

Producción y uso de haces radioactivos

Master Interuniversitario de Física Nuclear

Berta Rubio IFIC (CSIC-Univ. Valencia)

B. Rubio. Master FN, Valencia 2020

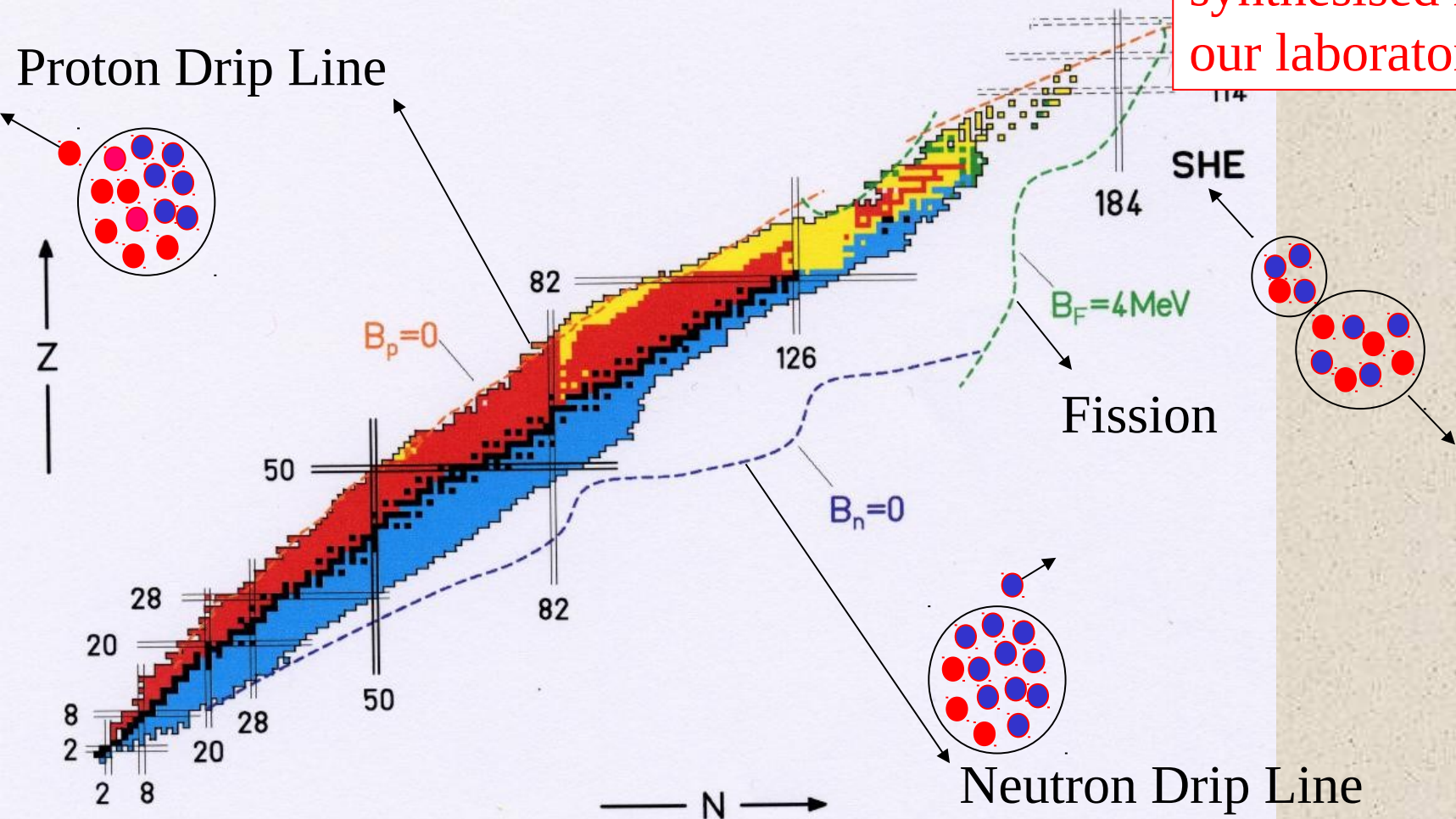
Porqué la producción de haces radioactivos ha marcado un antes y un después en la Física Nuclear

Reacciones nucleares y lo que nos enseñan de la estructura de los núcleos.

**Porqué necesitamos haces postacelerados
Porqué necesitamos alejarnos de la estabilidad.**

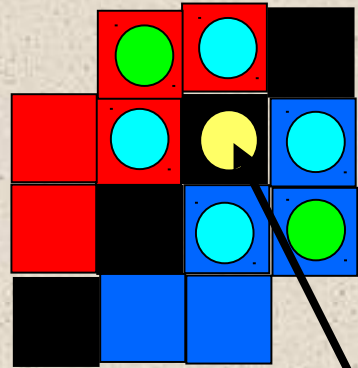
265 stable

About 3000 out
of 6000
synthesised in
our laboratories.



Close to the stability Direct Reactions with Light Projectiles, p , d , α , e^- , γ , etc.. give us detailed information on the structure of the target nucleus

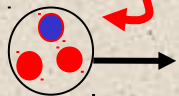
Z



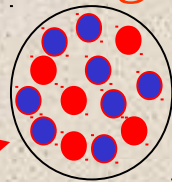
N

Target

interaction
(operator)



Projectile



Pick-up

Stripping

Charge-exchange

Coulomb Excitation

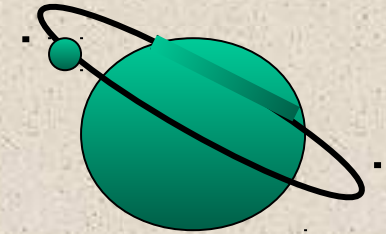
Elastic Scattering

Inelastic Scattering

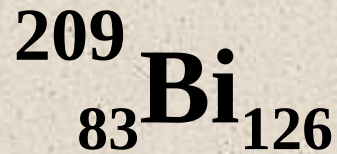
$|i\rangle \rightarrow$

$|f\rangle$

Let us use an illustrative example



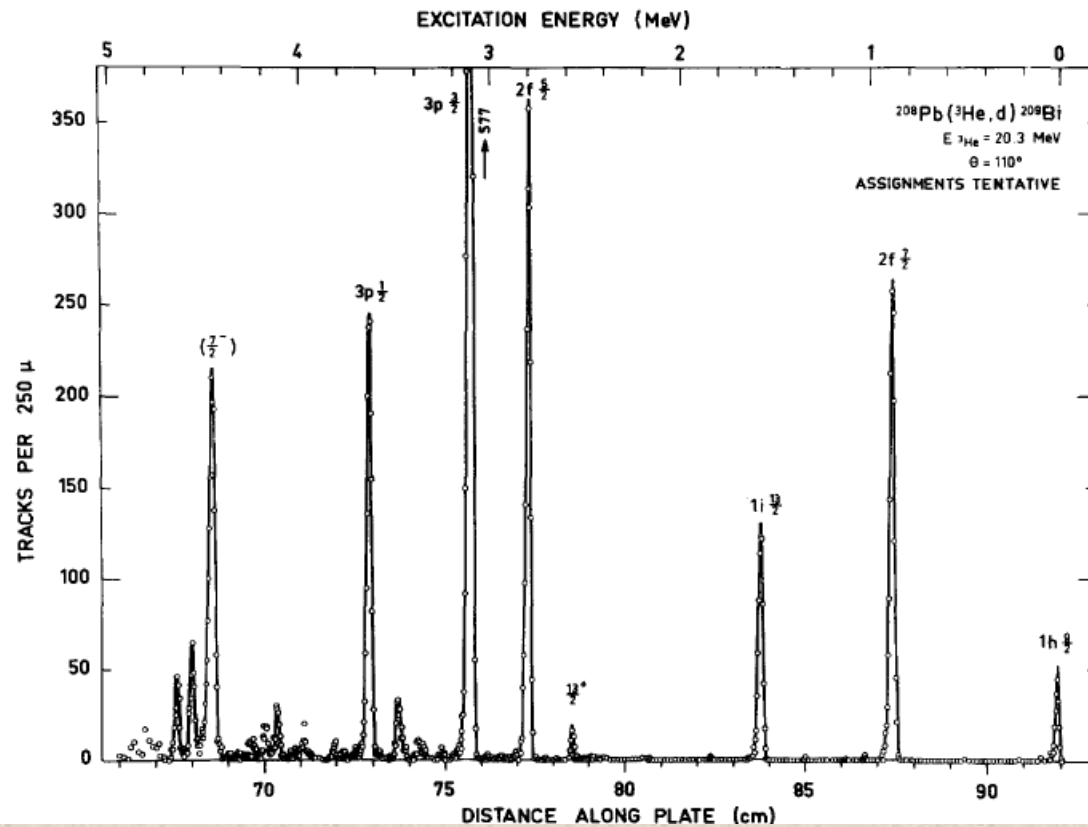
$$^{209}\text{Bi} = ^{208}\text{Pb} + p$$



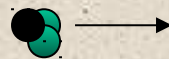
Closed shell + 1 p

Closed shell for neutrons

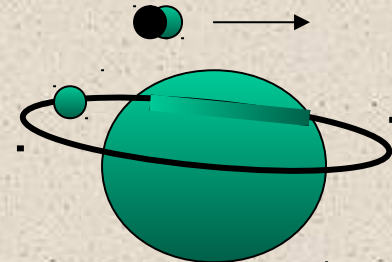
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Reaction:

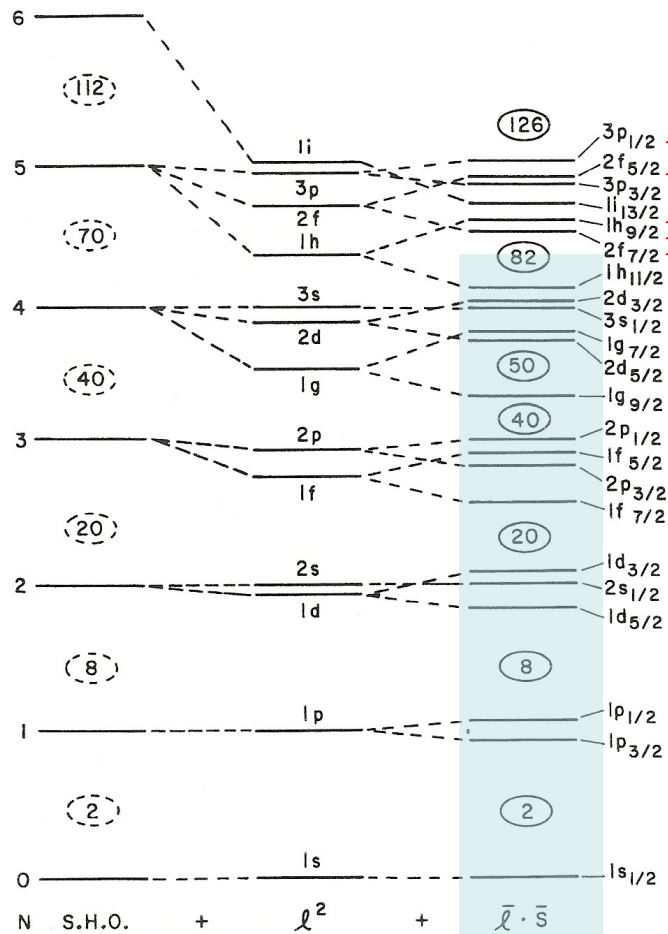
 $^{208}\text{Pb}(^3\text{He}, d)^{209}\text{Bi}$
 $E(^3\text{He}) = 20 \text{ MeV}$


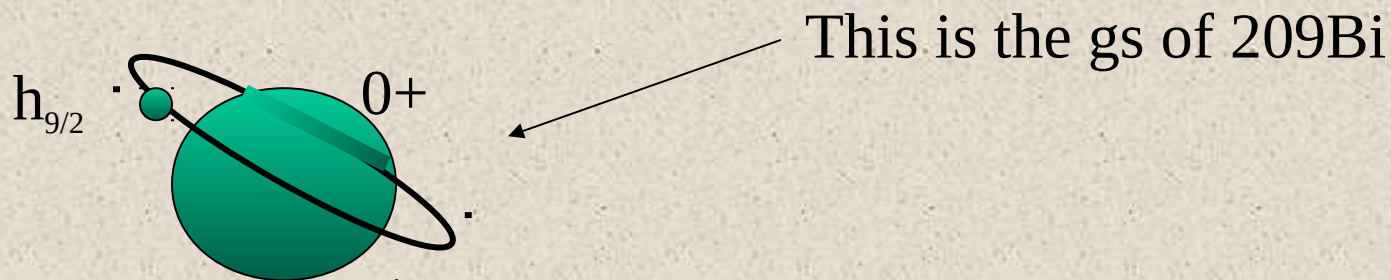
208Pb target



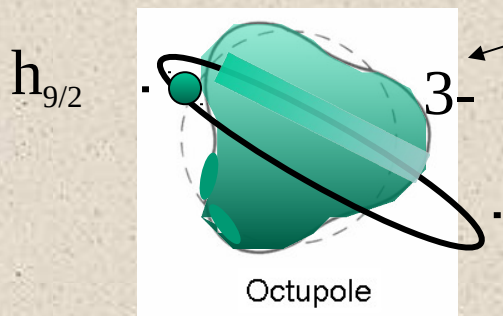
209Bi final nucleus

Shell Model and Residual Interactions





$^{209}\text{Bi} = ^{208}\text{Pb} + p$



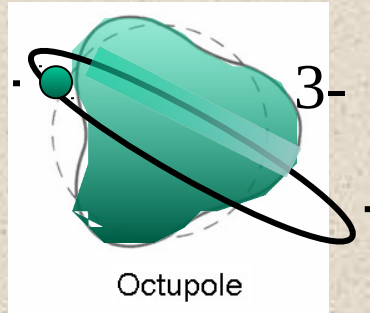
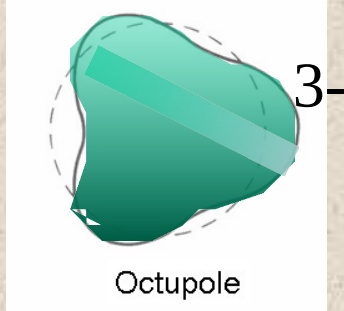
The first excited state

In ^{208}Pb is an octupole excitation, 3^- , at 2415 keV

One expects at the same energy in ^{209}Bi
A septuplet:

$$[3^-(^{208}\text{Pb}) \times h_{9/2}]_{3/2 \dots 15/2}^+$$

The appropriate reaction is inelastic scattering, for instance (d, d')



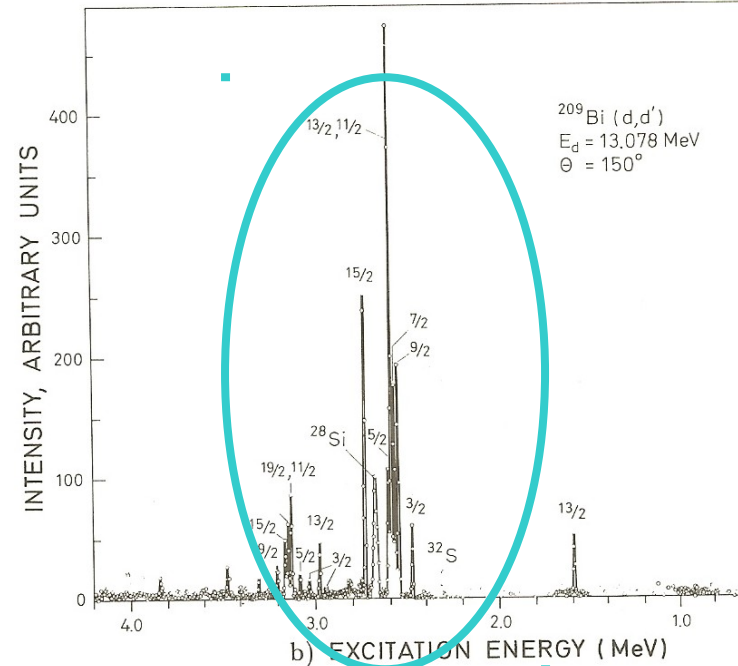
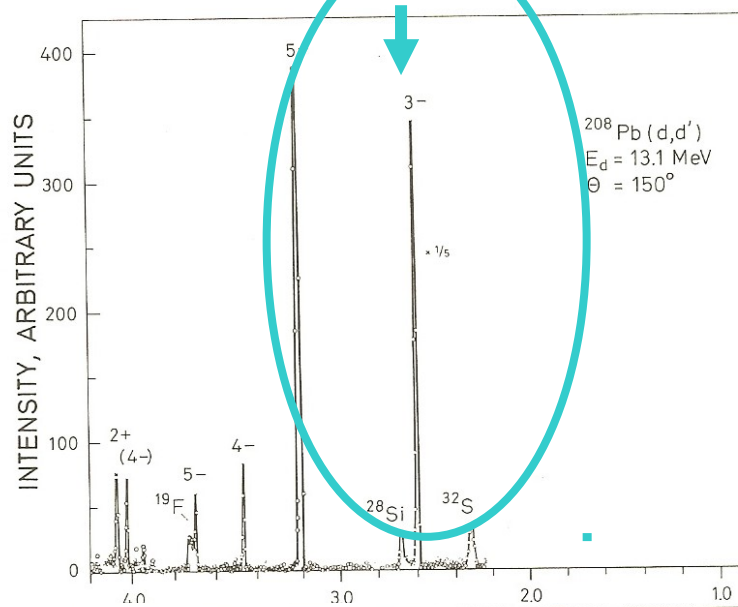
$h_{9/2}$

$209\text{Bi} = 208\text{Pb} + p$

572



VIBRATIONAL SPECTRA Ch. 6



39 2.6 46; 9...	α 6.040... γ 458; 571; 649; 73...	ε 5.783; 5.851... γ 674; 1363; 676...; g	α 6.264... γ	α 8.088; γ 540...
208 3 h	At 209 5.4 h	At 210 8.3 h	At 211 7.22 h	At 212 119 ms
ε 5.60; 560;	ε 5.647; γ 545; 782; 790...	ε 5.524; 5.442; 5.361... γ 1181; 245; 1483...	ε 5.867... γ (687...)	α 7.84; 7.90... γ 63... e-
207 5.84 h	Po 208 2.898 a	Po 209 102 a	Po 210 138.38 d	Po 211 25.2 s
ε 5.81... α 5.116... γ 992; 743; 912...; g	α 5.1152... γ (292; 571...) g	α 4.861... γ (895; 261; 263...)	α 5.28... γ (903); α < 0.005 α < 0.030; α < 0.002; α < 0.005	α 7.275; 8.883... γ 570; 1064... h
206 4 d	Bi 207 31.55 a	Bi 208 3.68 · 10 ⁵ a	Bi 209 100	Bi 210 0.10 ⁶ a
ε 5.81; 516; 37...	ε 5.70; 1064; γ 570; 1064; 1770...; α	ε 2615	α 0.011 + 0.023 α < 3E-7	α 4.946; 908... 266; 4... 0.054
205 10 ⁷ a	Pb 206 24.1	Pb 207 22.1	Pb 208 52.4	Pb 209 3.25
α 0.027	α 0.61	α 0.00023 α < 8E-6	β ⁻ 0.6 no γ	β ⁻ 0.6 no γ
204 8 a	Tl 205 70.48	Tl 206 3.7 m	Tl 207 4.33 s	Tl 208 3.05
ε	α 0.11	β ⁻ 1.5... γ (803...)	β ⁻ 1.4... γ (808...)	β ⁻ 1.8... γ 2615; 511; 860...
203 59 d	Hg 204 6.87	Hg 205 5.2 m	Hg 206 8.15 m	Hg 207 2.9
α 0.4	β ⁻ 1.5... γ 204...	β ⁻ 1.5... γ 305; 650... g	β ⁻ 1.8... γ 351; 9... 1637... m; g	β ⁻ 1.8... γ 351; 9... 1637... m; g
202 3 s	Au 203 60 s	Au 204 39.8 s	Au 205 31 s	
α 125; 204...	β ⁻ 2.0... γ 218; 44; 51; 318; 369	β ⁻ 437; 1511; 692; 723; 1392...	β ⁻ 379; 467; 946; 813...	
201 5 m	Pt 202 ~43.6 h			
β ⁻				

One could go to more sophisticated excitation modes such as

An isovector dipole resonance (E1)

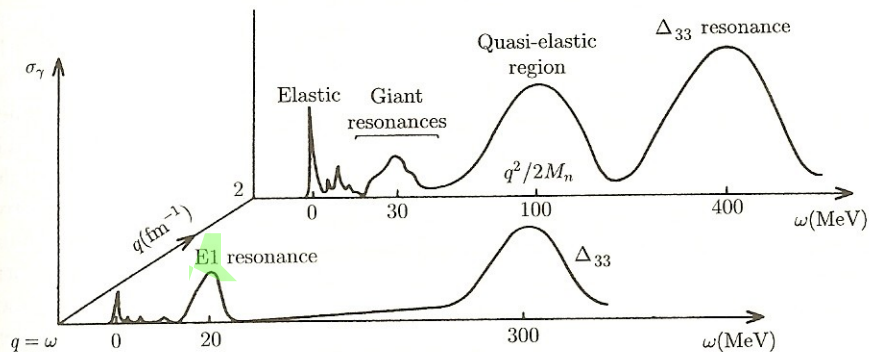
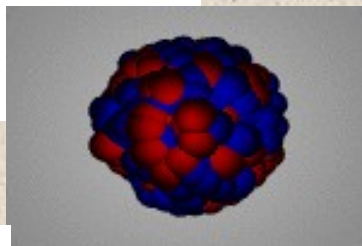
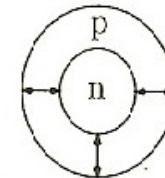


FIG. 1.1. Characteristic response (cross section σ) of an atomic nucleus as a function of the energy transfer ω and momentum transfer q . The lower curve is for photon absorption, the upper one for particle scattering with $q = 2 \text{ fm}^{-1} \neq \omega$.

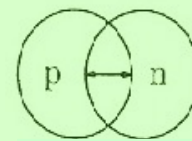


Isovector

$\Delta T = 0 \ 1$

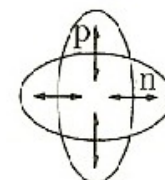


IVGMR



IVGDR

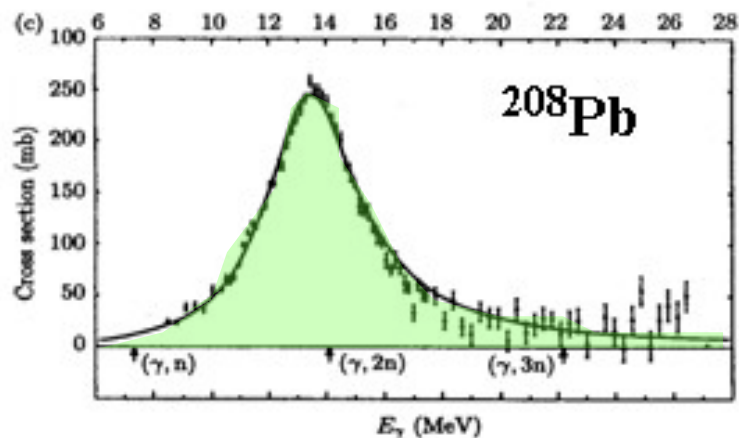
E1



IVGQR

$\Delta T = 1$

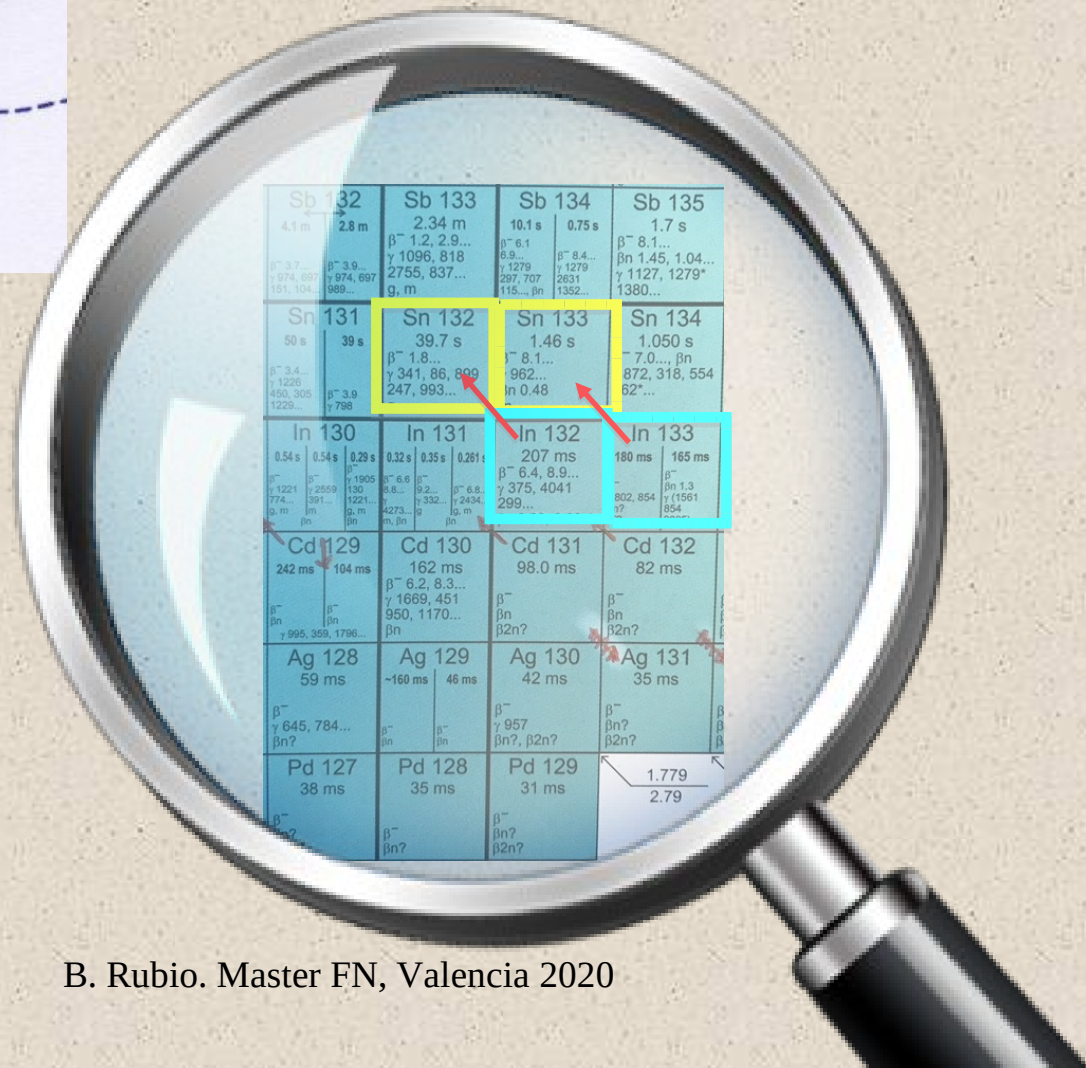
$\Delta S = 0$



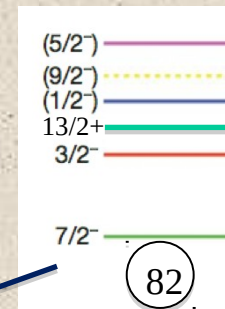
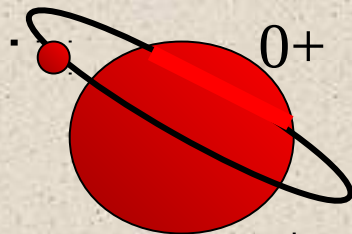
Qué ha sido necesario para poder hacer estos experimentos

1. Definir la reacción nuclear de interés.
2. Blancos estables: ^{208}Pb y ^{209}Bi en este caso
2. Un proyectil “sin estructura”
3. Un proyectil ligero (struggling)

Vamos a intentar aproximarnos a otro núcleo doblemente mágico:
 ^{132}Sn ($Z=50$, $N=82$)



Hasta qué punto es $^{133}\text{Sn} = ^{132}\text{Sn}(\text{core}) + n$?



Sb 132 4.1 m β^- 3.7... γ 974, 697 151, 104...	Sb 133 2.34 m β^- 1.2, 2.9... γ 1096, 818 2755, 837... g, m	Sb 134 10.1 s β^- 6.1... γ 1279 297, 707 115... βn 1352...	Sb 135 1.7 s β^- 8.1... βn 1.45, 1.04... γ 1127, 1279* 1380...
Sn 131 50 s β^- 3.4... γ 1226 480, 305 1229...	Sn 132 39 s β^- 1.8... γ 341, 86, 899 247, 993... m	Sn 133 1.46 s β^- 8.1... γ 962... βn 0.48 m	Sn 134 1.050 s β^- 7.0... βn 0.72, 318, 554 962*... m
In 130 0.54 s β^- 1221 γ 774... g, m	In 131 0.29 s β^- 1905 γ 2559 130 1221... g, m	In 132 207 ms β^- 6.4, 8.9... γ 375, 4041 299... βn 0.23, 0.33	In 133 180 ms β^- 1.3 βn 1.3 (1561) 854 2005)
Cd 129 242 ms β^- βn	Cd 130 162 ms β^- 6.2, 8.3... γ 1669, 451 950, 1170... βn	Cd 131 98.0 ms β^- βn	Cd 132 82 ms β^- βn
Ag 128 59 ms β^- γ 645, 784... βn ?	Ag 129 ~160 ms β^- βn		
Pd 127 38 ms β^- βn ?	Pd 128 35 ms β^- βn ?		

$^{133}_{50}\text{Sn}_{83}^{-1}$

From ENSDF - Evaluated October 2010

$^{133}\text{In} \beta^-$ decay [1996Ho16,2000Ho32](#)

Type	Author	History	Citation
Full Evaluation	Yu. Khazov and A. Rodionov, F. G. Kondev		NDS 112, 855 (2011)

Parent: ^{133}In : $E=0.0$; $J^\pi=(9/2^+)$; $T_{1/2}=165$ ms 3; $Q(\beta^-)=12917$ SY; % β^- decay=100.0
[1995JoZZ](#), [1996Ho16](#), [2000Ho32](#): $^{133}\text{In} \beta^-$ [from $^{238}\text{U}(p,f)$, $E=1$ GeV]; measured E_γ , I_γ , $\gamma\gamma$ -, $n\gamma$ -, $\beta\gamma$ -c
 $(\beta n)\gamma(t)$. ISOLDE facility; plastic scin β detector, two liquid scin n detectors, two Ge γ -ray detectors

^{133}Sn Levels

$E(\text{level})^\dagger$	J^π
0.0	$7/2^-$
853.7 3	$3/2^-$
1560.9 5	$(9/2^-)$
2004.6 10	$(5/2^-)$

Vidas medias, $T_{1/2}$
 207 ms !
 180/165 ms !

B. Rubio. Master FN,



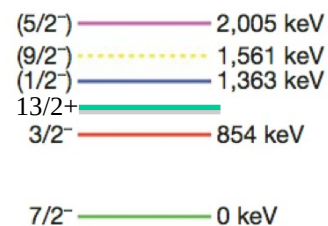
nature

Vol 465 | 27 May 2010 | doi:10.1038/nature09048

LETTERS

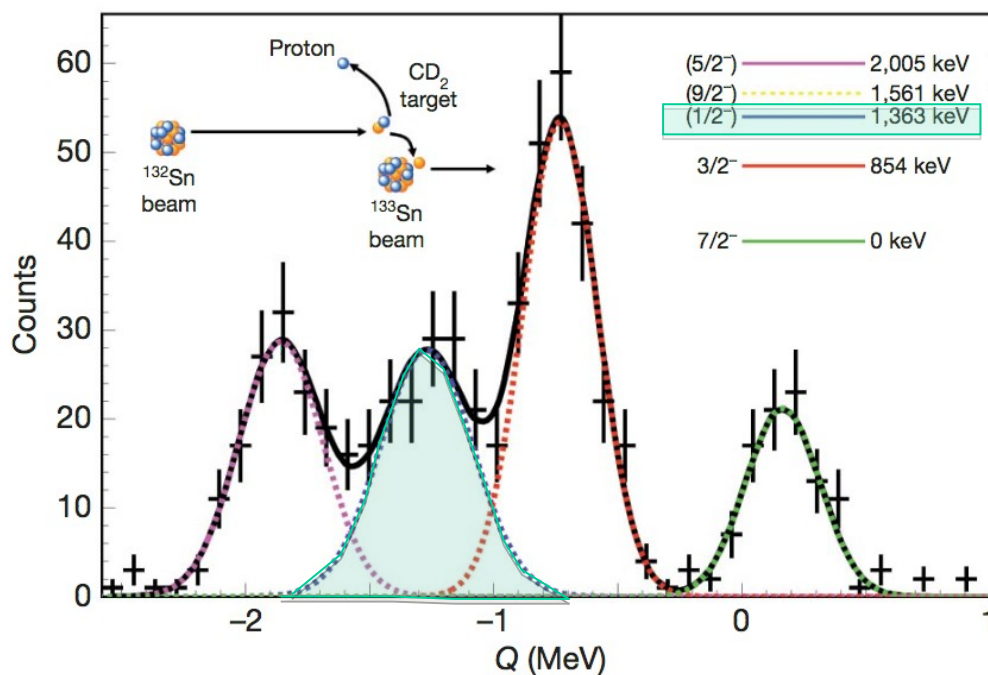
The magic nature of ^{132}Sn explored through the single-particle states of ^{133}Sn

K. L. Jones^{1,2}, A. S. Adekola³, D. W. Bardayan⁴, J. C. Blackmon⁴, K. Y. Chae¹, K. A. Chipps⁵, J. A. Cizewski², L. Erikson⁵, C. Harlin⁶, R. Hatarik⁷, R. Kapler¹, R. L. Kozub⁷, J. F. Liang⁴, R. Livesay², Z. Ma¹, B. H. Moazen¹, C. D. Nesaraja⁴, F. M. Nunes⁸, S. D. Pain², N. P. Patterson⁶, D. Shapira⁴, J. F. Shriner Jr⁷, M. S. Smith⁴, T. P. Swan^{2,6} & J. S. Thomas⁶



Año 2010!!

$^{132}\text{Sn}(d,p)^{133}\text{Sn}$



End of the story (so far):

PHYSICAL REVIEW C **99**, 024304 (2019)

β decay of ^{133}In : γ emission from neutron-unbound states in ^{133}Sn

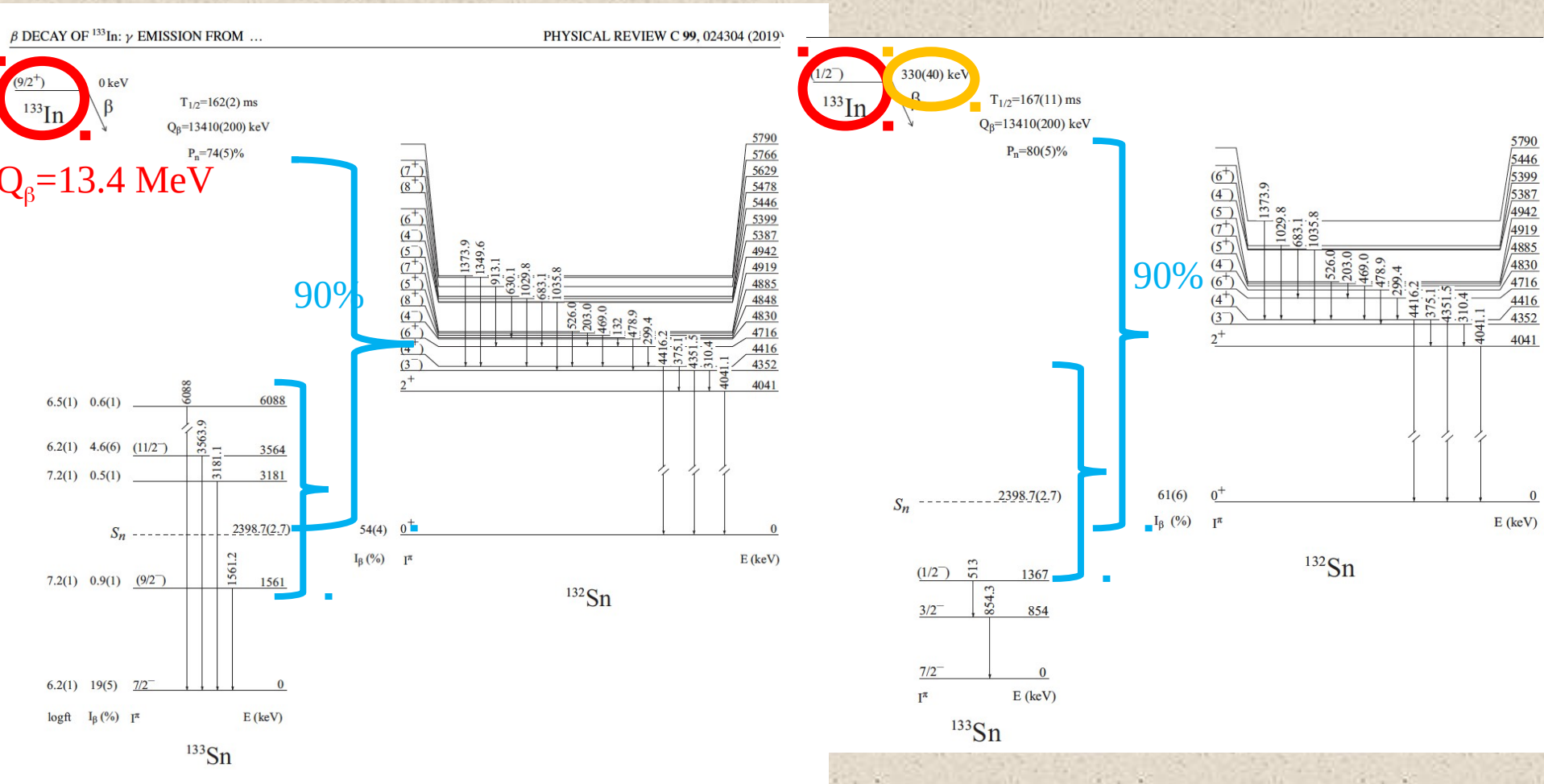
M. Piersa,^{1,*} A. Korgul,^{1,†} L. M. Fraile,^{2,3} J. Benito,² E. Adamska,¹ A. N. Andreyev,⁴ R. Álvarez-Rodríguez,⁵ A. E. Barzakh,⁶ G. Benzoni,⁷ T. Berry,⁸ M. J. G. Borge,^{3,9} M. Carmona,² K. Chrysalidis,³ J. G. Correia,^{3,10} C. Costache,¹¹ J. G. Cubiss,^{3,4} T. Day Goodacre,^{3,12} H. De Witte,¹³ D. V. Fedorov,⁶ V. N. Fedosseev,³ G. Fernández-Martínez,¹⁴ A. Fijałkowska,¹ M. Fila,¹ H. Fynbo,¹⁵ D. Galaviz,¹⁶ P. T. Greenlees,^{17,18} R. Grzywacz,^{19,20} L. J. Harkness-Brennan,²¹ C. Henrich,¹⁴ M. Huyse,¹³ A. Illana,¹³ Z. Janas,¹ K. Johnston,³ D. S. Judson,²¹ V. Karanyonchev,²² M. Kicińska-Habior,¹ J. Konki,^{17,18} J. Kurcewicz,³ I. Lazarus,²³ R. Lică,^{3,11} H. Mach,^{24,‡} M. Madurga,^{3,19} I. Marroquín,⁹ B. Marsh,³ M. C. Martínez,² C. Mazzocchi,¹ N. Mărginean,¹¹ R. Mărginean,¹¹ K. Miernik,¹ C. Mihai,¹¹ E. Nácher,⁹ A. Negret,¹¹ B. Olaizola,^{25,§} R. D. Page,²¹ S. Paulaskalas,¹⁹ S. Pascu,¹¹ A. Perea,⁹ V. Pucknell,²³ P. Rahkila,^{17,18} E. Rapisarda,³ J.-M. Régis,²² F. Rotaru,¹¹ S. Rothe,³ V. Sánchez-Tembleque,² G. Simpson,²⁶ Ch. Sotty,^{11,13} L. Stan,¹¹ M. Stănoiu,¹¹ M. Stryczyk,^{1,13} O. Tengblad,⁹ A. Turturica,¹¹ J. M. Udías,² P. Van Duppen,¹³ V. Vedia,² A. Villa,² S. Viñals,⁹ R. Wadsworth,⁴ W. B. Walters,²⁷ and N. Warr²²

(IDS Collaboration)

¹*Faculty of Physics, University of Warsaw, PL 02-093 Warsaw, Poland*

²*Grupo de Física Nuclear and UPARCOS, Universidad Complutense de Madrid, CEI Moncloa, E-28040 Madrid, Spain*

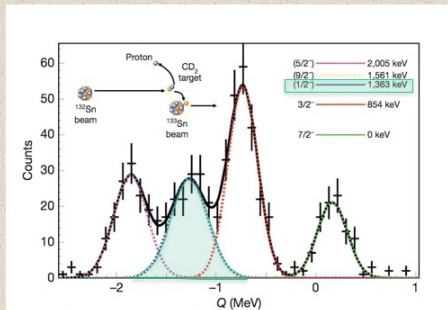
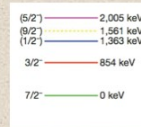
The location of the $i13/2$ single particle neutron state is still elusive, one possibility is that it is unbound state.



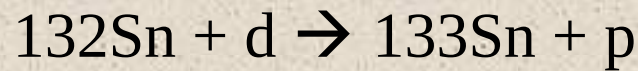
Conclusion of this part

It is important to have the possibility to perform reactions where either the target or the projectile are radioactive.

What is easier???



Haz radioactivo o target
Radioactivo?



To produce a beam of ^{132}Sn is possible (not easy),
To produce a target of ^{132}Sn ($T_{1/2}=40$ s) is impossible

We need to produce a beam of radioactive nuclei

The first thing to worry about is how to produce radioactive nuclei:
Tool: Reaction or Fission

The second thing to worry is to produce them in a “clean” way

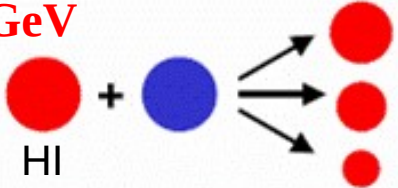
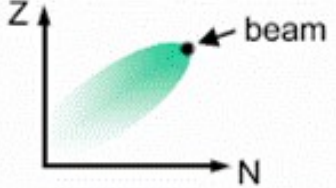
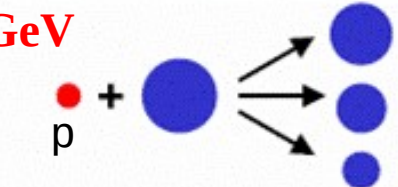
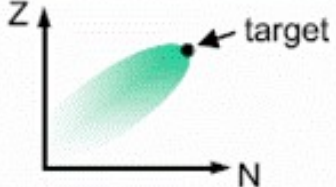
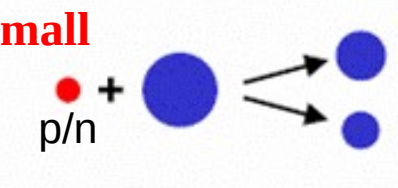
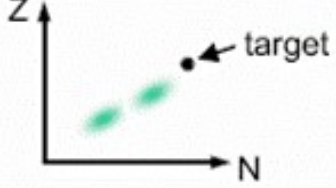
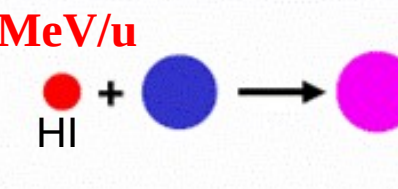
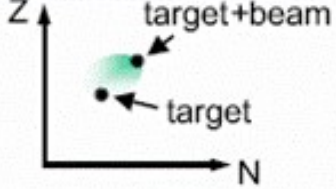
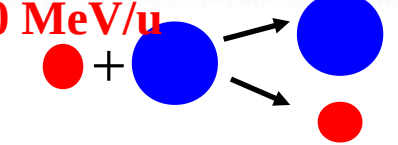
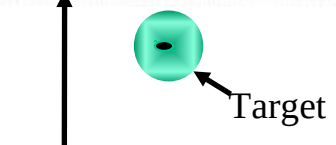
For that it is important to separate them from other radioactive products

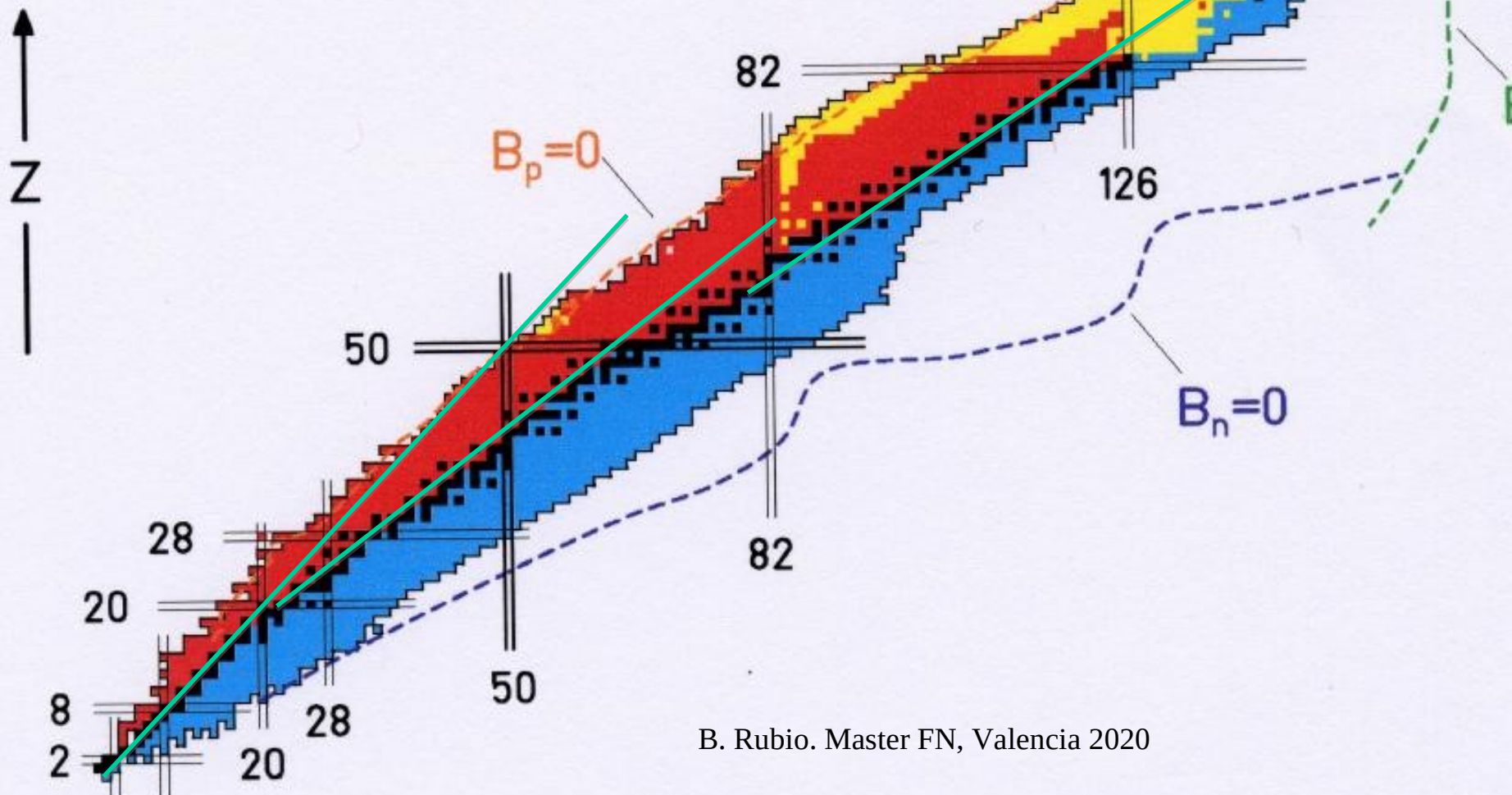
The third thing is to produce them at the right energy

There are two main ways: the Isol method and the fragmentation method

The first thing to worry about is how to produce radioactive nuclei.
Tool: Reaction or Fission

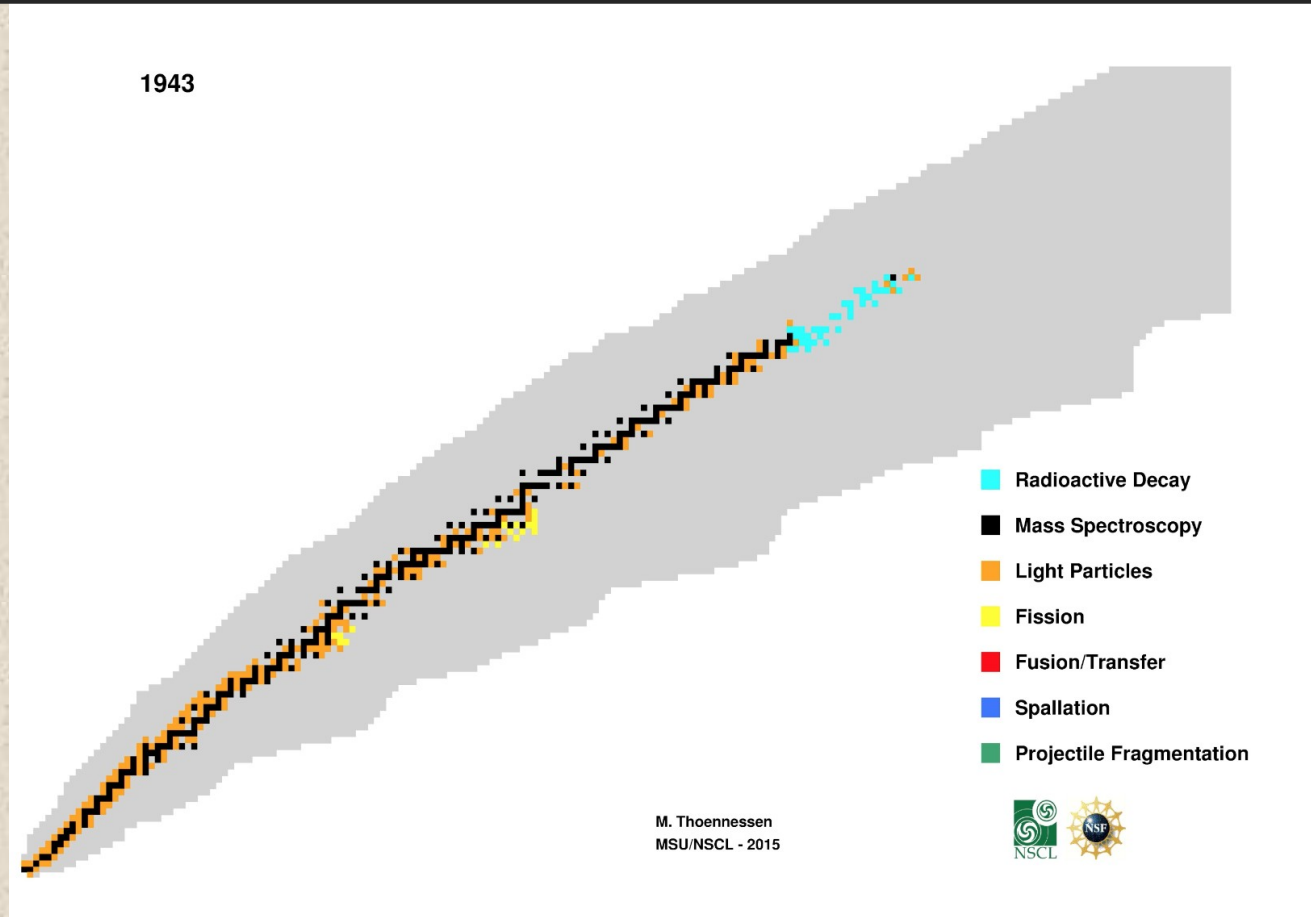
The reaction to produce Radioactive beams starting with stable beam and stable target

Beam + target → products		Product energy	Number of Products	
GeV  HI	fragmentation	“All nuclei” $V_{\text{product}} = V_{\text{beam}}$ RIKEN, GSI, MSU, FAIR, FRIB	up to 1000	
GeV  p	spallation	“All nuclei” few MeV/u ISOLDE	up to 1000	
small  p/n	fission	“mainly neutron rich” ~1 MeV/u Reactors, ISOLDE...	few 100	
5MeV/u  HI	fusion-evaporation	“mainly proton rich rich” $E_R = \frac{m_p}{m_p + m_t} E_p$ SPIRAL2	few (≤ 20)	
20 MeV/u 	Transfer reactions	“close to the stability” 5-100 MeV/u “Small facilities”	Very few: 2-5	



Historical view to the discovery of new isotopes

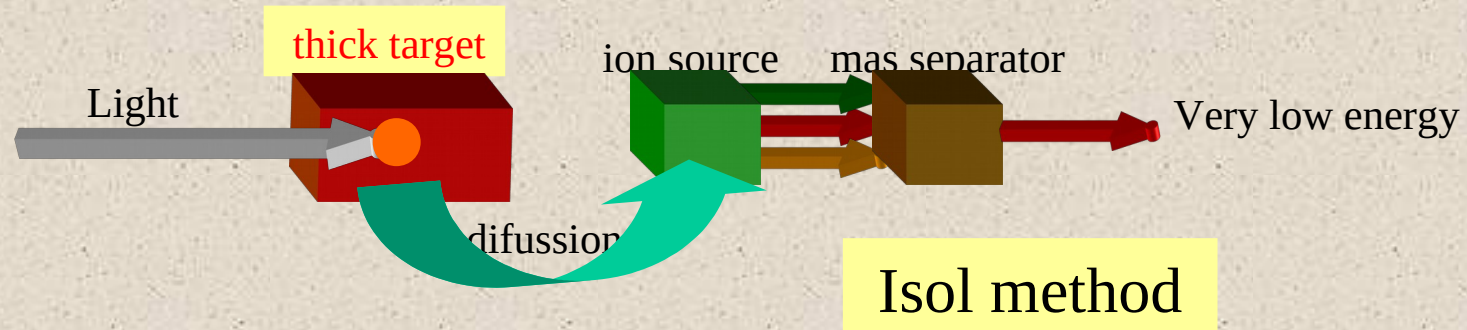
<https://people.nscl.msu.edu/~thoennes/isotopes/yearchart-2015.mp4>



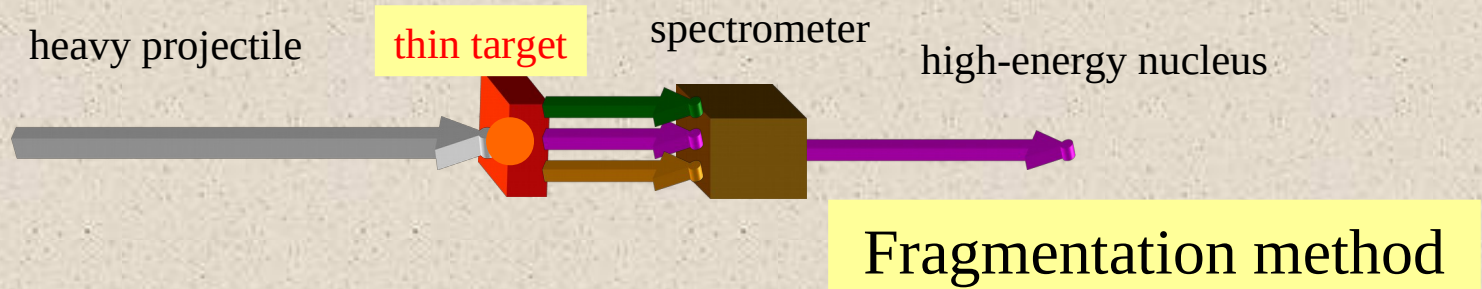
The second thing is to separate them from other radioactive products

Radioactive nuclei production techniques

Isotopic Separation On-Line (ISOL)



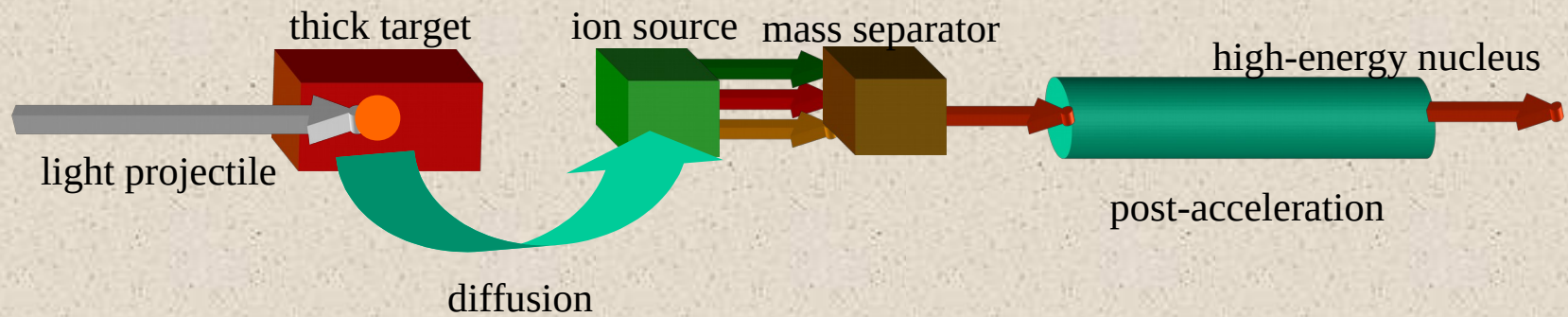
In-flight fragmentation



Production techniques Rad Beams: post acceleration

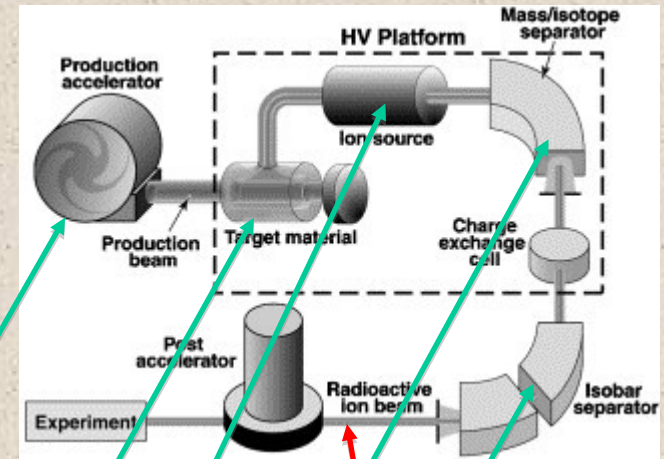
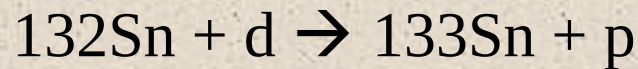
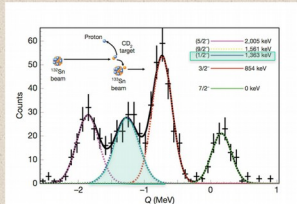
➤ Isotopic separation on-line (ISOL)

Isol method





LETTERS
The magic nature of ^{132}Sn explored through the single-particle states of ^{132}Sn



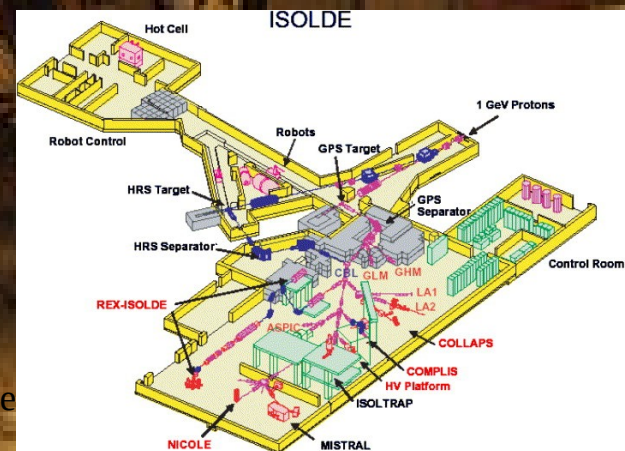
1. Protons accelerated in the cyclotron
2. Uranium Carbide power target
3. About 100 species produced, some of them ionised
4. ^{132}Sn Mass separated
5. ^{132}Sn separated
6. ^{132}Sn accelerated to 4.8 MeV x A
7. Experiment

On 16 October 1967, the first experiments were carried out at the Isotopic Separator On Line ISOLDE CERN.

CERN's longest serving experimental facility

Physics Reports 403–404 (2004) 459–469

Juha Äystö



B. Rubio. Master FN, Vale

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Dr. Karl Ove Nielsen, Aarhus
Prof. Jean Teillac, Orsay
Prof. W. Gentner, Heidelberg
Dr. Roald Tangen, Oslo
Dr. Göran Andersson, Gothenburg
Dr. Ingmar Bergström, Stockholm
cc. Prof. Preiswerk, for inform.

Meyrin-Genève, le 17 December, 1964

.....Furthermore, CERN is willing to contribute one sixth of the cost of the basic equipment needed as estimated in the proposal*, which corresponds to the participation of the CERN Nuclear Chemistry Group in the project, and one sixth of the cost of maintaining, running, and developing the above-mentioned basic equipment after the construction period. It is understood that the rest of the cost will have to be shared by the participating groups. These groups will also have to bring to CERN any specialized measuring equipment which they might wish to use in connection with the project.

Yours sincerely,

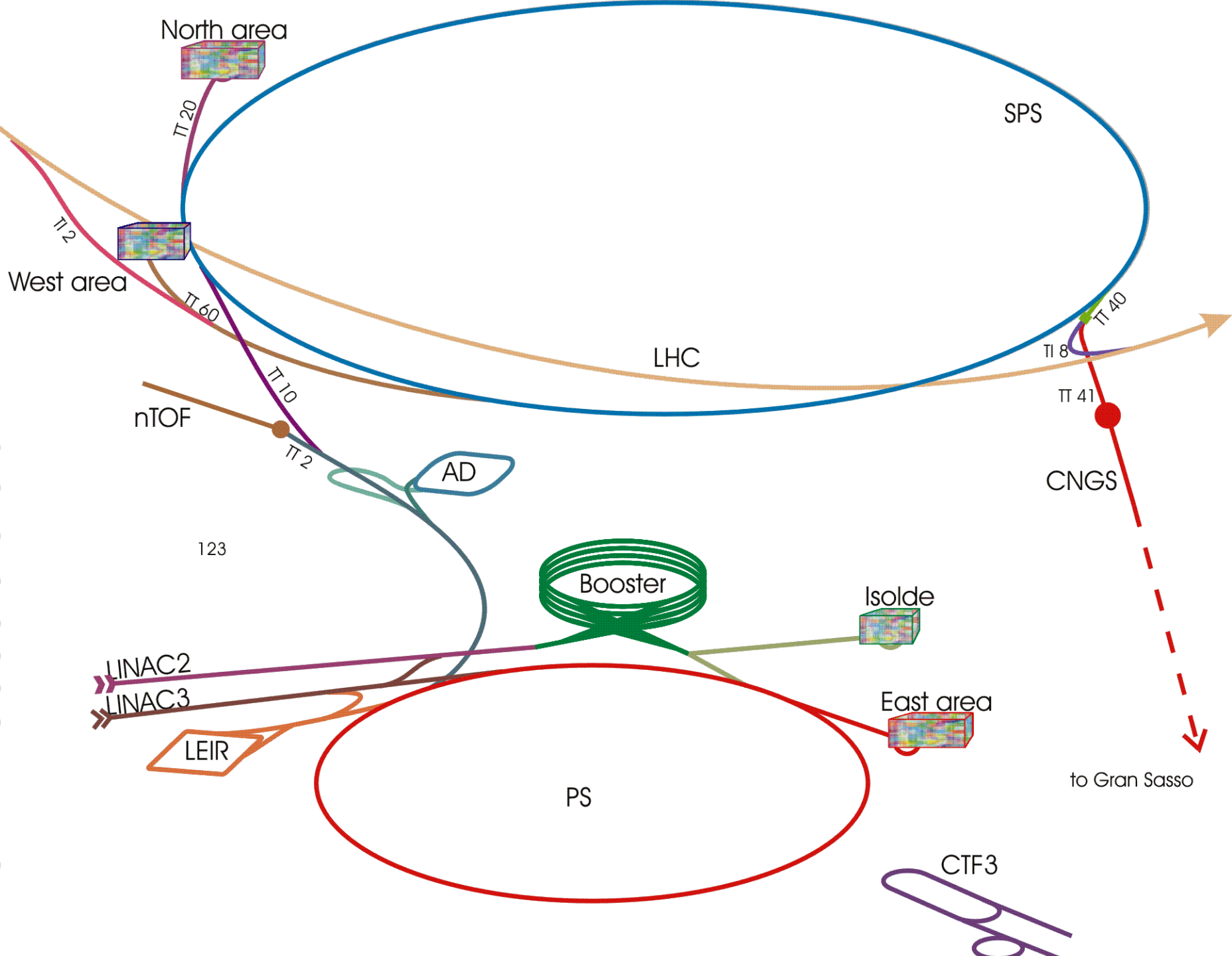
Viktor F. Weisskopf

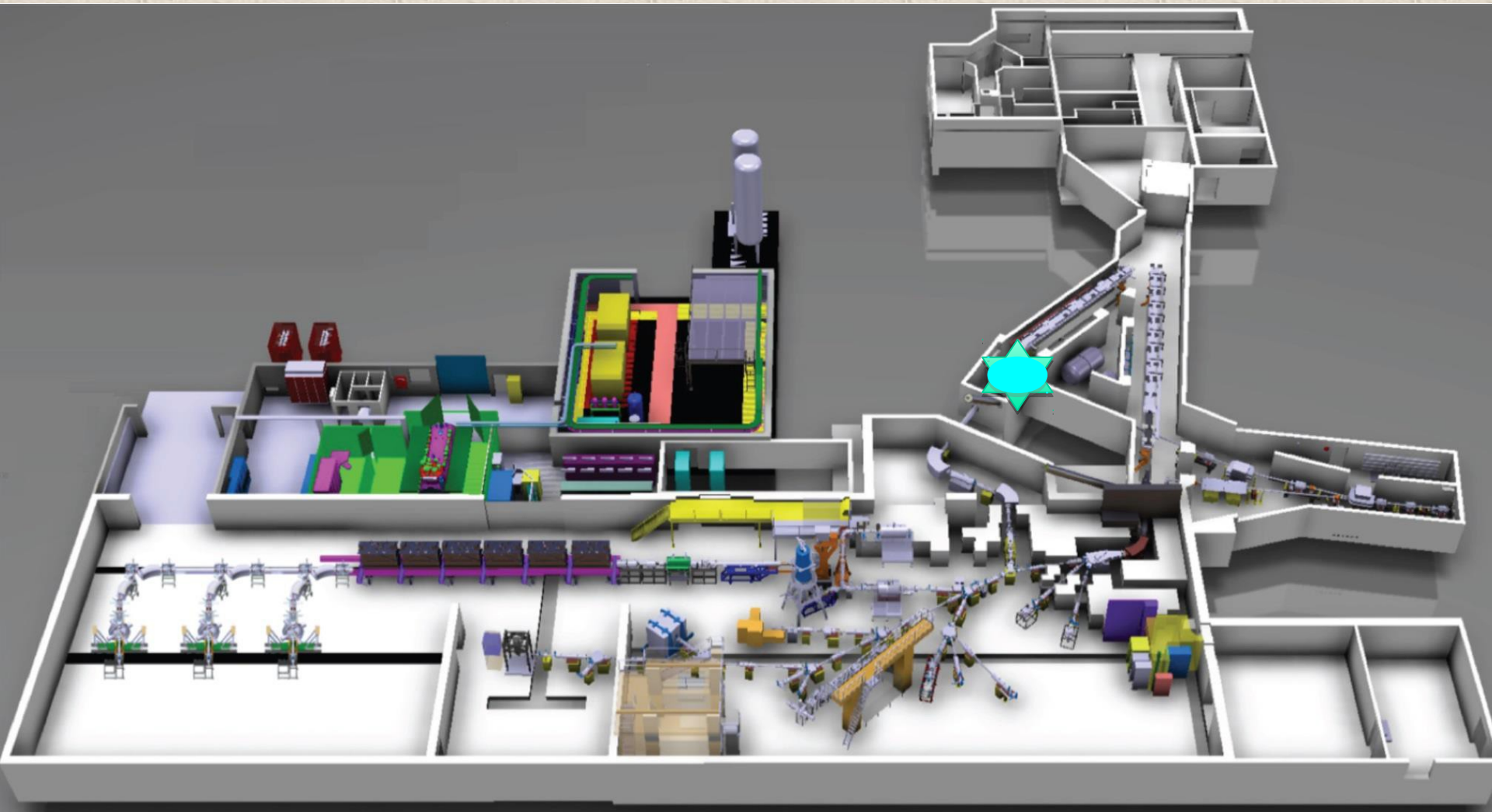


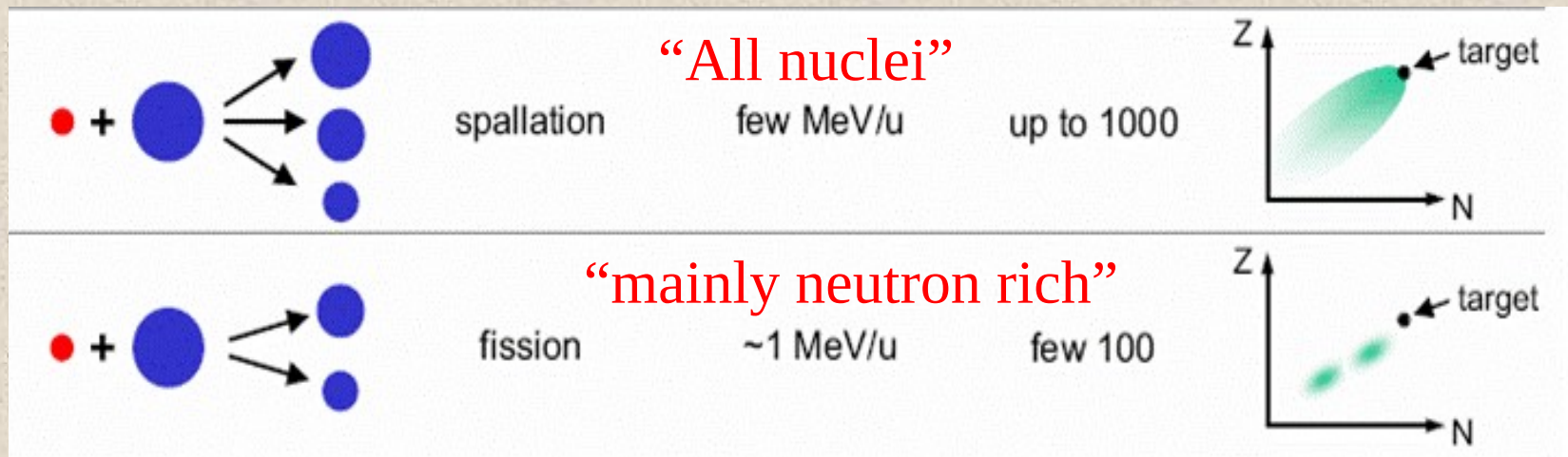
1967 First experiment
P of 600 MeV from the PSC

1992 moved to PS booster (today)
(in 2017, 50 years celebration)

CERN accelerators



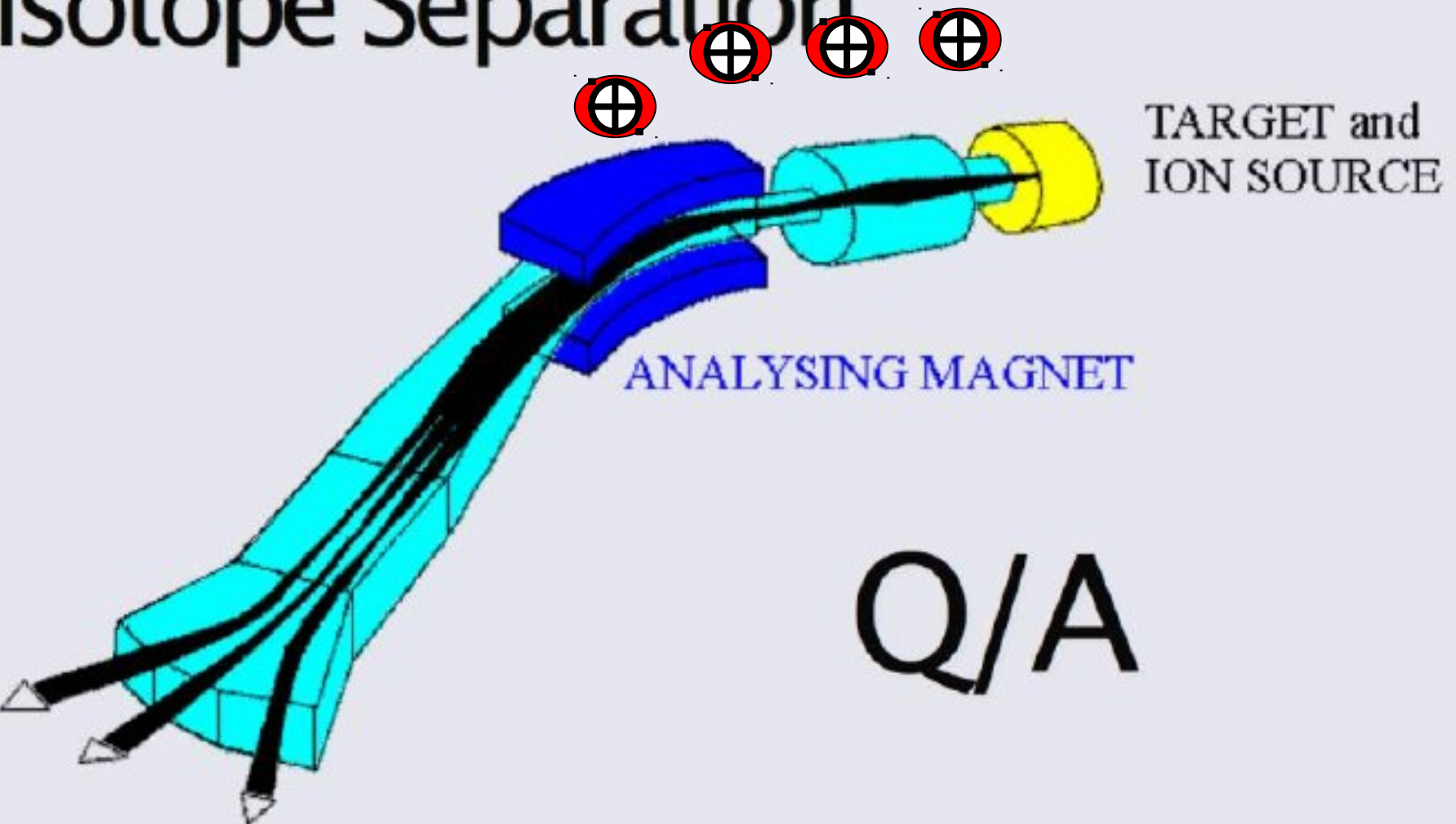




- ☺ High intensity p beam
- ☺ thick target (100% of range) => high beam current (upto 10^{16} s^{-1})
long extraction and ionization time (ms)
- ☹ Chemistry dependent

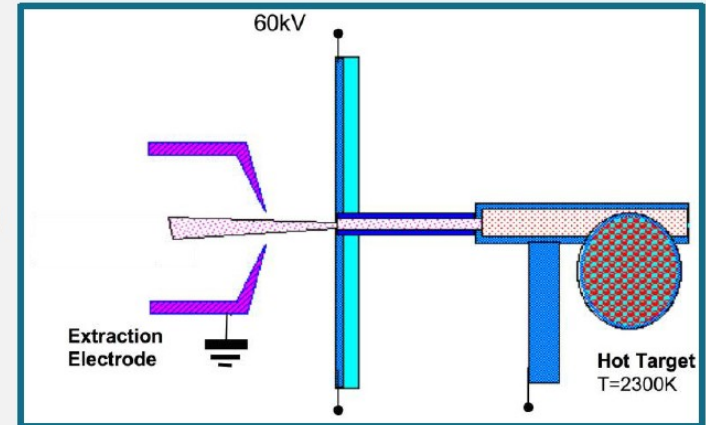
The Separator: in general, only
isobaric separation

Isotope Separation



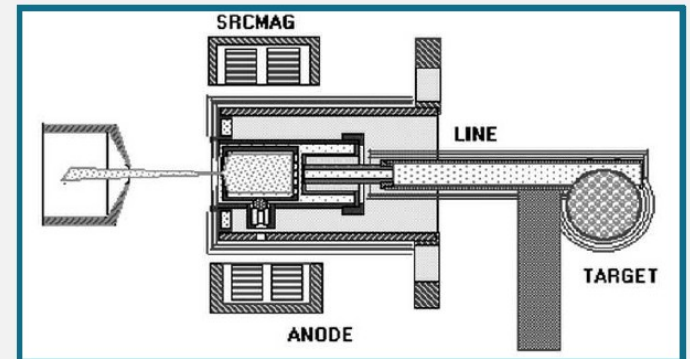
Hot surface ion source

- The ioniser is a hot tube
- Material with a higher work function than the element of interest
- Heated up to 2400 degrees



Plasma ion source

- Plasma: gas mixture (Ar and Xe) ionised by accelerated electrons
- Hot or cool transfer line

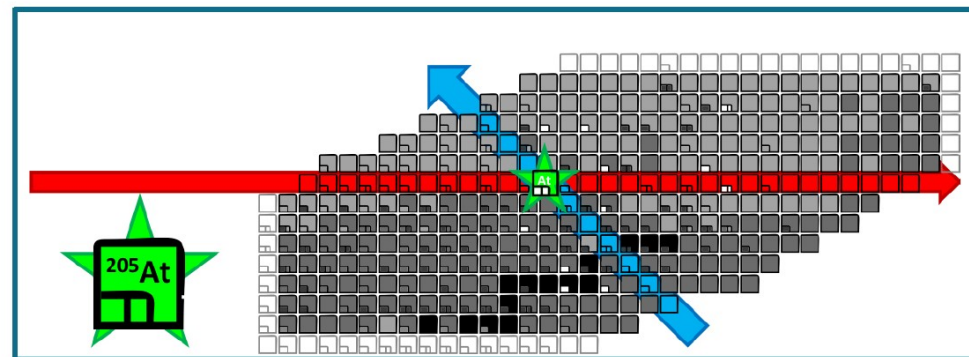
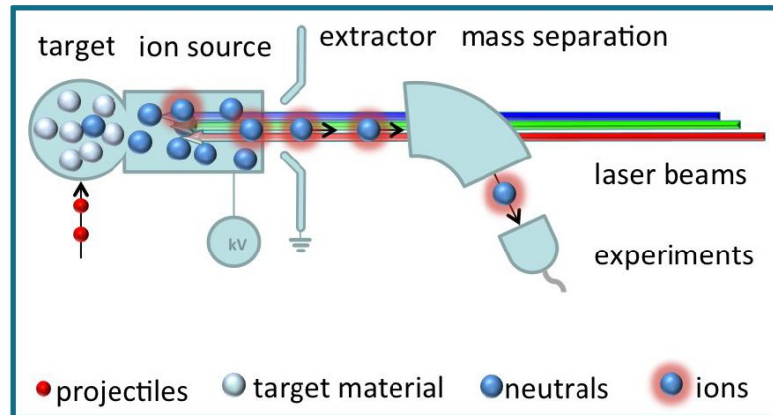
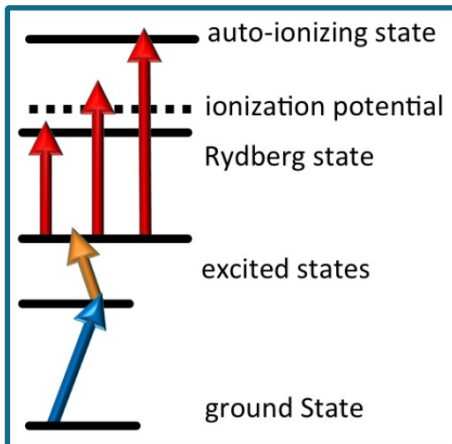




Ion sources

Laser ion source

- 2- or 3-step ionisation
- Isotope and isomer selection
- Universal (almost)



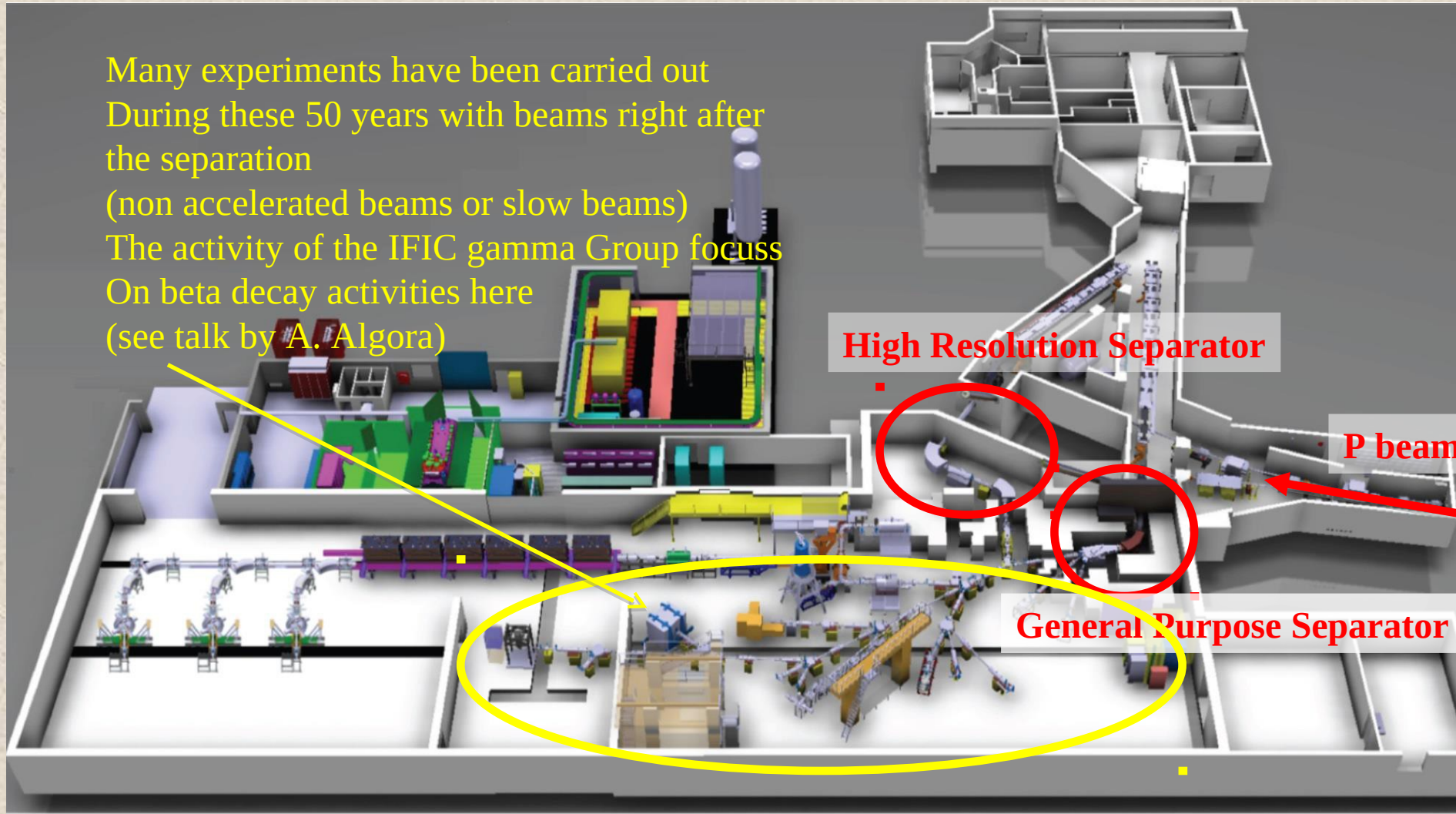
Isotopes produced at ISOLDE

H																	He												
Li	Be																	B	C	N	O	F	Ne						
Na	Mg																	Al	Si	P	S	Cl	Ar						
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr												
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe												
Cs	Ba	La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn												
Fr	Ra	Ac	Rf	Db	Sg	Bh	Hs	Mt	Ds	Rg	112	113	114	115															
			Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu													
			Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr													

ION SOURCE

+	SURFACE	-
hot	PLASMA	cooled
LASER		

Many experiments have been carried out
During these 50 years with beams right after
the separation
(non accelerated beams or slow beams)
The activity of the IFIC gamma Group focuss
On beta decay activities here
(see talk by A. Algora)



ISOLDE

The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

17 DECEMBER 1964

CERN approves the online separator project

View 



ISOLDE

The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

26 JUNE 1992

First experiment at the ISOLDE Proton-Synchrotron Booster

View 



ISOLDE

The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

31 OCTOBER 2001

New accelerator, REX-ISOLDE, goes live

View 



ISOLDE

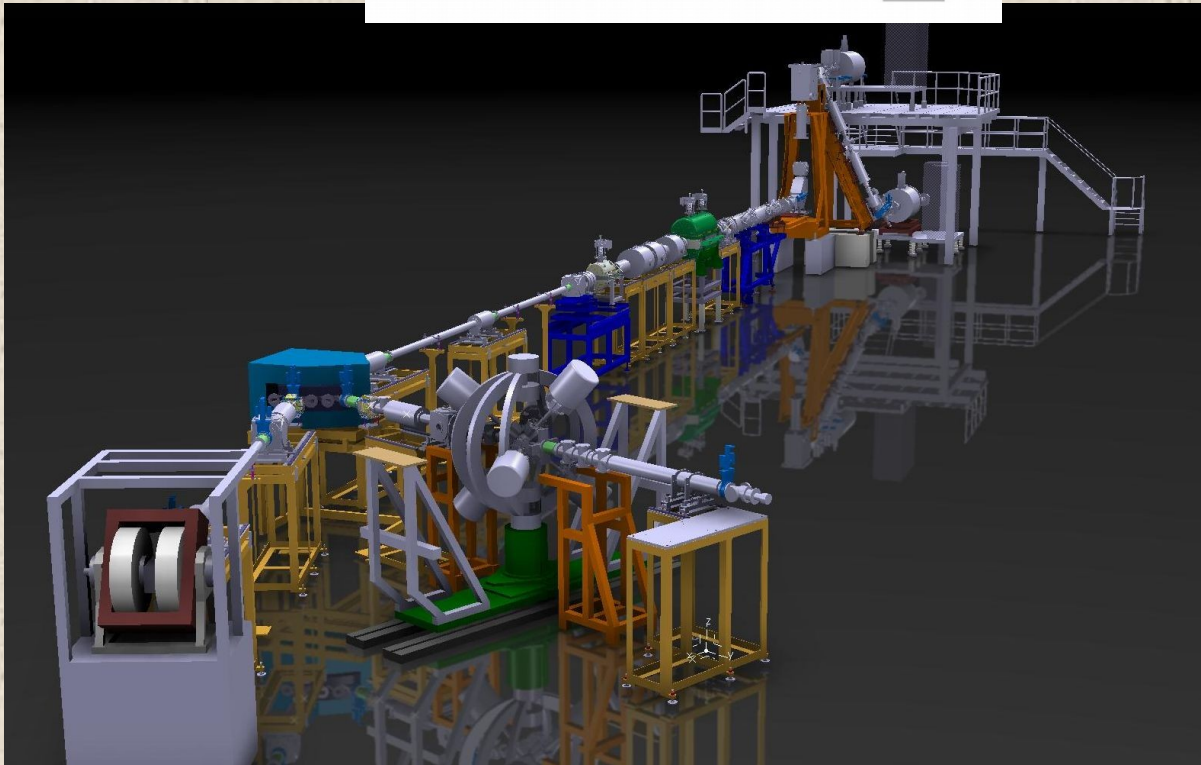
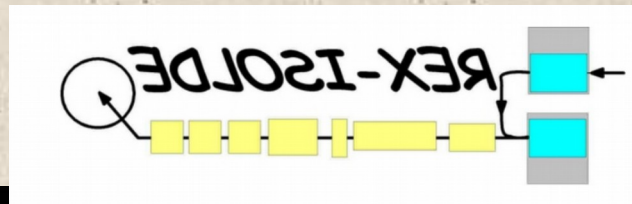
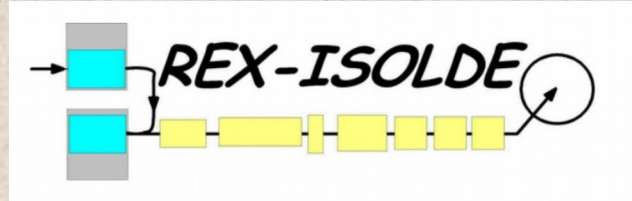
The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

22 OCTOBER 2015

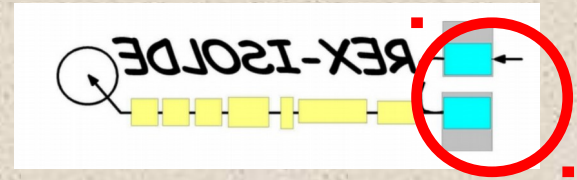
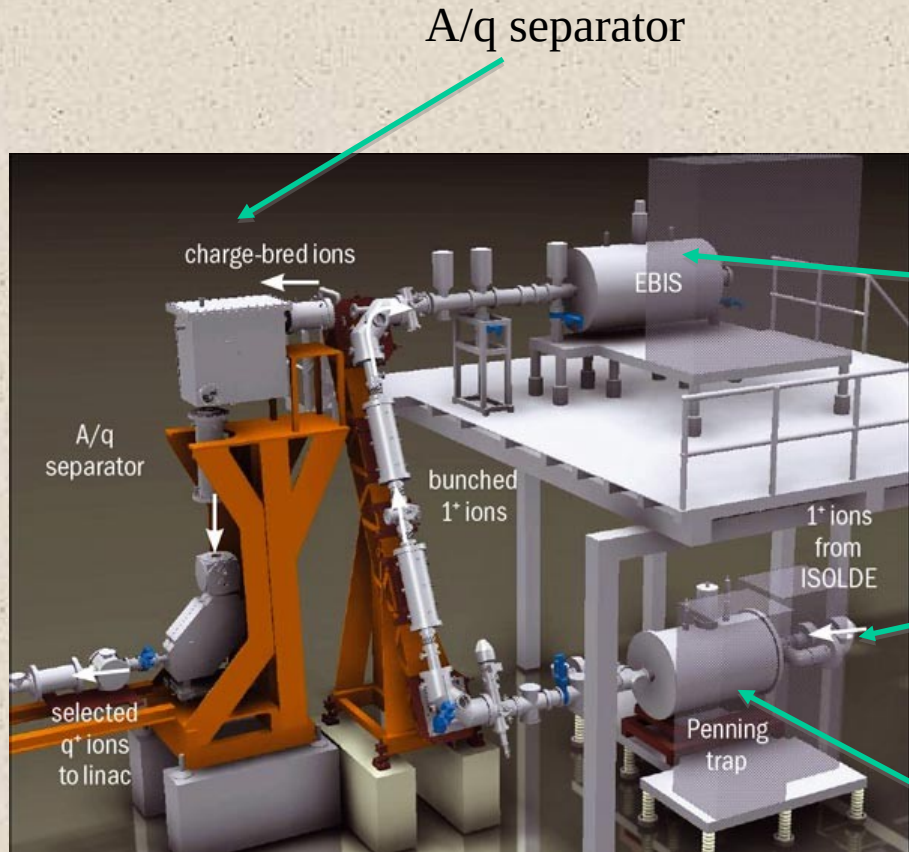
First radioactive isotope beam accelerated in HIE ISOLDE

View 





EBIS Ion Source and beam handling: 5-15% efficiency

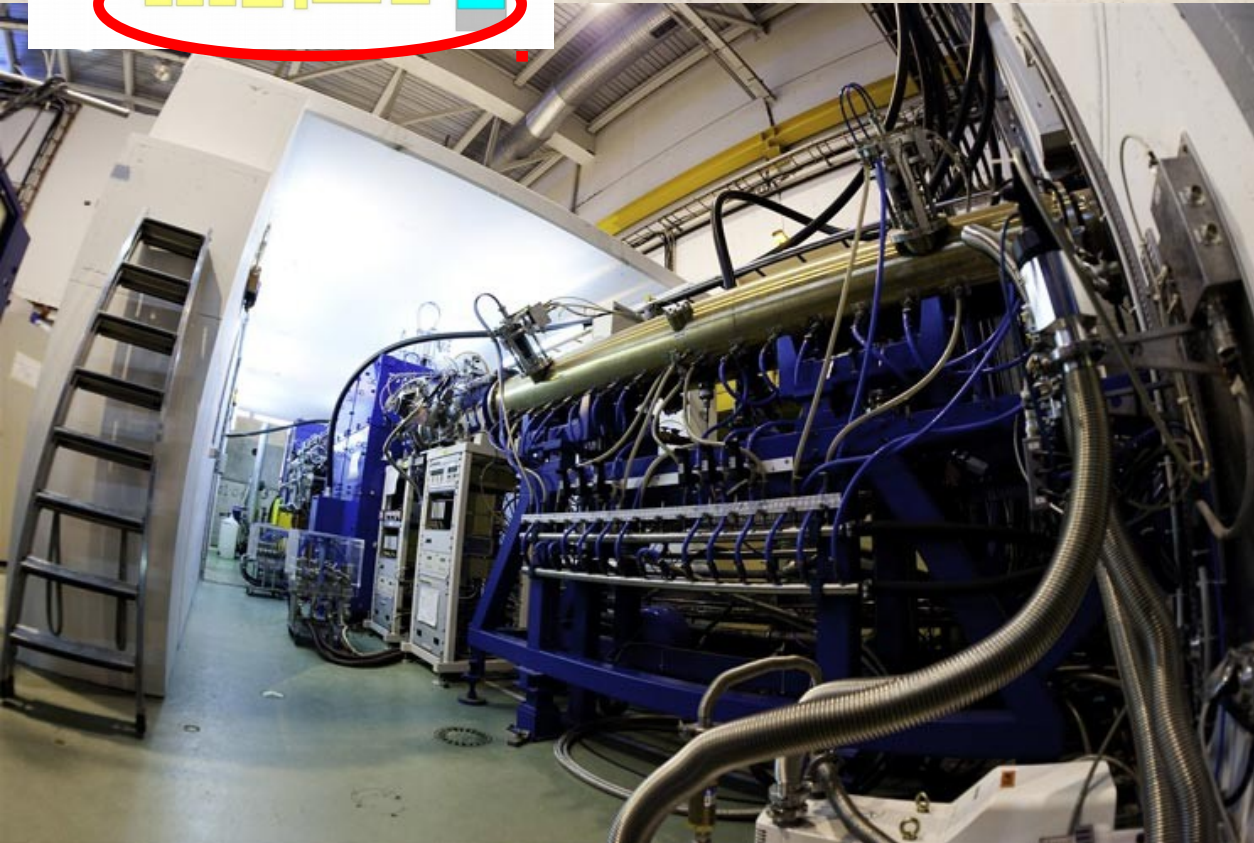
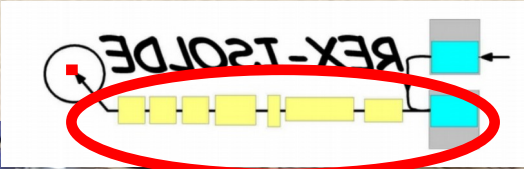


In the EBIS the 1^+ to n^+ ion conversion takes place

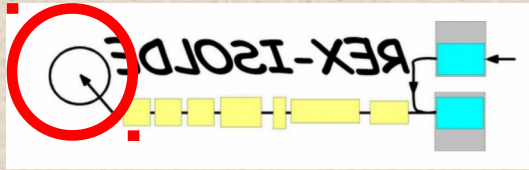
Beams from ISOLDE with 1^+ charge state

Preparation trap

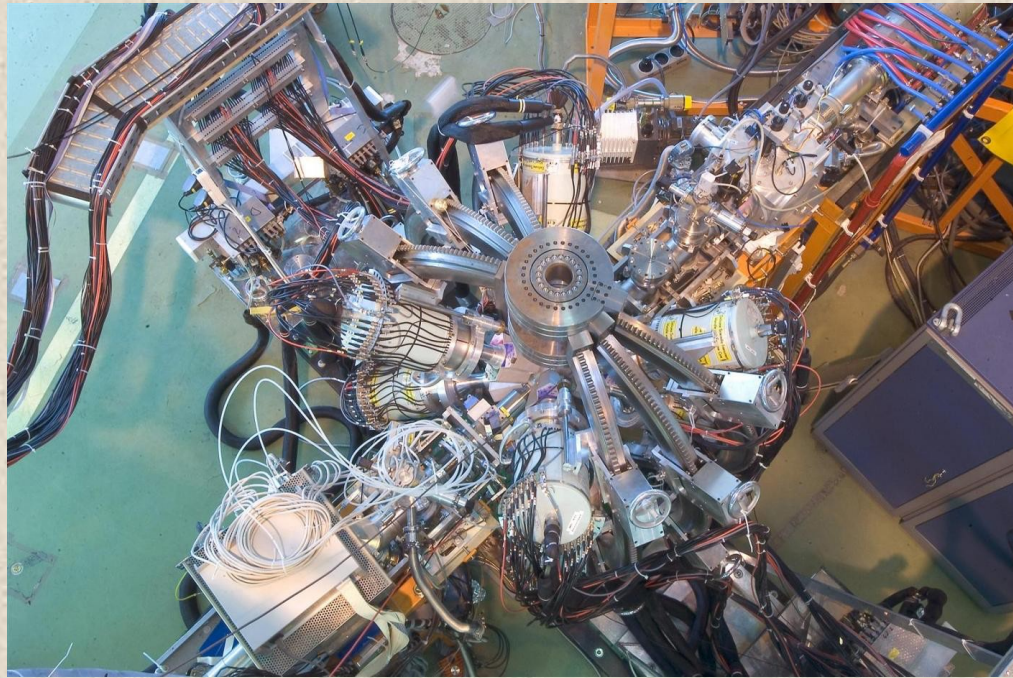
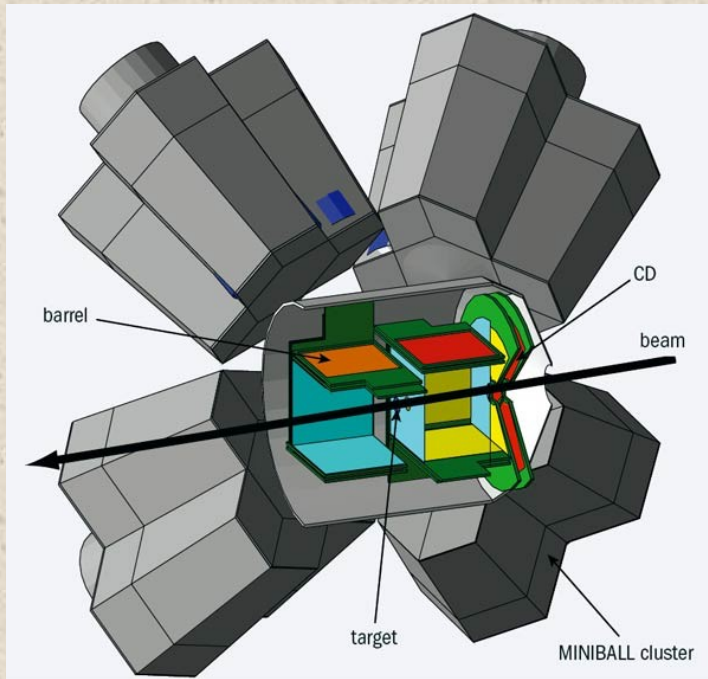
Post accelerator: Normal Conducting LINAC



However, energies only 0.8 to 3 MeV/u which are too low for Nuclear reactions



Miniball gamma array

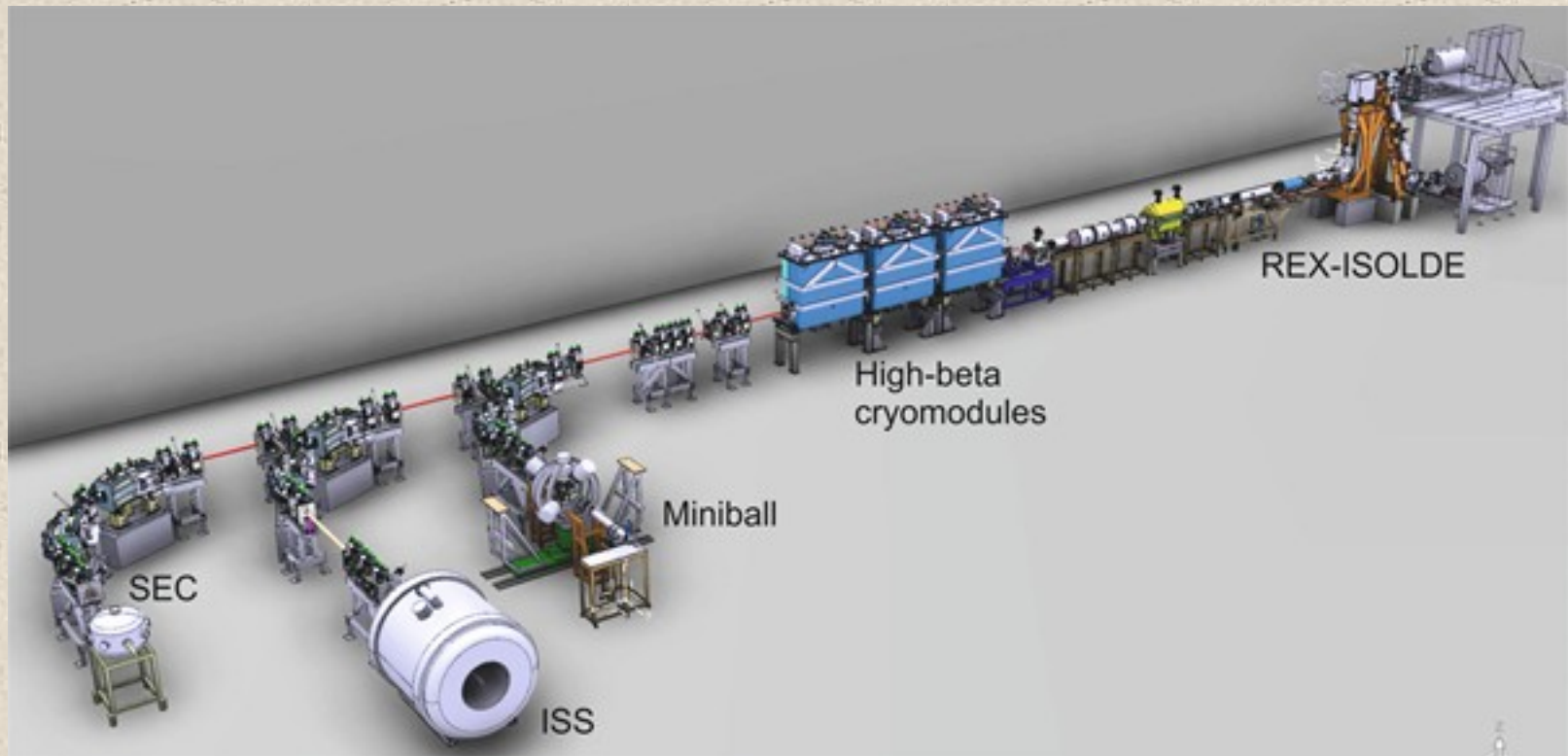


24 six-fold segmented, tapered, encapsulated high-purity germanium crystals

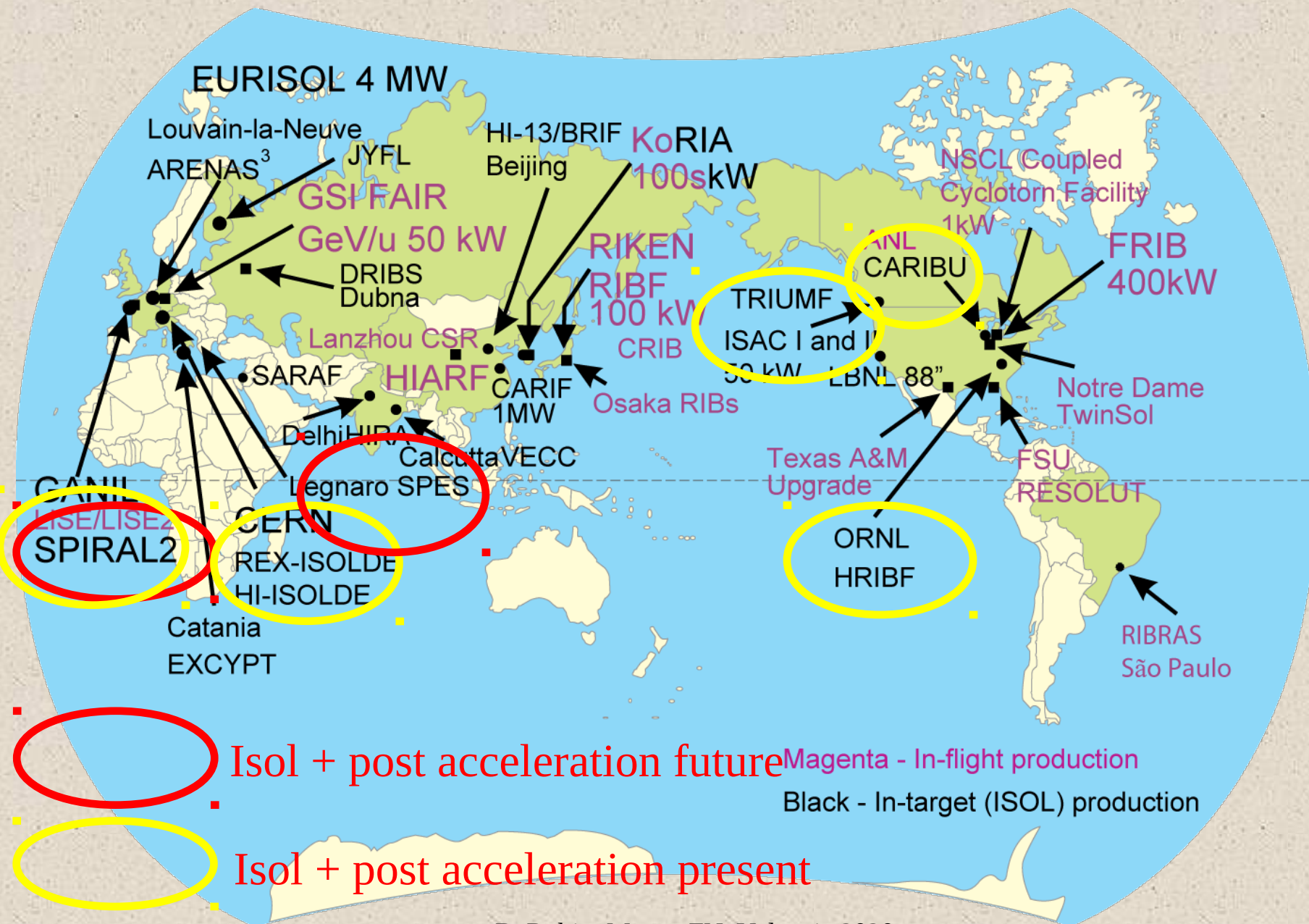
HIE ISOLDE

High Intensity and Energy ISOLDE (up to 10 MeV per nucleon)

4.3 MeV per nucleon



First experiments in 2015



Future ISOL facilities

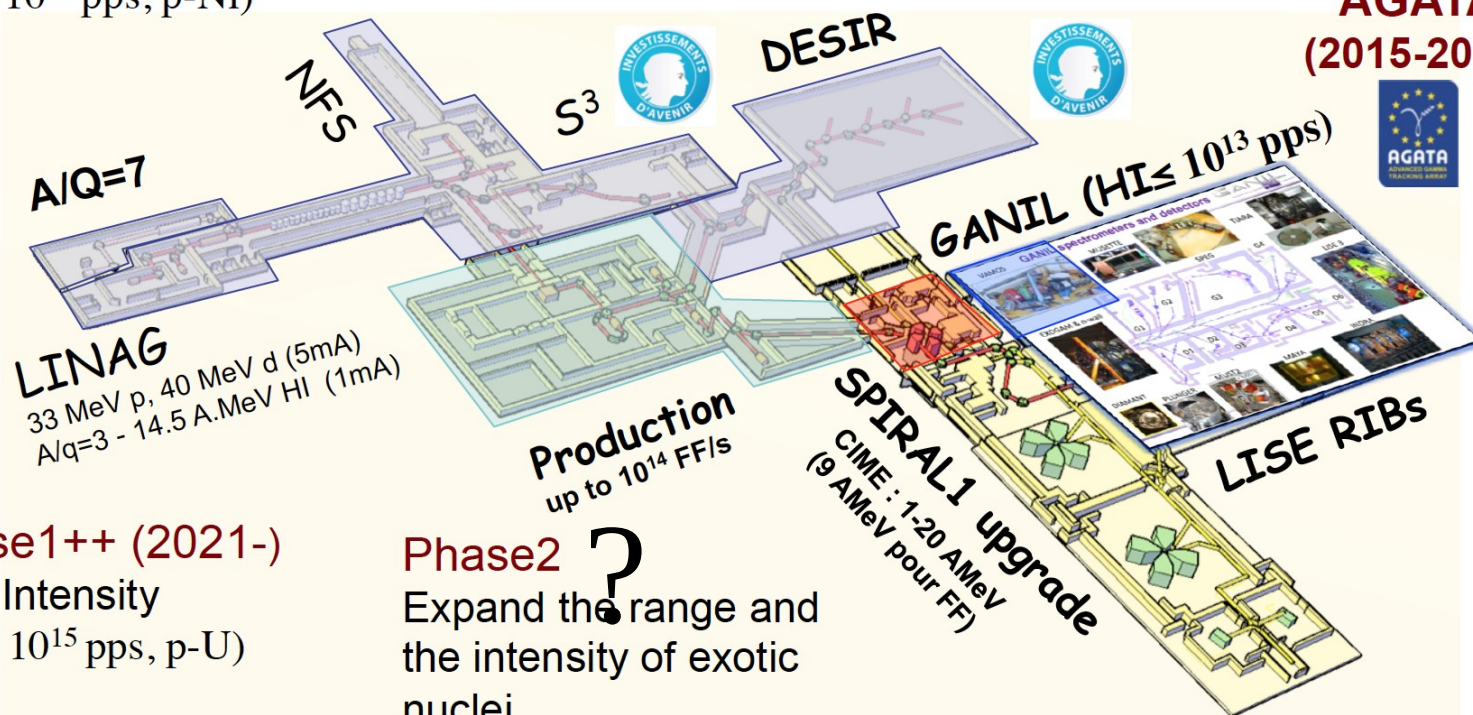
Phase1 (2016-)

Increase the intensity of stable beams
High intense neutron source
($HI \leq 10^{15}$ pps, p-Ni)

DESIR Phase1+ (2020-)

Low energy facility

AGATA (2015-2018)



Phase1++ (2021-)

High Intensity
($HI \leq 10^{15}$ pps, p-U)

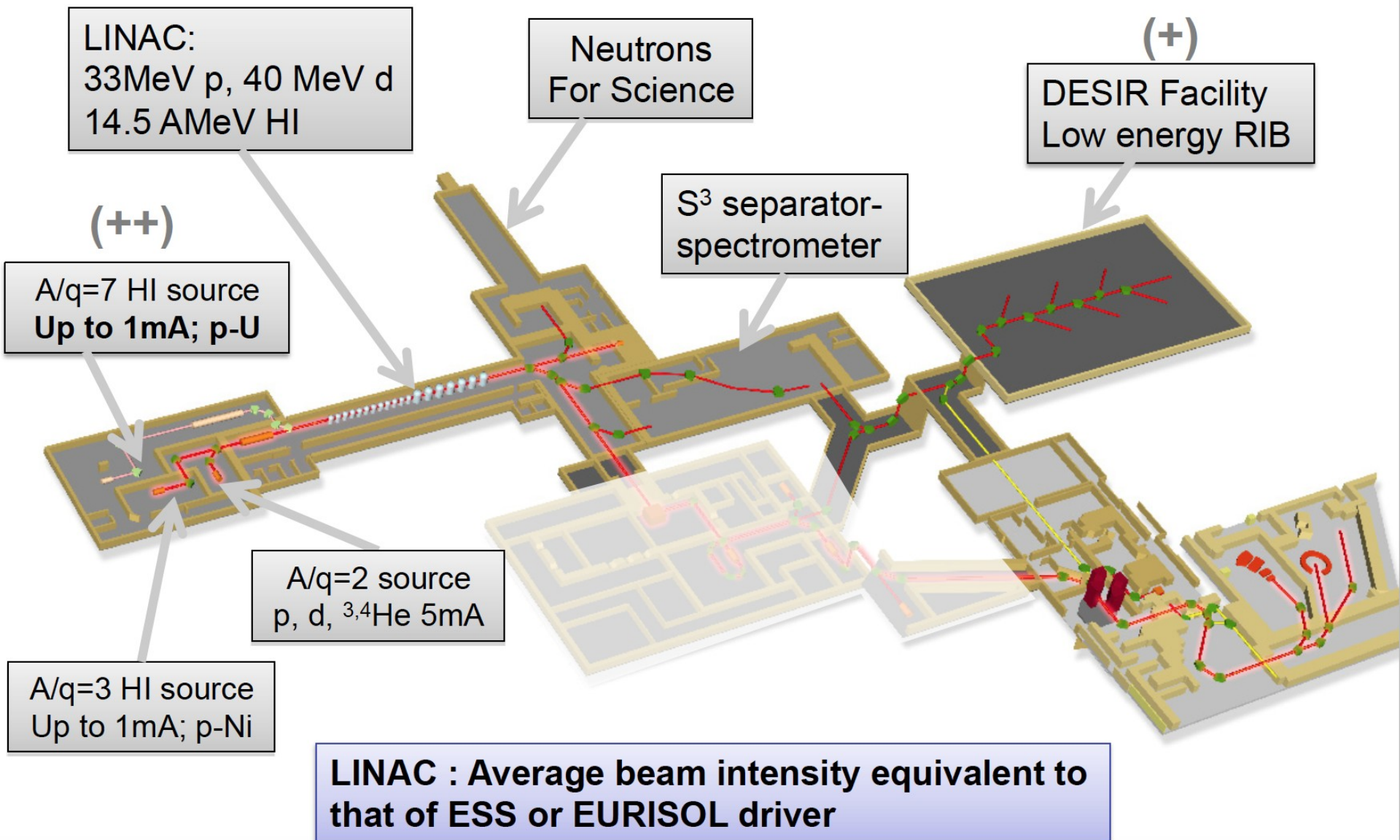
Phase2 ?

Expand the range and
the intensity of exotic
nuclei

SPIRAL1 Upgrade (2017-)

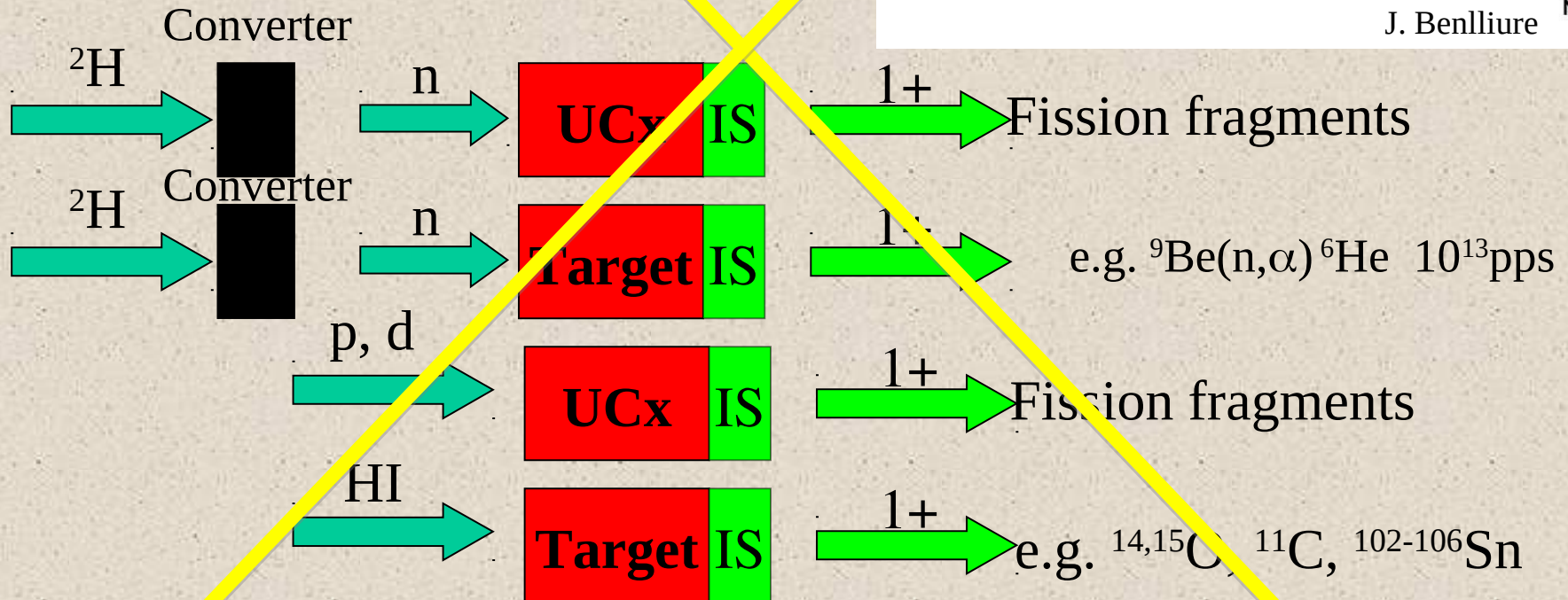
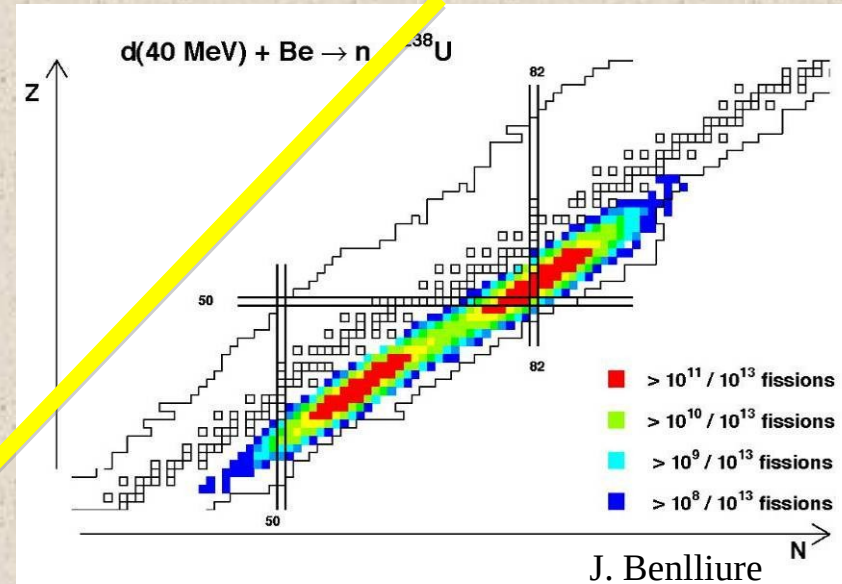
New light RIBs from
beam/target fragmentation

Near Future : SPIRAL2 Phase 1



ISOL Rare Isotope Beams at SPIRAL 2

Unfortunately the neutron rich side of The Spiral 2 Facility is not financed at The moment



Fortunately and strategically a new project for neutron rich (fission), has been approved in Legnaro (Italy), and is already under construction

SPES Selective Production of Exotic Species

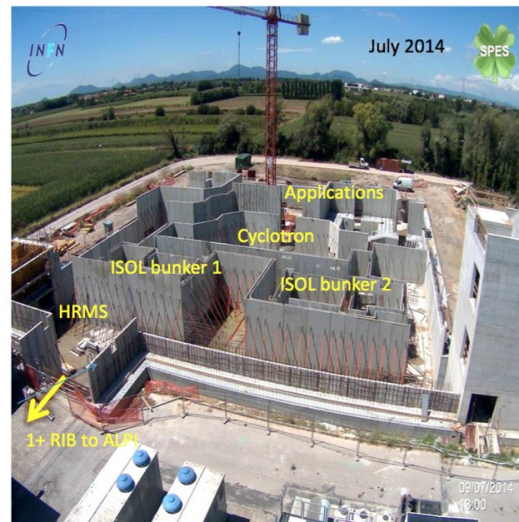


SPES Project

The SPES radioactive ion beam facility: status and perspectives



Giacomo de Angelis
INFN Laboratori Nazionali di Legnaro
for The SPES collaboration



July 2014



SPES building

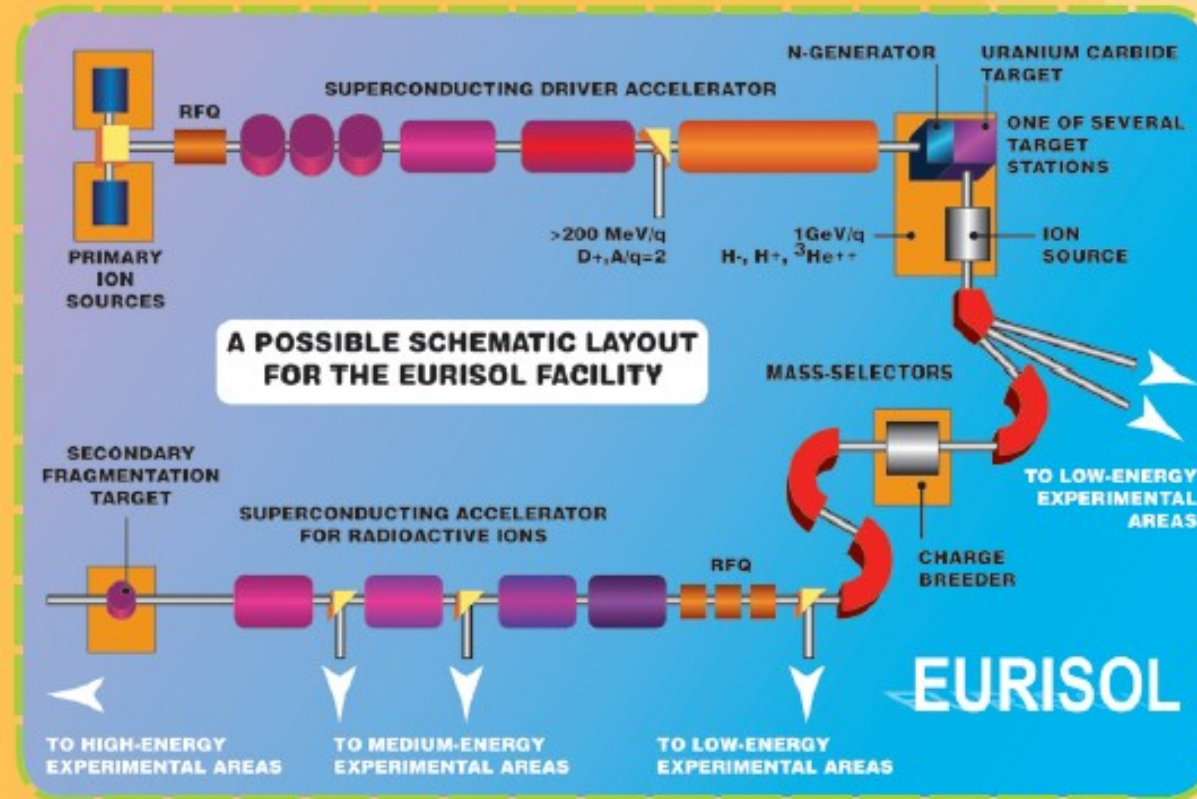
50x60 m²
-3 to +11 m height
24.000 m³ of concrete
1.150 tons iron
3-4 m shielding wall thick



Oct 2013

Far in the Future ISOL facility

Furthering target and accelerator technology



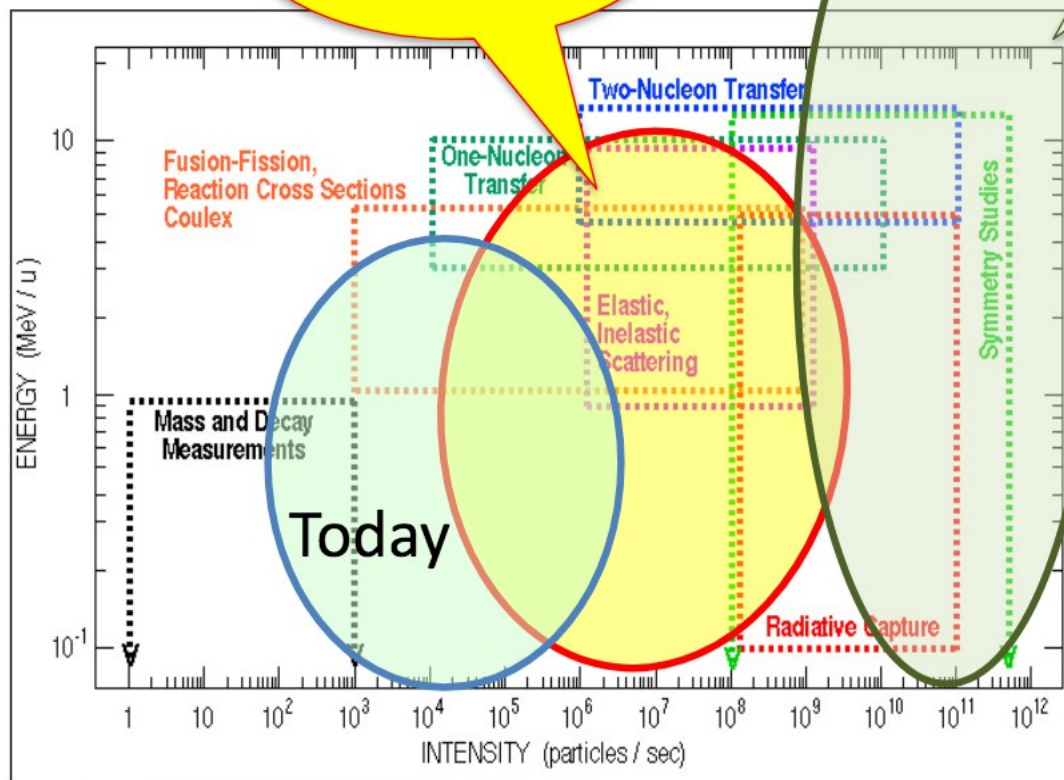
A 4-year Design Study began in 2005 in order to work on the technologically challenging aspects of the project, the instrumentation and the radiation safety issue. Synergy with other projects is being examined, including a feasibility study for the new 'beta-beam' neutrino proposal, forming an integral part of the Design Study. After this, possible sites will be evaluated, and the community will be ready for a full Engineering Design, to be followed by construction of the facility.

Researchers and engineers of several European laboratories are collaborating in twelve tasks to further the EURISOL design.

The EURISOL layout consists of a superconducting linear accelerator providing protons of energy 1 GeV and an impressive power of 5 MW, but also capable of accelerating deuterons, 3He and ions up to mass 40. The beams will impinge simultaneously on two types of targets, either directly or after conversion of the protons into neutrons through a loop containing 1 ton of mercury surrounded by kilograms of fissile material. The unstable nuclei produced diffuse out of the target, are ionized and selected, and can be used directly at low energy or reaccelerated by another linear accelerator to energies up to 150 MeV per nucleon in order to induce nuclear reactions.

Physics with ISOL RIB Intensity & Energy domains

Precision nuclear structure physics & applications



HIE-ISOLDE,
SPES, SPIRAL2,
ISOL@MYRRHA
EURISOL-DF

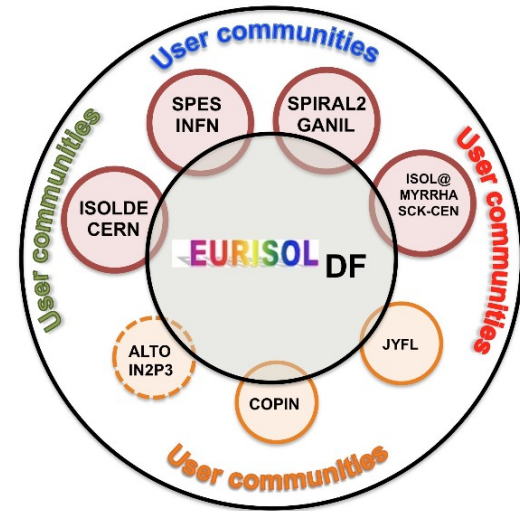
EURISOL



-> EURISOL-DF (Distributed Facility) Initiative from 2014 as an intermediate step towards EURISOL

EURISOL DF

EURISOL – Distributed Facility (DF)

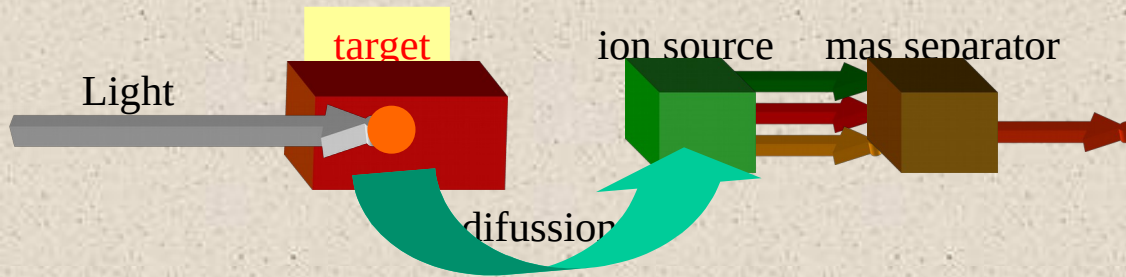


Core members :
HIE-ISOLDE/CERN
SPES-INFN
SPIRAL2-GANIL
ISOL@MYRRHA-SCK*CEN

Associated Members
JYFL, Finland
COPIN Consortium, Poland
(ALTO, Orsay)

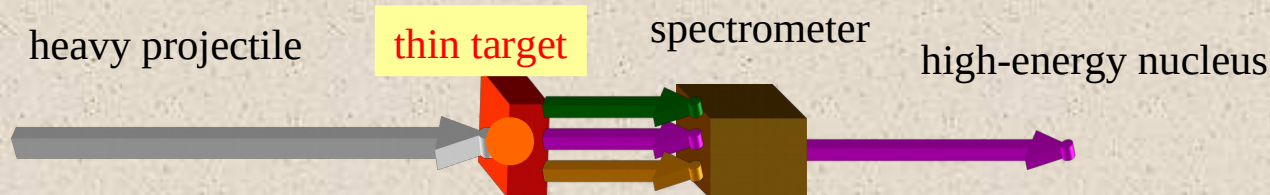
Radioactive nuclei production techniques

Isotopic separation on-line (ISOL)



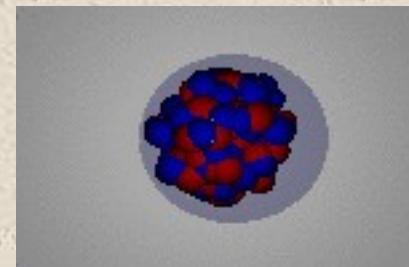
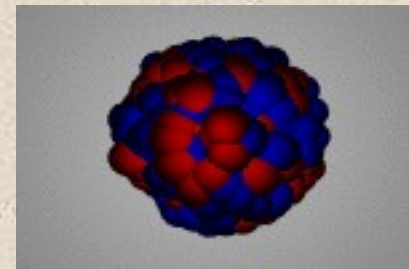
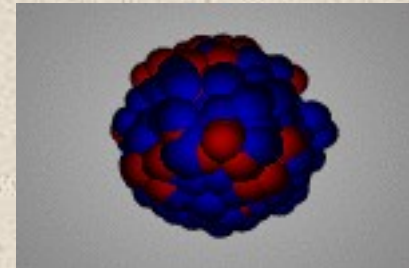
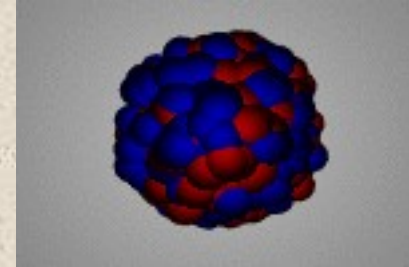
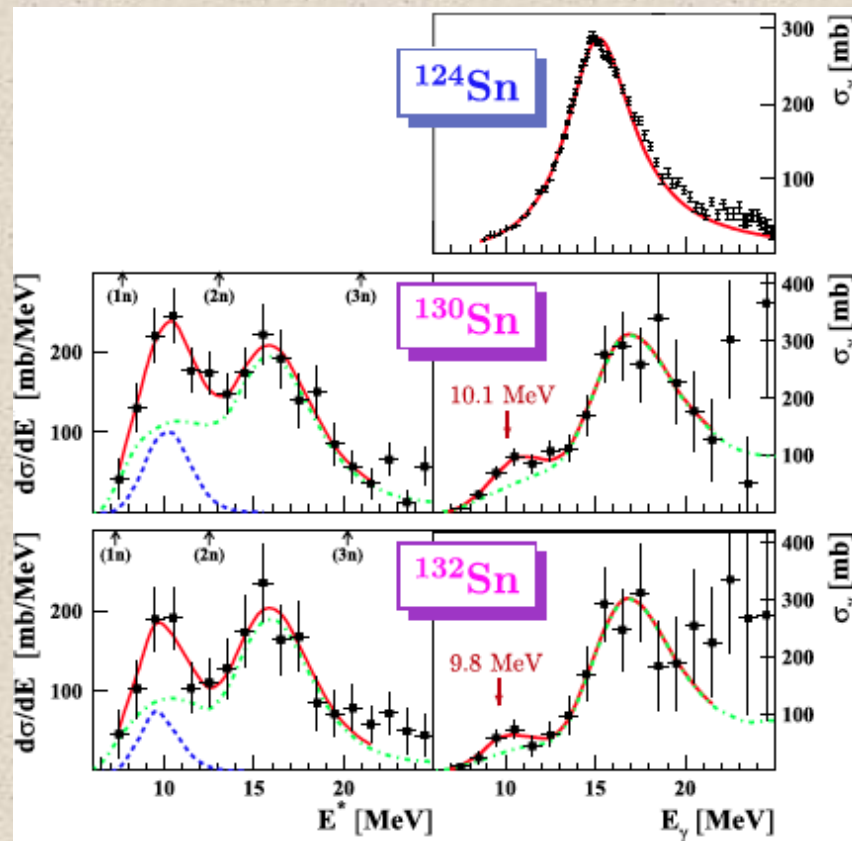
- ☺ We can use thick production targets (100% of energy range) => and high injector beam current (upto 10^{16} s^{-1})
- ☺ long extraction and ionization time (ms)
- ☺ chemistry dependent (disadvantages and some advantages)

In-flight fragmentation

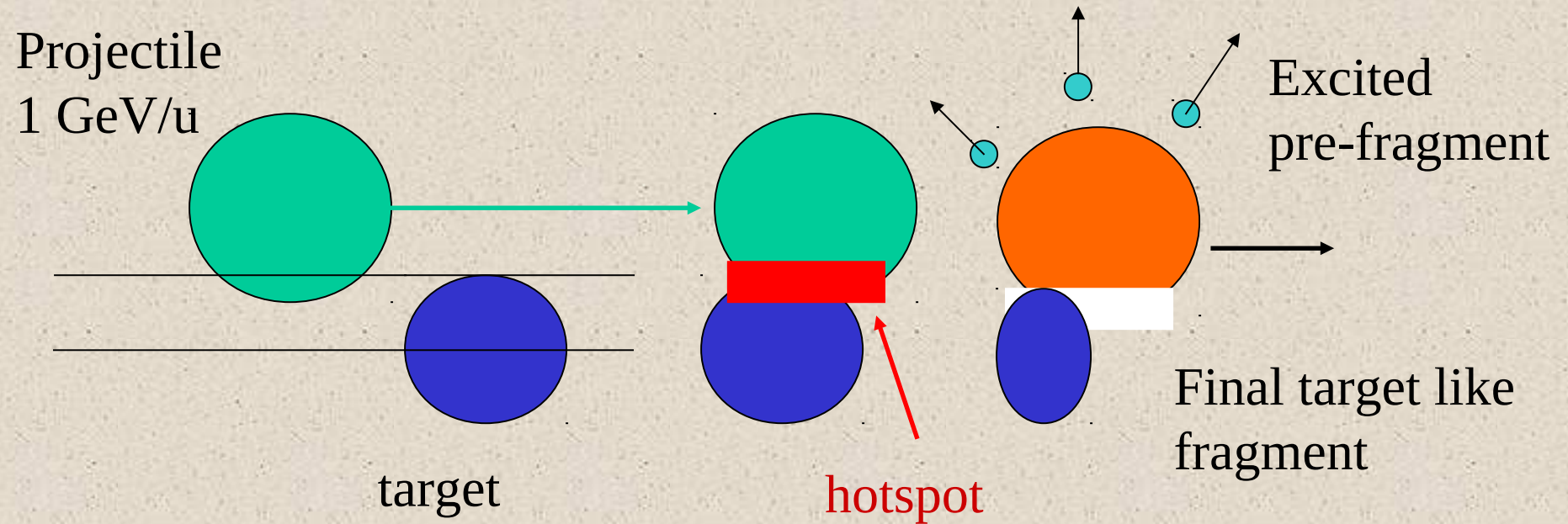


- ☺ short separation+identification time (100 ns)
- ☺ thinner targets (10% of range) => lower beam currents (upto 10^{12} s^{-1})
- ☺ chemistry independent

The energy issue: Some physics can only be reached with relativistic energies

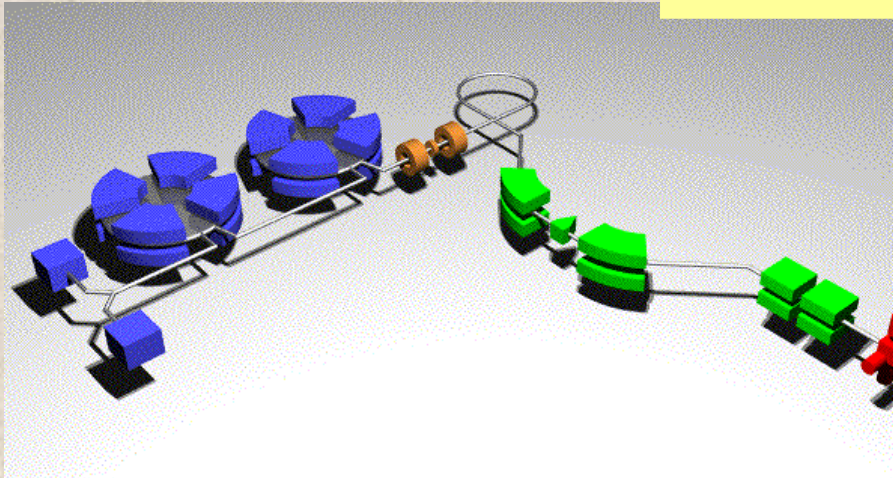


Projectile Fragmentation Reactions

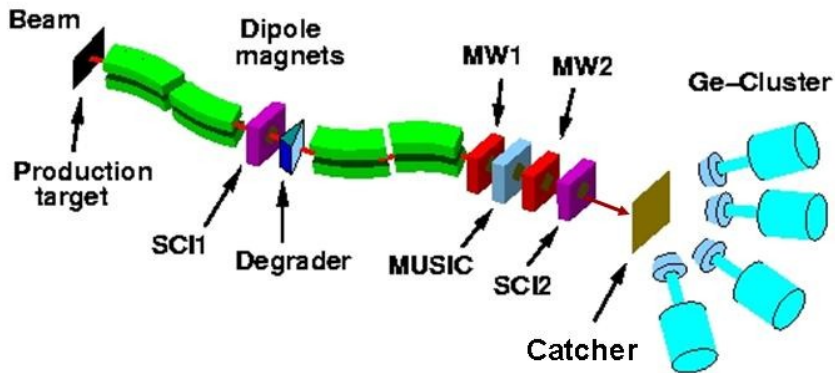


Energy (velocity) of beam > Fermi velocity inside nucleus ~ 30 MeV/u
Can 'shear off' different combinations of protons and neutrons.
Large variety of exotic nuclear species created, all at forward angles with $\sim$ beam velocity.

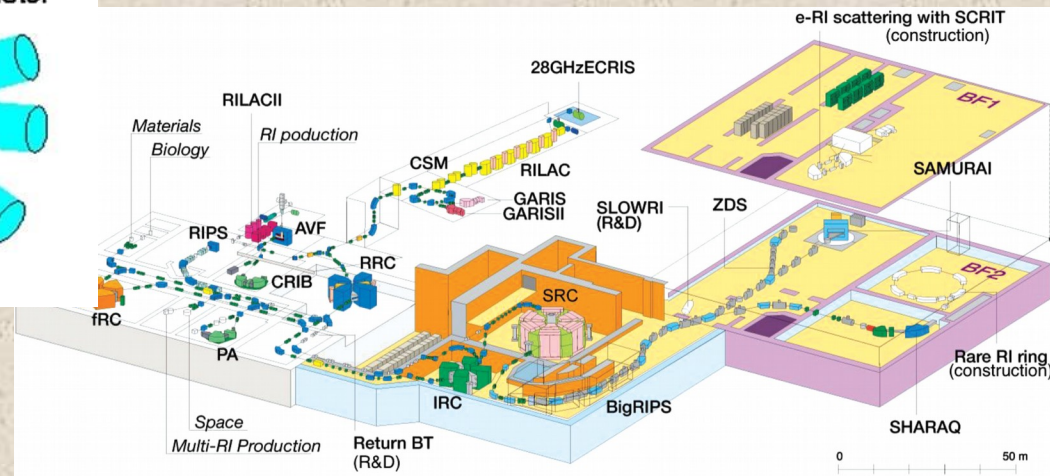
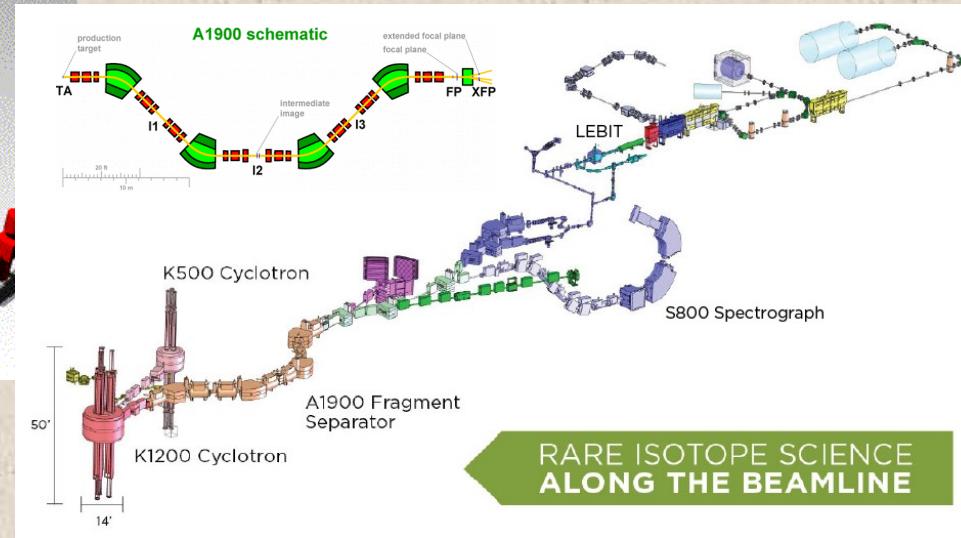
Present Fragment Separators

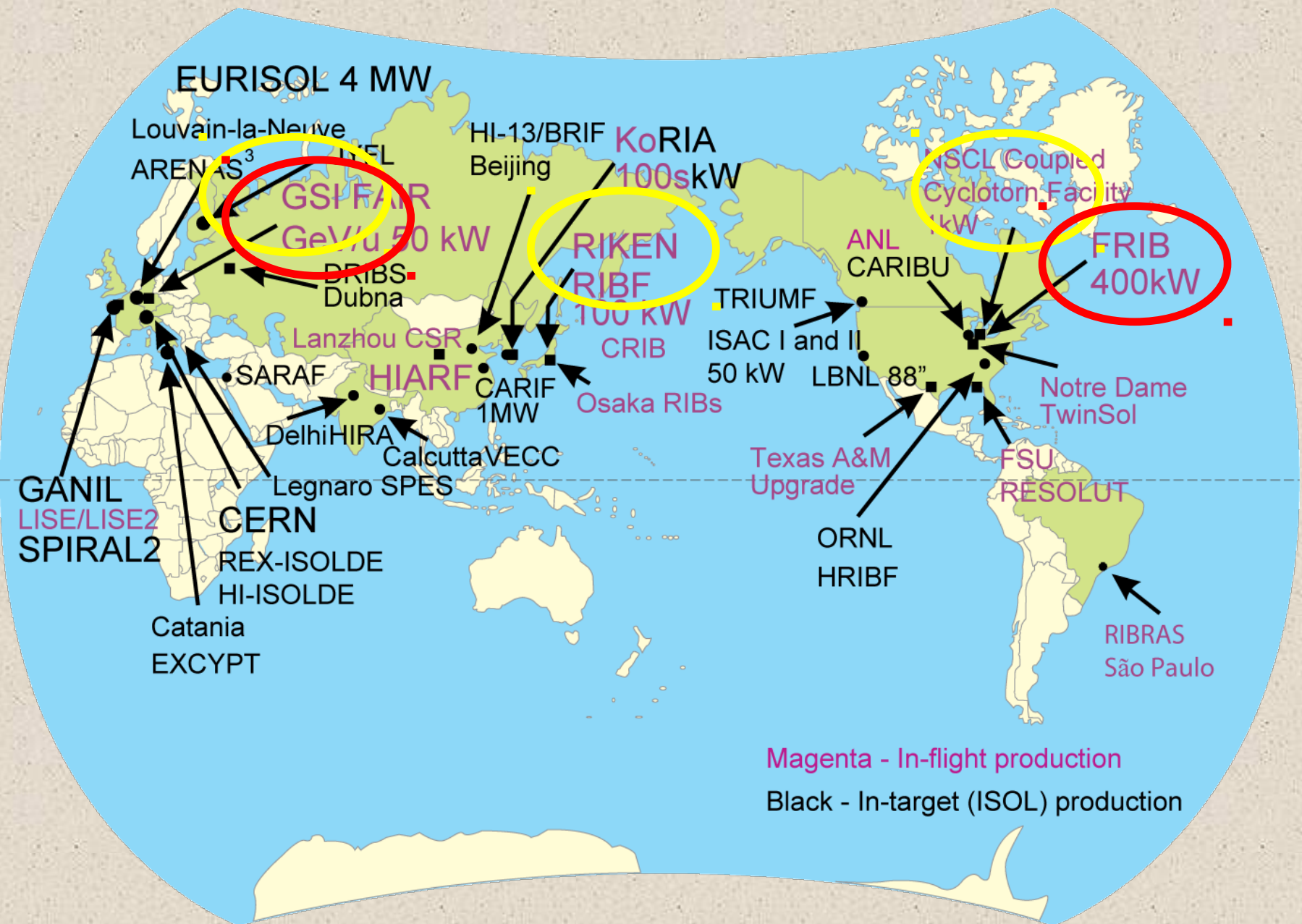


LISE-GANIL



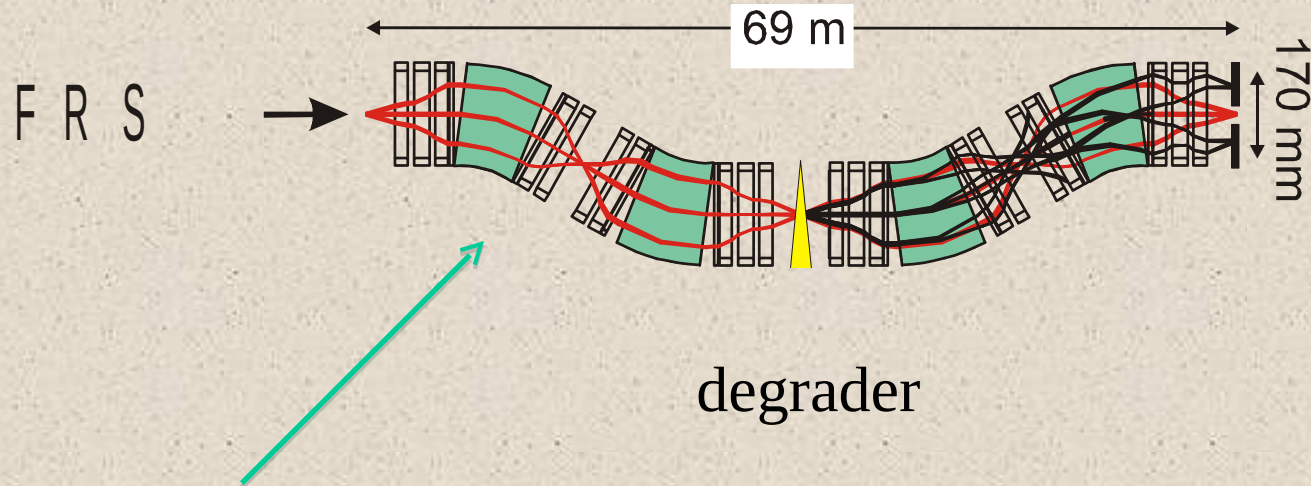
FRS GSI





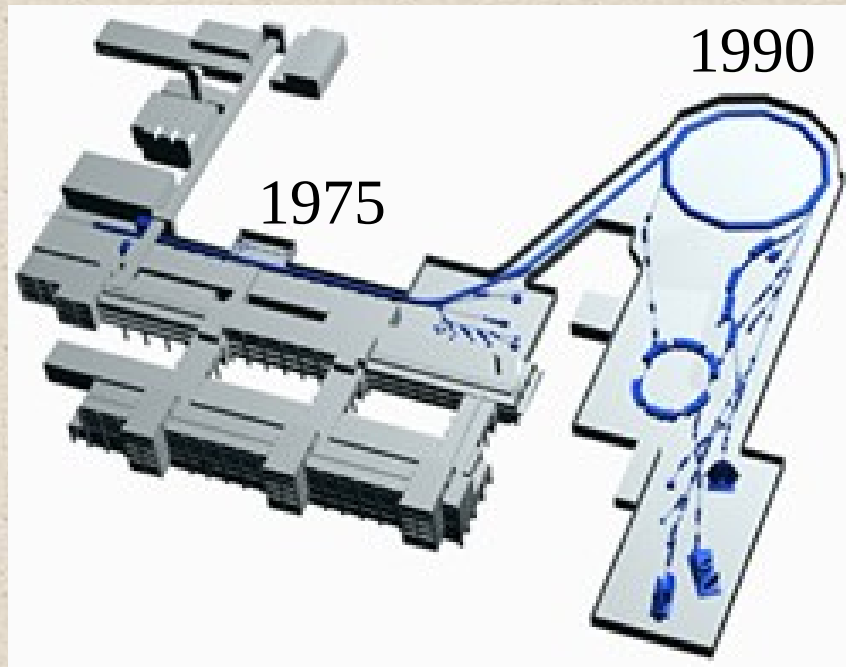
B ρ - Δ E-B ρ Separation Method

Optimised to separate
very fast fragment or
fission products



Ions are selected according
to the momentum to charge ratio
 $Mv/q = Mv/Z = B\rho$. Primary beam is rejected.

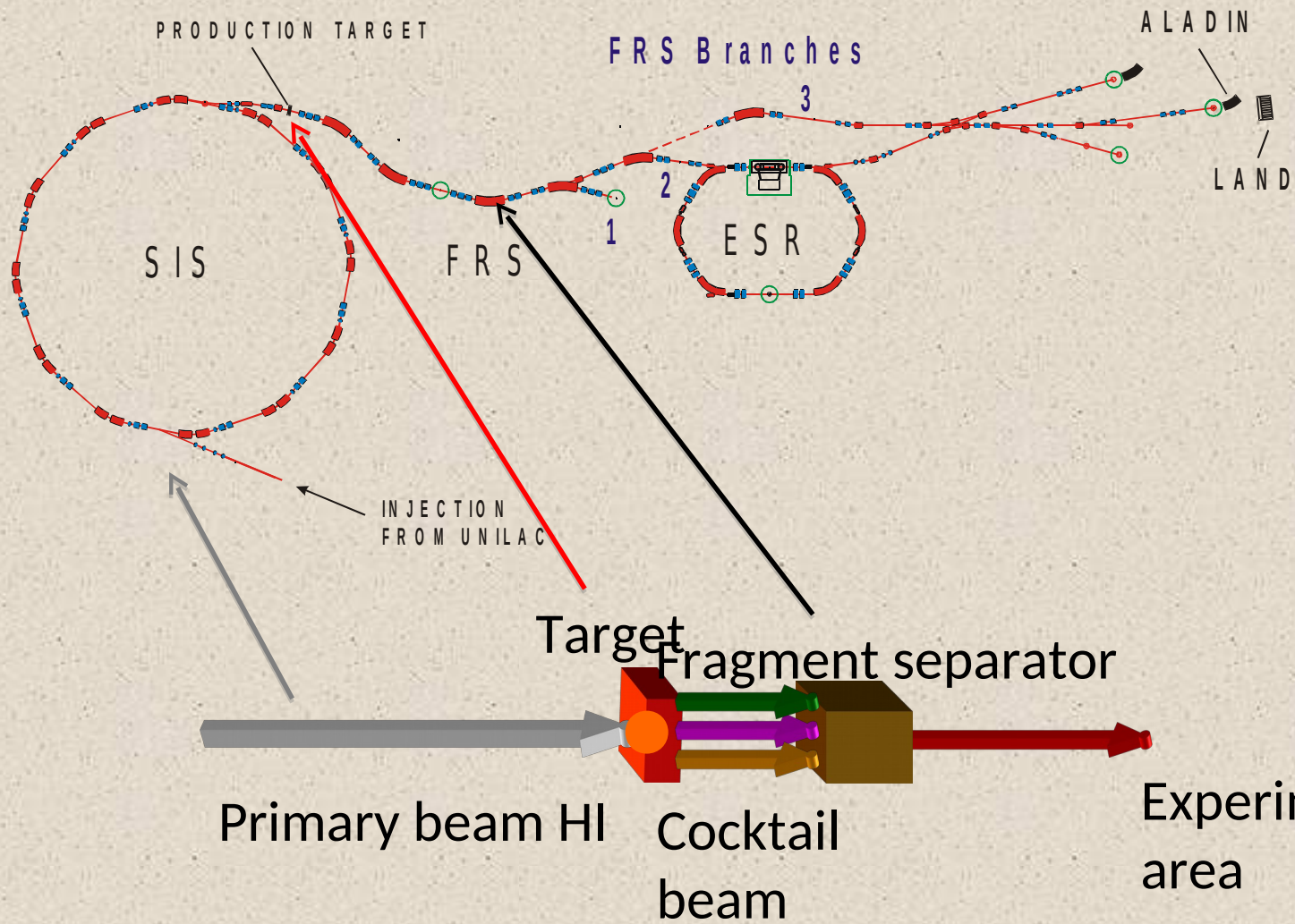
The GSI fragmentation facility



It is possible to accelerate stable Nuclei from H ($Z=1$) to U ($Z=92$), in the UNILAC + Synchrotron.

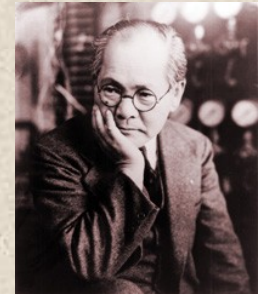
Fragmentation of the primary ion beam:
“cocktail” of ions



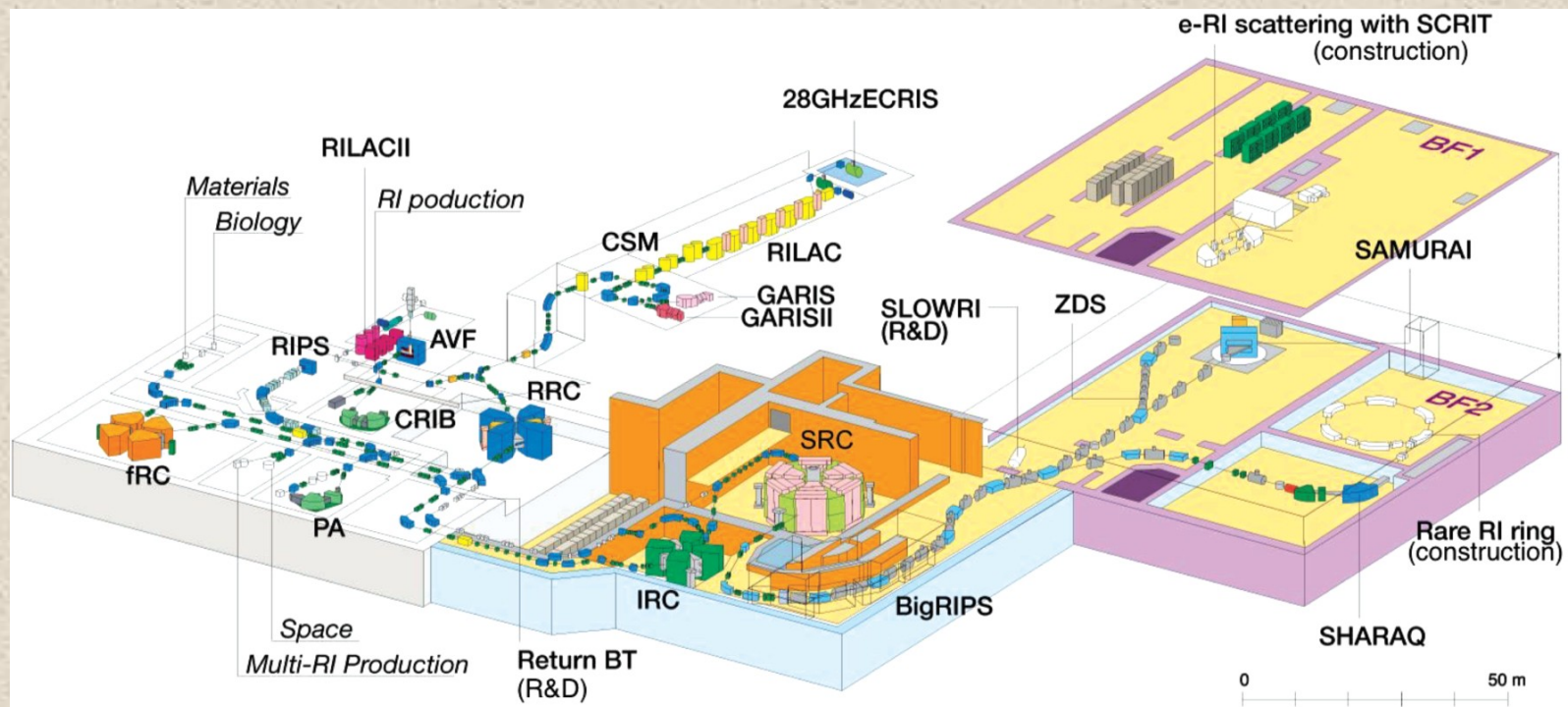


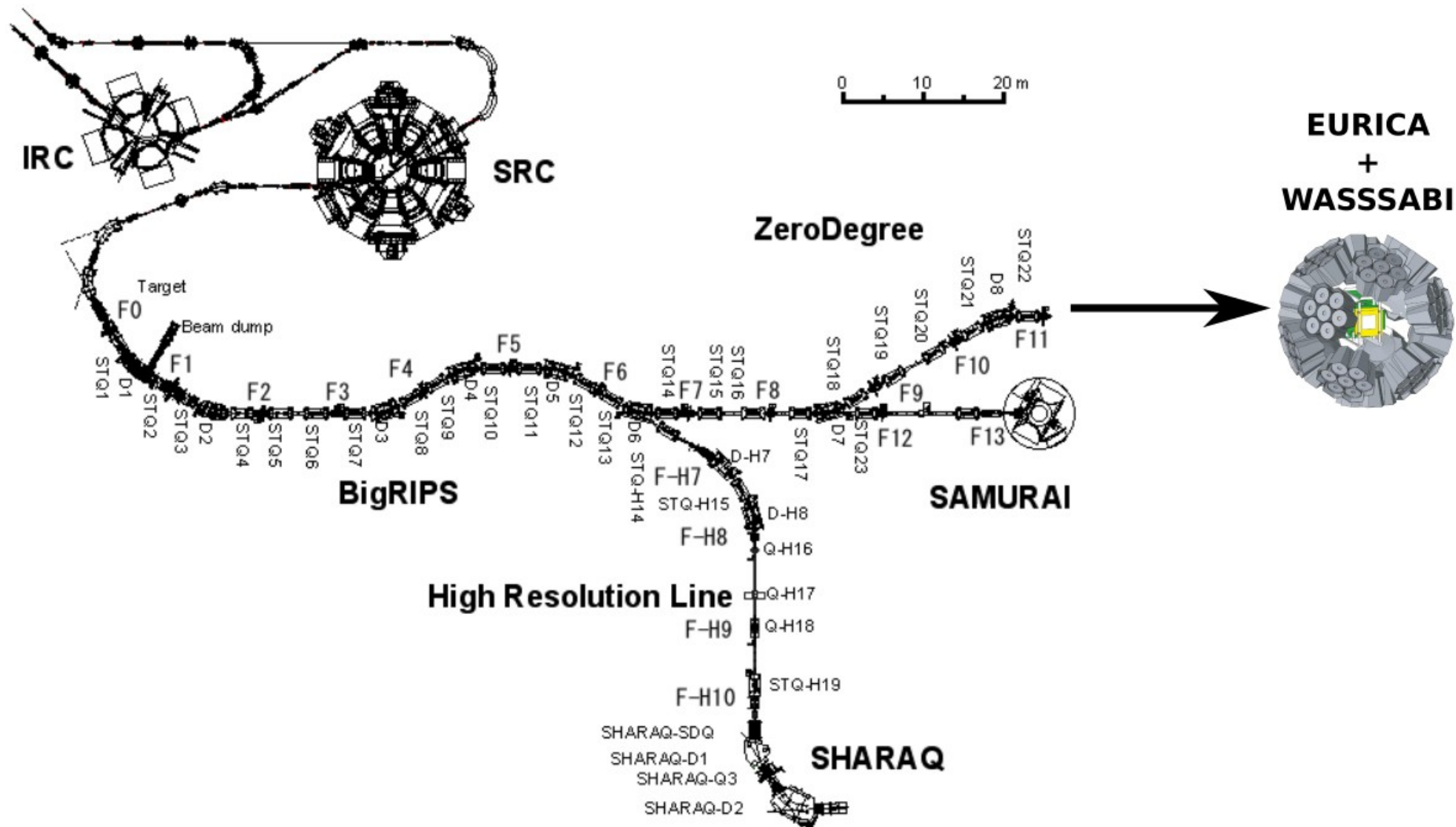
The highest intensity today we find at RIKEN, Nishina Center, this is partially due to the implementation of the superconducting technology

Yoshio Nishina

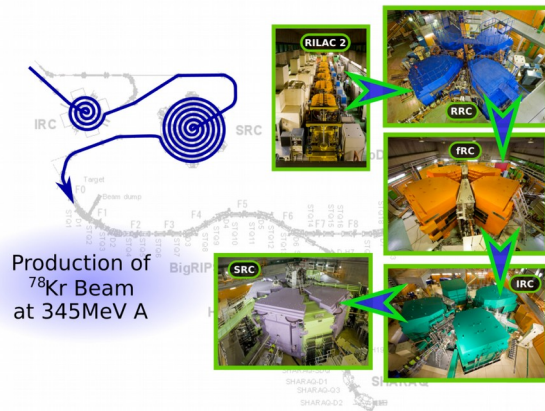


1937, first cyclotron in Japan and second in the World)

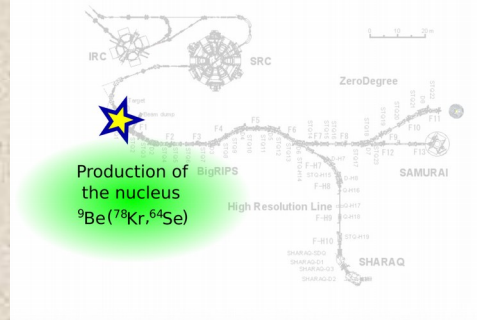




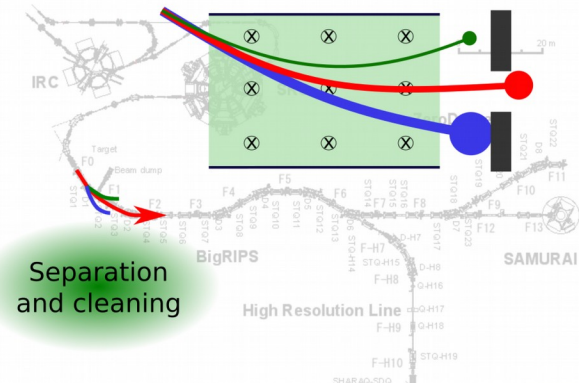
Experimental Setup BigRIPs In-Flight Separator



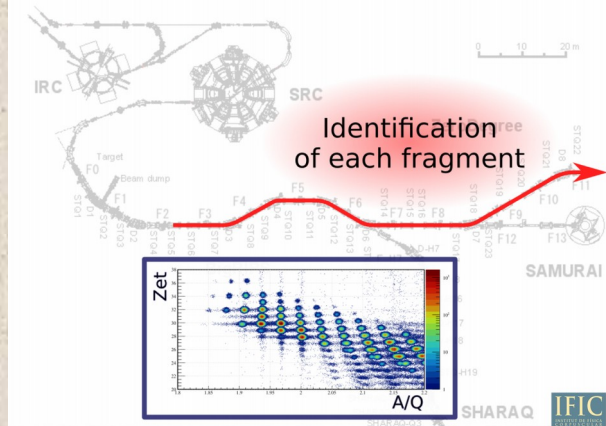
Experimental Setup BigRIPs In-Flight Separator



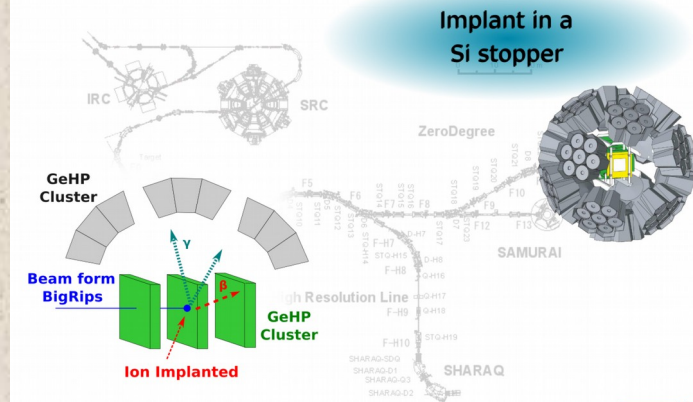
Experimental Setup BigRIPs In-Flight Separator

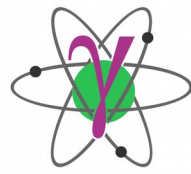


Experimental Setup BigRIPs In-Flight Separator



Experimental Setup BigRIPs In-Flight Separator





Estudios de núcleos exóticos en RIKEN (Japón)

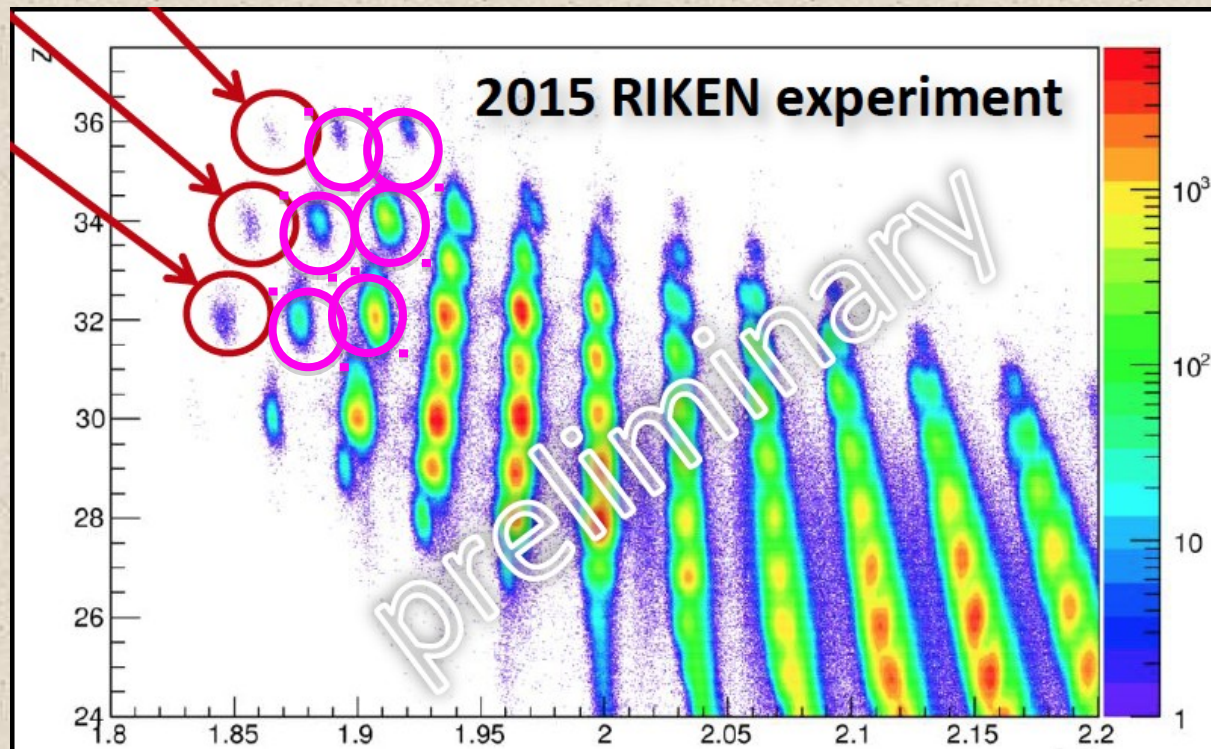
A. Algora et al.

B. Blank et al.

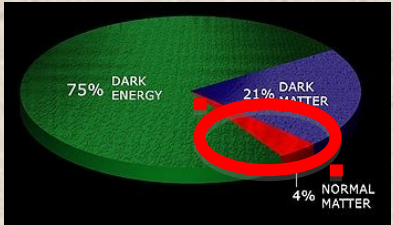
B. Rubio et al.

Descubrimiento de nuevos isótopos

Estudio de los núcleos de Ge, Se, y Kr más exóticos



Alcanzado el límite de Existencia de los núcleos



										Krypton Z=36														69Kr	70Kr	71Kr	72Kr	73Kr	74Kr	75Kr	76Kr	77Kr	78Kr
										Bromine Z=35															70Br	71Br	72Br	73Br	74Br	75Br	76Br	77Br	
Selenium Z=34														64Se	65Se	66Se	67Se	68Se	69Se	70Se	71Se	72Se	73Se	74Se	75Se	76Se							
Arsenic Z=33																		64As	65As	66As	67As	68As	69As	70As	71As	72As	73As	74As	75As				
Germanium Z=32												59Ge	60Ge	61Ge	62Ge	63Ge	64Ge	65Ge	66Ge	67Ge	68Ge	69Ge	70Ge	71Ge	72Ge	73Ge	74Ge						
Gallium Z=31																		60Ga	61Ga	62Ga	63Ga	64Ga	65Ga	66Ga	67Ga	68Ga	69Ga	70Ga	71Ga	72Ga	73Ga		
Zinc Z=30	54Zn	55Zn	56Zn	57Zn	58Zn	59Zn	60Zn	61Zn	62Zn	63Zn	64Zn	65Zn	66Zn	67Zn	68Zn	69Zn	70Zn	71Zn	72Zn														
52Cu	53Cu	54Cu	55Cu	56Cu	57Cu	58Cu	59Cu	60Cu	61Cu	62Cu	63Cu	64Cu	65Cu	66Cu	67Cu	68Cu	69Cu	70Cu	71Cu														
51Ni	52Ni	53Ni	54Ni	55Ni	56Ni	57Ni	58Ni	59Ni	60Ni	61Ni	62Ni	63Ni	64Ni	65Ni	66Ni	67Ni	68Ni	69Ni	70Ni														

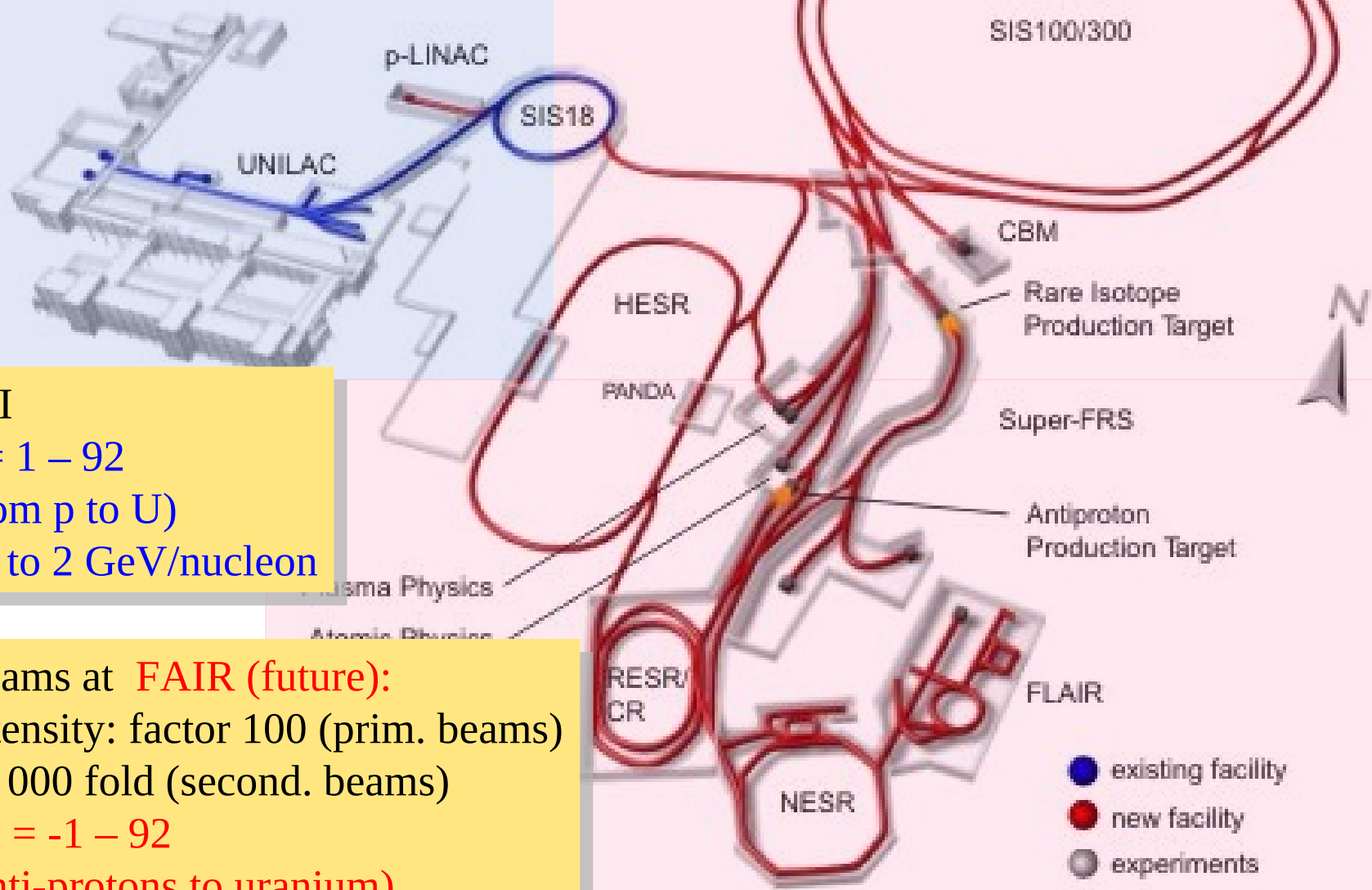
Future (in Europe)



B. Rubio. Master FN, Valencia 2020

Existing GSI facility

New facility



GSI

$Z = 1 - 92$

(from p to U)

Up to 2 GeV/nucleon

Beams at **FAIR (future):**

Intensity: factor 100 (prim. beams)

10 000 fold (second. beams)

$Z = -1 - 92$

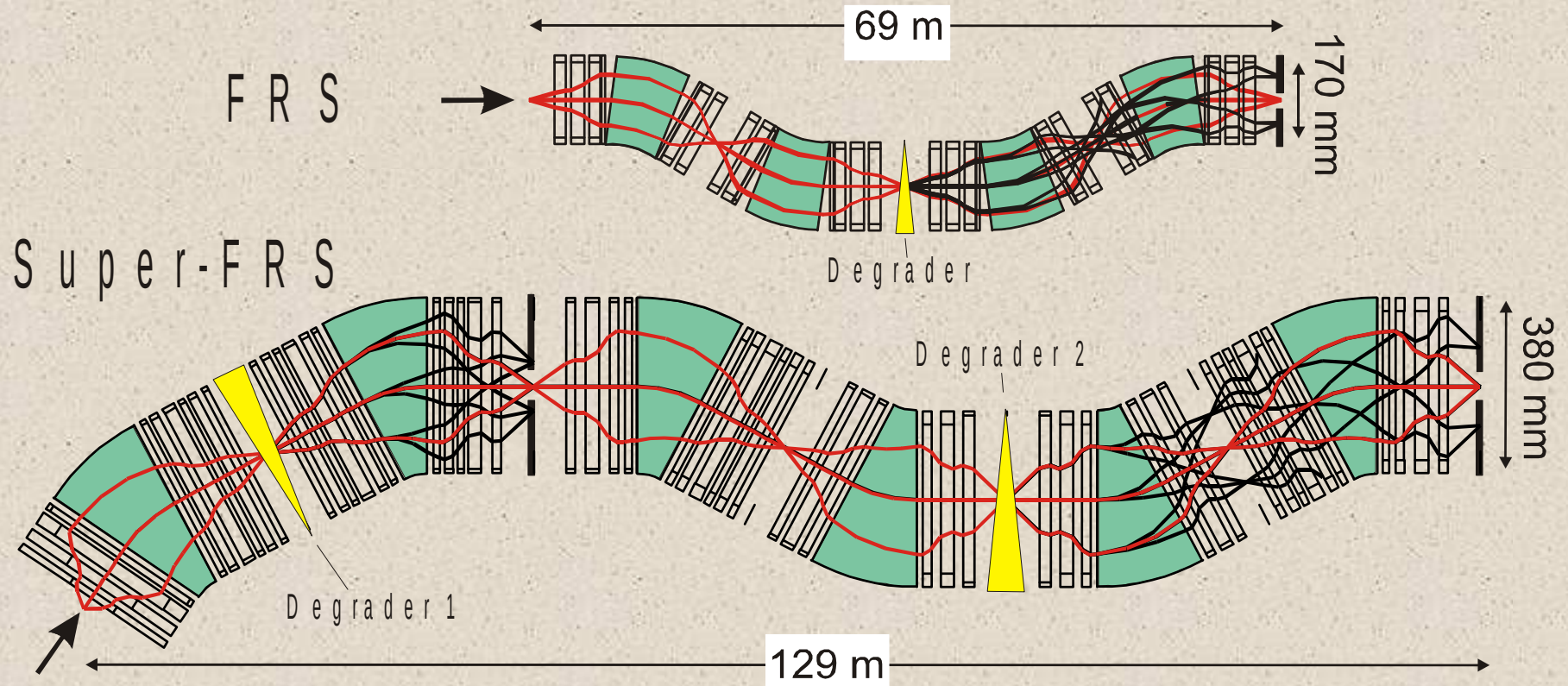
(anti-protons to uranium)

Up to 35 - 45 GeV/u

„full beam cooling“

FRS to Super-FRS

$B\rho$ - ΔE - $B\rho$ Separation Method



Optimised to separate
very fast fragment or
fission products

FAIR Project Progress – Civil Construction



SIS 18 in direction of pLinac
Preparation of western trans
channel (WTK)

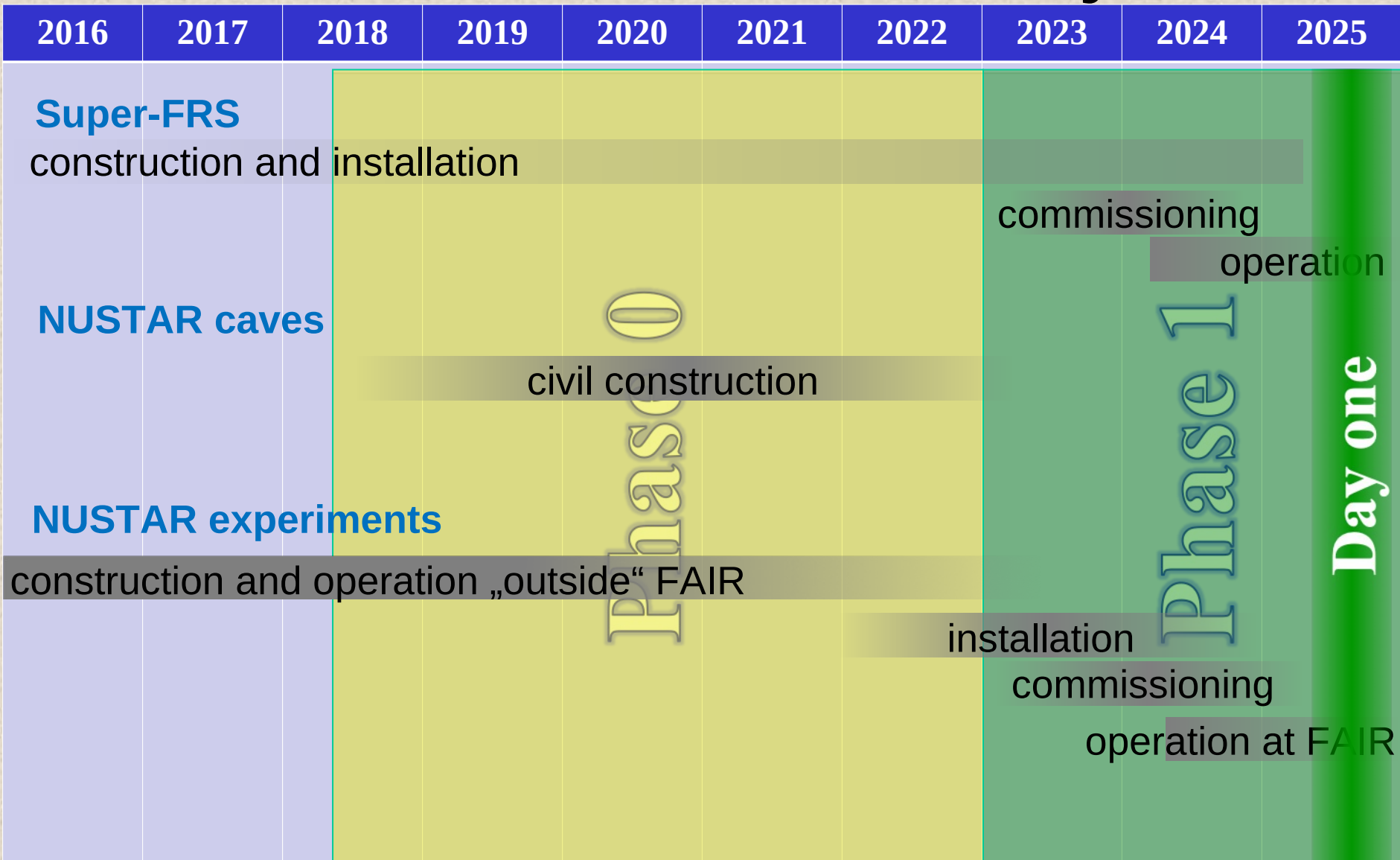
SIS 100:
trench sheeting
(first 8 meters
sheet piling)



FAIR Construction Field



Scenario for NUSTAR: Day one



ISOL and In-Flight facilities-Partners

It is probably true to say that both types of facility but they are complementary.

ISOL

- High intensity beams with ion optics comparable to stable beams
- Easy to manipulate beam energies from keV to 10s of MeV
- Cleaner

In-Flight

- Relativistic beams
- Universal in Z
- Down to very short $T_{1/2}$
- More exotic beams

FIN

difficulties

- extremely low production cross sections
- overwhelming production of unwanted species in the same nuclear reaction,
- very short half lives of the nuclei of interest.

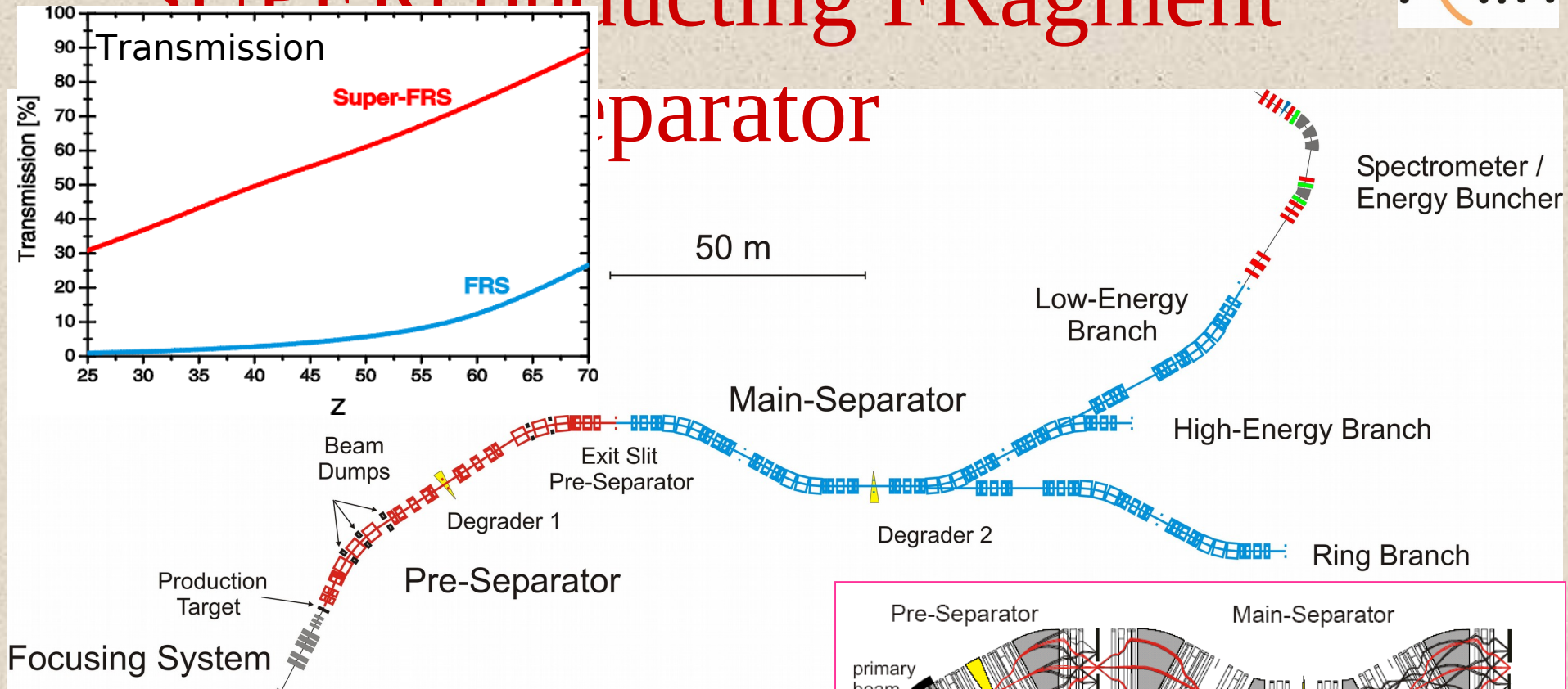
challenges

- As much intensity as possible (good driver, targets to stand the intensity)
- Good separation
- Bring the species to a “clean environment” . Make the process as fast as possible

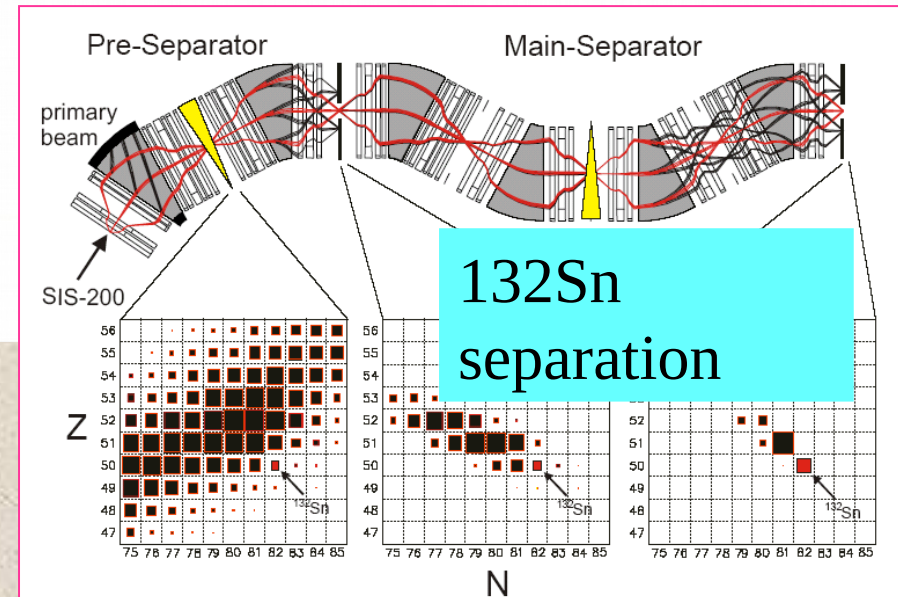
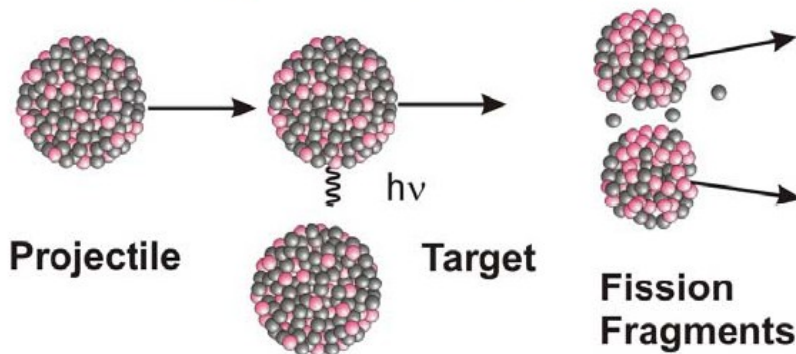


B. Rubio. Master FN, Valencia 2018

STIPER conducting FRagment Separator



Projectile Fission



A good (although a bit outdated) article



<http://iopscience.iop.org/article/10.1088/0031-8949/2013/T152/014023/pdf>

Phys. Scr. **T152** (2013) 014023

Y Blumenfeld *et al*

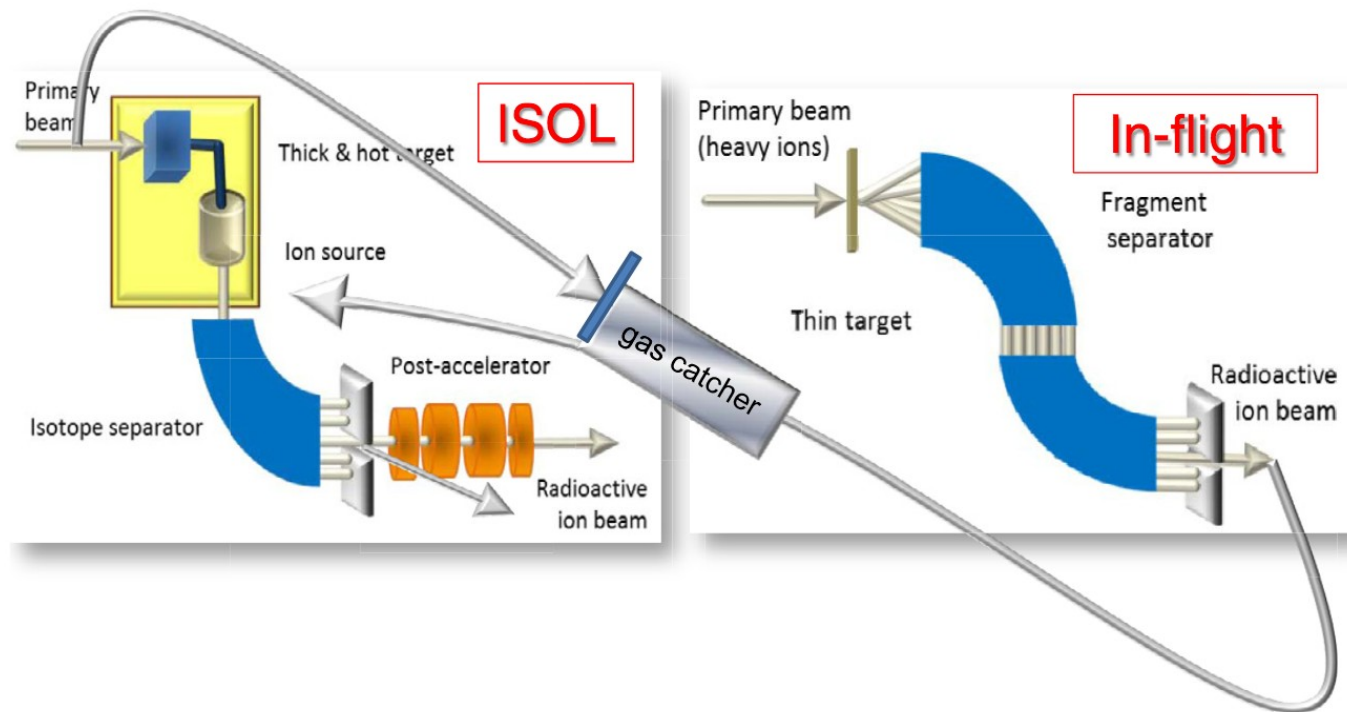


Figure 1. A schematic drawing of the ISOL and in-flight based production methods for RIBs.