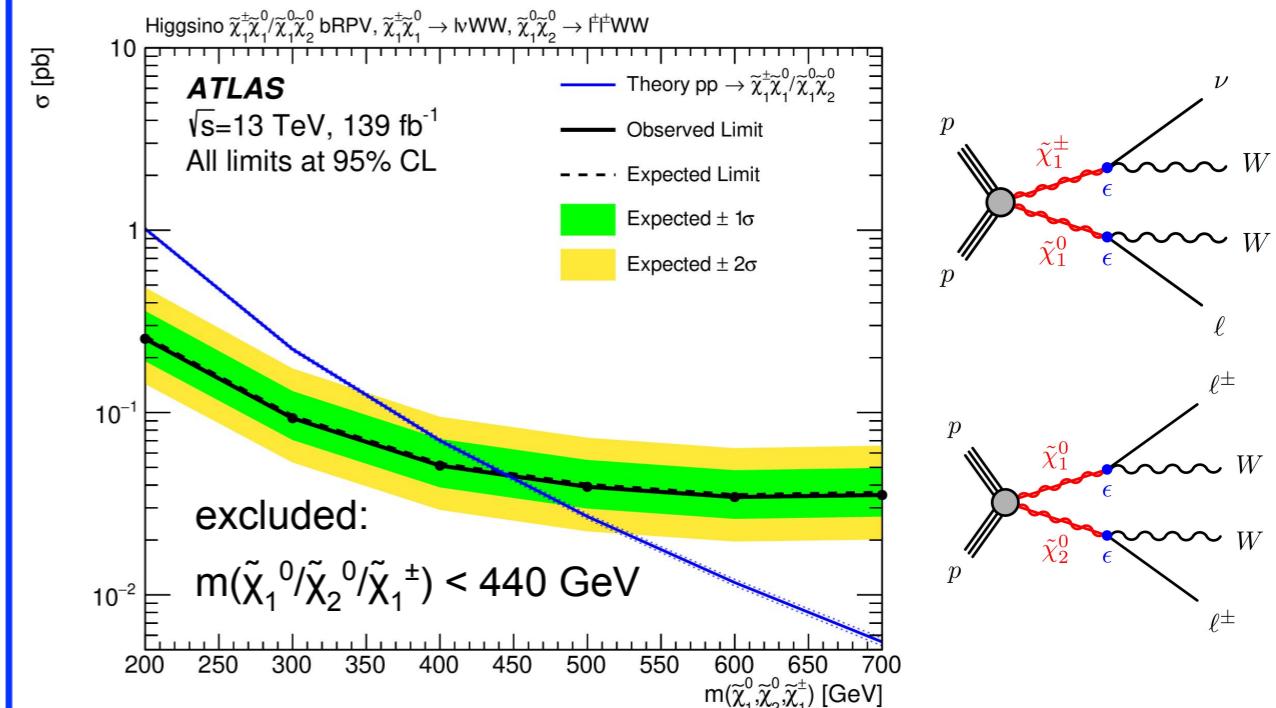
AITANA

SUSY searches

- Peak or edge in $m_{\ell\ell}$ distribution
- Exclude regions of SUSY mass parameter space by “stitching” SRs
- Also considered electroweakino production



- Final state with 2 same-sign or 3 leptons
- **Bilinear R-parity violation (bRPV)**
- Higgsino lightest SUSY partner
- First constraints on bRPV in electroweak production!



- IFIC: Analysis contact & contact editor

ATLAS, JHEP 11 (2023) 150

NA64: Latest results on Light Dark Matter

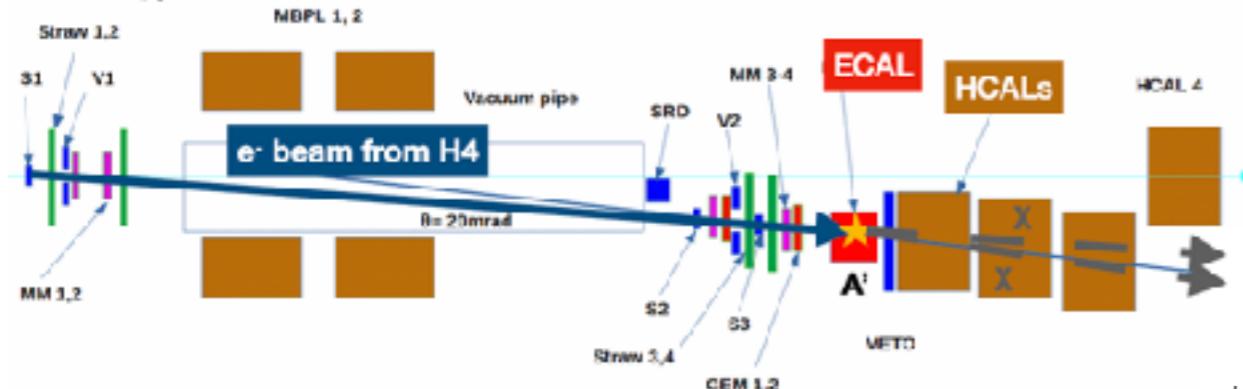
Editors' Suggestion



Search for Light Dark Matter with NA64 at CERN

Yu. M. Andreev *et al.* (NA64 Collaboration)

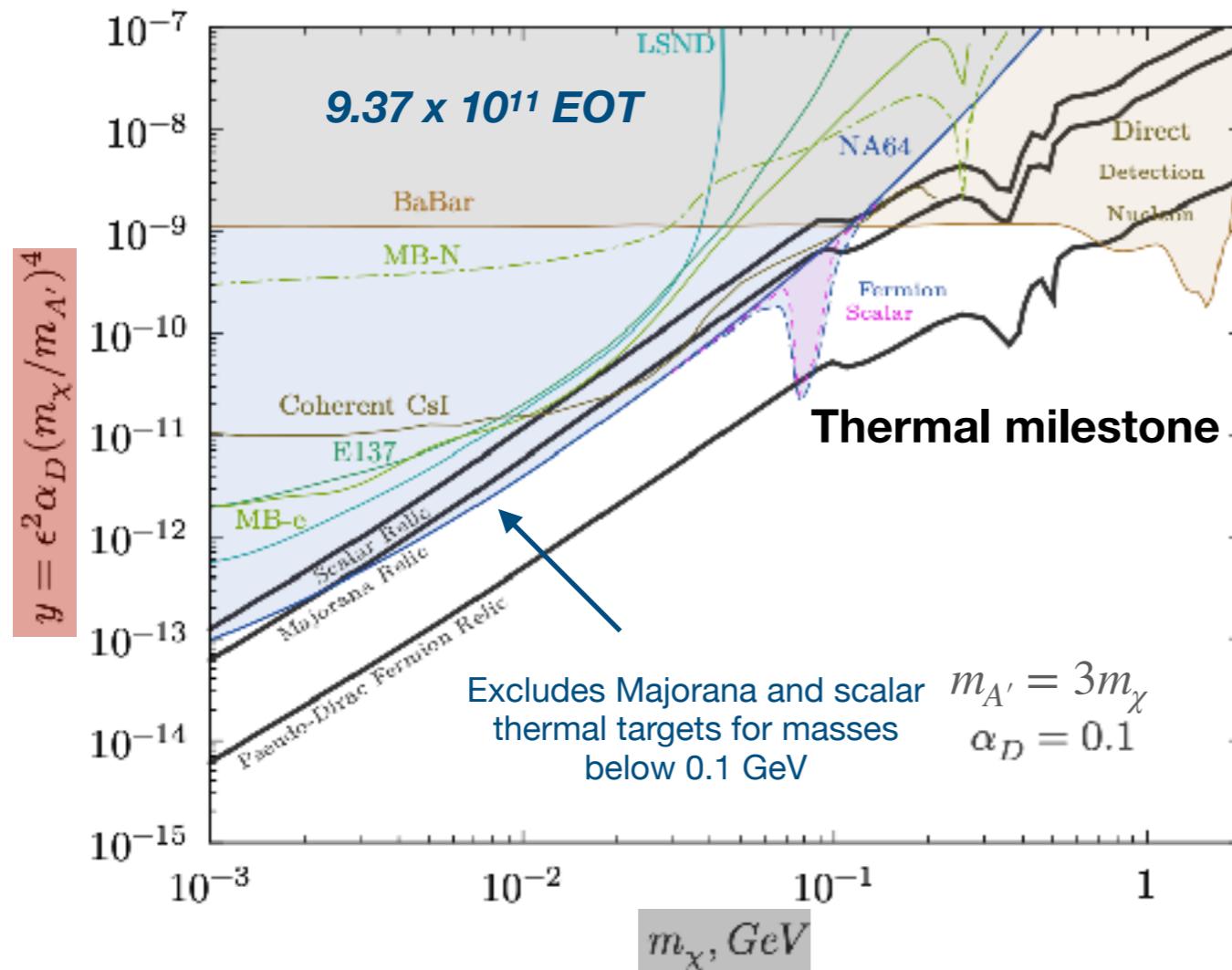
Phys. Rev. Lett. **131**, 161801 (2023) – Published 16 October 2023



An experiment at CERN performed by scattering 100 GeV electrons off a fixed target results in limits on thermal dark matter—that requires interactions mediated by hypothetical dark photons—in the MeV mass range.

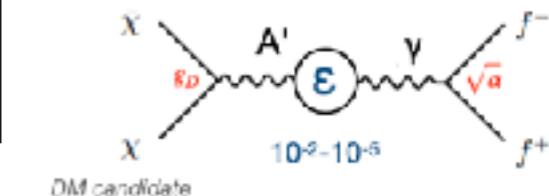
Show Abstract +

June 2023 CERN EP Newsletter: <https://ep-news.web.cern.ch/content/lighting-dark-na64-news-light-dark-matter-searches>



Thermal milestone

$$\Omega_\chi \propto \frac{1}{<\sigma v>} \sim \frac{m_\chi^2}{g_\gamma^4}$$



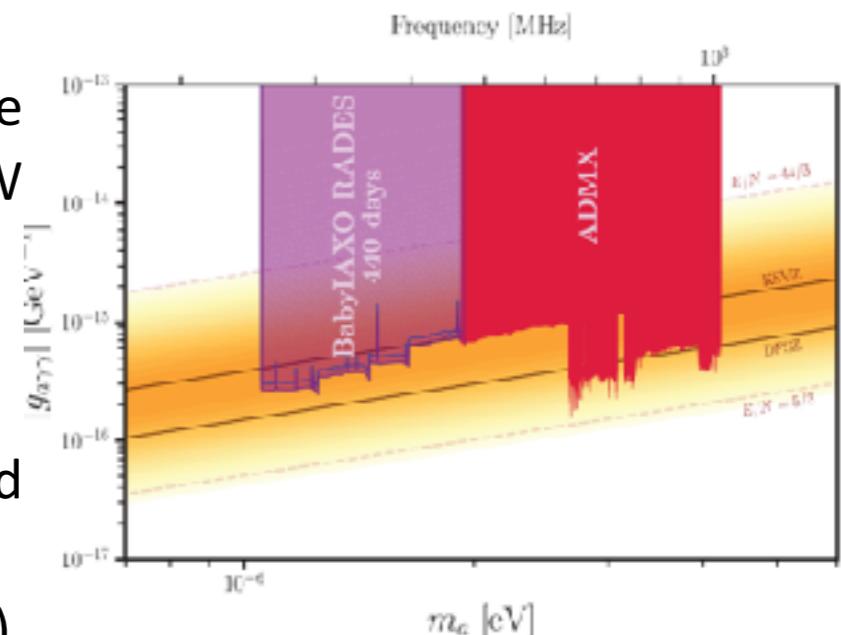
$$\sigma v(\chi\chi \rightarrow A' \rightarrow ff) \propto \epsilon^2 \alpha_D \frac{m_\chi^2}{m_{A'}^4} = \frac{y}{m_\chi^2}$$

A' $\rightarrow \chi\bar{\chi}$: NA64 Results start probing for the first time the LDM benchmark model parameter space

□ Detection of dark matter axions with microwave and millimeter-waves haloscopes:

- ✓ Development of full-wave advanced modal techniques for the rigorous and accurate analysis of the coupling between the axions and the resonant modes of the cavities (axion-photon coupling).
- ✓ Electromagnetic design of different resonators for detection in three frequency bands: UHF-VHF bands (300 MHz), X band (8 GHz) and W band (90 GHz)
- ✓ Participation in three international collaborations:
 - RADES (Relic Axion Detector Exploratory Setup) at CERN (X-band)
 - CADEx (Canfranc Axion Detection Experiment) at Underground Laboratory of Canfranc, Huesca (W band)
 - Baby-IAXO (International Axion Experiment) at DESY (UHF-VHF bands)

Baby-IAXO sensitivity



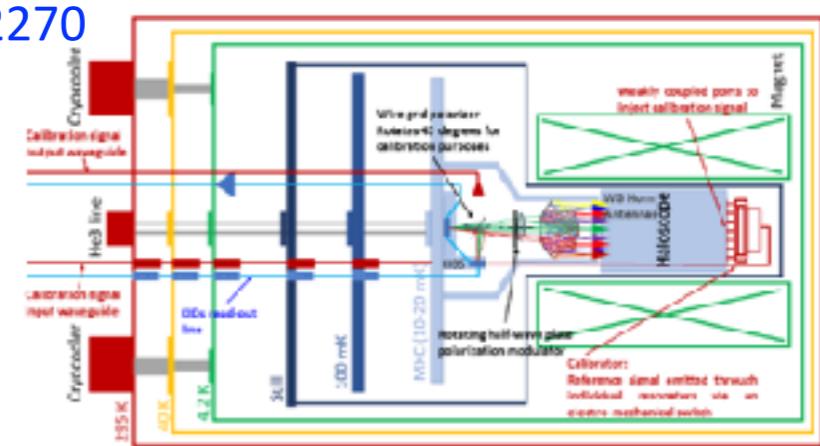
□ Detection of High-Frequency Gravitational Waves (HF-GW) in the microwave frequency band.

Some recent works by IFIC people:

[V. Domcke, C. Garcia-Cely, S. M. Lee and N. L. Rodd, arXiv:2306.03125](#)

[P. Navarro, B. Gimeno, J. Monzón-Cabrera, A. Díaz-Morillo and D. Blas, arXiv:2312.02270](#)

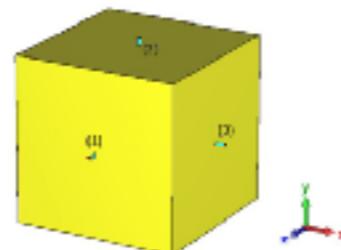
Scheme of the CADEx experiment



RADES cavity



Cubic resonator for HF-GWs exploration



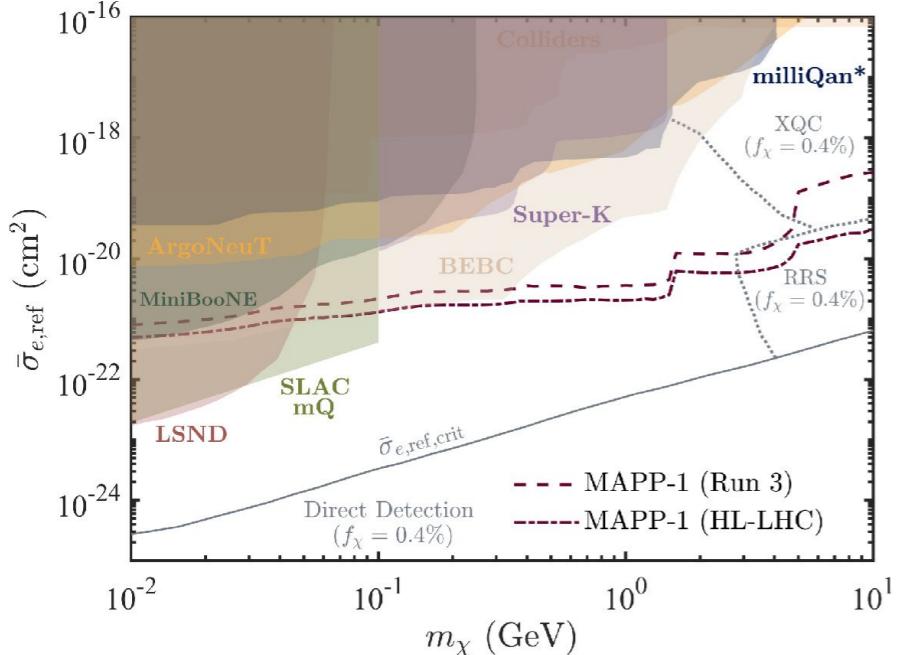
Other searches for exotic physics



MoEDAL

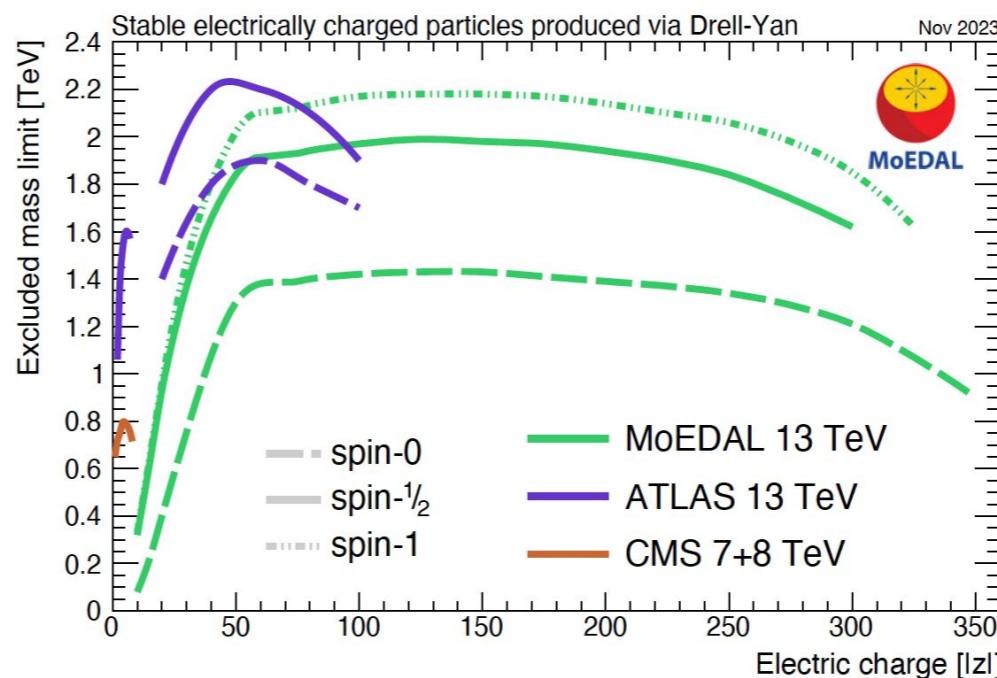
AITANA

- Prospects for detecting millicharged particles @ MAPP
- Strongly interacting DM



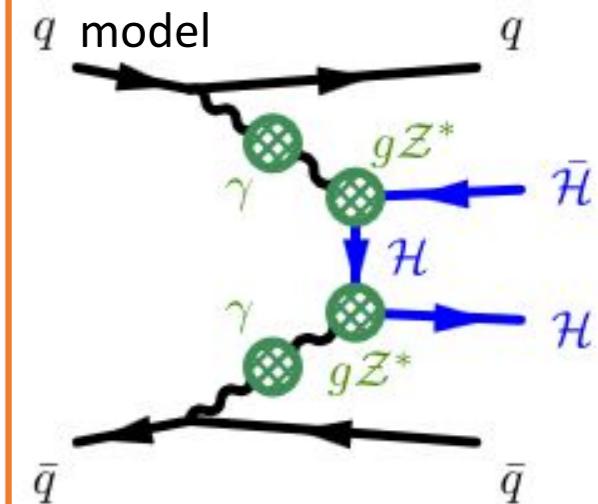
Mitsou, Montigny, Mukhopadhyay, Ouimet, Pinfold, Shaa, Staelens, [arXiv:2311.02185](https://arxiv.org/abs/2311.02185) [hep-ph], submitted to JHEP

- Full Run-2 MoEDAL analysis
- Magnetic monopole sensitivity to **10 Dirac charges**
- Electric charges up to **350e**
- **Highest charges probed in colliders!**

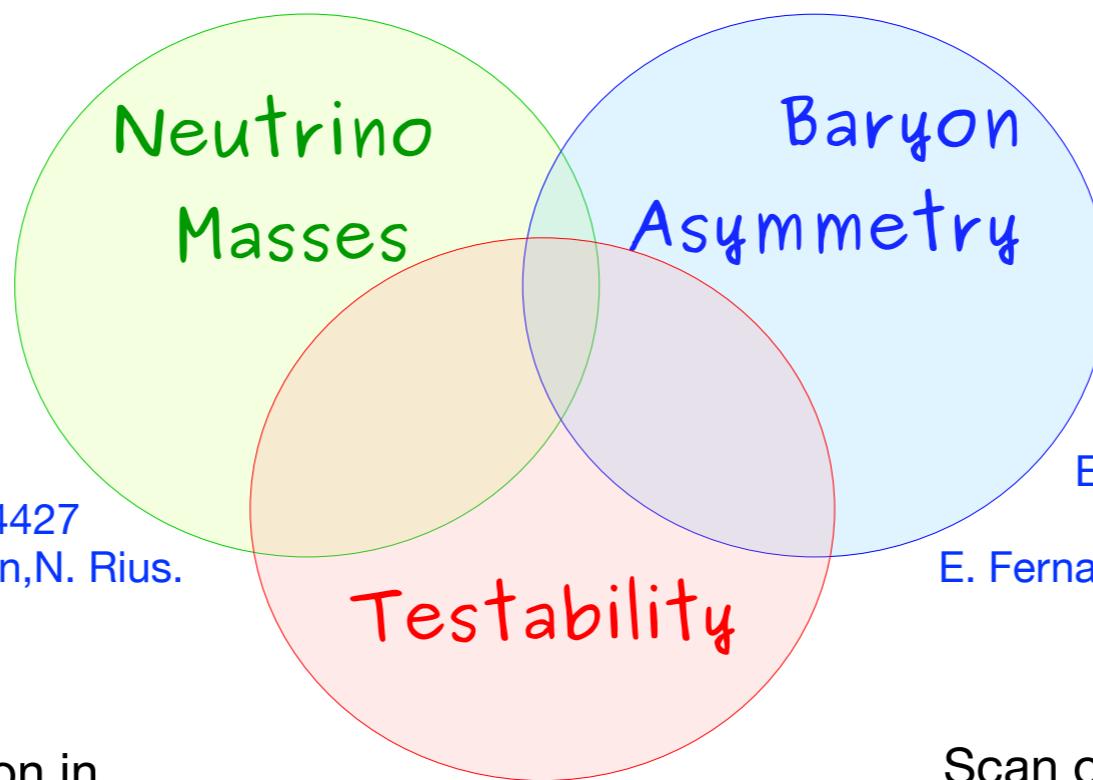


MoEDAL, [arXiv:2311.06509](https://arxiv.org/abs/2311.06509) [hep-ex], submitted to PRL

- Resummation Dyson-Schwinger on high-electric-charge particles to tackle large (non-perturbative) couplings to photons and Z bosons
- Implemented in MadGraph via UFO model



Alexandre, Mavromatos, Mitsou, Musumeci, [arXiv:2310.17452](https://arxiv.org/abs/2310.17452) [hep-ph], submitted to PRD
Musumeci, LHCP2023



JHEP 11 (2023), 153. ArXiv:2305.14427

S. Sandner, P. Hernandez, J Lopez-Pavon, N. Rius.

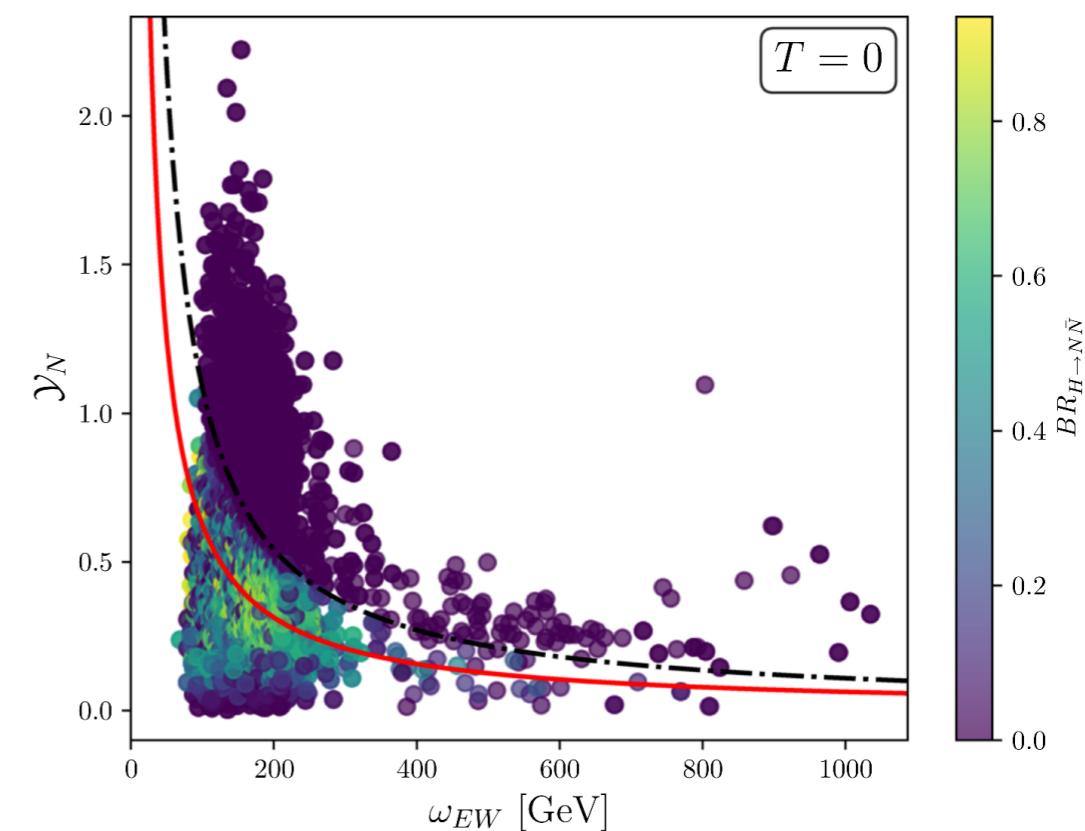
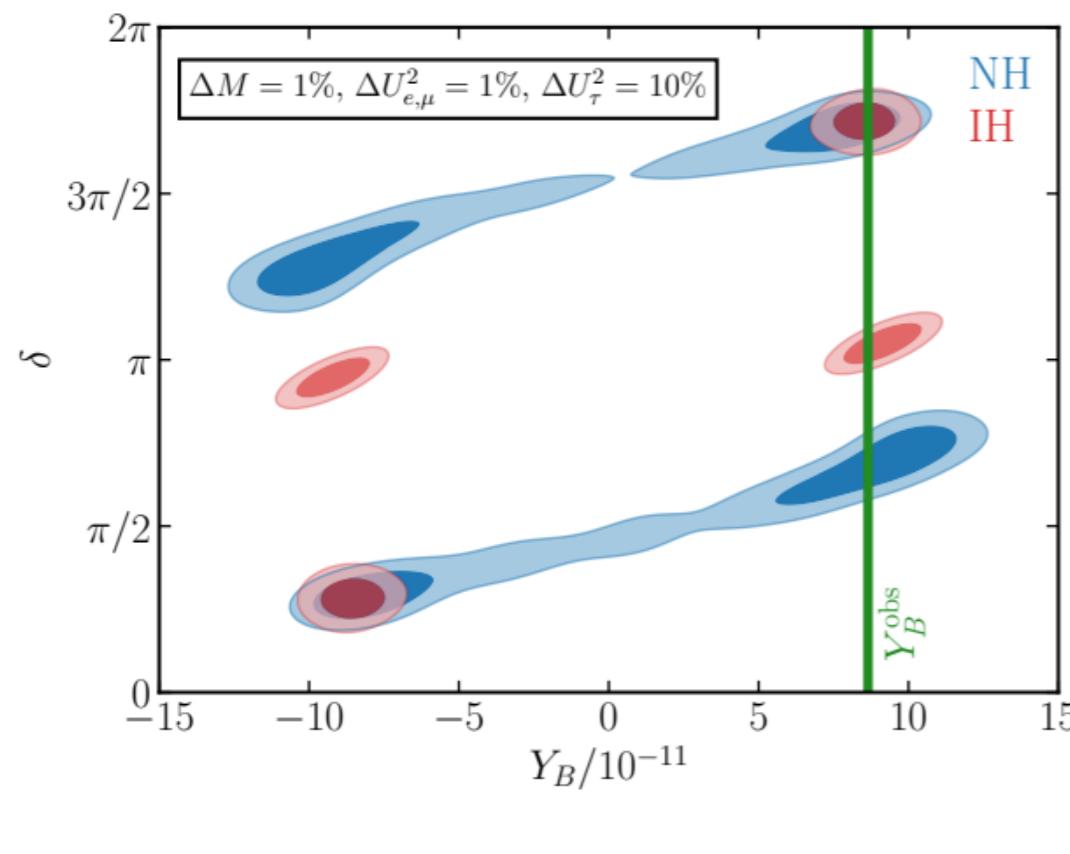
Eur. Phys. J. C 83 (2023) no.8, 715.

ArXiv:2210.16279

E. Fernandez-Martinez, J. Lopez-Pavon, J.M. No,
T. Ota and S. Rosauro-Alcaraz

Measurement of CP violation in neutrino oscillations, HNL masses and mixings with electron, muon and tau flavours can suffice to pin down matter-antimatter asymmetry from laboratory measurements.

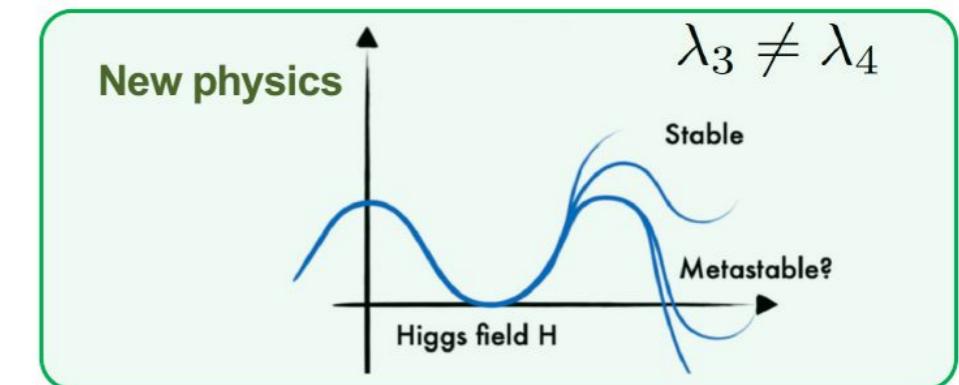
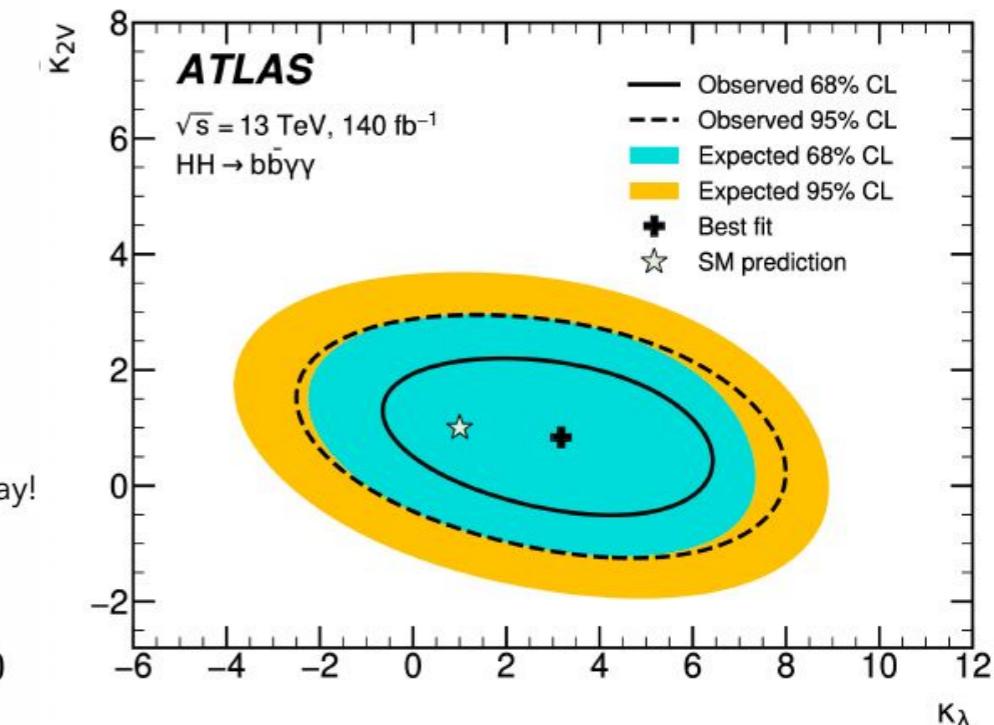
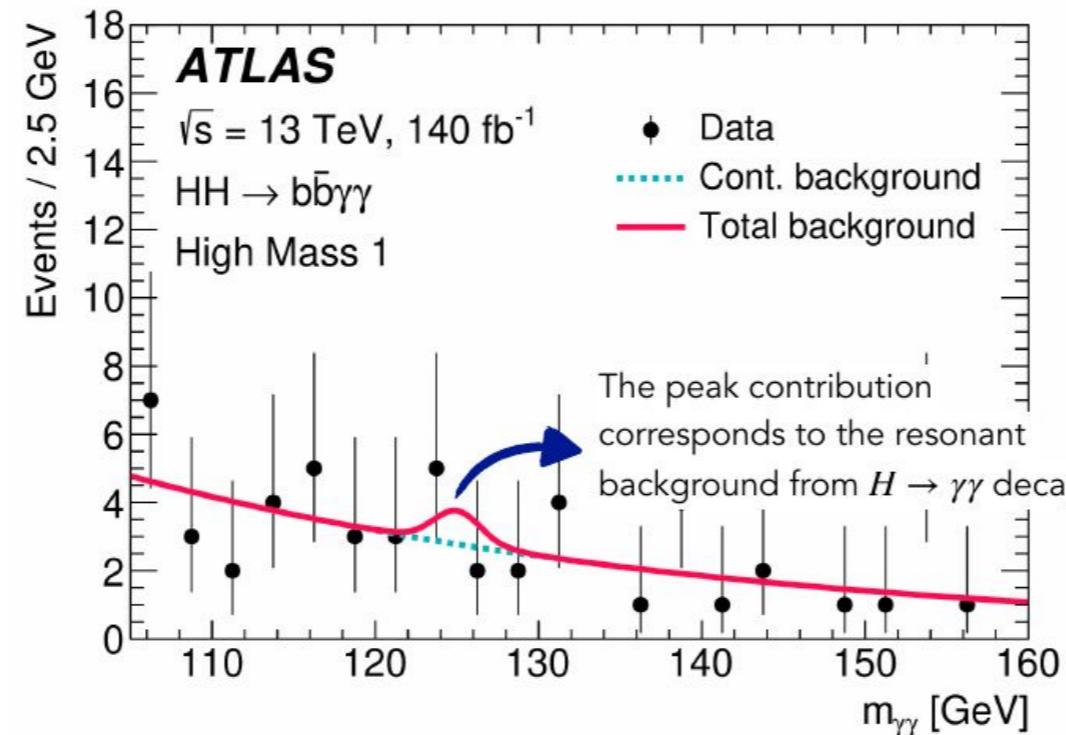
Scan of the parameter space of general singlet scalar BSM extension to identify regions which can lead to strong first-order phase transition, as required by electroweak baryogenesis, taking into account bubble nucleation as fundamental constraint.



Search for Higgs boson pair production in the $b\bar{b}\gamma\gamma$ final state with the ATLAS detector

<https://arxiv.org/abs/2310.12301>

- Shape of the Higgs potential together with m_H can have big consequences for the Universe.
- Searching for Higgs boson pair production constitutes the only direct probe to trilinear Higgs self coupling (k_λ), linked to the the Higgs potential shape.
- The $b\bar{b}\gamma\gamma$ is one of the golden channels for HH searches: Large $H \rightarrow b\bar{b}$ branching fraction + very clean di-photon signature.
- No deviations from SM found, but constraints on self coupling are getting closer to SM expectation.
- Effective Field Theory coefficients constrained too.

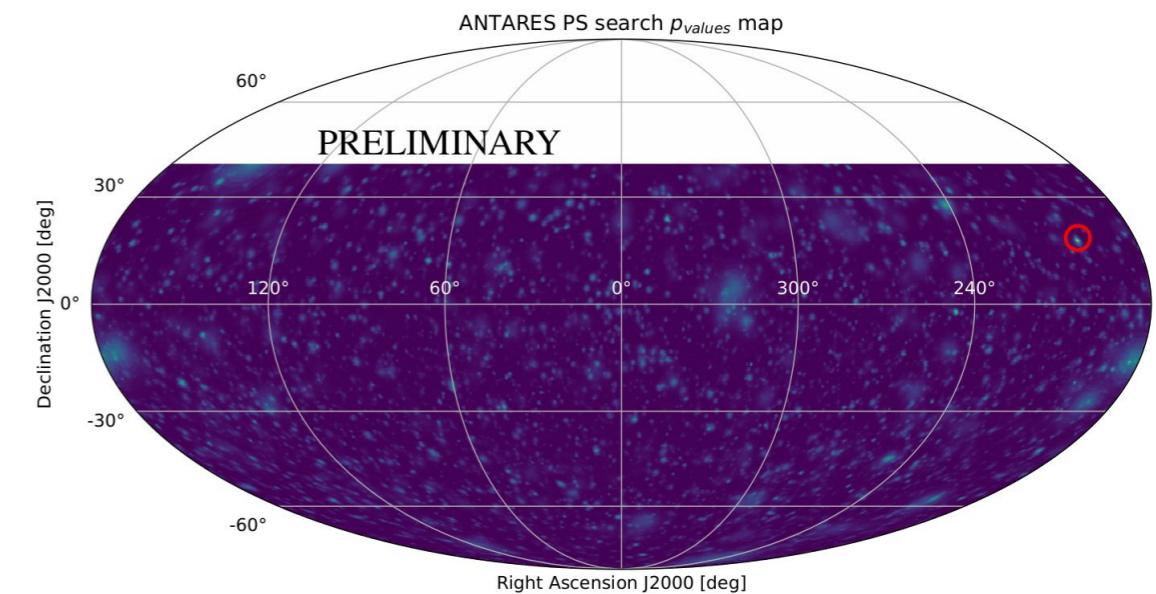


ANTARES: Search for cosmic neutrino sources

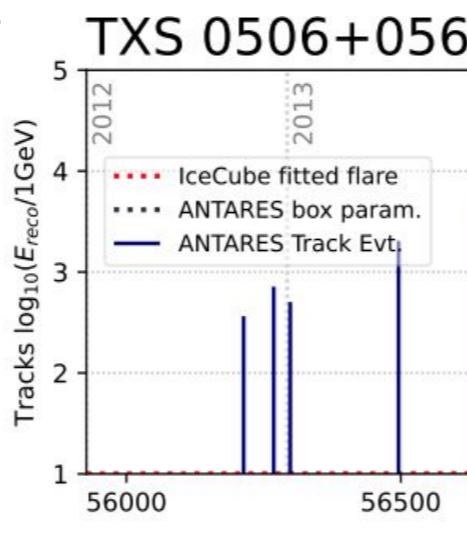
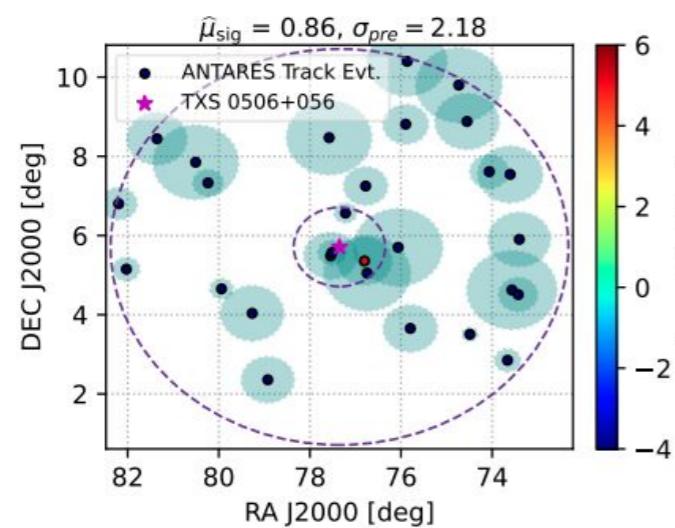


- ANTARES decommissioned in 2022: **final analyses being published**.
- IFIC led the **search for cosmic neutrino sources**, both time-integrated and time-dependent [S. Alves, F. Salesa, A. Sánchez].
- Best candidates compatible with background fluctuations.
- Some **interesting upper fluctuations** observed:
 - TXS0506+056 in the “top 5” most promising candidates (2.4σ).
 - Neutrino event compatible with TXS0506+056 detected in time coincidence with a flaring period reported by IceCube (2.2σ).

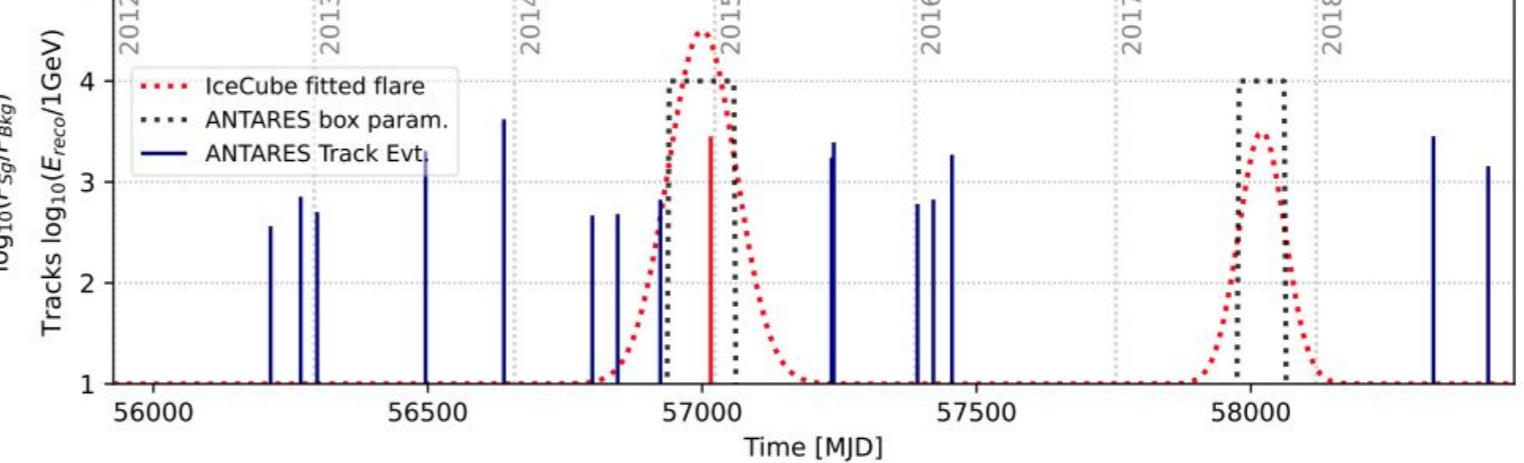
[PoS ICRC 2023 \(1128\)](#)



[PoS ICRC 2023 \(1480\)](#)



ANTARES PRELIMINARY

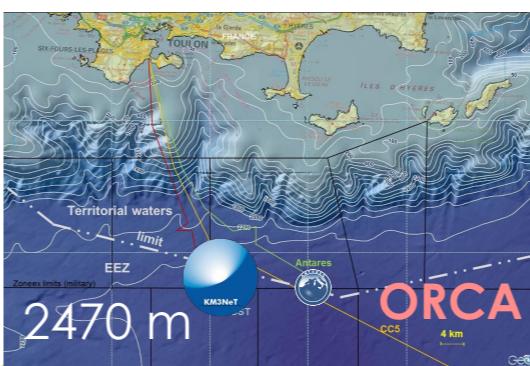
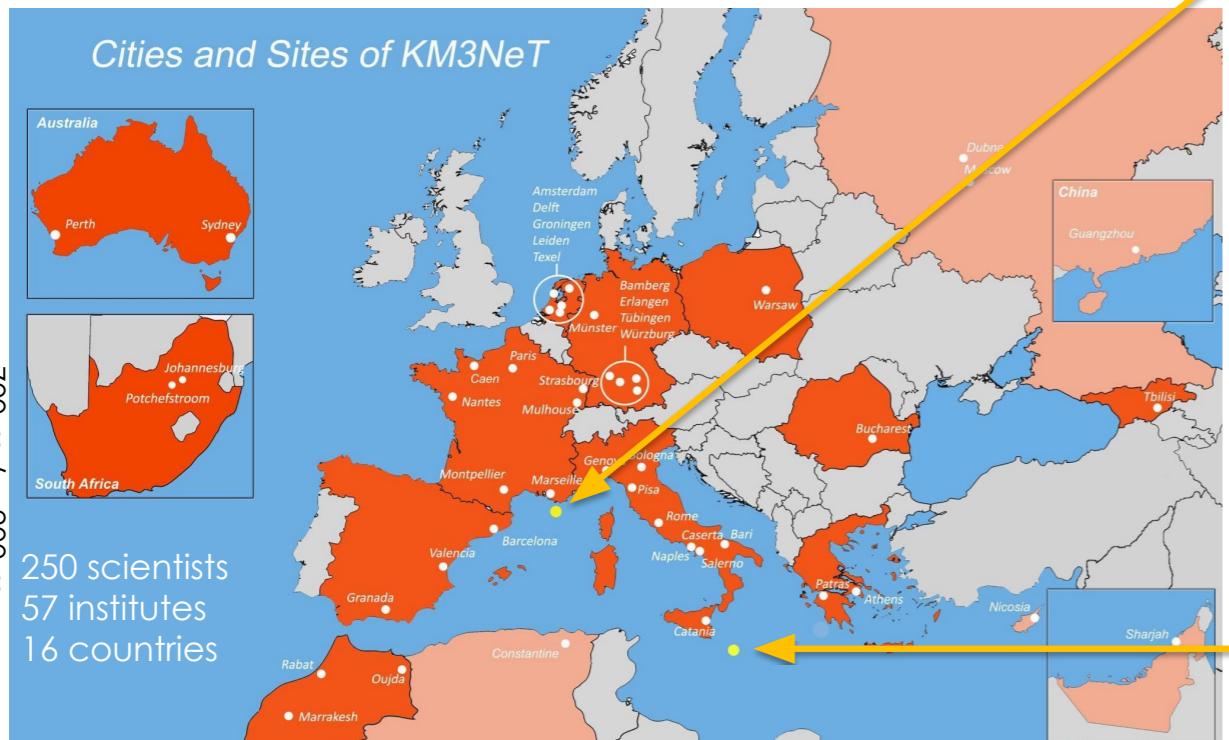




KM3NeT

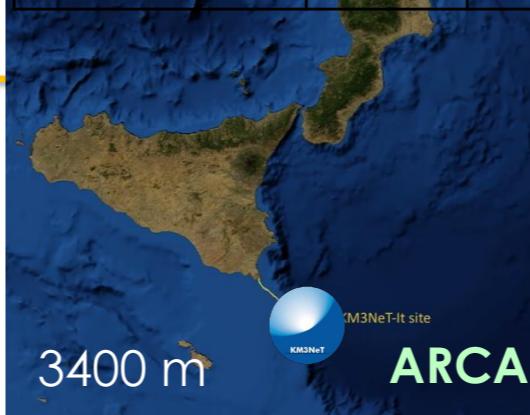


Multi-site, deep-sea infrastructure
Single collaboration, Single technology
Two outstanding physics cases

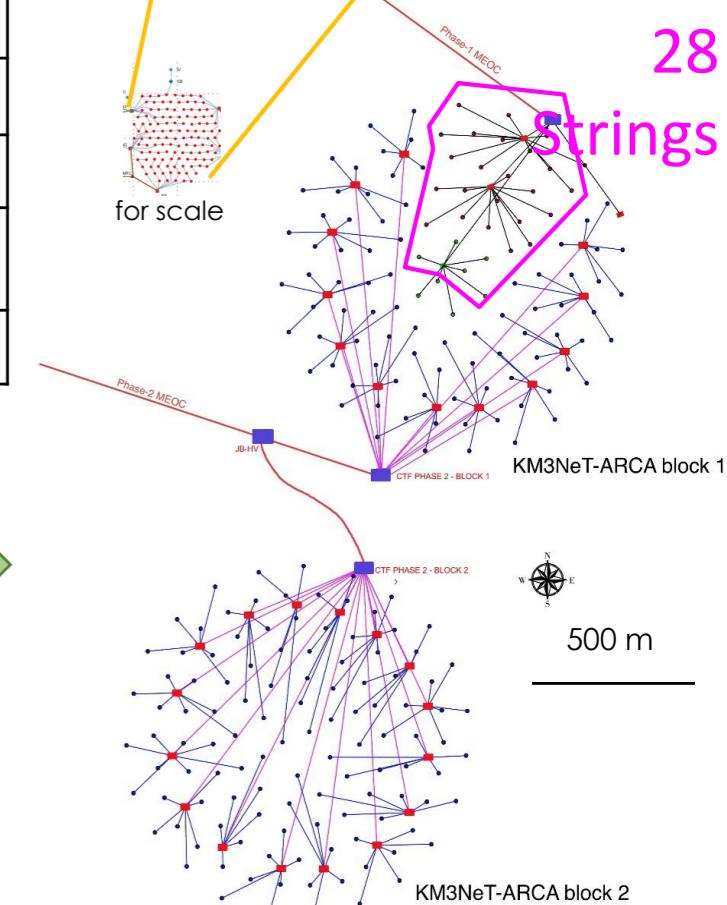
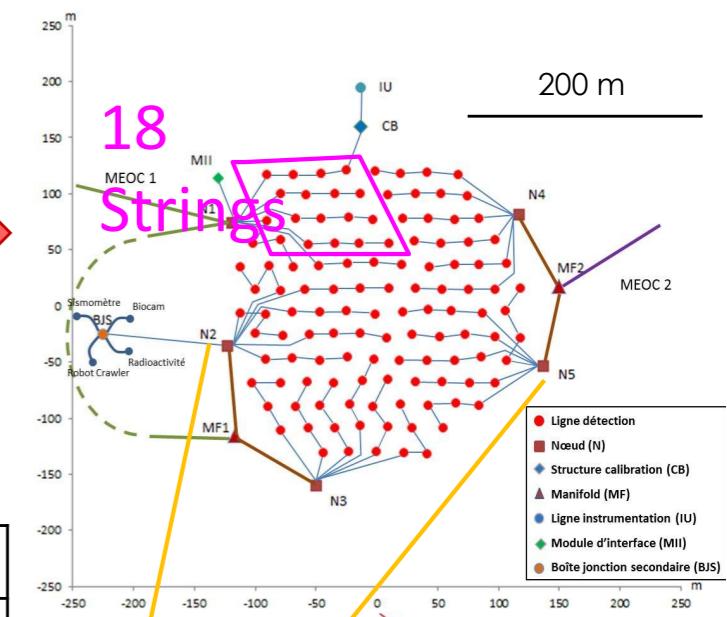


Oscillation Research with Cosmics in the Abyss

	ORCA	ARCA
Strings	115	115×2
String spacing	20 m	90 m
DOM spacing	9 m	36 m
Instrumented mass	7 Mton	500×2 Mton
Energy range	GeV	TeV – PeV



Astroparticle Research with Cosmics in the Abyss

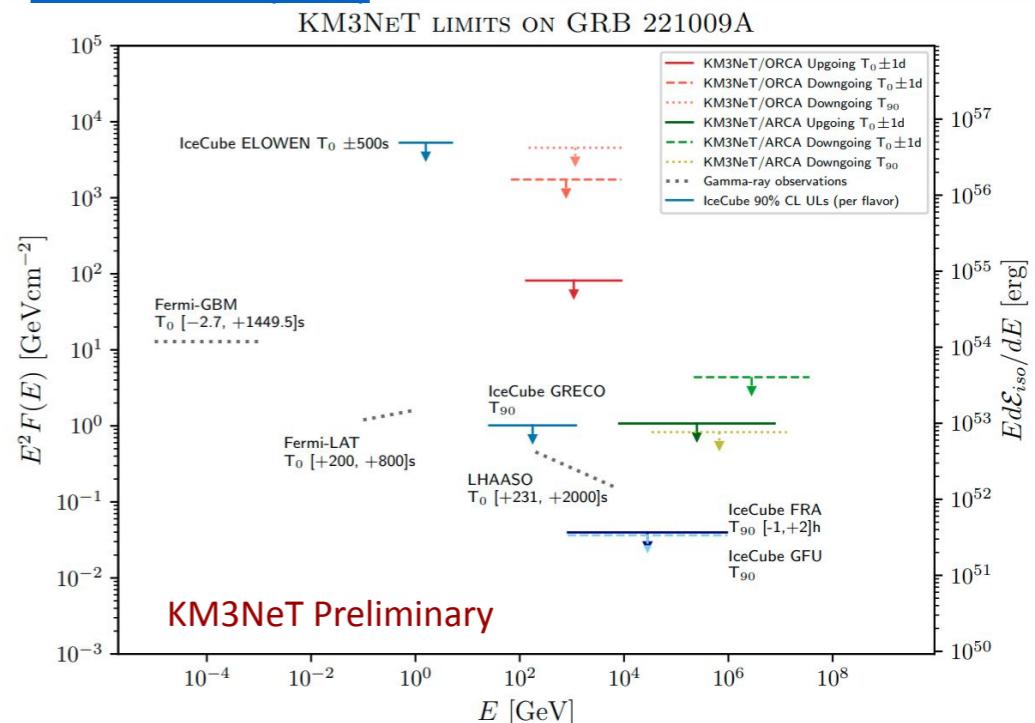


KM3NeT: Search for cosmic neutrino sources

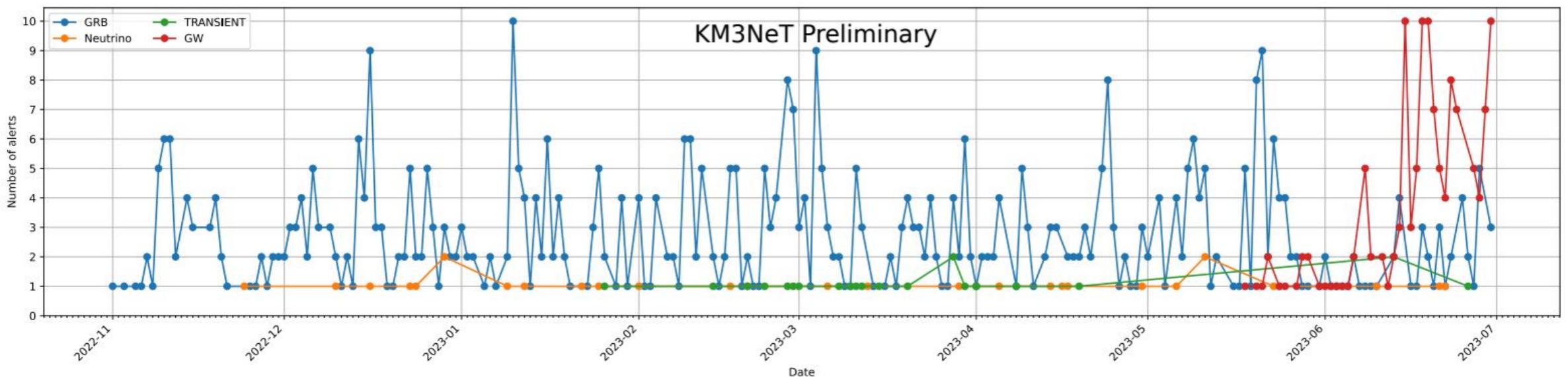


- ANTARES upper fluctuations mean **KM3NeT potential for discoveries**.
- IFIC deeply involved in **cosmic neutrino source searches**:
 - Time-dependent searches [**E. Pastor**, AS].
 - Transients searches: **GRB 221009A** (brightest GRB ever detected) [**J. Palacios**, FS, AS].
 - Real-time transient searches [**JP, V. Cecchini**, FS].
 - Development of the “Online Analysis Framework” for multi-messenger astronomy [**VC, JP**].

[PoS ICRC 2023 \(1503\)](#)



[PoS ICRC 2023 \(1521\)](#)



Using other messengers (gamma rays) for point-source searches

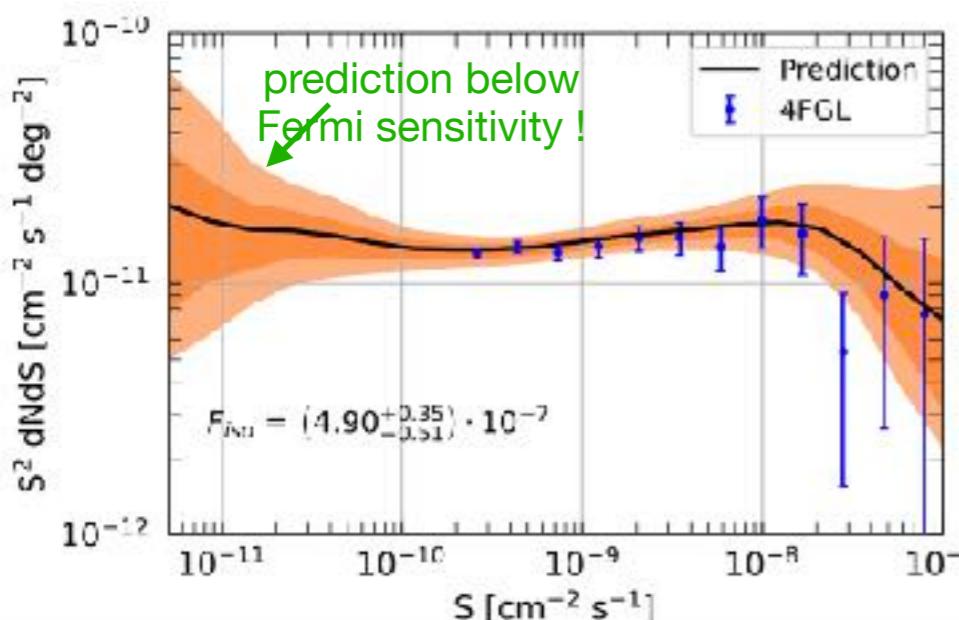
Deepening gamma-ray point-source catalogues with sub-threshold information

Aurelio Amerio, Francesca Calore (Annecy), Pasquale Serpico (Annecy), Bryan Zaldívar arXiv: 2306.16483

Idea: Providing a “probabilistic catalog” of gamma-ray sources going also below the current Fermi-LAT sensitivity

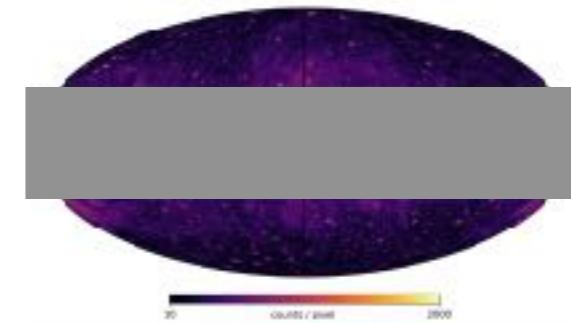
Methodology

- ▶ Gamma-ray sky map simulator using a source count distribution dN/dS given by a pre-trained ML model (Amerio et al, 2302.01947)

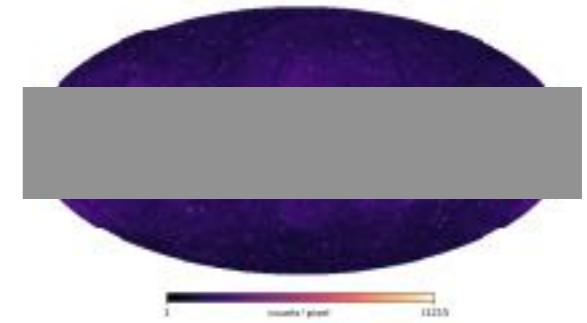


(Latest Fermi-LAT catalog contains ~ 6000 point sources)

- ▶ Comparing the measured map with simulated maps



A simulated sky



The measured sky (Fermi-LAT)

- ▶ $TS_i = \frac{(x_i - \lambda_i)^2}{\lambda_i}$ A distribution of test-statistic for pixel i , comparing simulated/ measured count x_i with diffuse model prediction λ_i

- ▶ TS-distribution of simulated sky and measured sky compatible up to a certain threshold TS_* according to Kolmogorov-Smirnov test

- ▶ Fermi sky pixels having $TS > TS_*$ are candidates of new sources (even below Fermi sensitivity)

Results

Public code available on GitHub: delivers a source catalog on demand for a given confidence level and source reliability (QF)

α	N^{cat}	$N^{\text{sim}}(QF = 0.9)$	$N^{\text{sim}}(QF = 0.8)$	$N^{\text{sim}}(QF = 0.5)$
0.01	7408	4641	8111	10003
0.05	6692	2745	5790	9532
0.10	6549	1855	4063	8684

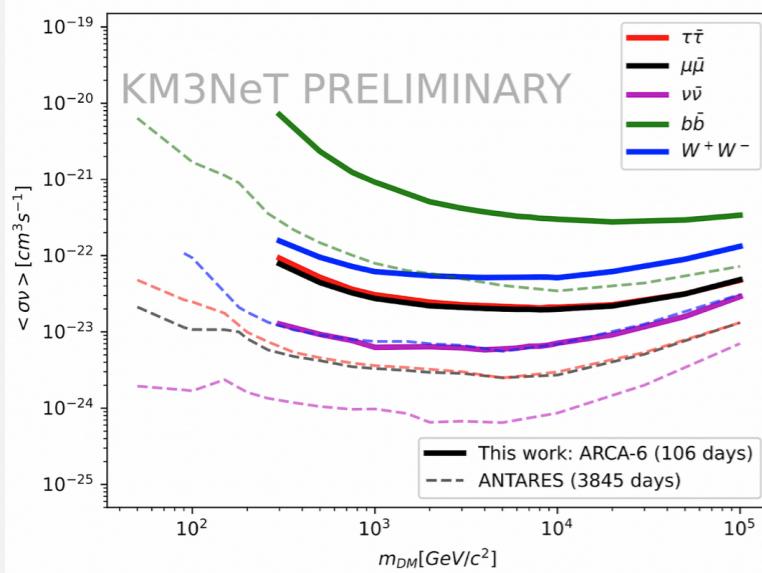
Dark matter indirect searches with neutrinos: from the galactic center, the Sun...



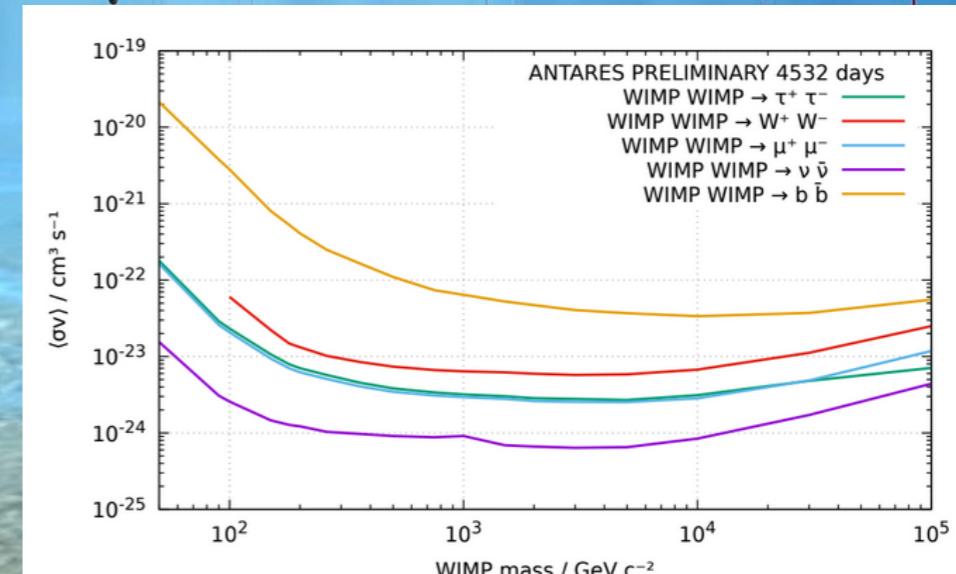
Physics beyond the Standard Model expectations: dark matter

ANTARES and first KM3NeT data investigated looking for indirect dark matter signatures

First KM3NeT data (8 and 21 lines)
[PoS(ICRC2023)1377]



Final ANTARES data set unblinded - test statistics is consistent with background hypothesis [PoS(ICRC2023)1375]

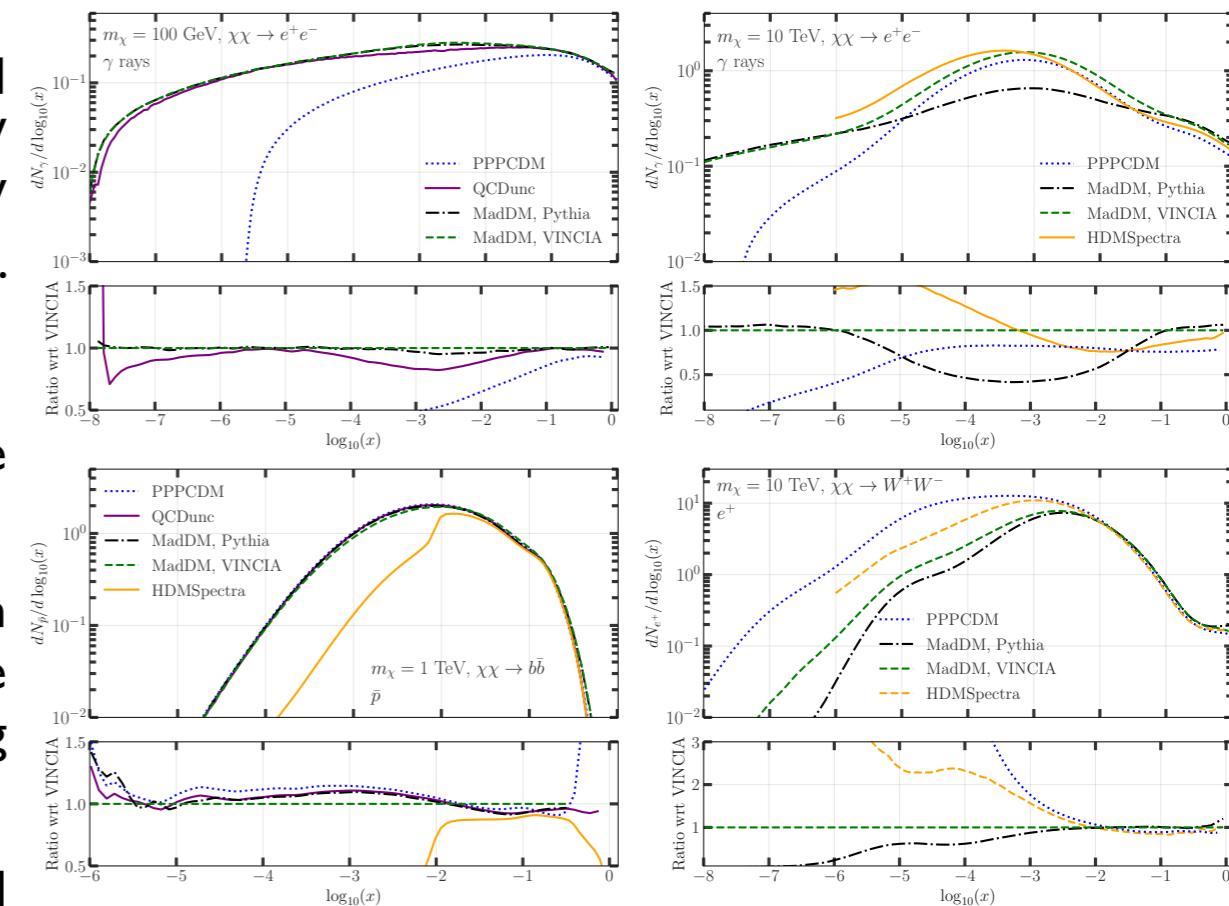


Also: combined search for dark matter joining neutrinos and γ rays ongoing

Novelties with respect to **PPPCDM** (standard in the community)

- Polarisation effects taken into account during the whole shower evolution.
- All the helicity information of the produced particles in DM annihilation/decay during the whole process. Therefore, EW corrections are properly taken in account where all the possible EW branchings are included (even those involving triple-gauge couplings).
- EW corrections are properly resummed through Sudakov factors.
- Decays of heavy resonances are interleaved with the rest of the shower evolution.
- Off-shell effects are taken into account for the annihilation channels: WW/ZZ/HZ. For this case, we produce the spectra for the four-body channels, e.g. DM DM \rightarrow WW \rightarrow 4f. Therefore spanning masses from 5 GeV to about the mass of the gauge boson.
- Spectra for two new channels (never considered before): HZ and gamma Z.
- Full loop effects for loop-induced annihilation channels: gamma gamma, gamma Z and gluon gluon.
- 5M annihilation events for each channel and each DM mass (reducing statistical uncertainties in the tails below 5%).
- The spectra calculated for DM masses between 5 GeV and 100 TeV.

Our results: **MadDM + VINCIA**



Large differences in spectra like gamma rays!

Tables available on <https://github.com/ajueid/CosmiXs>

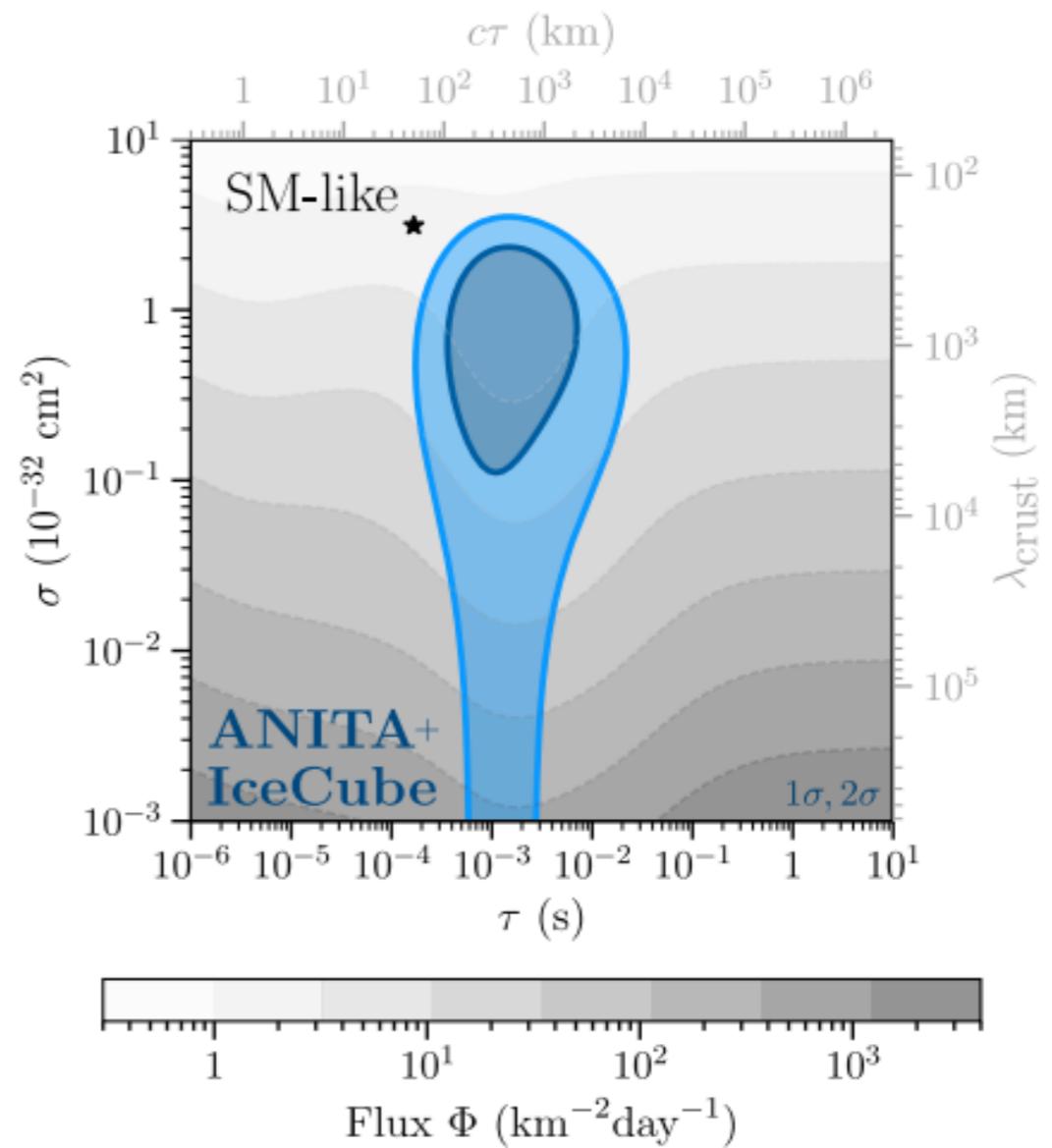
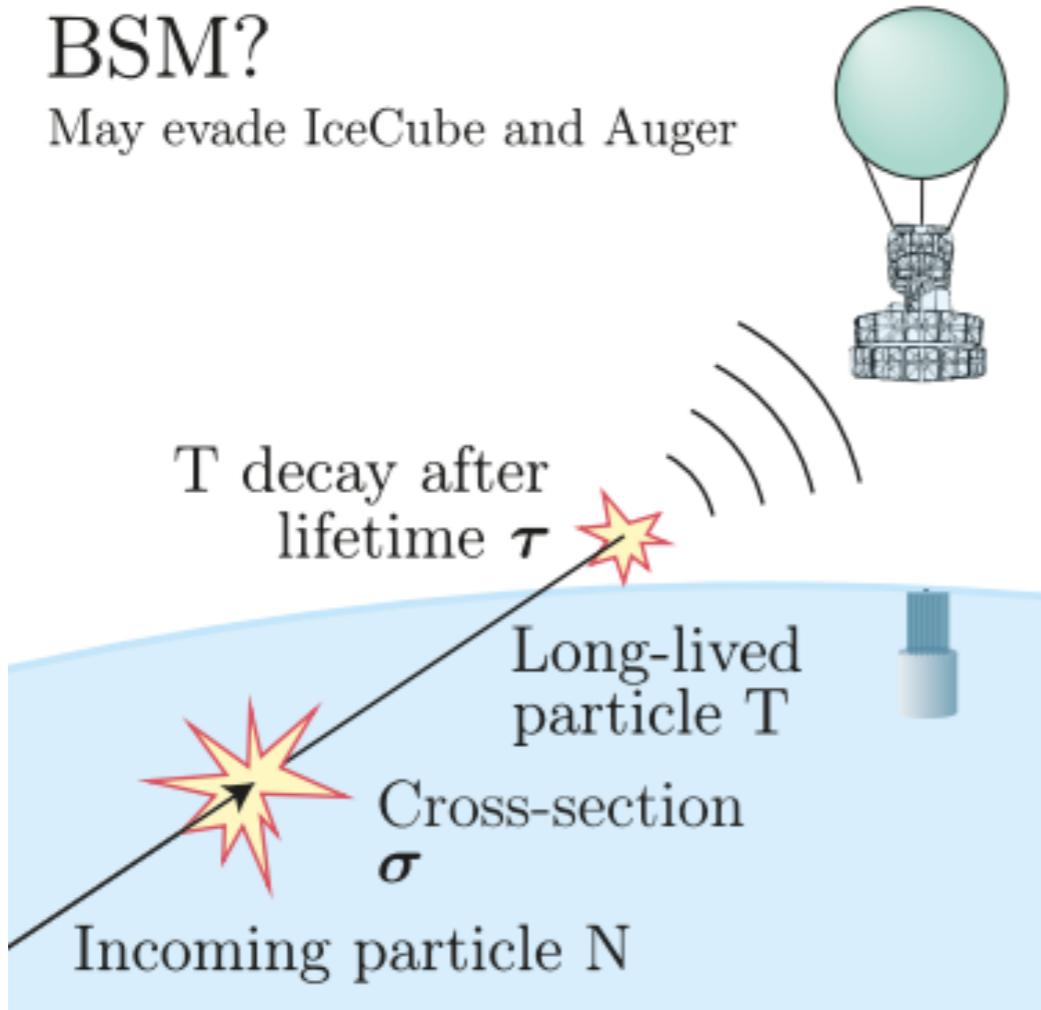


Ultra high-energy anomalous events: BSM?

JHEP 07 (2023), 005. arXiv:2305.03746

T. Bertolez-Martinez, C. Arguelles, I. Esteban, J. Lopez-Pavon, I. Martinez-Soler, J. Salvado.

ANITA-IV recently announced detection of new, unsettling UHE events. Their internal consistency and the implications of the lack of similar events in IceCube in SM-like and BSM scenarios is studied.



The analysis highlights the importance of simultaneous observations by high-energy optical neutrino telescopes and new UHE radio detectors to uncover cosmogenic neutrinos or discover new physics.

L4: Baryogenesis, dark matter and cosmic messengers



A bit of a hodgepodge of astro-particle related topics



A lot of activity at IFIC,
both from theory and experimental sides



MATNEA



Baby 





FELIZ AÑO NUEVO
2024