

## **Cosmic Ray Light Nuclei Isotopes Measured by the Alpha Magnetic Spectrometer on the ISS**

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Precise knowledge of the spectra of secondary cosmic-ray isotopes provides a powerful tool to constrain cosmic-ray propagation models. In particular,  $^3\text{He}$  isotopes in cosmic rays are produced by interactions of primary  $^4\text{He}$  with the interstellar medium; similarly, lithium and beryllium are mainly produced by the fragmentation of heavier primary cosmic-ray nuclei. Among light isotopes, the radioactive  $^{10}\text{Be}$ , which has half-life comparable to the cosmic-ray residence time in the Galaxy, provides crucial complementary information. We present precision measurements of the fluxes of helium isotopes,  $^3\text{He}$  and  $^4\text{He}$ , lithium isotopes,  $^6\text{Li}$  and  $^7\text{Li}$ , and beryllium isotopes,  $^7\text{Be}$ ,  $^9\text{Be}$  and  $^{10}\text{Be}$ , as well as the time dependence of the  $^3\text{He}/^4\text{He}$  flux ratio, based on data collected by AMS on the International Space Station.

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