

Charged Lepton Flavor Violation in low scale seesaw models

G. H. Tomé*, J. I. Illana*, M. Masip*
G. López†, and P. Roig.†

* CAFPE and Departamento de Física Teórica y del Cosmos, Universidad de Granada

† Departamento de Física, CINVESTAV México.

October/2019

XI CPAN DAYS (OVIEDO)



- Original formulation of the SM (massless neutrinos)

$$-\mathcal{L}_Y = (Y_u)_{ij} \bar{Q}_{L_i} u_{R_j} \tilde{\Phi} + (Y_d)_{ij} \bar{Q}_{L_i} d_{R_j} \Phi + (Y_e)_{ij} \bar{L}_{L_i} e_{R_j} \Phi + h.c. \quad (1)$$

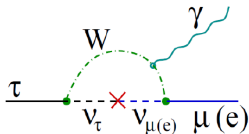
No incorporation of right-handed neutrinos

The minimality construction of eq. (1) implies massless neutrinos and the conservation of both lepton flavor and lepton number.

- **Experimental evidence of neutrino oscillation \Rightarrow LF numbers are not conserved, being required an extended model with tiny neutrino mass.**



- Minimal extension (νSM) \Rightarrow already results in cLFV transitions at one loop level, although strongly suppressed



- $BR(\mu \rightarrow e\gamma) \simeq \frac{\Gamma(\mu \rightarrow e\gamma)}{\Gamma(\mu \rightarrow e\nu\bar{\nu})} = \frac{3\alpha}{32\pi} \left| \sum_{k=1,3} \frac{U_{\mu k} U_{ek}^* m_{\nu k}^2}{m_W^2} \right|^2 \sim 10^{-54}$.
 - T. P. Cheng and L. F. Li, S. T. Petcov, Sov. J. Nucl. Phys. 25, 340 (1977).
- $BR(Z \rightarrow \ell'\ell) \sim 10^{-54}$
 - J. I. Illana and T. Riemann, Phys. Rev. D 63, 053004 (2001).
- $BR(h \rightarrow \ell'\ell) \sim 10^{-55}$
 - E. Arganda, A. M. Curiel, M. J. Herrero and D. Temes, Phys. Rev. D 71, 035011 (2005).
- $BR(\mu^\pm \rightarrow e^\pm e^\pm e^\mp) \sim 10^{-53}$
 - S. T. Petcov, Sov. J. Nucl. Phys. 25, 340 (1977).
- $BR(\tau^\pm \rightarrow \mu^\pm \ell^\pm \ell^\mp) \sim 10^{-54}$
 - G. Hernández-Tomé, G. López Castro and P. Roig, Eur. Phys. J. C 79. (2018)

Far away from any possible detection!



- The νSM incorporates further features
 - **Very tiny Yukawa couplings in order to explain the observed neutrino masses**
 - **Possible invariant Majorana mass terms $\frac{1}{2}MN_R N_R$? ($\Delta L = 2$)**

Type-I Seesaw (High-scale scenario $M \gg m_D$)

- Natural explanation for the tiny ν masses via new dynamics with a heavy scale

Seesaw Relations

In the case of only one generation of heavy neutrinos, the light-heavy mixings are fixed to be very small

$$m_\nu (\sim 0,05 \text{ eV}) \simeq \frac{m_D^2 (\sim 10^4 \text{ GeV}^2)}{M (\sim 10^{14} \text{ GeV})}, \quad \theta (\sim 10^{-6}) \simeq \sqrt{\frac{m_\nu}{M}}.$$

with $m_D = y_\nu v$.

New Heavy effects decoupled!
Again, cLFV very suppressed!

Low Scale Seesaw Models

- Inverse Seesaw
 - [PhysRevD.34.1642](#)
- Linear Seesaw
 - [PhysRevLett.95.161801](#)

Based on approximate symmetries
or ansatz on the neutrino matrix.

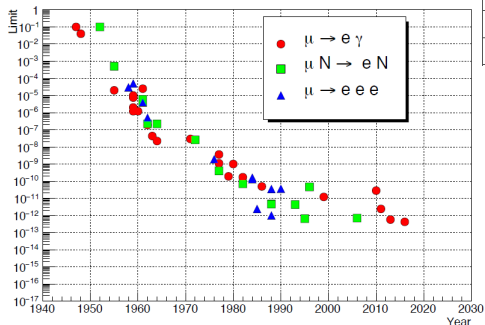
More than one sterile neutrino!

Masses and mixings can be taken as free parameters, constrained
only by experimental limits!

A new window to new Physics via cLFV transitions and $\Delta L = 2$
processes!



So far, no evidence of cLFV!



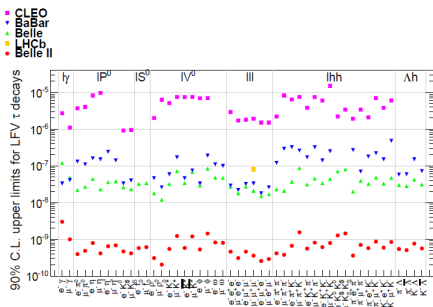
Reaction	Present Limit 90% C.L	Future Sensitivity
$\mu \rightarrow e \gamma$	4.2×10^{-13}	4×10^{-14}
$\mu - e$ (Au)	7.0×10^{-13}	-
$\mu - e$ (Ti)	4.3×10^{-12}	10^{-18}
$\mu \rightarrow e e \bar{e}$	1.0×10^{-12}	10^{-16}
$\tau \rightarrow \mu \gamma$	4.4×10^{-8}	1×10^{-9}

Future experiments MEG-II, Mu3e, PRISM and COMET

- A. M. Baldini *et al.*, arXiv:1301.7225
- A. Blondel *et al.*, arXiv:1301.6113.
- A. Alekou *et al.*, arXiv:1310.0804
- Y. Kuno [COMET Collaboration], PTEP **2013**, 022C01 (2013).



Experimental limits



Belle-II will be improved until two orders of magnitude when they achieve their maximum luminosity

- E. Kou *et al.* [Belle-II Collaboration], arXiv:1808.10567 [hep-ex]

Reaction	Present Limit 90% C.L	Future Sensitivity
$\tau \rightarrow e\gamma$	3.3×10^{-8}	3×10^{-9}
$\tau \rightarrow \mu\gamma$	4.4×10^{-8}	1×10^{-9}
$\tau \rightarrow ee\bar{e}$	2.7×10^{-8}	$\sim (2 - 5) \times 10^{-10}$
$\tau \rightarrow \mu\mu\bar{\mu}$	2.1×10^{-8}	$\sim (2 - 5) \times 10^{-10}$
$\tau \rightarrow \mu e\bar{e}$	1.8×10^{-8}	$\sim (2 - 5) \times 10^{-10}$
$\tau \rightarrow e\mu\bar{\mu}$	2.7×10^{-8}	$\sim (2 - 5) \times 10^{-10}$
$\tau \rightarrow \mu\mu\bar{e}$	1.7×10^{-8}	$\sim (2 - 5) \times 10^{-10}$
$\tau \rightarrow ee\bar{\mu}$	1.5×10^{-8}	$\sim (2 - 5) \times 10^{-10}$

● Other channels

Reaction	Present Limit 95% C.L.	Future Sensitivity
$Z \rightarrow \mu e$	7.3×10^{-7} [34]	10^{-13} [44]
$Z \rightarrow \tau e$	9.8×10^{-6} [35]	-
$Z \rightarrow \tau \mu$	1.2×10^{-5} [36]	-

FCC-ee would improve the current limits to $Z \rightarrow \bar{\ell}\ell'$ by about four orders of magnitude

- Phys. Rev. D **95**, no. 7, 075028 (2017)

cLFV in low Scale Seesaw Models



We consider an interfamily seesaw scenario where the neutrino sector is formed by

- n_R of right-handed (Majorana) neutrino singlets
- n_G left-handed neutrinos ($\nu_L^0 = \nu_e, \nu_\mu, \nu_\tau, \dots$)

Charged and neutral currents involving neutrinos

$$\mathcal{L}_{W^\pm} = -\frac{g}{\sqrt{2}} W_\mu^- \sum_{i=1}^{n_G} \sum_{j=1}^{n_G+n_R} B_{\ell_{ij}} \bar{\ell}_i \gamma^\mu P_L \nu_j + \text{h.c.}, \quad (2)$$

$$\mathcal{L}_Z = -\frac{g}{4c_W} Z_\mu \sum_{i,j=1}^{n_G+n_R} \bar{\nu}_i \gamma^\mu (C_{ij} P_L - C_{ij}^* P_R) \nu_j, \quad (3)$$

$$\mathcal{L}_{G^\pm} = -\frac{g}{\sqrt{2}m_W} G^- \sum_{i=1}^{n_G} \sum_{j=1}^{n_G+n_R} B_{\ell_{ij}} \bar{\ell}_i (m_{\ell_i} P_L - m_j P_R) \nu_j + \text{h.c.}, \quad (4)$$

- [A. Ilakovac and A. Pilaftsis, Nucl. Phys. B 437, 491 \(1995\)](#)
- [F. del Aguila and J. A. Aguilar-Savedra, Nucl. Phys. B 813, 22 \(2009\)](#)



In the case of $n_R = 2$, the neutrino mass spectrum consists of three light Majorana neutrinos which have been identified with the three known neutrinos, ν_e , ν_μ and ν_τ and two heavy ones denoted by N_1 and N_2 . The light-heavy mixings correspond to

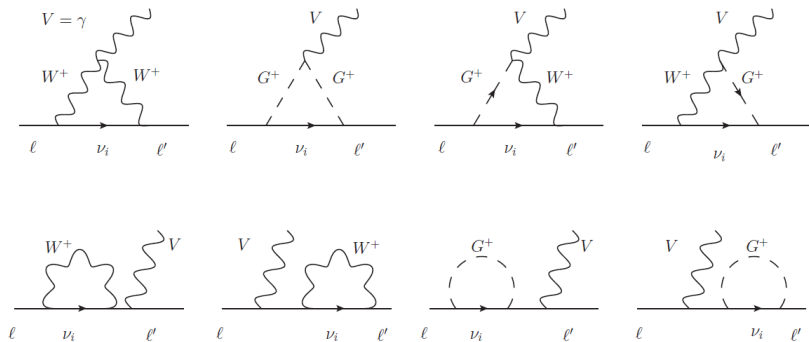
$$s_{\nu k}^2 = \sum_i |B_{\ell_k N_i}|^2. \quad (5)$$

In the simple of $n_R = 2$, and considering the light neutrinos massless

$$B_{\ell_k N_1} = \frac{r^{\frac{1}{4}}}{\sqrt{1+r^{\frac{1}{2}}}} s_{\nu k}, \quad B_{\ell_k N_2} = \frac{i}{\sqrt{1+r^{\frac{1}{2}}}} s_{\nu k}, \quad (6)$$

$$C_{N_1 N_1} = \frac{r^{\frac{1}{2}}}{1+r^{\frac{1}{2}}} \sum_{k=1}^{n_G} s_{\nu k}^2, \quad C_{N_1 N_2} = -C_{N_2 N_1} = \frac{i r^{\frac{1}{4}}}{1+r^{\frac{1}{2}}} \sum_{k=1}^{n_G} s_{\nu k}^2, \quad (7)$$

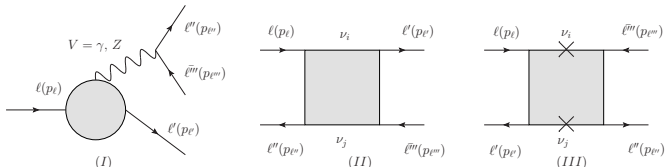
$$C_{N_2 N_2} = \frac{1}{1+r^{\frac{1}{2}}} \sum_{k=1}^{n_G} s_{\nu k}^2, \quad r = \left(\frac{m_{N_2}}{m_{N_1}} \right)^2. \quad (8)$$



- $\ell \rightarrow \ell' \gamma$ ($\mu \rightarrow e \gamma$, $\tau \rightarrow e \gamma$, $\tau \rightarrow \mu \gamma$)

$$\Gamma(\ell \rightarrow \ell' \gamma) = \frac{\alpha}{2} m_\ell^3 (|F_M^\gamma|^2 + |F_E^\gamma|^2). \quad (9)$$

• $\ell \rightarrow \ell' \ell'' \bar{\ell}'''$



	$\ell \rightarrow \ell' \ell'' \bar{\ell}'''$
Type 1 ($\ell' = \ell''$ and $\ell'' = \ell'''$)	$\mu \rightarrow ee\bar{e}$ $\tau \rightarrow \mu\mu\bar{\mu}$ $\tau \rightarrow ee\bar{e}$
Type 2 ($\ell' \neq \ell''$ and $\ell'' = \ell'''$)	$\tau \rightarrow \mu e\bar{e}$ $\tau \rightarrow e\mu\bar{\mu}$
Type 3 ($\ell' = \ell''$ and $\ell'' \neq \ell'''$)	$\tau \rightarrow \mu\mu\bar{e}$ $\tau \rightarrow ee\bar{\mu}$

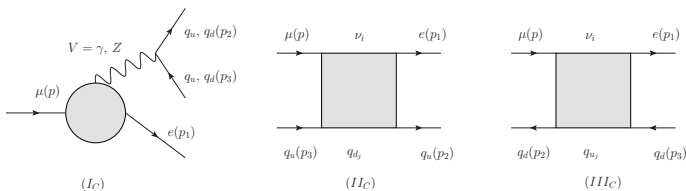
• Type-1

$$\Gamma = \frac{\alpha^2 m_\ell^5}{96\pi} \left\{ 3|A_L|^2 + 2|A_R|^2 \left(8 \log \left(\frac{m_\ell}{m_{\ell''}} \right) - 13 \right) + 2|F_{LL}|^2 + |F_{LR}|^2 + \frac{1}{2}|B_L|^2 \right. \\
 \left. - \left[6A_L A_R^* - (A_L - 2A_R)(2F_{LL}^* + F_{LR}^* + B_L^*) - F_{LL} B_L^* + \text{h.c.} \right] \right\}, \quad (10)$$

- $Z \rightarrow \bar{\ell}\ell'$ ($Z \rightarrow \mu\bar{e}$, $Z \rightarrow \tau\bar{\mu}$, $Z \rightarrow \tau\bar{e}$)

$$\Gamma(Z \rightarrow \bar{\ell}\ell') = \frac{\alpha}{3} m_Z |\mathcal{A}_L^Z|^2. \quad (11)$$

- $\mu N - eN$ Conversion



$$\Gamma(\mu \rightarrow e) = \frac{\alpha^5 Z_{\text{eff}}^4}{Z} F(q)^2 m_\mu^5 \left| 2Z \left(A_1^L + A_2^R \right) - (2Z + N) \bar{B}_{1u}^L - (Z + 2N) \bar{B}_{1d}^L \right|^2, \quad (12)$$

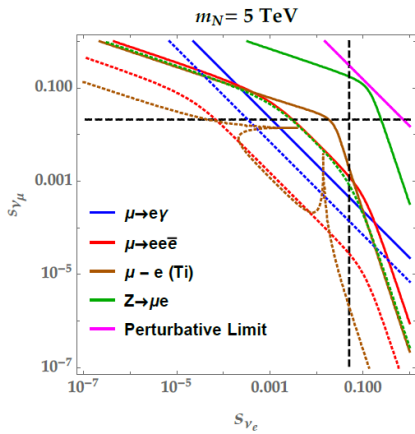
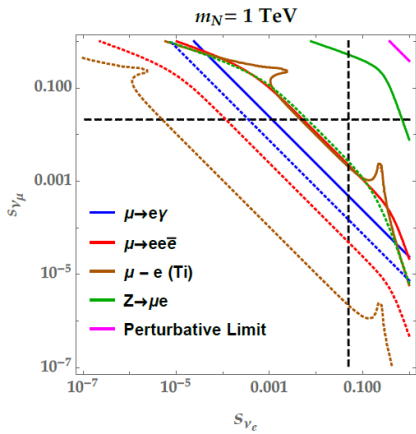
Limits on light heavy-mixings

- Indirect limits on the light heavy mixing angles

$$s_{\nu_e} < 0,050 (0,031), \quad s_{\nu_\mu} < 0,021 (0,011), \quad s_{\nu_\tau} < 0,075 (0,044). \quad (13)$$

- E. Fernández-Martínez, J. Hernández-García and J. López-Pavón, JHEP **1608**, 033 (2016)

- Direct limits from cLFV

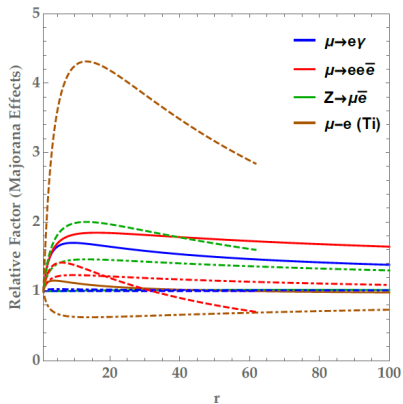
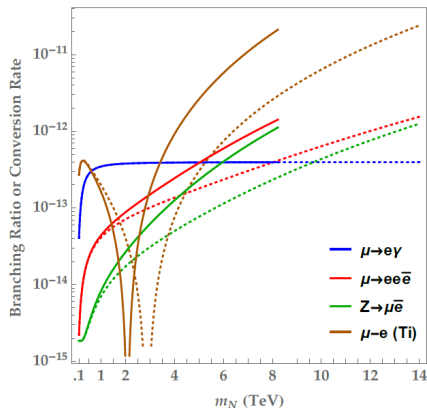


Genuine Majorana effects $\mu - e$ transitions

Perturbative Unitarity condition

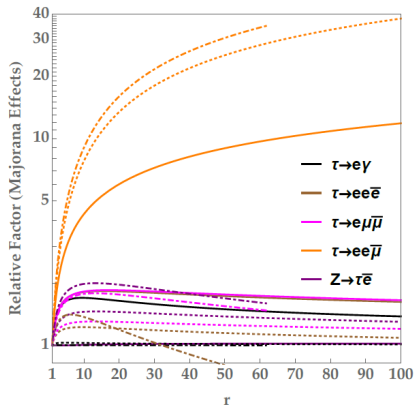
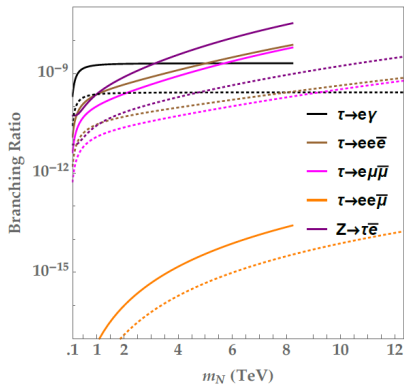
$$Y_{N_\alpha} = \frac{\sqrt{2m_{N_1}m_{N_2}}}{v} s_{\nu_\alpha} < \sqrt{4\pi}.$$

Maximum predictions for $\mu - e$ transitions



Genuine Majorana effects $\tau - e$ transitions

- Maximum predictions for $\tau - e$ transitions



- CLFV processes are strongly suppressed in the νSM and the usual *Type – I* seesaw model. However low scale seesaw models offer a possibility for those effects.
- We have obtained updated estimations for some of the most attractive cLFV processes in low scale seesaw models. Some of these values are at the reach of future experiments.
- We focused on the differences between the degenerate and non-degenerate (Genuine Majorana effects) case of two Majorana neutrinos.
- The no observation of such effects set strong limits on the parameter-space of these hypothetical scenarios.

Thank you!

Back up



The dimension of the rectangular B mixing matrix is $n_G \times (n_G + n_R)$, whereas the C matrix has dimension $(n_G + n_R) \times (n_G + n_R)$; they are determined as follows

$$B_{\ell ij} = \sum_{k=1}^{n_G} (U_{ki}^\ell)^* U_{kj}^\nu, \quad C_{ij} = \sum_{k=1}^{n_G} (U_{ki}^\nu)^* U_{kj}^\nu, \quad (14)$$

where U_{ij}^ℓ and $B_{\ell ij}$ represent the PMNS mixing matrix and its generalized version, respectively.

In the simple case of $n_R = 2$,

$$B_{\ell_k N_1} = \frac{r^{\frac{1}{4}}}{\sqrt{1+r^{\frac{1}{2}}}} s_{\nu k}, \quad B_{\ell_k N_2} = \frac{i}{\sqrt{1+r^{\frac{1}{2}}}} s_{\nu k}, \quad (15)$$

$$C_{N_1 N_1} = \frac{r^{\frac{1}{2}}}{1+r^{\frac{1}{2}}} \sum_{k=1}^{n_G} s_{\nu k}^2, \quad C_{N_1 N_2} = -C_{N_2 N_1} = \frac{ir^{\frac{1}{4}}}{1+r^{\frac{1}{2}}} \sum_{k=1}^{n_G} s_{\nu k}^2, \quad (16)$$

$$C_{N_2 N_2} = \frac{1}{1+r^{\frac{1}{2}}} \sum_{k=1}^{n_G} s_{\nu k}^2. \quad (17)$$

The matrices B and C satisfy the following identities, which are crucial to maintaining the renormalizability of the model

$$\sum_{k=1}^{n_G+n_R} B_{\ell_1 k} B_{\ell_2 k}^* = \delta_{\ell_1 \ell_2}, \quad \sum_{k=1}^{n_G+n_R} C_{ik} C_{jk}^* = \sum_{k=1}^{n_G} B_{\ell_k i} B_{\ell_k j}^* = C_{ij}, \quad (18)$$

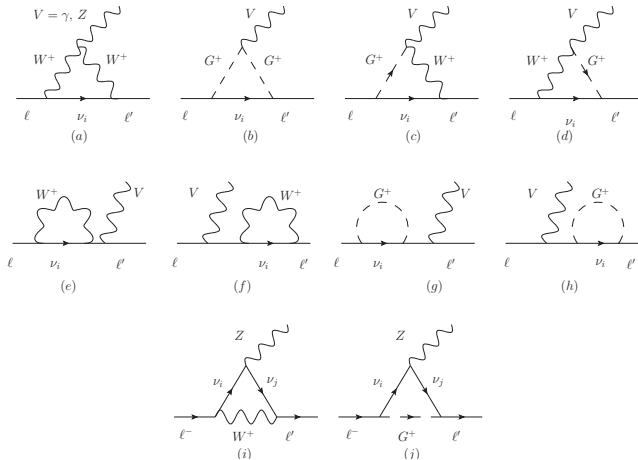
$$\sum_{k=1}^{n_G+n_R} B_{\ell_i k} C_{kj} = B_{\ell_i j} \quad (19)$$

$$\sum_{k=1}^{n_G+n_R} m_k C_{ik} C_{jk} = \sum_{k=1}^{n_G+n_R} m_k B_{\ell_k i} C_{kj}^* = \sum_{k=1}^{n_G+n_R} m_k B_{\ell_1 k} B_{\ell_2 k} = 0. \quad (20)$$

- A. Ilakovac and A. Pilaftsis, Nucl. Phys. B 437, 491 (1995)



● Effective $V\ell\ell'$ vertex generated in the presence of massive neutrinos



The dipole form factors relevant in the $\ell \rightarrow \ell' \gamma$ decay are given by

$$\begin{aligned}
 F_M^\gamma(0) = -iF_E^\gamma(0) &= \frac{m_\ell \alpha_W}{16\pi m_W^2} \sum_i^{n_G+n_R} B_{\ell i}^* B_{\ell' i} f_M^\gamma(x_i), \\
 &= \frac{m_\ell \alpha_W}{16\pi m_W^2} \sum_{N_i}^{n_R} B_{\ell N_i}^* B_{\ell' N_i} [f_M^\gamma(x_{N_i}) - f_M^\gamma(0)], \quad (21)
 \end{aligned}$$

where we have defined $x_i = m_i^2/m_W^2$, and

$$f_M^\gamma(x) = \frac{3x^3 \log(x)}{2(x-1)^4} - \frac{2x^3 + 5x^2 - x}{4(x-1)^3} + \frac{5}{6}, \quad f_M^\gamma(0) = \frac{5}{6}. \quad (22)$$



In the $\ell \rightarrow \ell' \ell'' \bar{\ell}'''$ decays, the photon is off-shell, as a consequence, they also involve the vector and axial-vector F_V^γ and F_A^γ form factors, but since $F_V = F_A$ it turns out convenient to define

$$\begin{aligned}
 2F_V^\gamma = F_L^\gamma(q^2) &= \frac{\alpha_W}{8\pi m_W^2} \sum_i^{n_G+n_R} B_{\ell i}^* B_{\ell' i} f_L^\gamma(q^2, x_i), \\
 &= \frac{\alpha_W}{8\pi m_W^2} \sum_{N_i}^{n_R} B_{\ell N_i}^* B_{\ell' N_i} [f_L^\gamma(q^2, x_i) - f_L^\gamma(q^2, 0)], \quad (23)
 \end{aligned}$$

where

$$f_L^\gamma(q^2, x) = \frac{q^2 x^2 (x^2 - 10x + 12) \log(x)}{6(x-1)^4} + \frac{q^2 (7x^3 - x^2 - 12x)}{12(x-1)^3} - \frac{5q^2}{9} + 2m_W^2 \Delta, \quad (24)$$

$$f_L^\gamma(q^2, 0) = -\frac{5q^2}{9} + 2m_W^2 \Delta, \quad (25)$$

and $\Delta = \frac{1}{\epsilon} - \gamma_E + \log(4\pi) + \log\left(\frac{\mu^2}{m_W^2}\right)$ stands for an ultraviolet divergence.

In the approximation where masses and momenta of the external particles are neglected, the effective $Z\ell\ell'$ is given in terms of only one form factor associated with a $V - A$ Lorentz structure

$$A_L^Z = \frac{\alpha_W}{8\pi s_W c_W} \sum_{i,j}^{n_G+n_R} B_{\ell i}^* B_{\ell' j} [F_Z(x_i)\delta_{ij} + C_{ij}^* G_Z(x_i, x_j) + C_{ij}\sqrt{x_i x_j} H_Z(x_i, x_j)], \quad (26)$$

$$F_Z(x) = \frac{5x^2 \log(x)}{2(x-1)^2} - \frac{5x}{2(x-1)} + \frac{1}{4} - \Delta \left(\frac{5}{2} - 2s_W^2 \right), \quad (27)$$

$$G_Z(x, y) = \frac{1}{2(x-y)} \left(\frac{x^2(y-1)\log(x)}{(x-1)} - \frac{(x-1)y^2\log(y)}{(y-1)} \right) + \frac{1}{2} \left(\Delta - \frac{1}{2} \right), \quad (28)$$

$$H_Z(x, y) = \frac{1}{4(x-y)} \left(\frac{(x-4)x\log(x)}{x-1} - \frac{(y-4)y\log(y)}{y-1} \right) - \frac{1}{4} \left(\Delta + \frac{1}{2} \right). \quad (29)$$