

Dynamical gravitational corrections to the conformal anomaly

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Invisibles Workshop 2019

13th of June of 2019

Work in collaboration with
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Idea

Can we tell if **gravity is dynamical or not** by computing the conformal anomaly?

Is gravity a fundamental interaction?

Existence of gravitons

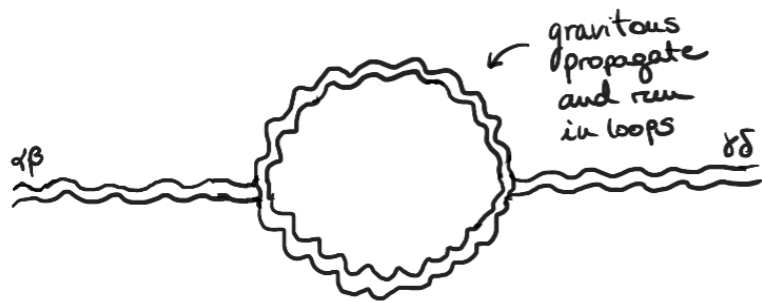
Entropic origin

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Einstein's general relativity (GR) \rightarrow Theory for the gravitational field

Fierz-Pauli lagrangian for free spin 2 particles



When interactions are included in a consistent way \rightarrow GR

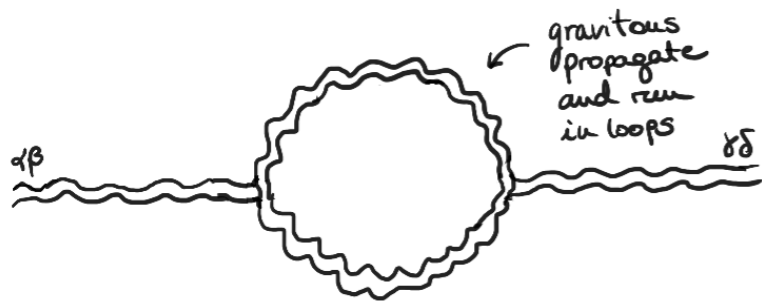
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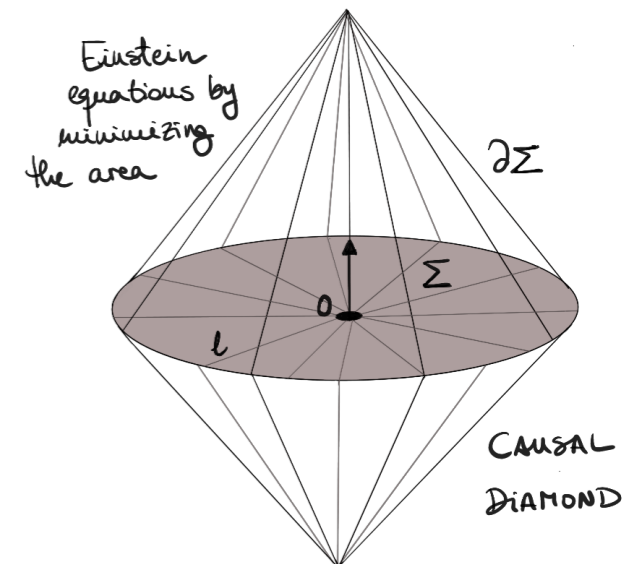
Dictionary between BH laws and thermodynamic laws

$$\delta S = \delta A$$

Instead of thinking in the fundamental dof → Entropic force

Einstein equations arise from minimizing the entropy (the area) with a fixed volume

Jacobson, Padmanabhan, Verlinde,...



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Scale invariance
(flat spacetime)

$$x^\mu \rightarrow \lambda x^\mu$$

Conformal invariance
(flat spacetime)

Conformal group

Weyl invariance
(**curved** spacetimes)

$$g_{\mu\nu} \rightarrow \Omega(x)^2 g_{\mu\nu}$$

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$$\langle T^\mu_\mu \rangle = a_2 \quad \text{Proportional to the **one-loop counterterm**}$$

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If gravity is **dynamical or just a background** is important in the path integral

$$a_2 \sim \int \mathcal{D}\phi_i \mathcal{D}g_{\mu\nu} \quad \text{Sensitive to the **graviton dynamics**}$$

Background vs. dynamical gravity

We expand the gravitational field in a background and a perturbation

$$g_{\mu\nu} = \underbrace{\bar{g}_{\mu\nu}}_{\text{Background field}} + \kappa \underbrace{h_{\mu\nu}}_{\text{Quantum field}}$$

(constant curvature spaces for simplicity)

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We take for example an scalar field coupled to gravity

Lagrangian	Anomaly
$\sqrt{\bar{g}} \bar{\nabla}_{\mu} \phi \bar{\nabla}^{\mu} \phi$	$\frac{\sqrt{\bar{g}}}{(4\pi)^2} \frac{29}{135} \lambda^2$
$\sqrt{\bar{g}} \left(\bar{\nabla}_{\mu} \phi \bar{\nabla}^{\mu} \phi + \frac{1}{6} \phi^2 \bar{R} \right)$	$-\frac{\sqrt{\bar{g}}}{(4\pi)^2} \frac{\lambda^2}{135}$
$\sqrt{g} \left(\nabla_{\mu} \phi \nabla^{\mu} \phi + \frac{1}{6} \phi^2 R \right)$	0
$\sqrt{g} \left(\nabla_{\mu} \phi \nabla^{\mu} \phi + R_{\mu\nu}^2 - \frac{1}{2} R^2 \right)$	$\frac{\sqrt{g}}{(4\pi)^2} \frac{1078}{135} \lambda^2$

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	Lagrangian	Anomaly
Non-dynamical	$\sqrt{\bar{g}} \bar{\nabla}_\mu \phi \bar{\nabla}^\mu \phi$	$\frac{\sqrt{\bar{g}}}{(4\pi)^2} \frac{29}{135} \lambda^2$ (curvature of CCS)
	$\sqrt{\bar{g}} \left(\bar{\nabla}_\mu \phi \bar{\nabla}^\mu \phi + \frac{1}{6} \phi^2 \bar{R} \right)$	$-\frac{\sqrt{\bar{g}}}{(4\pi)^2} \frac{\lambda^2}{135}$
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Non-dynamical	$\sqrt{\bar{g}} \bar{\nabla}_\mu \phi \bar{\nabla}^\mu \phi$	$\frac{\sqrt{\bar{g}}}{(4\pi)^2} \frac{29}{135} \lambda^2$ <p>(curvature of CCS)</p>
	$\sqrt{\bar{g}} \left(\bar{\nabla}_\mu \phi \bar{\nabla}^\mu \phi + \frac{1}{6} \phi^2 \bar{R} \right)$ <p>Non-minimal coupling</p>	$-\frac{\sqrt{\bar{g}}}{(4\pi)^2} \frac{\lambda^2}{135}$
Dynamical	$\sqrt{g} \left(\nabla_\mu \phi \nabla^\mu \phi + \frac{1}{6} \phi^2 R \right)$	0
	$\sqrt{g} \left(\nabla_\mu \phi \nabla^\mu \phi + \underbrace{R^2_{\mu\nu} - \frac{1}{2} R^2}_{\text{Quadratic gravity}} \right)$	$\frac{\sqrt{g}}{(4\pi)^2} \frac{1078}{135} \lambda^2$

Outlook

Can we measure the **conformal anomaly** and be **sensitive to the character of the gravitational field**?

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Starobinsky's inflation driven by the trace anomaly of conformally coupled matter fields

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Transport effects due to conformal anomaly (similar to the ones due to chiral anomalies)

Measurements of scale transport, **SEE** and **SME**, in Weyl semimetals

Work ongoing regarding same type of computations for **fermions**

Thank you!

Backup

Let us go back to see the origin of the anomaly

$$\underline{\Gamma_0[\bar{g}_{\mu\nu}, \bar{\phi}]} = \lim_{n \rightarrow 4} \left\{ \underline{\Gamma[\bar{g}_{\mu\nu}, \bar{\phi}; n]} - \Gamma_\infty[\bar{g}_{\mu\nu}, \bar{\phi}; n] \right\}$$

Renormalized effective action

One-loop effective action

Divergent piece

The **full on-shell effective action** has the same symmetries as the original theory in any dimension

Breitenlohner, Maison

$$\mathfrak{D}\Gamma[\bar{g}_{\mu\nu}, \bar{\phi}; n] = \left[2\bar{g}^{\mu\nu} \frac{\delta}{\delta\bar{g}^{\mu\nu}} + \frac{n-2}{2} \bar{\phi} \frac{\delta}{\delta\bar{\phi}} \right] \Gamma[\bar{g}_{\mu\nu}, \bar{\phi}; n] = 0$$

But we have **evanescent operators** coming from the counterterm

$$\mathfrak{D}\Gamma_0[\bar{g}_{\mu\nu}, \bar{\phi}] = \mathfrak{D} \left[\frac{1}{(n-4)} \mathcal{B}[\bar{g}_{\mu\nu}, \bar{\phi}] \right] = \mathcal{A}[\bar{g}_{\mu\nu}, \bar{\phi}]$$

Anomaly unless we find a proper counterterm