

# Exploring Dark matter, Neutrino mass and $R_{K^{(*)}}$ anomalies in $L_\mu - L_\tau$ model

Rukmani Mohanta

University of Hyderabad, Hyderabad-500046, India

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# Outline of Talk

- 1  $L_\mu - L_\tau$  model with Scalar Leptoquark
- 2 Symmetry breaking and mass spectrum
- 3 Dark Matter Phenomenology
- 4 Constraints from Flavour Sector
- 5 Implications
- 6 Conclusion

# Motivation

- No direct evidence of NP either in Energy frontier or Intensity frontier (except Neutrino Oscillation)
- However, in last few years several anomalies are reported in in  $b \rightarrow sll$  FCNC transitions, which might possibly suggest the presence of New Physics
- **Interesting one: Lepton Non-Universality Parameter** Sizable discrepancies reported by the LHCb Collaboration in the ratio  $R_{K^{(*)}}$  of

$$R_{K^{(*)}} = \frac{\text{Br}(B \rightarrow K^{(*)} \mu \mu)}{\text{Br}(B \rightarrow K^{(*)} e e)}$$

LNU parameters	SM predictions	Expt. values	Deviations
$R_K  _{q^2 \in [1.0, 6.0]}$	$1.0003 \pm 0.0001$	$0.745^{+0.090}_{-0.074} \pm 0.036$	$2.6\sigma$
$R_{K^*}  _{q^2 \in [0.045, 1.1]}$	$0.92 \pm 0.02$	$0.660^{+0.110}_{-0.070} \pm 0.024$	$2.2\sigma$
$R_{K^*}  _{q^2 \in [1.1, 6.0]}$	$1.00 \pm 0.01$	$0.685^{+0.113}_{-0.007} \pm 0.047$	$2.4\sigma$

R. Aaij et al. [LHCb Collab], Phys. Rev. Lett. 113, 151601 (2014); JHEP 08, 055 (2017).

- Another famous anomaly observed in the angular analyses of  $B \rightarrow K^* \mu \mu$  is in the FFI  $P_5'$  observable which has around  $3\sigma$  deviation in the interval  $4.3 < q^2 < 8.68 \text{ GeV}^2$
- Also the differential branching fraction  $B_s \rightarrow \phi \mu \mu$  process, in the  $1 < q^2 < 6 \text{ GeV}^2$  bin is more than  $3\sigma$  below the SM predictions
- Interesting to note individual deviations in the observables mentioned above are in the same direction, i.e. destructive with the SM.
- The effective Hamiltonian describing  $b \rightarrow s \ell \ell$  is given as

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \sum_i \left( C_i^\ell O_i^\ell + C_i^{\prime\ell} O_i^{\prime\ell} \right), \quad C_i^{(\prime)\ell} = C_i^{(\prime)\text{SM}} + C_i^{(\prime)\ell, \text{NP}}$$

- Global fit shows that sizable NP contributions to  $C_{9(10)}^\mu$  can accommodate the data
- **Bottom Line: NP could couple primarily couple to muon sector**

## Model description (Gauged $L_\mu - L_\tau$ )

- The SM has accidental  $U(1)$  global symmetries like  $B$  and  $L$  no. conservation, they become anomalous if converted into a local one
- The anomaly free situation can be obtained if instead of considering  $B$  and  $L$  separately, one uses some combinations between them, e.g.,  $B - L$ ,  $L_e - L_\mu$ ,  $L_e - L_\tau$  or  $L_\mu - L_\tau$
- For the anomaly cancellation of local  $B - L$  models, one requires 3 RHNs with appropriate  $B - L$  charges
- Unlike  $B - L$  case, the anomaly cancellation does not require any extra chiral fermionic degrees of freedom for  $L_\alpha - L_\beta$ , as anomalies cancel between different leptonic generations.
- $U(1)_{L_\mu - L_\tau}$  is less constrained, as the extra  $Z'$  does not couple to electrons and quarks, so free from any constraints coming from lepton and hadron colliders, such as LEP and LHC.
- Another theoretical motivation: it can explain the muon ( $g - 2$ ) anomaly

# Particle Content of $L_\mu - L_\tau$ model

	Field	$SU(3)_C \times SU(2)_L \times U(1)_Y$	$U(1)_{L_\mu - L_\tau}$	$Z_2$
Fermions	$Q_L \equiv (u, d)_L^T$	(3, 2, 1/6)	0	+
	$u_R$	(3, 1, 2/3)	0	+
	$d_R$	(3, 1, -1/3)	0	+
	$e_L \equiv (\nu_e, e)_L^T$	(1, 2, -1/2)	0	+
	$e_R$	(1, 1, -1)	0	+
	$\mu_L \equiv (\nu_\mu, \mu)_L^T$	(1, 2, -1/2)	1	+
	$\mu_R$	(1, 1, -1)	1	+
	$\tau_L \equiv (\nu_\tau, \tau)_L^T$	(1, 2, -1/2)	-1	+
	$\tau_R$	(1, 1, -1)	-1	+
	$N_e$	(1, 1, 0)	0	-
	$N_\mu$	(1, 1, 0)	1	-
	$N_\tau$	(1, 1, 0)	-1	-
	Scalars	$H$	(1, 2, 1/2)	0
$\eta$		(1, 2, 1/2)	0	-
$\phi_2$		(1, 1, 0)	2	+
$S_1$		(3, 1, 1/3)	-1	-

Table: Fields and their charges of the proposed  $U(1)_{L_\mu - L_\tau}$  model.

# Lagrangian of the Model

The Lagrangian of the present model can be written as

$$\begin{aligned}
 \mathcal{L} = & \mathcal{L}_{\text{SM}} - \frac{1}{4} Z'_{\mu\nu} Z'^{\mu\nu} - g_{\mu\tau} \bar{\mu}_L \gamma^\mu \mu_L Z'_\mu - g_{\mu\tau} \bar{\mu}_R \gamma^\mu \mu_R Z'_\mu \\
 & + g_{\mu\tau} \bar{\tau}_L \gamma^\mu \tau_L Z'_\mu + g_{\mu\tau} \bar{\tau}_R \gamma^\mu \tau_R Z'_\mu + \bar{N}_e i \not{\partial} N_e + \bar{N}_\mu (i \not{\partial} - g_{\mu\tau} Z'_\mu \gamma^\mu) N_\mu \\
 & + \bar{N}_\tau (i \not{\partial} + g_{\mu\tau} Z'_\mu \gamma^\mu) N_\tau - \frac{f_\mu}{2} (\bar{N}_\mu^c N_\mu \phi_2^\dagger + \text{h.c.}) - \frac{f_\tau}{2} (\bar{N}_\tau^c N_\tau \phi_2 + \text{h.c.}) \\
 & - \frac{1}{2} M_{ee} \bar{N}_e^c N_e - \frac{1}{2} M_{\mu\tau} (\bar{N}_\mu^c N_\tau + \bar{N}_\tau^c N_\mu) - \sum_{q=d,s,b} (y_{qR} \bar{d}_{qR}^c S_1 N_\mu + \text{h.c.}) \\
 & - \sum_{i=e,\mu,\tau} Y_{\alpha i} (\bar{\ell}_L)_\alpha \tilde{\eta} N_{iR} + |(i\partial_\mu - 2g_{\mu\tau} Z'_\mu) \phi_2|^2 \\
 & + \left| \left( i\partial_\mu - \frac{g}{2} \tau^a \cdot \mathbf{W}_\mu^a - \frac{g'}{2} B_\mu \right) \eta \right|^2 + \left| \left( i\partial_\mu - \frac{g'}{3} B_\mu + g_{\mu\tau} Z'_\mu \right) S_1 \right|^2 \\
 & - V(H, \eta, \phi_2, S_1),
 \end{aligned}$$

# Scalar potential

The scalar potential  $V$  is expressed as

$$\begin{aligned} V(H, \eta, \phi_2, S_1) &= \mu_H^2 H^\dagger H + \lambda_H (H^\dagger H)^2 + \mu_\eta (\eta^\dagger \eta) + \lambda_{H\eta} (H^\dagger H) (\eta^\dagger \eta) \\ &+ \lambda_\eta (\eta^\dagger \eta)^2 + \lambda'_{H\eta} (H^\dagger \eta) (\eta^\dagger H) + \frac{\lambda''_{H\eta}}{2} [(H^\dagger \eta)^2 + \text{h.c.}] \\ &+ \mu_2^2 (\phi_2^\dagger \phi_2) + \lambda_2 (\phi_2^\dagger \phi_2)^2 + \mu_S^2 (S_1^\dagger S_1) + \lambda_S (S_1^\dagger S_1)^2 \\ &+ [\lambda_{H2} (\phi_2^\dagger \phi_2) + \lambda_{HS} (S_1^\dagger S_1)] (H^\dagger H) + \lambda_{S2} (\phi_2^\dagger \phi_2) (S_1^\dagger S_1) \\ &+ \lambda_{\eta 2} (\phi_2^\dagger \phi_2) (\eta^\dagger \eta) + \lambda_{S\eta} (S_1^\dagger S_1) (\eta^\dagger \eta). \end{aligned}$$



# Mass Spectra

- Spontaneous symmetry breaking pattern:

$$SU(2)_L \times U(1)_Y \times U(1)_{L_\mu - L_\tau} \implies SU(2)_L \times U(1)_Y \implies SU(2)_L \times U(1)_{em}$$

$$\phi_2^0 = \frac{1}{\sqrt{2}}(v_2 + h_2) + \frac{i}{\sqrt{2}}A_2, \quad H^0 = \frac{1}{\sqrt{2}}(v + h) + \frac{i}{\sqrt{2}}A^0.$$

- Fermion and Scalar Mass Matrices

$$M_N = \begin{pmatrix} \frac{1}{\sqrt{2}}f_\mu v_2 & M_{\mu\tau} \\ M_{\mu\tau} & \frac{1}{\sqrt{2}}f_\tau v_2 \end{pmatrix}, \quad M_S = \begin{pmatrix} 2\lambda_H v^2 & \lambda_{H2} v v_2 \\ \lambda_{H2} v v_2 & 2\lambda_2 v_2^2 \end{pmatrix}.$$

One can diagonalize the above mass matrices using a  $2 \times 2$  rotation matrix

$$U_{\alpha(\zeta)}^T M_{N(S)} U_{\alpha(\zeta)} = \text{diag} [M_{N_-(H_1)}, M_{N_+(H_2)}],$$

$$\text{with } \zeta = \frac{1}{2} \tan^{-1} \left( \frac{\lambda_{H2} v v_2}{\lambda_2 v_2^2 - \lambda_H v^2} \right), \quad \alpha = \frac{1}{2} \tan^{-1} \left( \frac{M_{\mu\tau}}{(f_\tau - f_\mu)(v/\sqrt{2})} \right).$$

- The lightest fermion mass eigenstate  $N_-$  considered as probable DM candidate, and  $M_{H_1}$  as the SM Higgs

- Inert doublet components  $(\eta^+, \eta^0)^T$  with  $\eta^0 = (\eta_e + \eta_o)/\sqrt{2}$ :

$$M_{\eta^+}^2 = \mu_\eta^2 + \frac{\lambda_{H\eta}}{2} v^2 + \frac{\lambda_{\eta^2}}{2} v_2^2,$$

$$M_{\eta_{e,o}}^2 = \mu_\eta^2 + \frac{\lambda_{\eta^2}}{2} v_2^2 + (\lambda_{H\eta} + \lambda'_{H\eta} \pm \lambda''_{H\eta}) \frac{v^2}{2}.$$

- Scalar Letpoquark:  $M_{S_1}^2 = 2\mu_S^2 + \lambda_{HS} v^2 + \lambda_{S2} v_2^2$
- New  $U(1)$  gauge boson:  $M_{Z'} = 2v_2 g_{\mu\tau}$ .
- Benchmark values of scalar spectra:  $(M_{S_1}, M_{\eta^+}, M_{\eta_{e,o}}) = (1.2, 2, 1.5)$  TeV

# Light Neutrino mass generation (Scotogenic model)

- Scotogenic model presents an attractive example of the generation of light  $\nu$ -mass via radiative mechanism in the dark sector
- The model enlarges the SM field content by: (a) 3 RHN fields  $N_k$   
(b) a second  $SU(2)_L$  scalar doublet  $\eta$
- The scalar potential is

$$\begin{aligned} V(\phi, \eta) &= m_\phi^2 \phi^\dagger \phi + m_\eta^2 \eta^\dagger \eta + \frac{\lambda_1}{2} (\phi^\dagger \phi)^2 + \frac{\lambda_2}{2} (\eta^\dagger \eta)^2 \\ &+ \lambda_3 (\phi^\dagger \phi)(\eta^\dagger \eta) + \lambda_4 (\phi^\dagger \eta)(\eta^\dagger \phi) + \lambda_5 \left( (\phi^\dagger \eta)^2 + (\eta^\dagger \phi)^2 \right) \end{aligned}$$

- Inert scalar sector :  $\eta^\pm, \eta^0 = (\eta_R \pm i\eta_I)/\sqrt{2}$  with mass

$$\begin{aligned} m_{\eta^\pm}^2 &= m_\eta^2 + \lambda_3 \frac{v^2}{2}, \quad m_{\eta^0}^2 = m_\eta^2 + (\lambda_3 + \lambda_4 \pm \lambda_5) \frac{v^2}{2} \\ \implies m_{\eta_R}^2 - m_{\eta_I}^2 &= \lambda_5 v^2 \end{aligned}$$

# Radiative Neutrino mass generation

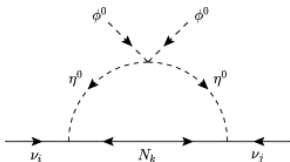
- The structure of light neutrino mass would be

$$m_\nu \sim \frac{1}{16\pi^2} (\text{favour structure}) \times (\text{mass scale}) \times (\text{loop function})$$

- The active neutrino mass matrix obtained from the diagram (below)

$$\begin{aligned}
 (\mathcal{M}_\nu)_{\alpha\beta} &= \sum_{k=1}^3 \frac{h_{k\alpha} h_{k\beta}}{16\pi^2} M_k \left[ \frac{m_{\eta_R}^2}{m_{\eta_R}^2 - M_k^2} \ln \left( \frac{m_{\eta_R}^2}{M_k^2} \right) - \frac{m_{\eta_I}^2}{m_{\eta_I}^2 - M_k^2} \ln \left( \frac{m_{\eta_I}^2}{M_k^2} \right) \right] \\
 &\simeq \sum_{k=1}^3 \frac{\lambda_5 h_{k\alpha} h_{k\beta} v^2}{16\pi^2} M_k \left[ 1 - \frac{M_k^2}{(m_0^2 - M_k^2)} \ln \left( \frac{m_0^2}{M_k^2} \right) \right]
 \end{aligned}$$

where  $m_{\eta_R}^2 - m_{\eta_I}^2 = \lambda_5 v^2 \ll m_0^2 = (m_{\eta_R}^2 + m_{\eta_I}^2)/2$ .



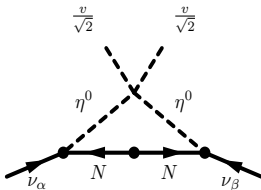
- **Light neutrino mass** : From the Yukawa interaction term

$$\sum_{i=e,\mu,\tau} Y_{\alpha i} (\bar{\ell}_L)_\alpha \tilde{\eta} N_{iR}$$

one can obtain

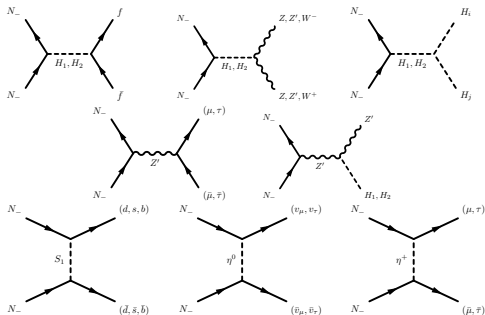
$$(\mathcal{M}_\nu)_{\alpha\beta} = \frac{\lambda''_{H\eta} v^2}{16\pi^2} \sum_{i=e,\mu,\tau} \frac{Y_{\alpha i} Y_{\beta i} M_{Di}}{m_0^2 - M_{Di}^2} \left[ 1 - \frac{M_{Di}^2}{m_0^2 - M_{Di}^2} \ln \frac{m_0^2}{M_{Di}^2} \right].$$

With sample parameter space  $(Y_{\alpha i}, \lambda''_{H\eta}) \sim (10^{-2}, 10^{-5})$  and  $(m_0, M_-, M_{ee}, M_+) \sim (1.5, 0.4, 3, 3)$  TeV, one can explain the light neutrino mass  $m_\nu$ , near eV scale



# Relic density

- (Scalar Leptoquark,  $\eta$ ,  $H_1$ ,  $H_2$ ,  $Z'$ ) - portal



- DM relic density is computed by

$$\Omega h^2 = \frac{2.14 \times 10^9 \text{GeV}^{-1}}{g_*^{1/2} M_{Pl}} \frac{1}{J(x_f)}, \quad J(x_f) = \int_{x_f}^{\infty} \frac{\langle \sigma v \rangle(x)}{x^2} dx.$$

where  $x = M_-/T$  and  $x_f$  is the freeze out parameter.

# Relic density

- Relevant parameters:  $M_-, g_{\mu\tau}, M_{Z'}, Y_{\alpha i}, y_{qR}$
- $g_{\mu\tau}$  controls the s-channel contribution whereas  $Y_{\alpha i}, y_{qR}$  relevant for t-channel

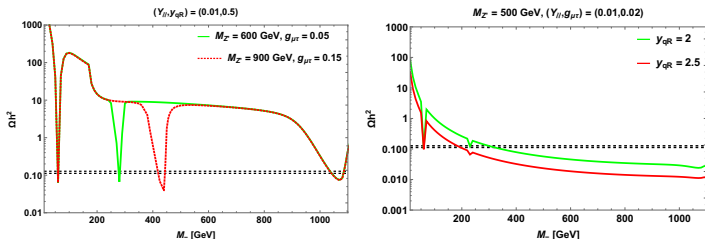


Figure: Behaviour of Relic density with DM mass

- Relic density for s-channel meet Planck limit near the resonance in propagator ( $H_1, H_2, Z'$ ) near  $M_- = M_{\text{PROP}}/2$  ( $80 < M_- < 1000$ )GeV.

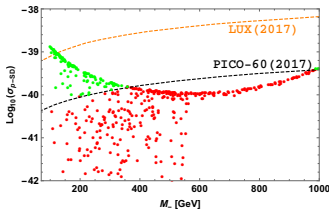
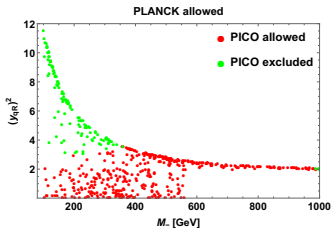
# Direct Detection: Scalar LQ portal

The effective interaction Lagrangian is given by

$$\mathcal{L}_{\text{eff}} \simeq \frac{y_{qR}^2 \cos^2 \alpha}{4(M_{S_1}^2 - M_-^2)} \bar{N}_- \gamma^\mu \gamma^5 N_- \bar{q} \gamma_\mu \gamma^5 q.$$
$$\Rightarrow \sigma_{S_1}^{SD} = \frac{M_-^2 M_n^2}{\pi(M_- + M_n)^2} \frac{\cos^4 \alpha}{(M_{S_1}^2 - M_-^2)^2} \left[ y_{dR}^2 \Delta_d + y_{sR}^2 \Delta_s \right]^2 J_N (J_N + 1).$$

where the angular momentum  $J_N = \frac{1}{2}$ ,  $M_n \simeq 1$  GeV for nucleon.

\* DM study has no significant impact on  $M_{Z'}$  –  $g_{\mu\tau}$  parameters.

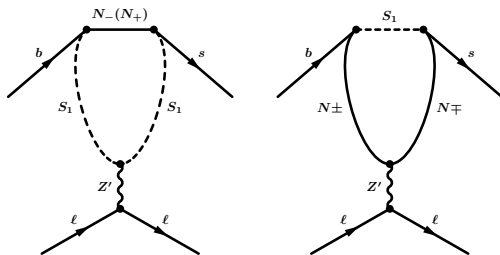




# Constraints from Flavor Sector

## A. $B \rightarrow K\ell^+\ell^-$ :

We use  $\text{Br}(B^+ \rightarrow K^+\tau^+\tau^-)$  and  $R_K$  observables, which will put constraint on all the four parameters, i.e.,  $(y_{qR})^2$ ,  $g_{\mu\tau}$ ,  $M_{Z'}$  and  $M_-$ .



**Figure:** Penguin diagram of  $b \rightarrow s\ell\ell$  processes, where  $\ell = \mu, \tau$  with leptoquark in the loop.

- Computing the penguin loop and comparing with SM effective Hamiltonian gives

$$C_9^{\text{NP}} = \frac{\sqrt{2}}{2^4 \pi G_F \alpha_{\text{em}} V_{tb} V_{ts}^*} \frac{y_{bR} y_{sR} g_{\mu\tau}^2}{M_{Z'}^2} \mathcal{V}_{sb} (\chi_-, \chi_+).$$

- Expt. data used to obtain the constrained parameters:

$$\text{Br}(B^+ \rightarrow K^+ \tau^+ \tau^-) \Big|_{\text{Expt}} < 2.5 \times 10^{-3},$$

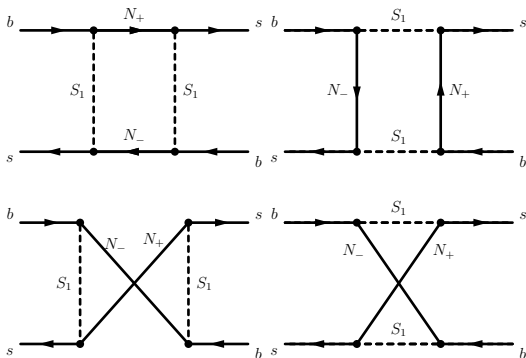
whereas

$$\text{Br}(B^+ \rightarrow K^+ \tau^+ \tau^-) \Big|_{\text{SM}} = (1.486 \pm 0.12) \times 10^{-7}.$$

- Since  $Z'$  doesn't couple to electron, the branching ratio of  $B^+ \rightarrow K^+ e^+ e^-$  process is considered to be SM like.
- The anomalies of  $b \rightarrow sll$  decay modes can put constraint on all the four parameters, i.e.,  $(y_{qR})^2$ ,  $g_{\mu\tau}$ ,  $M_{Z'}$  and  $M_-$ .

## B. $B_s - \bar{B}_s$ mixing:

- $\Delta M_s^{\text{SM}} = (17.426 \pm 1.057) \text{ ps}^{-1}$ ,  $\Delta M_s^{\text{Expt}} = 17.761 \pm 0.022 \text{ ps}^{-1}$ .
- $B_s - \bar{B}_s$  mixing will put bound on  $(y_{qR})^2$  and  $M_-$  parameters.



$$\Delta M_s^{\text{NP}} = \frac{(y_{sR} y_{bR})^2}{48\pi^2 M_{S_1}^2} \cos^2 \alpha \sin^2 \alpha C_{B_s}^{\text{NP}} \eta_B \hat{B}_{B_s} f_{B_s}^2 M_{B_s}.$$

C.  $B \rightarrow X_s \gamma$  and  $\tau \rightarrow \mu \nu_\tau \bar{\nu}_\mu$  :

- $B \rightarrow X_s \gamma$  will constrain the  $(y_{qR})^2$  and  $M_-$  parameters
- $\text{Br}(\tau \rightarrow \mu \nu_\tau \bar{\nu}_\mu)$  will put bounds on  $M_{Z'} - g_{\mu\tau}$  parameter space.

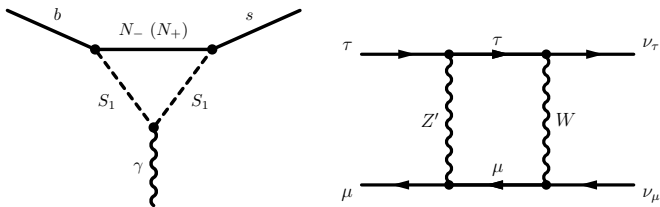
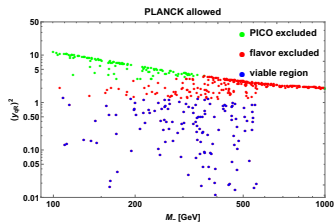
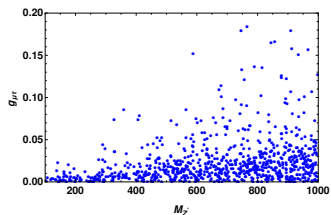


Figure: Feynman diagram of  $b \rightarrow s \gamma$  (left panel) and  $\tau \rightarrow \mu \nu_\tau \bar{\nu}_\mu$  (right panel) processes in the presence of scalar leptoquark.

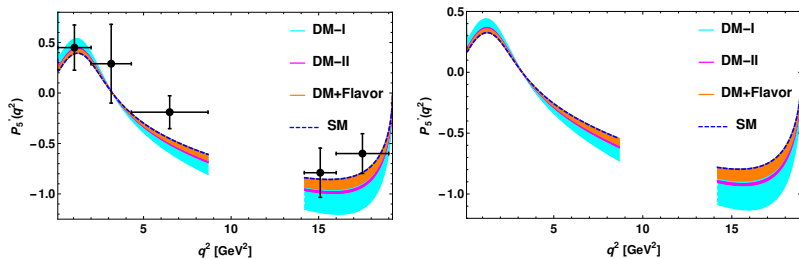
# Constraints on the parameter space from Flavor and Dark Matter sectors



Parameters	DM-I	DM-II	DM+Flavor
$M_-$ [GeV]	103 – 560	561 – 988	103 – 560
$(y_{qR})^2$	0 – 3.51	1.94 – 2.56	0 – 1.26

Table: Predicted allowed range of parameters  $M_-$  and  $(y_{qR})^2$ .

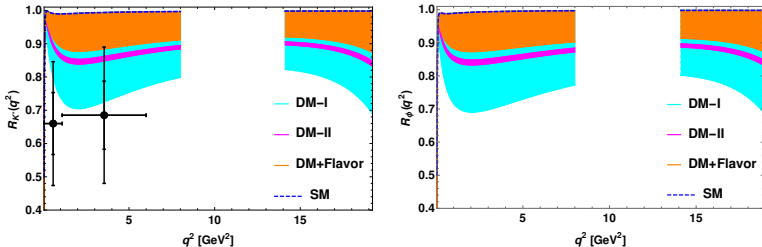
# Implication on $B_{(s)} \rightarrow K^*(\phi)\mu^+\mu^-$ Processes



**Figure:** Left panel corresponds to  $P'_5$  observable of  $B \rightarrow K^*\mu^+\mu^-$ , and right panel corresponds to  $B_s \rightarrow \phi\mu^+\mu^-$ .

# Implication on $B_{(s)} \rightarrow K^*(\phi)\mu^+\mu^-$ Processes

We have tested the LNU parameters ( $R_{K^*}$ ,  $R_\phi$ ) and the optimized angular observables,  $P'_{4,5}$ .



**Figure:** The  $q^2$  variation of  $R_{K^*}$  (left panel) and  $R_\phi$  (right panel) LNU parameters in the  $L_\mu - L_\tau$  model.

# Summary

- We explored Majorana dark matter in a  $U(1)_{L_\mu-L_\tau}$  gauge extension of SM with an additional  $(\bar{\mathbf{3}}, \mathbf{1}, 1/3)$  Scalar LQ.
- Relic density is investigated in SLQ,  $Z'$  as well as scalar portal and the spin-dependent WIMP-nucleon cross section in SLQ-portal.
- In flavor sector, the model parameters are constrained from experimental limits on  $R_K$ ,  $B \rightarrow K\tau^+\tau^-$ ,  $B_s - \bar{B}_s$  mixing,  $B \rightarrow X_s\gamma$  and  $\tau \rightarrow \mu\nu_\tau\bar{\nu}_\mu$  processes.
- Using the allowed parameters space, the LNU parameters ( $R_{K^*}$ ,  $R_\phi$ ), and the  $P'_5$  observables of  $B_{(s)} \rightarrow K^*(\phi)\mu^+\mu^-$  decay modes are estimated.
- This simple gauge extension provides a platform to address dark matter, neutrino and flavor sectors in phenomenological perspective.

Thank You