

# **Producción y uso de haces radioactivos**

**Master Interuniversitario de Física Nuclear**

**Berta Rubio IFIC (CSIC-Univ. Valencia)**

B. Rubio. Master FN, Valencia 2018

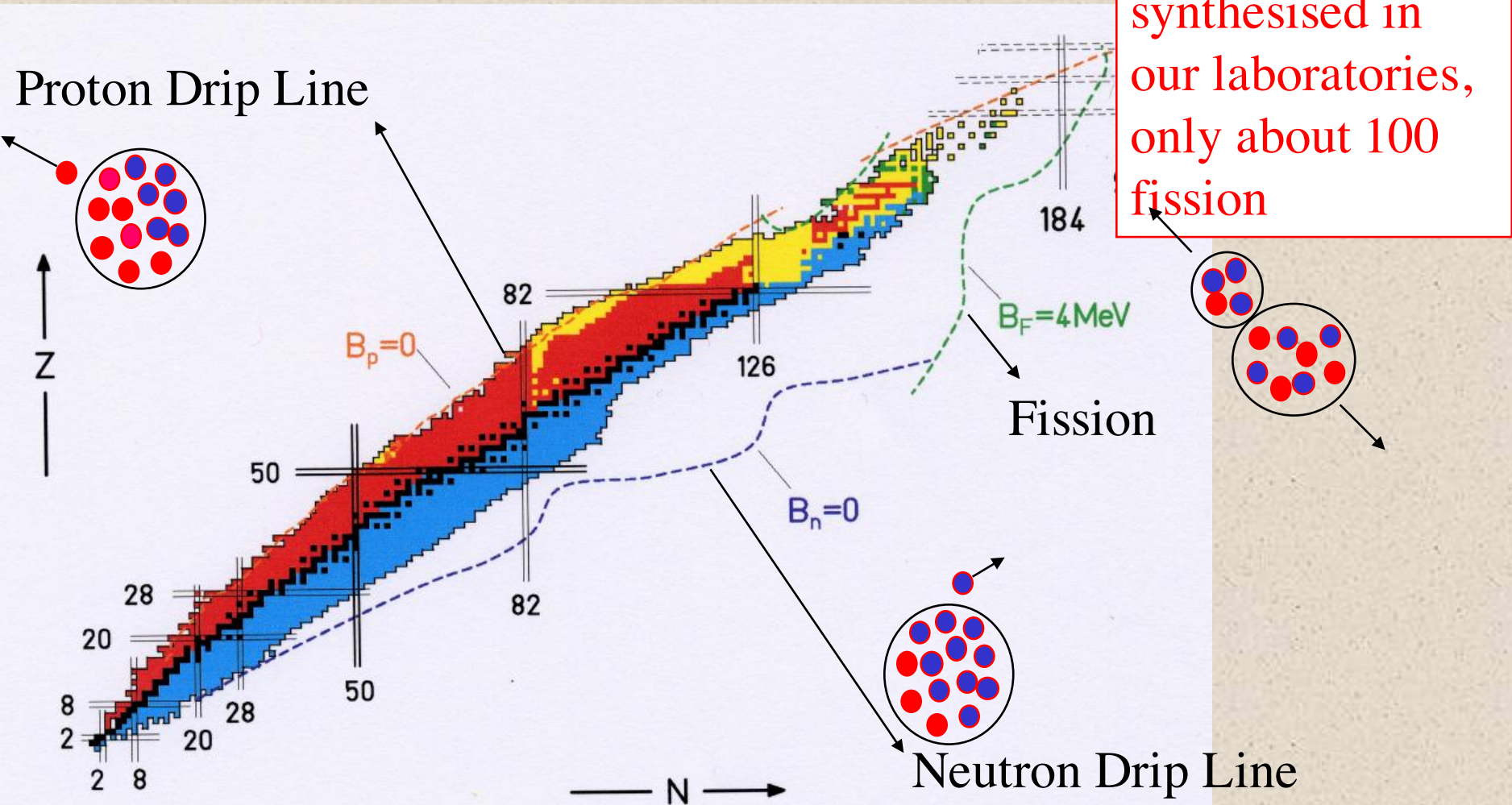
**Porqué la producción de haces radioactivos ha marcado un antes y un después en la Física Nuclear**

**Reacciones nucleares y lo que nos enseñan de la estructura de los núcleos.**

**Porqué necesitamos haces postacelerados  
Porqué necesitamos alejarnos de la estabilidad.**

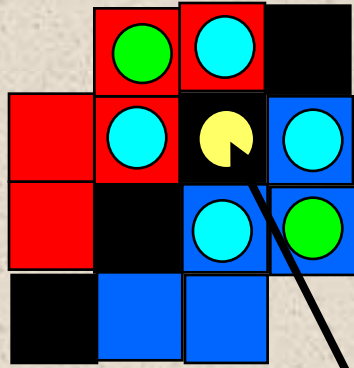
265 stable

About 3000 out of 6000 synthesised in our laboratories, only about 100 fission



Close to the stability Direct Reactions with Light Projectiles,  $p$ ,  $d$ ,  $\alpha$ ,  $e^-$ ,  $\gamma$ , etc.. give us detailed information on the structure of the target nucleus

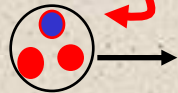
Z



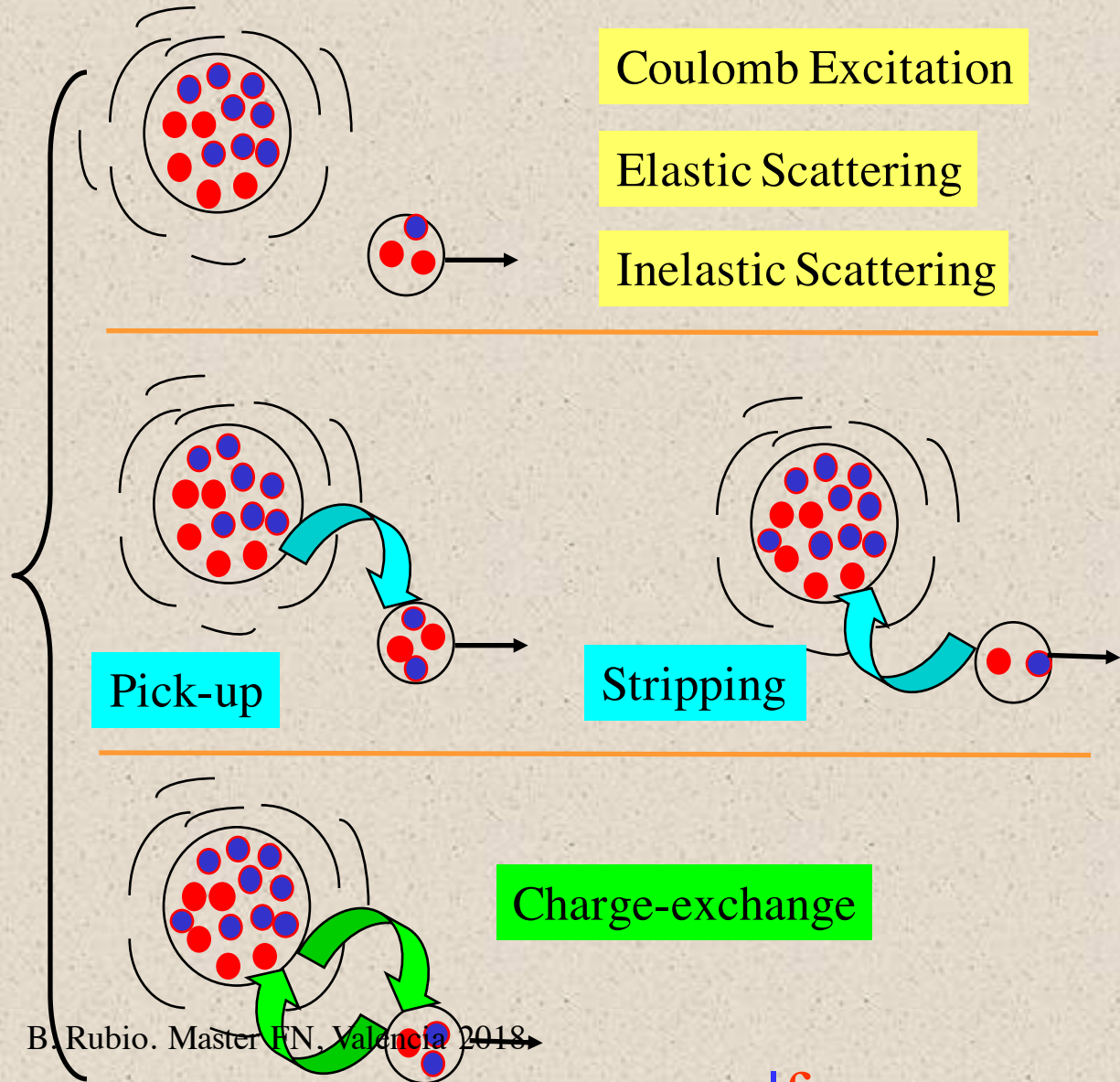
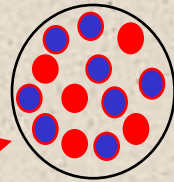
N

Target

interaction  
(operator)



Projectile

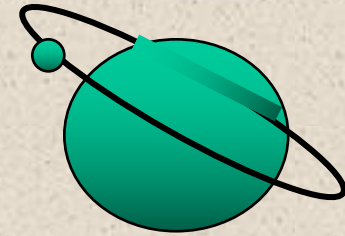


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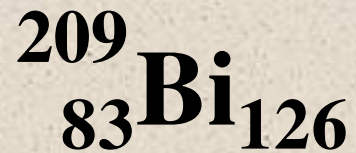
$|i\rangle \rightarrow$

$|f\rangle$

Let us use an illustrative example



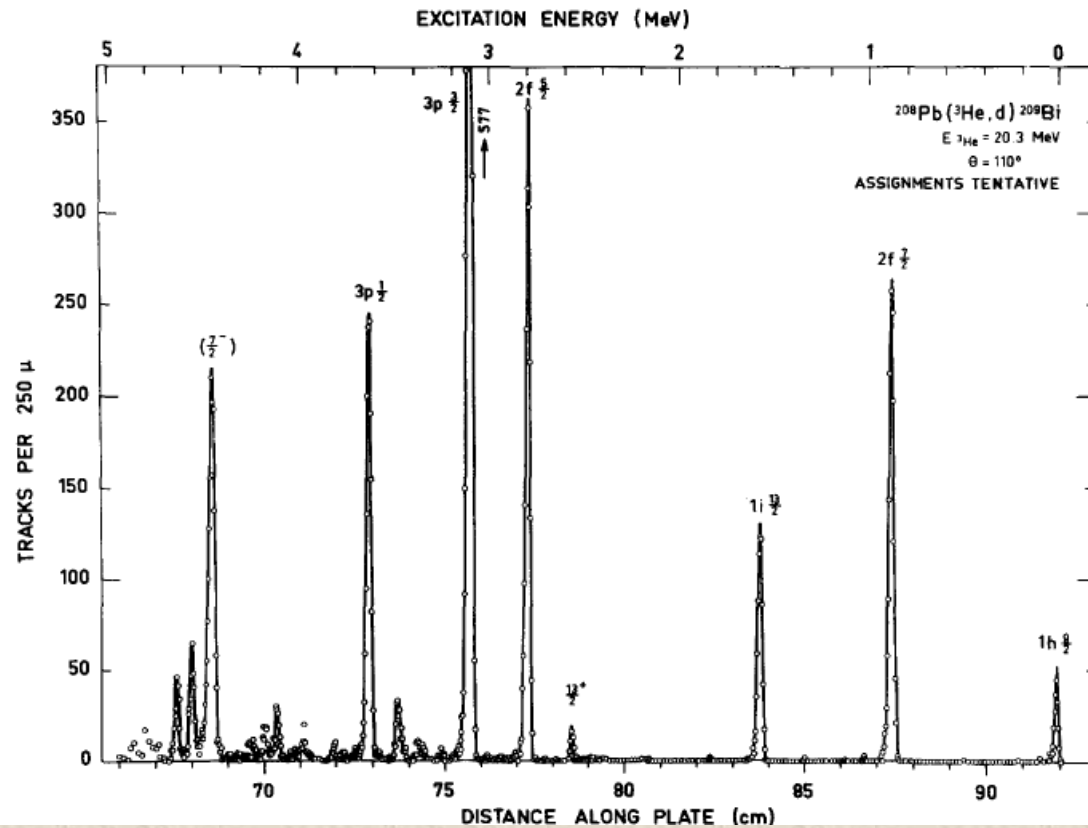
$$^{209}\text{Bi} = ^{208}\text{Pb} + p$$



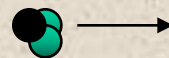
Closed shell + 1 p

Closed shell for neutrons

Z

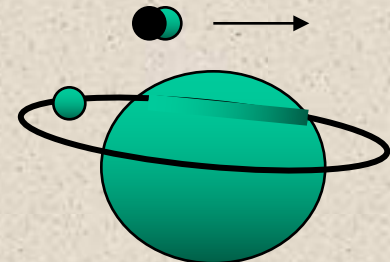


Reaction:

 $^{208}\text{Pb}(^3\text{He}, d)^{209}\text{Bi}$   
 $E(^3\text{He}) = 20 \text{ MeV}$ 


208Pb target

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209Bi final nucleus

## Shell Model and Residual Interactions

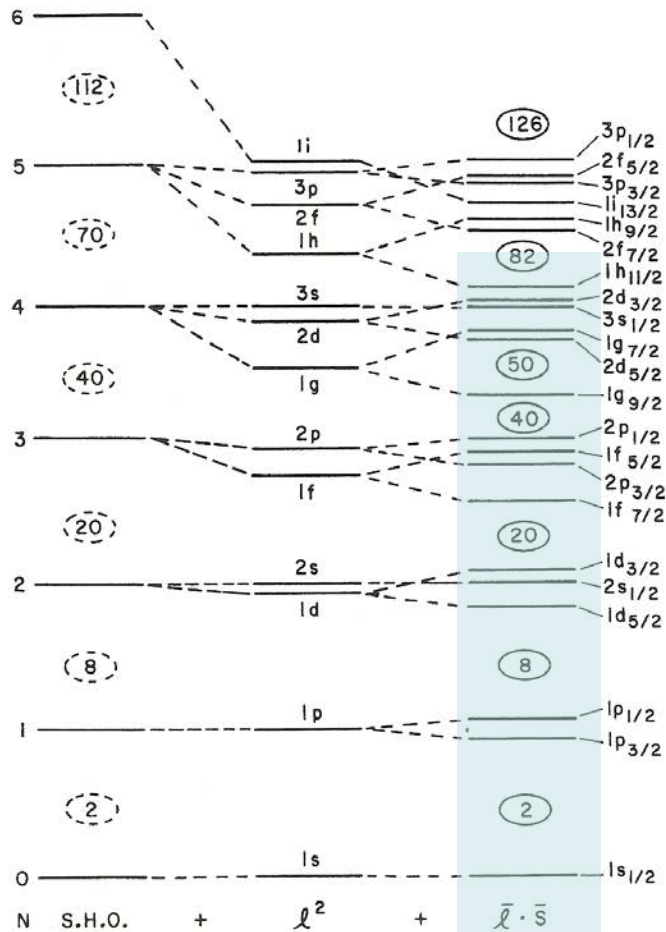
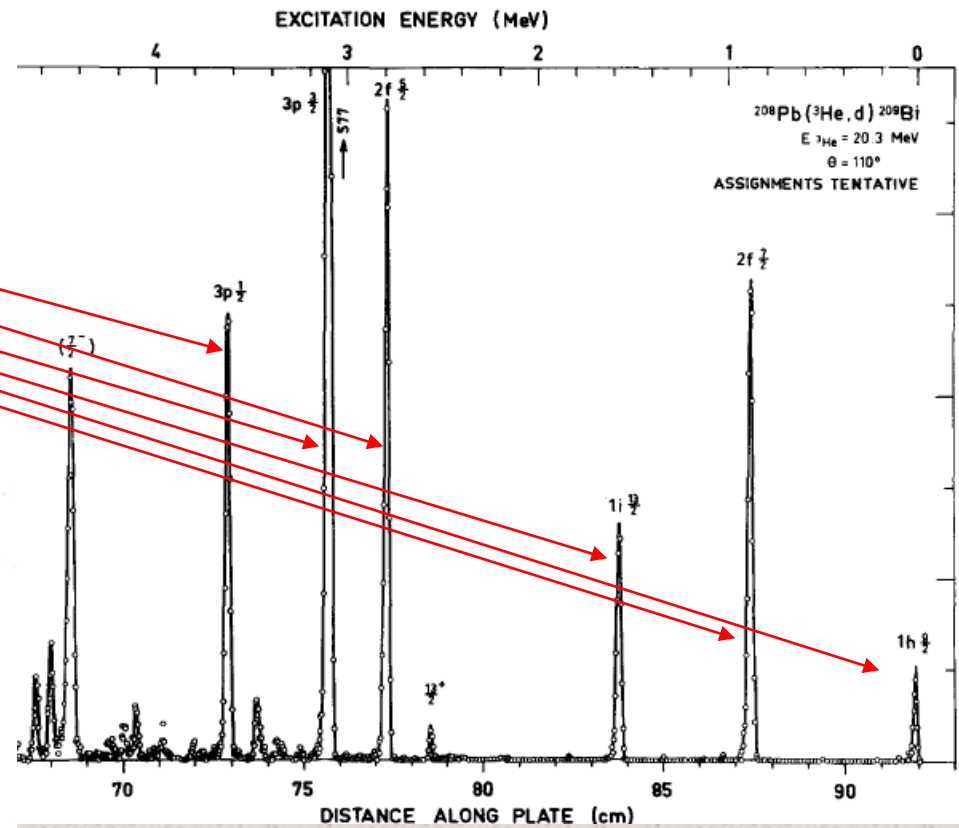


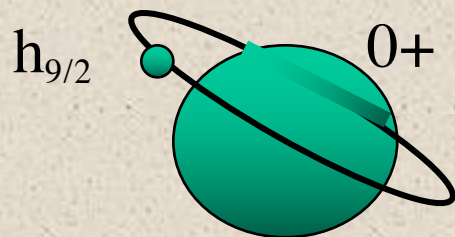
Fig. 3.2. Single-particle energies for a simple harmonic oscillator (S.H.O.), a modified harmonic oscillator with  $l^2$  term, and a realistic shell model potential with  $l^2$  and spin orbit ( $l \cdot s$ ) terms.



$^{208}\text{Pb}(^3\text{He}, d)^{209}\text{Bi}$   
 $E(^3\text{He}) = 20 \text{ MeV}$

$^{209}\text{Bi} = ^{208}\text{Pb} + p$

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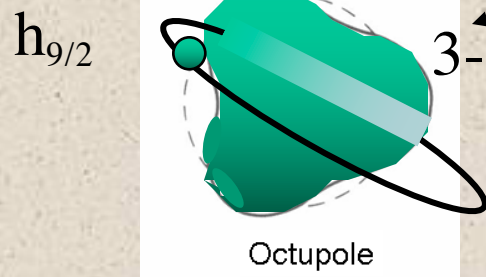


This is the gs of  $^{209}\text{Bi}$

$^{209}\text{Bi} = ^{208}\text{Pb} + p$

The first excited state

In  $^{208}\text{Pb}$  is an octupole excitation,  $3-$ , at 2415 keV

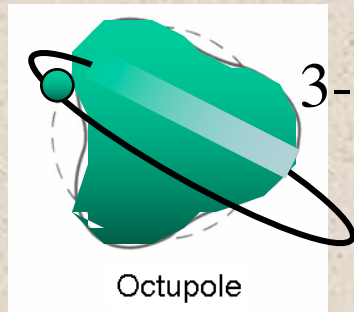
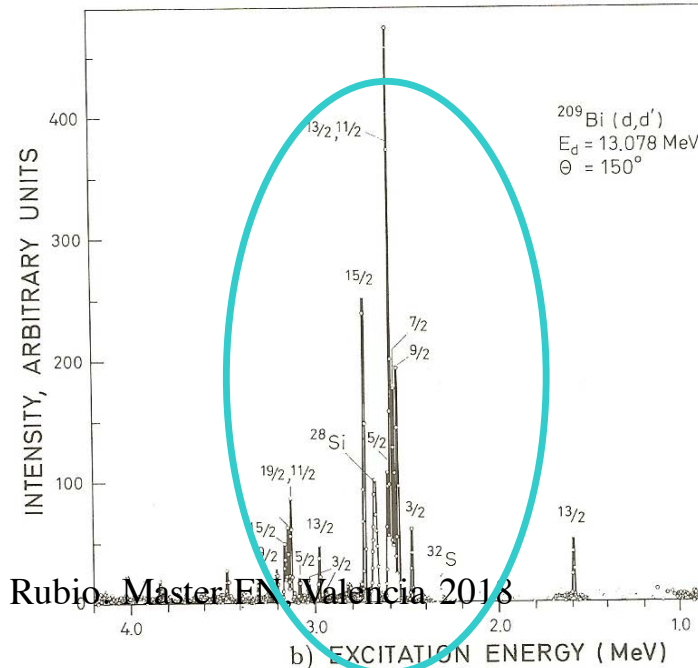
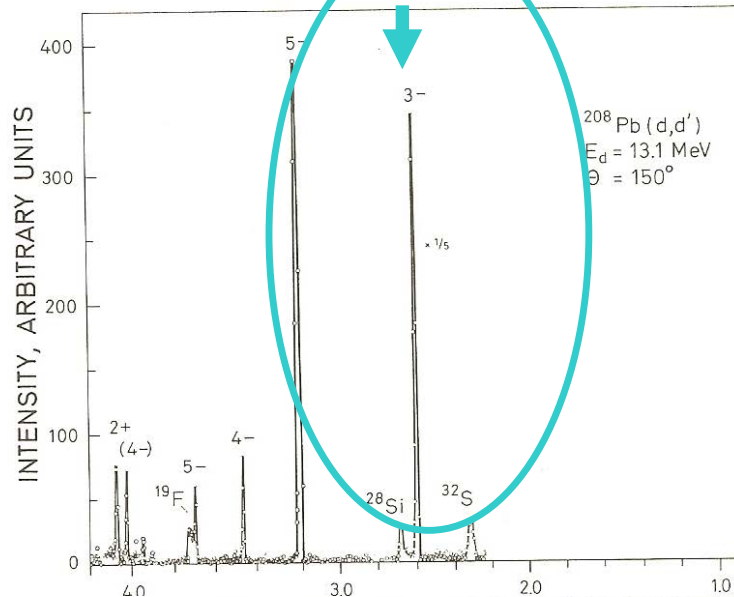


One expects at the same energy in  $^{209}\text{Bi}$

A septuplet:

$$[3-(^{208}\text{Pb}) \times h^{9/2}]_{3/2 \dots 15/2}^+$$

The appropriate reaction is inelastic scattering, for instance  $(d, d')$


 $h_{9/2}$ 
 $209\text{Bi} = 208\text{Pb} + p$ 


|                          |                   |                               |                  |                              |
|--------------------------|-------------------|-------------------------------|------------------|------------------------------|
| 39<br>2.6<br>46;<br>9... | At 209<br>5.4 h   | At 210<br>8.3 h               | At 211<br>7.22 h | At 212<br>119 ms             |
| 208<br>3 h               | 207<br>5.84 h     | Po 208<br>2.898 a             | Po 209<br>102 a  | Po 210<br>138.38 d           |
| 206<br>4 d               | Bi 207<br>31.55 a | Bi 208<br>$3.68 \cdot 10^5$ a | Bi 209<br>100    | Bi 210<br>$1.0 \cdot 10^6$ a |
| 205<br>$10^7$ a          | Pb 206<br>24.1    | Pb 207<br>22.1                | Pb 208<br>52.4   | Pb 209<br>3.25               |
| 204<br>8 a               | Tl 205<br>70.48   | Tl 206<br>3.7 m               | Tl 207<br>3.3 s  | Tl 208<br>3.05               |
| 203<br>59 d              | Hg 204<br>6.87    | Hg 205<br>5.2 m               | Hg 206<br>8.15 m | Hg 207<br>2.9                |
| 202<br>3 s               | Au 203<br>60 s    | Au 204<br>39.8 s              | Au 205<br>31 s   | Au 206<br>...                |
| 201<br>5 m               | Pt 202<br>~43.6 h |                               |                  |                              |

# One could go to more sophisticated excitation modes such as An isovector dipole resonance (E1)

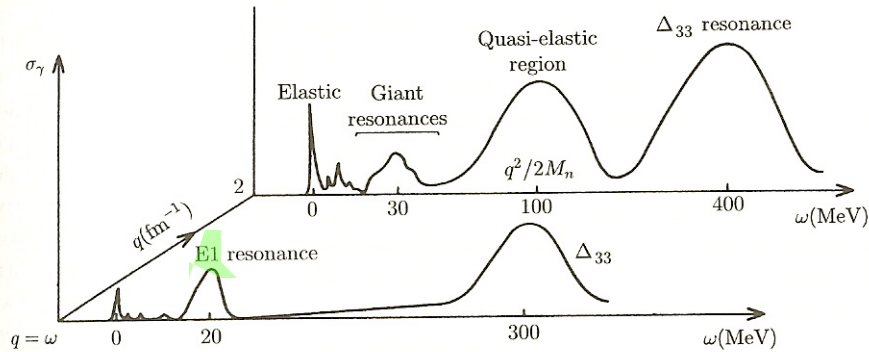
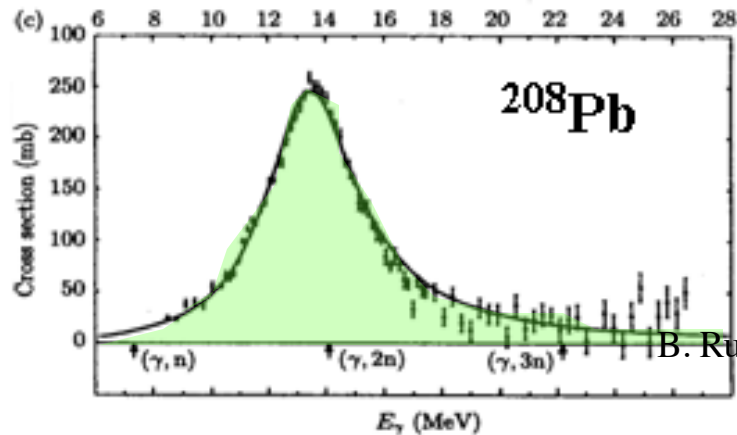
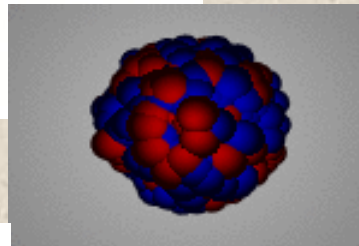
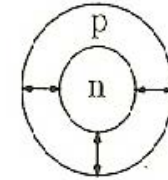


FIG. 1.1. Characteristic response (cross section  $\sigma$ ) of an atomic nucleus as a function of the energy transfer  $\omega$  and momentum transfer  $q$ . The lower curve is for photon absorption, the upper one for particle scattering with  $q = 2 \text{ fm}^{-1} \neq \omega$ .

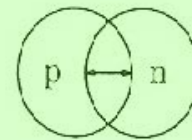


Isovector

$\Delta T=0 \ 1$

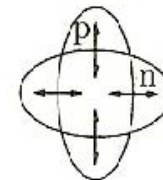


IVGMR



IVGMR

E1



IVGQR

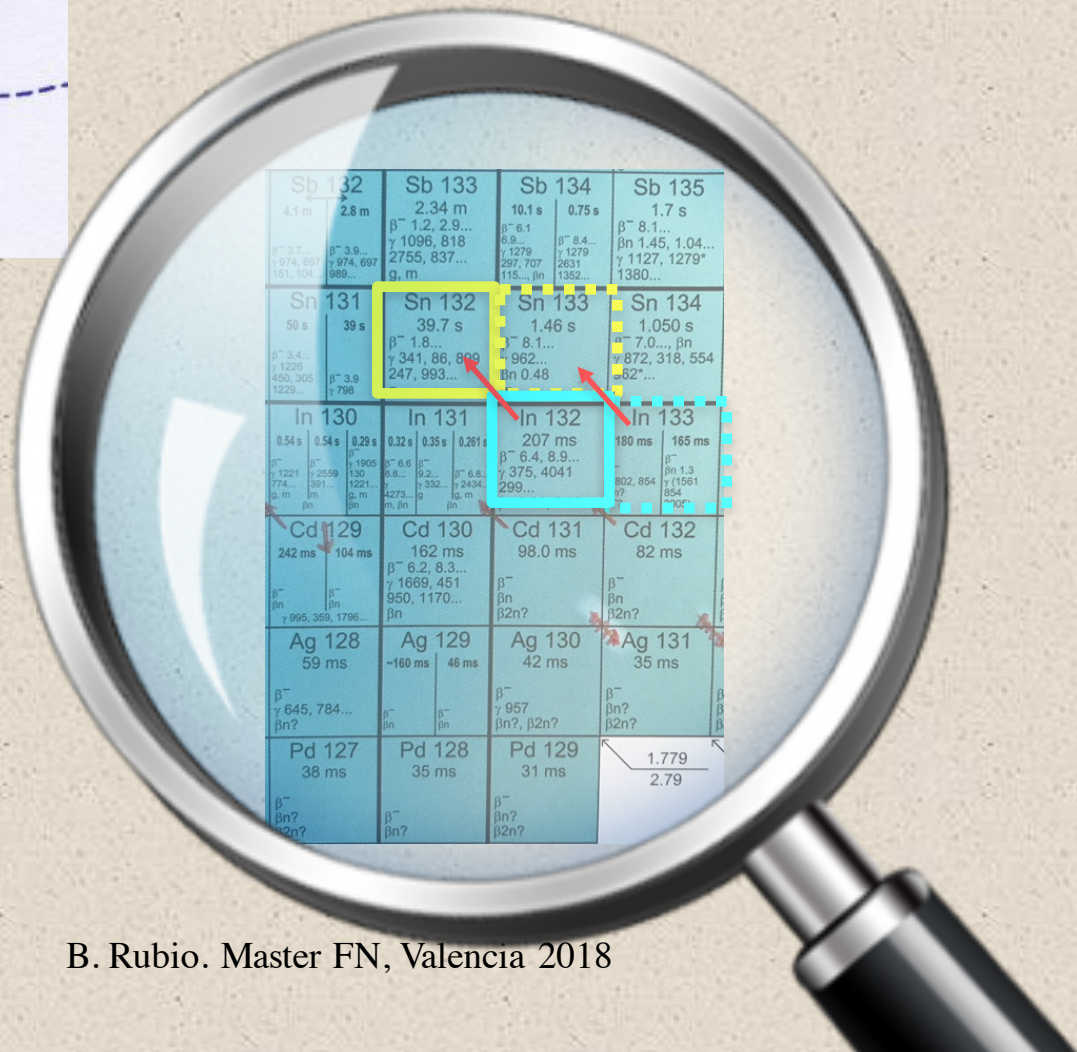
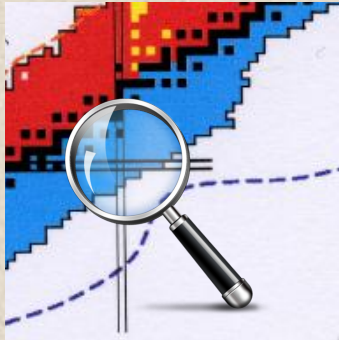
$\Delta T = 1$

$\Delta S = 0$

# Qué ha sido necesario para poder hacer estos experimentos

1. Definir la reacción nuclear de interés.
2. Blanco estable:  $^{208}\text{Pb}$  y  $^{209}\text{Bi}$  en este caso
2. Un proyectil “sin estructura”
3. Un proyectil ligero (struggling)

# Vamos a intentar aproximarnos a otro núcleo doblemente mágico: 132Sn



# Vamos a intentar aproximarnos a otro núcleo doblemente mágico: $^{132}\text{Sn}$

Vidas medias,  $T_{1/2}$   
207 ms !  
180/165 ms !

|                                                                                                             |                                                                                                                                                |                                                                                                      |                                                                                                          |
|-------------------------------------------------------------------------------------------------------------|------------------------------------------------------------------------------------------------------------------------------------------------|------------------------------------------------------------------------------------------------------|----------------------------------------------------------------------------------------------------------|
| <b>Sb 132</b><br>4.1 m 2.8 m<br>$\beta^-$ 3.7...<br>$\gamma$ 974, 697<br>151, 104...<br>989...              | <b>Sb 133</b><br>2.34 m<br>$\beta^-$ 1.2, 2.9...<br>$\gamma$ 1096, 818<br>2755, 837...<br>g, m                                                 | <b>Sb 134</b><br>10.1 s 0.75 s<br>$\beta^-$ 6.1...<br>$\gamma$ 1279<br>297, 707<br>115...<br>1352... | <b>Sb 135</b><br>1.7 s<br>$\beta^-$ 8.1...<br>$\beta n$ 1.45, 1.04...<br>$\gamma$ 1127, 1279*<br>1380... |
| <b>Sn 131</b><br>50 s 39 s<br>$\beta^-$ 3.4...<br>$\gamma$ 1226<br>480, 305<br>1229...                      | <b>Sn 132</b><br>39.7 s<br>$\beta^-$ 1.8...<br>341, 86, 899<br>47, 993...                                                                      | <b>Sn 133</b><br>1.46 s<br>$\beta^-$ 8.1...<br>962...<br>0.48                                        | <b>Sn 134</b><br>1.050 s<br>$\beta^-$ 7.0...<br>$\beta n$ 1.2, 318, 554<br>972...                        |
| <b>In 130</b><br>0.54 s 0.54 s 0.29 s<br>$\beta^-$ 1221<br>1774...<br>391...<br>1221...<br>g, m m g, m g, m | <b>In 131</b><br>0.32 s 0.35 s 0.261 s<br>$\beta^-$ 6.6...<br>$\gamma$ 1905<br>130<br>8.8...<br>9.2...<br>1221...<br>4273...<br>g, m g, m g, m | <b>In 132</b><br>207 ms<br>$\beta^-$ 6.4, 8.9...<br>375, 4041<br>39...<br>1.023, 0.33                | <b>In 133</b><br>165 ms<br>$\beta^-$ 1.3<br>$\beta n$ 1.561<br>854<br>2005                               |
| <b>Cd 129</b><br>242 ms 104 ms<br>$\beta^-$ 995, 359, 1796...                                               | <b>Cd 130</b><br>162 ms<br>$\beta^-$ 6.2, 8.3...<br>$\gamma$ 1669, 451<br>950, 1170...<br>$\beta n$                                            | <b>Cd 131</b><br>98.0 ms                                                                             | <b>Cd 132</b><br>82 ms                                                                                   |
| <b>Ag 128</b><br>59 ms<br>$\beta^-$ 645, 784...<br>$\beta n$ ?                                              | <b>Ag 129</b><br>~160 ms 46 ms<br>$\beta^-$ $\beta n$ $\beta^-$ $\beta n$                                                                      |                                                                                                      |                                                                                                          |
| <b>Pd 127</b><br>38 ms<br>$\beta^-$ $\beta n$ ? $\beta 2n$ ?                                                | <b>Pd 128</b><br>35 ms<br>$\beta^-$ $\beta n$ ?                                                                                                |                                                                                                      |                                                                                                          |

$(5/2^-)$  ———  
 $(9/2^-)$  - - - - -  
 $(1/2^-)$  ———  
  
 $3/2^-$  ———  
  
 $7/2^-$  ———

$^{133}_{50}\text{Sn}_{83-1}$ 
From ENSDF - Evaluated October 2010

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$^{133}\text{In} \beta^-$  decay
 1996Ho16,2000Ho32

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| Type            | Author                                   | History | Citation            |
|-----------------|------------------------------------------|---------|---------------------|
| Full Evaluation | Yu. Khazov and A. Rodionov, F. G. Kondev |         | NDS 112, 855 (2011) |

Parent:  $^{133}\text{In}$ :  $E=0.0$ ;  $J^\pi=(9/2^+)$ ;  $T_{1/2}=165$  ms 3;  $Q(\beta^-)=12917$  SY; % $\beta^-$  decay=100.0  
 1995JoZZ, 1996Ho16, 2000Ho32:  $^{133}\text{In} \beta^-$  [from  $^{238}\text{U}(p,f)$ ,  $E=1$  GeV]; measured  $E\gamma$ ,  $I\gamma$ ,  $\gamma\gamma$ -,  $n\gamma$ -,  $\beta\gamma$ -c  
 ( $\beta n$ ) $\gamma$ (t). ISOLDE facility; plastic scin  $\beta$  detector, two liquid scin n detectors, two Ge  $\gamma$ -ray detectors

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$^{133}\text{Sn}$  Levels

| $E(\text{level})^\dagger$ | $J^\pi$   |
|---------------------------|-----------|
| 0.0                       | $7/2^-$   |
| 853.7 3                   | $3/2^-$   |
| 1560.9 5                  | $(9/2^-)$ |
| 2004.6 10                 | $(5/2^-)$ |



## LETTERS

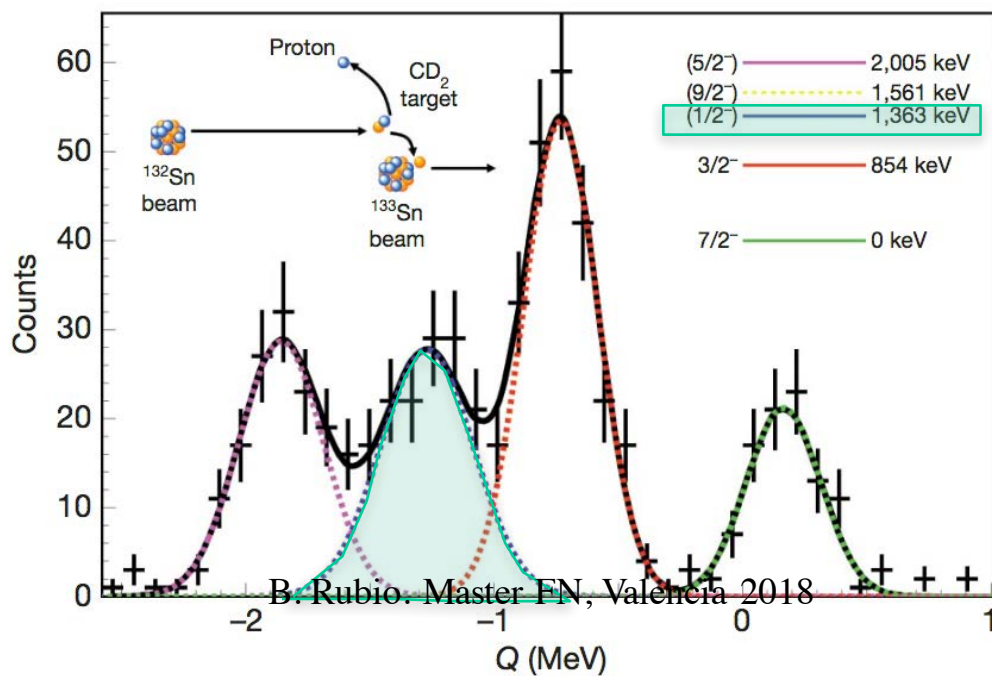
### The magic nature of $^{132}\text{Sn}$ explored through the single-particle states of $^{133}\text{Sn}$

K. L. Jones<sup>1,2</sup>, A. S. Adekola<sup>3</sup>, D. W. Bardayan<sup>4</sup>, J. C. Blackmon<sup>4</sup>, K. Y. Chae<sup>1</sup>, K. A. Chipps<sup>5</sup>, J. A. Cizewski<sup>2</sup>, L. Erikson<sup>5</sup>, C. Harlin<sup>6</sup>, R. Hatarik<sup>7</sup>, R. Kapler<sup>1</sup>, R. L. Kozub<sup>7</sup>, J. F. Liang<sup>4</sup>, R. Livesay<sup>2</sup>, Z. Ma<sup>1</sup>, B. H. Moazen<sup>1</sup>, C. D. Nesaraja<sup>4</sup>, F. M. Nunes<sup>8</sup>, S. D. Pain<sup>2</sup>, N. P. Patterson<sup>6</sup>, D. Shapira<sup>4</sup>, J. F. Shriner Jr<sup>7</sup>, M. S. Smith<sup>4</sup>, T. P. Swan<sup>2,6</sup> & J. S. Thomas<sup>6</sup>

(5/2<sup>-</sup>) — 2,005 keV  
(9/2<sup>-</sup>) — 1,561 keV  
(1/2<sup>-</sup>) — 1,363 keV

3/2<sup>-</sup> — 854 keV

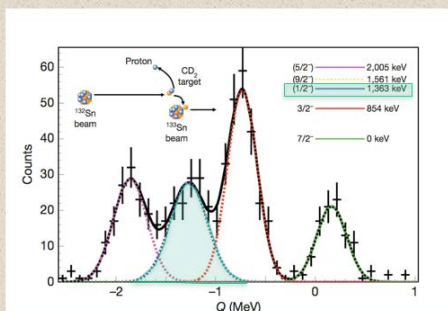
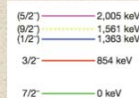
7/2<sup>-</sup> — 0 keV



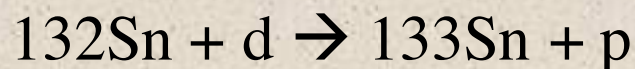
## Conclusion of this part

It is important to go as far as possible from the stability because The primary information on the structure of a nucleus normally comes from the study of its decay.

It is important to have the possibility to perform reactions where either the target or the projectile are radioactive.



Haz radioactivo o target  
Radioactivo?



To produce a beam of  $^{132}\text{Sn}$  is possible (not easy),  
To produce a target of  $^{132}\text{Sn}$  ( $T_{1/2}=40$  s) is impossible

We need to produce a beam of radioactive nuclei

The first thing to worry about is how to produce radioactive nuclei:  
Tool: Reaction or Fission

The second thing to worry is to produce them in a “clean” way

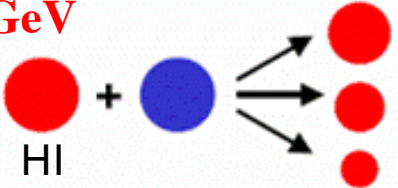
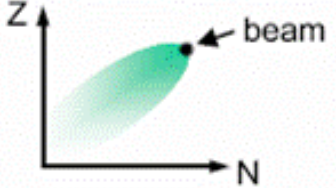
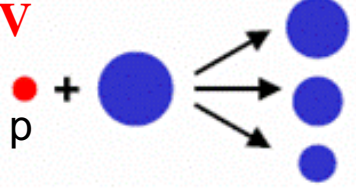
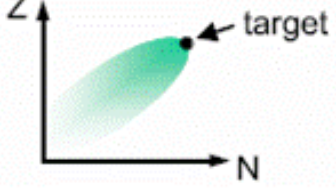
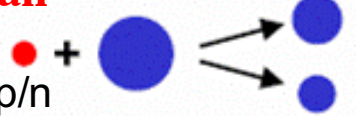
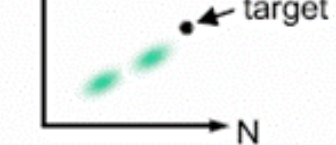
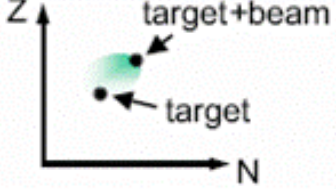
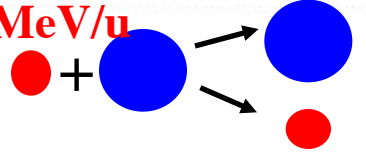
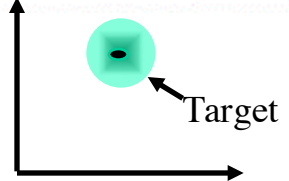
For that it is important to separate them from other radioactive products

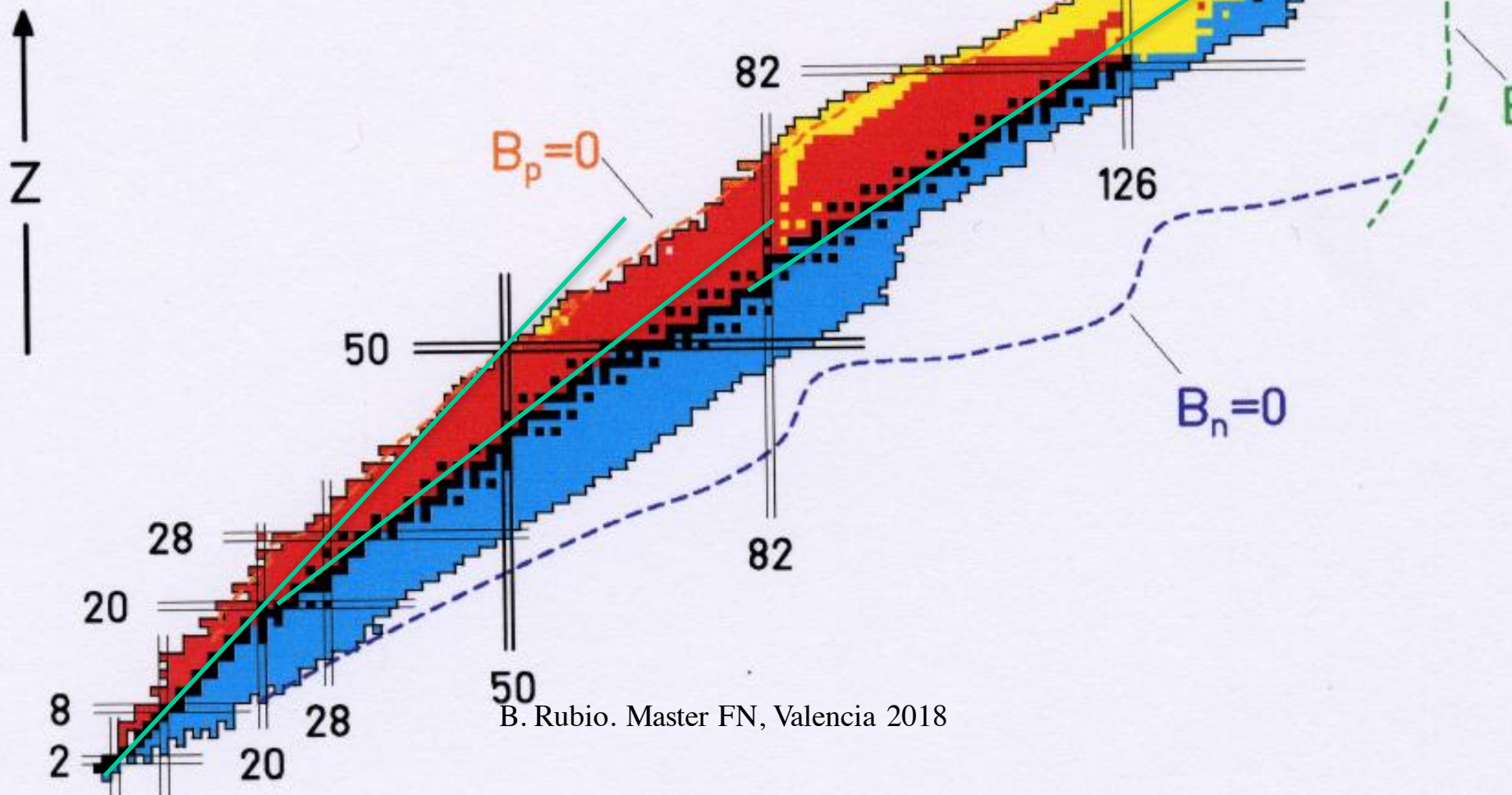
There are two main ways: the Isol method and the fragmentation method

Postacceleration

The first one is known since many years ago

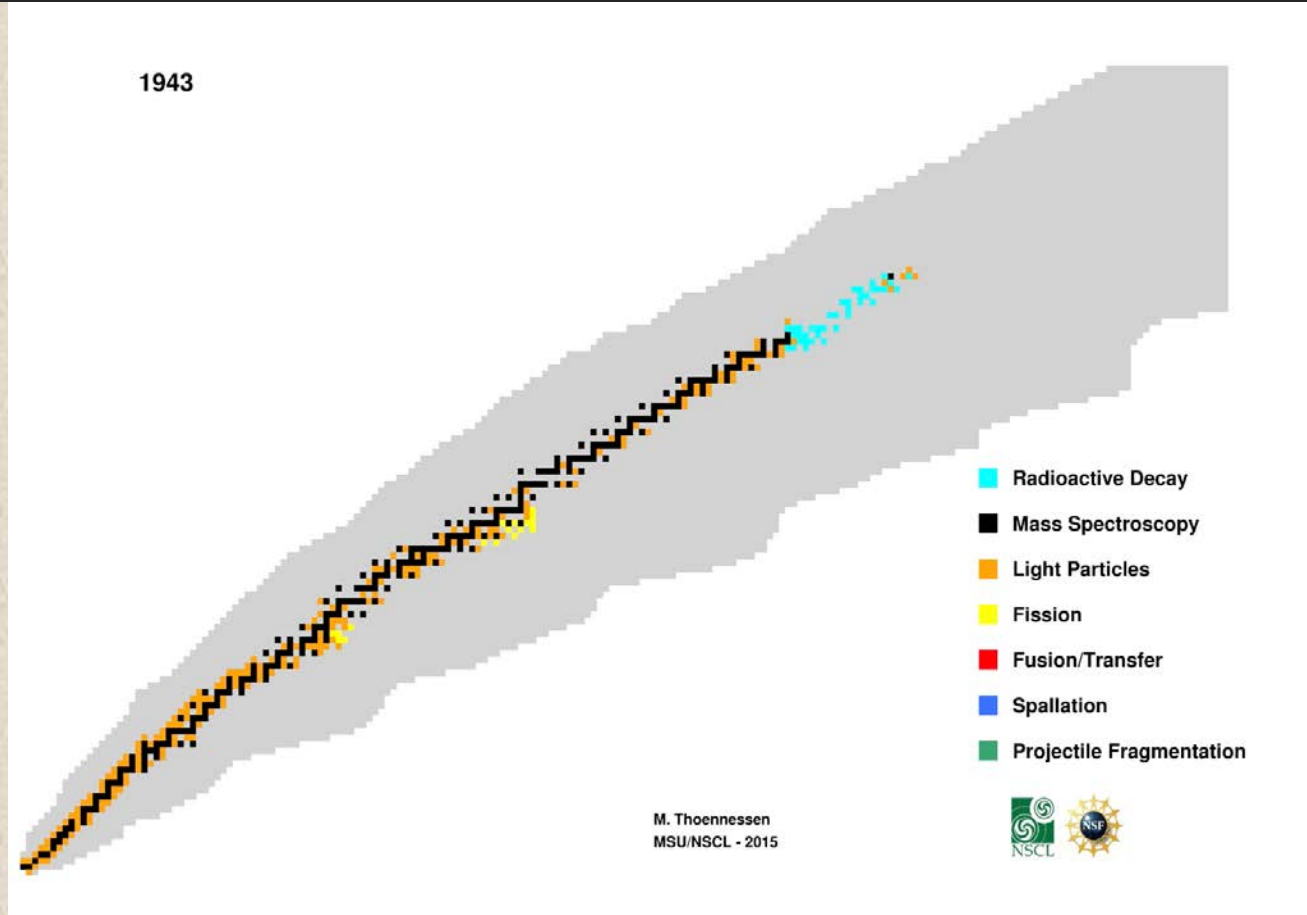
# The reaction to produce Radioactive beams

| Beam + target → products                                                                                   |                    | Product energy                                                        | Number of Products |                                                                                       |
|------------------------------------------------------------------------------------------------------------|--------------------|-----------------------------------------------------------------------|--------------------|---------------------------------------------------------------------------------------|
| <b>GeV</b><br><br>HI      | fragmentation      | <b>“All nuclei”</b><br>$v_{\text{product}} = v_{\text{beam}}$         | up to 1000         |    |
| <b>GeV</b><br><br>p       | spallation         | <b>“All nuclei”</b><br>few MeV/u                                      | up to 1000         |    |
| <b>small</b><br><br>p/n   | fission            | <b>“mainly neutron rich”</b><br>~1 MeV/u                              | few 100            |   |
| <b>5MeV/u</b><br><br>HI | fusion-evaporation | <b>“mainly proton rich rich”</b><br>$E_R = \frac{m_p}{m_p + m_t} E_P$ | few ( $\leq 20$ )  |  |
| <b>20 MeV/u</b><br>     | Transfer reactions | <b>“close to the stability”</b><br>5-100 MeV/u                        | Very few:<br>2-5   |  |



# Historical view to the discovery of new isotopes

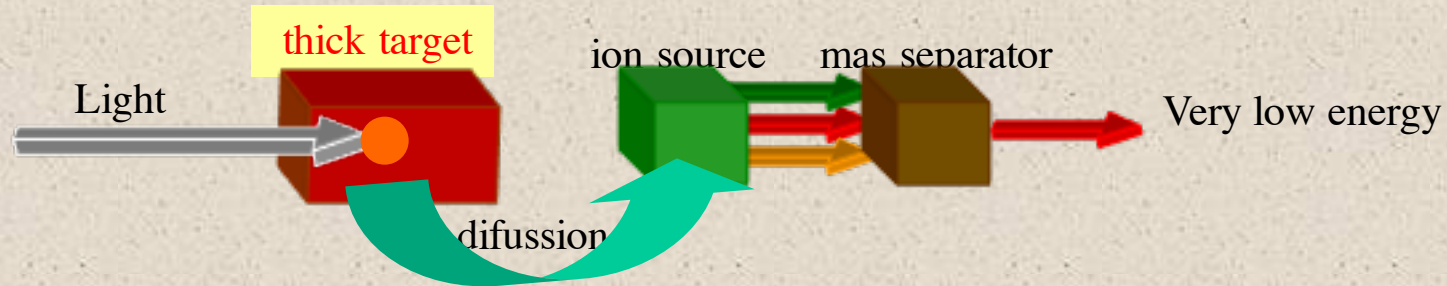
<https://people.nscl.msu.edu/~thoennes/isotopes/yearchart-2015.mp4>



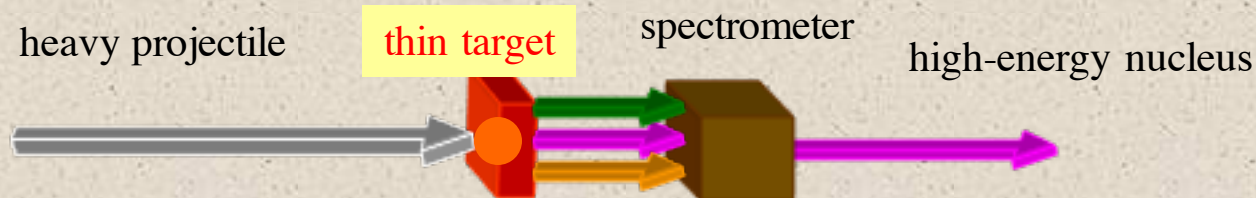
The second thing is to separate them from other radioactive products

## Radioactive nuclei production techniques

### Isotopic separation on-line (ISOL)

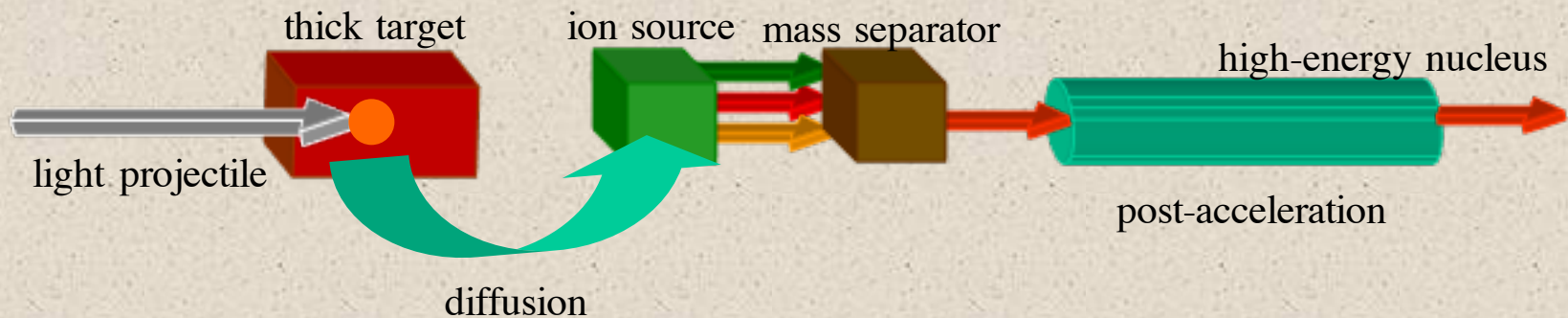


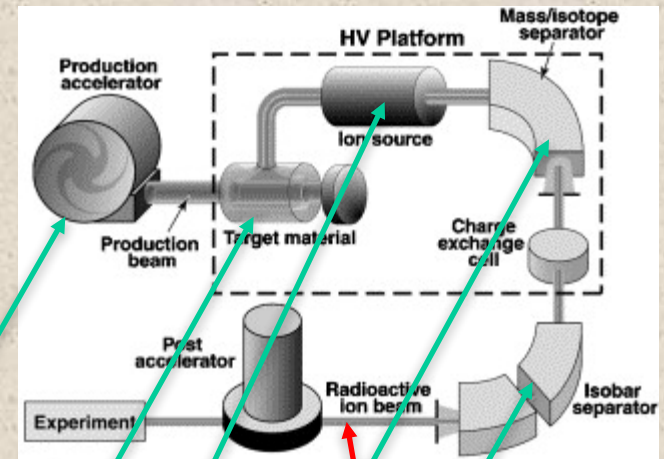
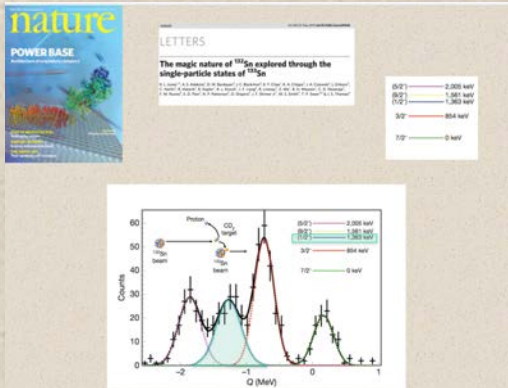
### In-flight fragmentation



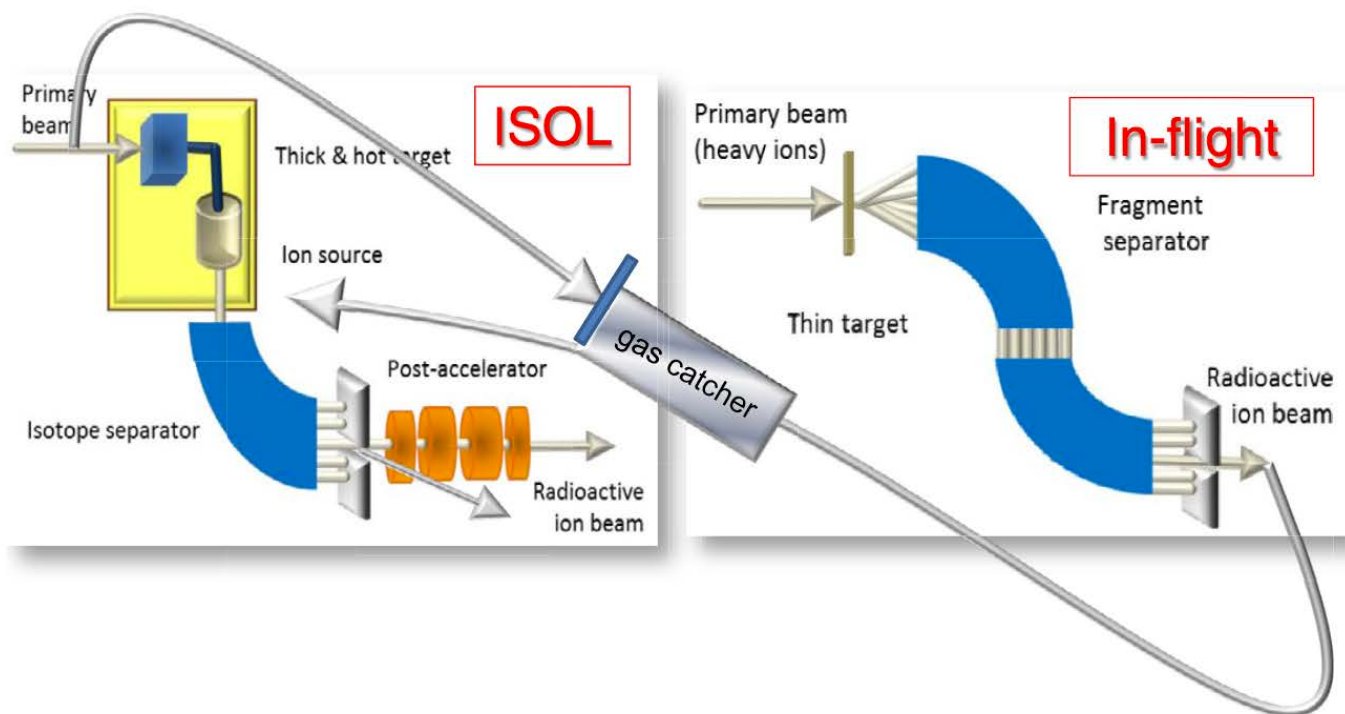
# Production techniques Rad Beams: post acceleration

## ➤ Isotopic separation on-line (ISOL)





1. Protons accelerated in the cyclotron
2. Uranium Carbide power target
3. About 100 species produced, some of them ionised
4.  $^{132}\text{Sn}$  Mass separated
5.  $^{132}\text{Sn}$  separated
6.  $^{132}\text{Sn}$  accelerated to 4.8 MeV x A
7. Experiment



**Figure 1.** A schematic drawing of the ISOL and in-flight based production methods for RIBs.

<http://iopscience.iop.org/article/10.1088/0031-8949/2013/T152/014023/pdf>

# difficulties

- extremely low production cross sections,
- overwhelming production of unwanted species in the same nuclear reaction,
- very short half lives of the nuclei of interest.

# challenges

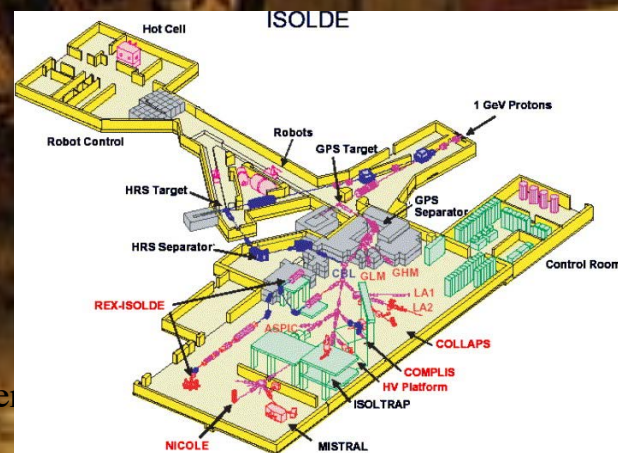
- As much intensity as possible (good driver, targets to stand the intensity)
- Good separation
- Bring the species to a “clean environment” Make the process as fast as possible

On 16 October 1967, the first experiments were carried out at the Isotopic Separator On Line ISOLDE CERN.

CERN's longest serving experimental facility

Physics Reports 403–404 (2004) 459–469

Juha Äystö



B. Rubio. Master FN, Valen

ORGANISATION EU  
**CERN** EUROPE

Téléphone : (022) 45 98 11

Télégramme : CERILAB-GENEVE

Télex : GENEVE - 218 48

Adresse la correspondance :

CERN  
1211 Genève 23

Votre référence :

Notre référence CERN/9750

Dr. A. Bohr, Copenhagen  
Dr. Karl Ove Nielsen, Aarhus  
Prof. Jean Teillac, Orsay  
Prof. W. Gentner, Heidelberg  
Dr. Roald Tangen, Oslo  
Dr. Göran Andersson, Gothenburg  
Dr. Ingmar Bergström, Stockholm  
cc. Prof. Preiswerk, for inform.

Meyrin-Geneve, le 17 December, 1964

.....Furthermore, CERN is willing to contribute one sixth of the cost of the basic equipment needed as estimated in the proposal\*, which corresponds to the participation of the CERN Nuclear Chemistry Group in the project, and one sixth of the cost of maintaining, running, and developing the above-mentioned basic equipment after the construction period. It is understood that the rest of the cost will have to be shared by the participating groups. These groups will also have to bring to CERN any specialized measuring equipment which they might wish to use in connection with the project.

Yours sincerely,

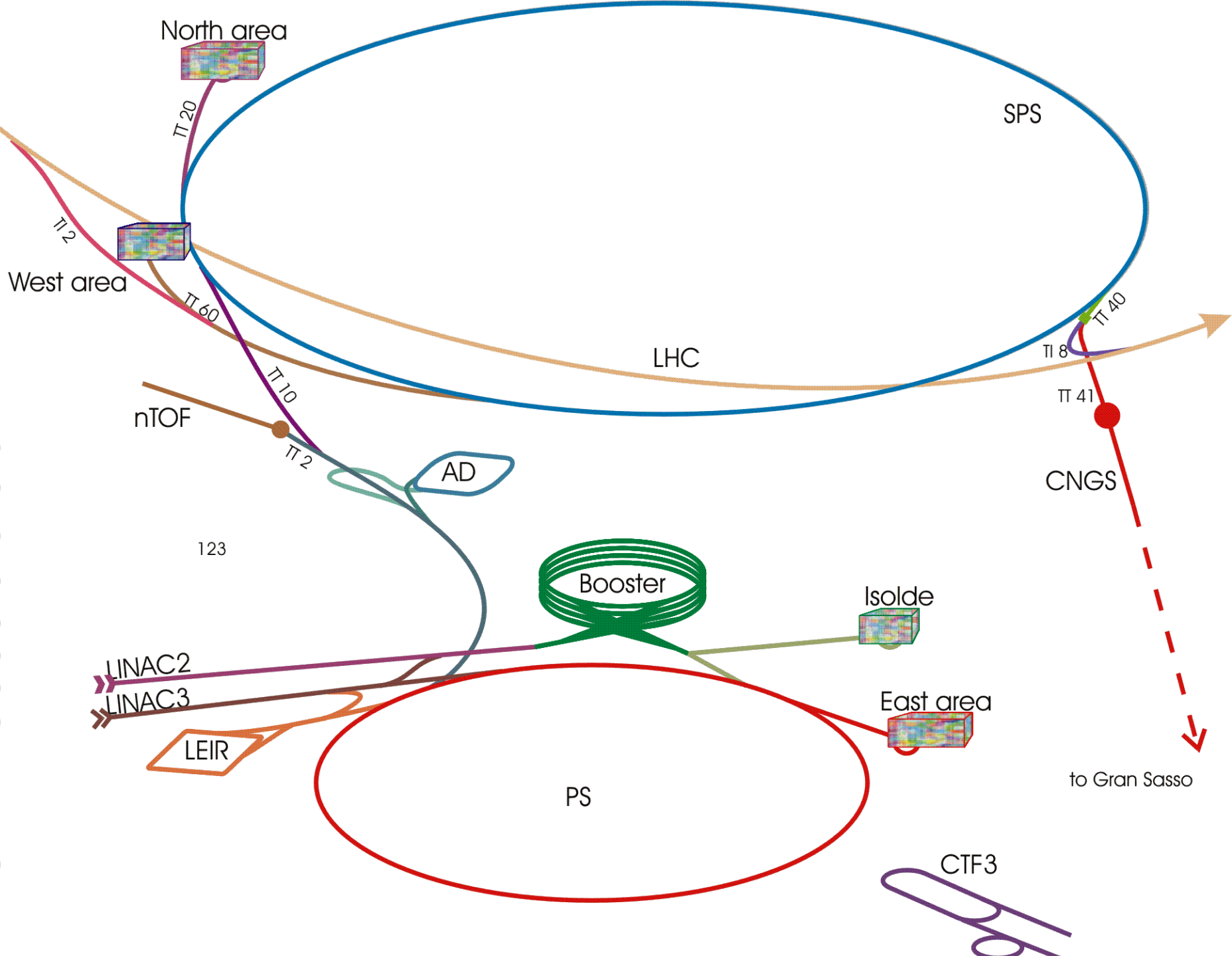
*Viktor F. Weisskopf*

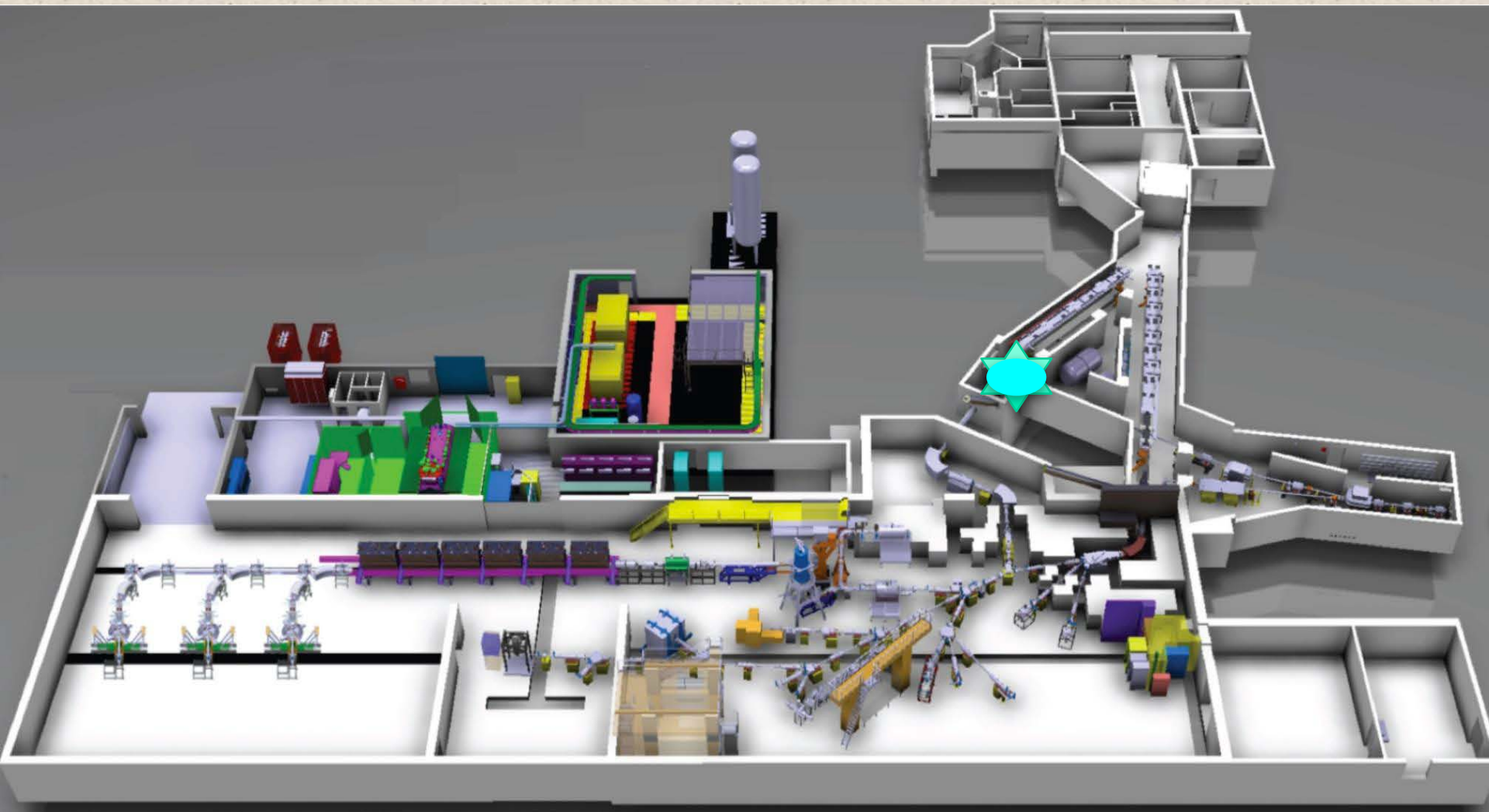


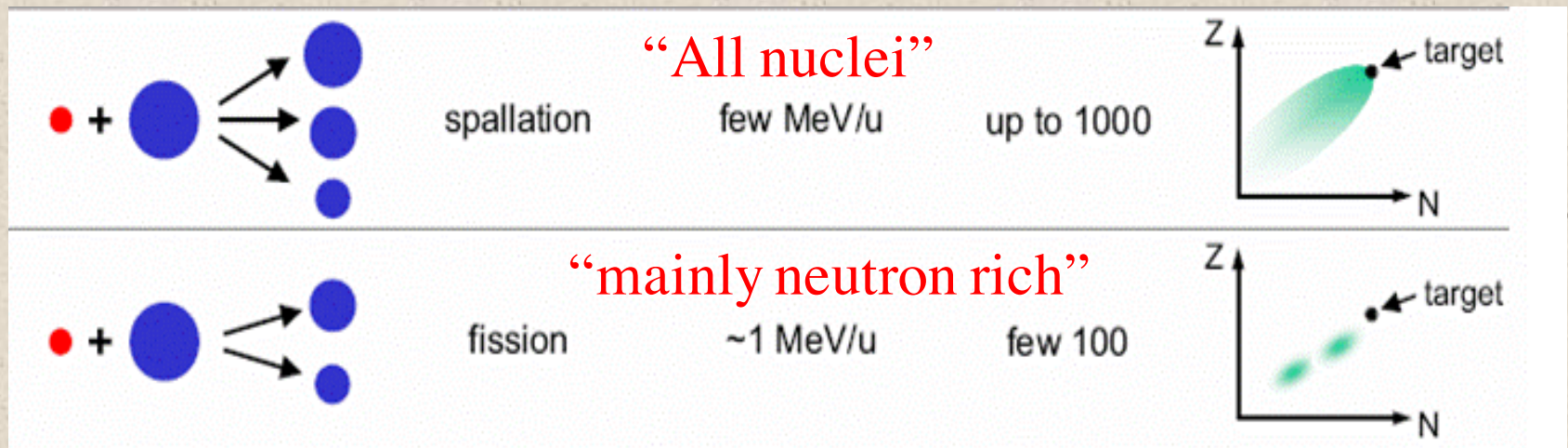
1967 First experiment  
P of 600 MeV from the PSC

1992 moved to PS booster (today)  
(last year 50 years celebration)

# CERN accelerators



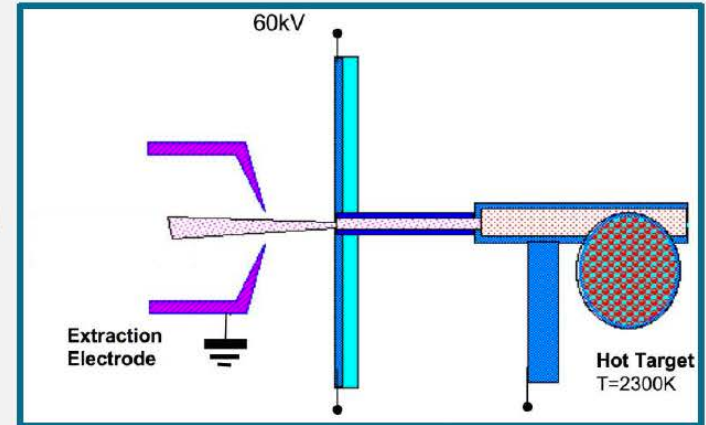




- 😊 High intensity p beam
- 😊 thick target (100% of range) => high beam current (upto  $10^{16} \text{ s}^{-1}$ )
- 😞 long extraction and ionization time (ms)
- 😞 Chemistry dependent

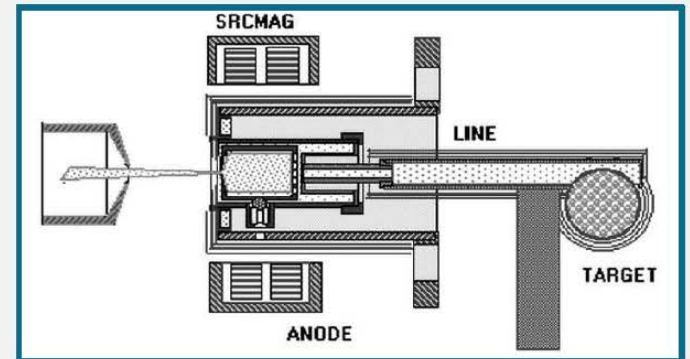
## Hot surface ion source

- The ioniser is a hot tube
- Material with a higher work function than the element of interest
- Heated up to 2400 degrees



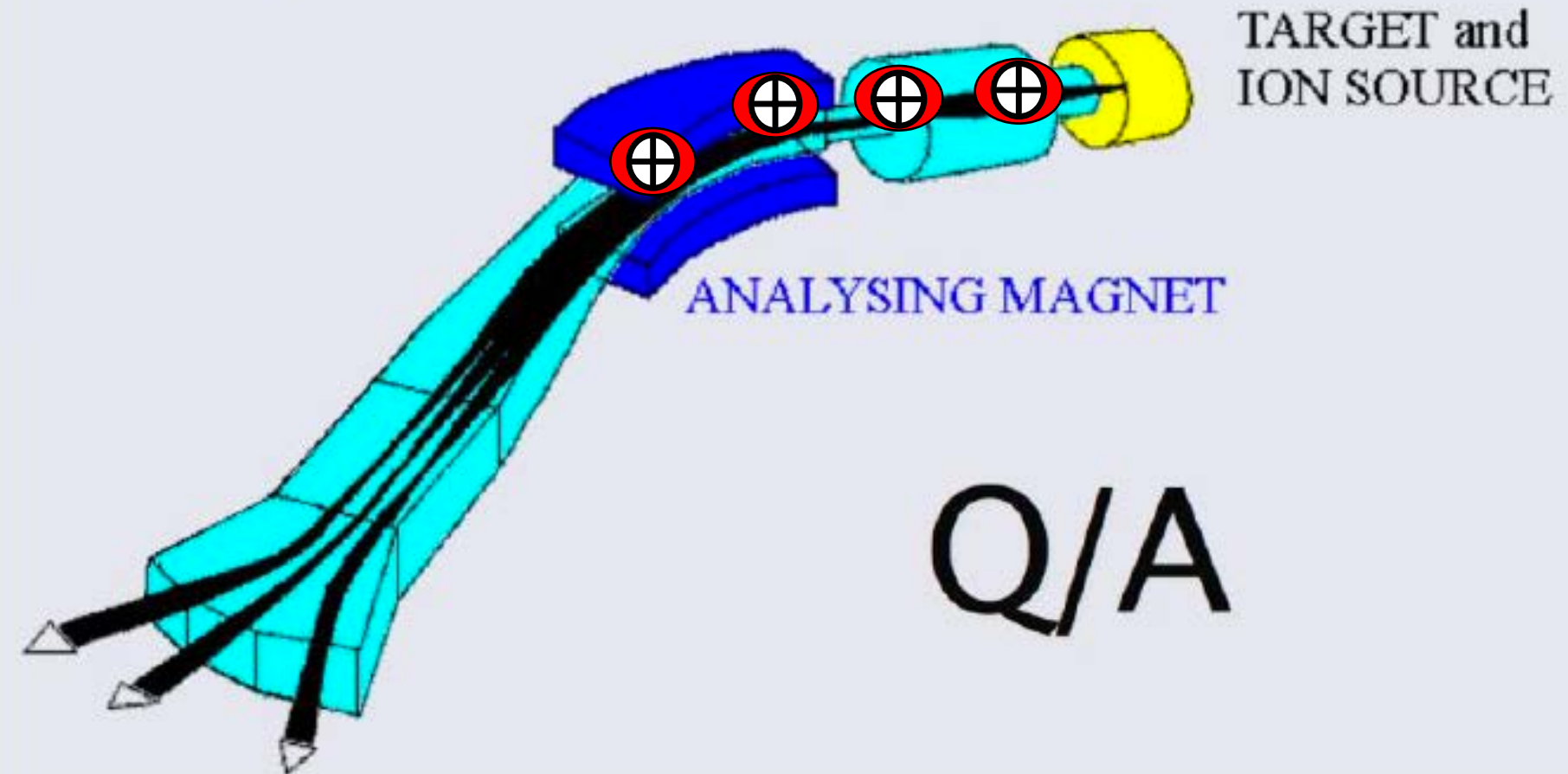
## Plasma ion source

- Plasma: gas mixture (Ar and Xe) ionised by accelerated electrons
- Hot or cool transfer line



# The Separator

## Isotope Separation

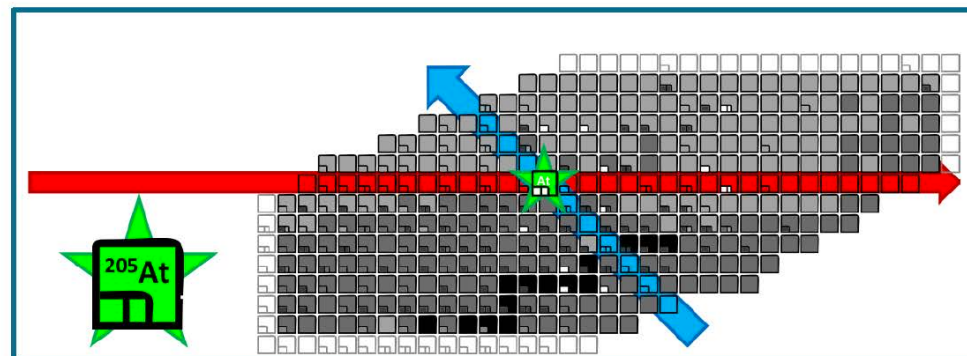
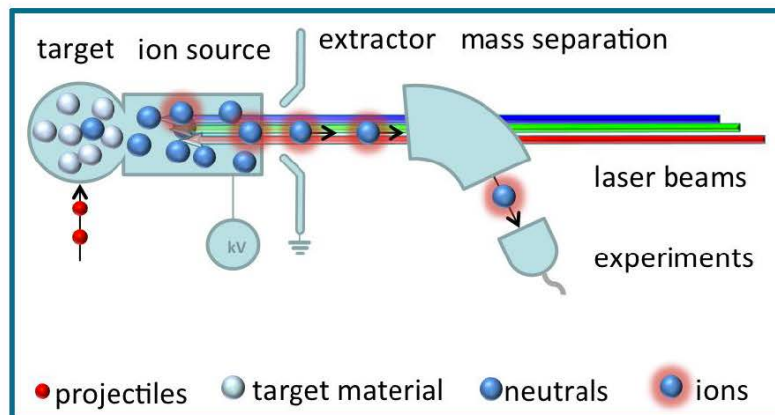
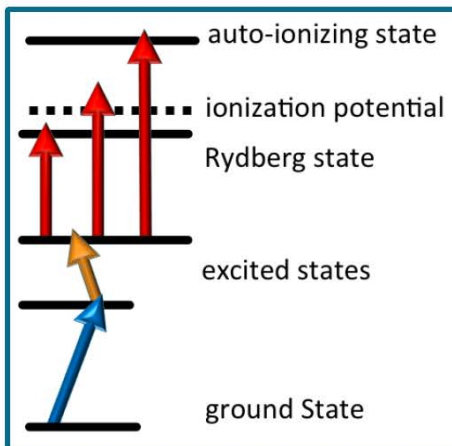




## Ion sources

### Laser ion source

- 2- or 3-step ionisation
- Isotope and isomer selection
- Universal (almost)



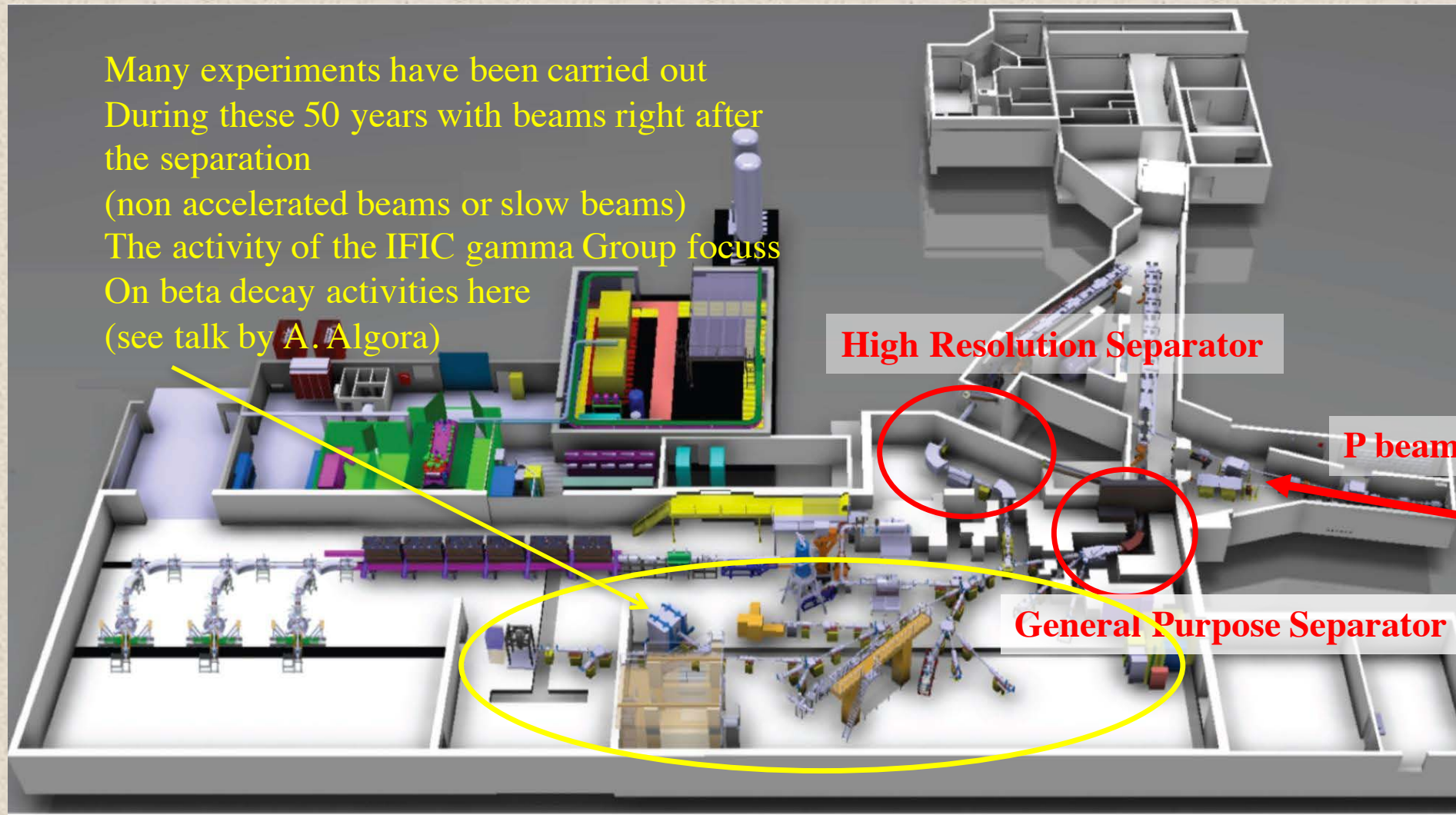
# Isotopes produced at ISOLDE

|    |    |    |    |    |    |    |    |    |    |    |     |     |     |     |    |    |    |    |    |   |   |    |    |  |  |
|----|----|----|----|----|----|----|----|----|----|----|-----|-----|-----|-----|----|----|----|----|----|---|---|----|----|--|--|
| H  |    |    |    |    |    |    |    |    |    |    |     |     |     |     |    |    | He |    |    |   |   |    |    |  |  |
| Li | Be |    |    |    |    |    |    |    |    |    |     |     |     |     |    |    |    | B  | C  | N | O | F  | Ne |  |  |
| Na | Mg |    |    |    |    |    |    |    |    |    |     |     |     |     |    |    |    | Al | Si | P | S | Cl | Ar |  |  |
| K  | Ca | Sc | Ti | V  | Cr | Mn | Fe | Co | Ni | Cu | Zn  | Ga  | Ge  | As  | Se | Br | Kr |    |    |   |   |    |    |  |  |
| Rb | Sr | Y  | Zr | Nb | Mo | Tc | Ru | Rh | Pd | Ag | Cd  | In  | Sn  | Sb  | Te | I  | Xe |    |    |   |   |    |    |  |  |
| Cs | Ba | La | Hf | Ta | W  | Re | Os | Ir | Pt | Au | Hg  | Tl  | Pb  | Bi  | Po | At | Rn |    |    |   |   |    |    |  |  |
| Fr | Ra | Ac | Rf | Db | Sg | Bh | Hs | Mt | Ds | Rg | 112 | 113 | 114 | 115 |    |    |    |    |    |   |   |    |    |  |  |
|    |    |    |    |    |    |    |    |    |    |    |     |     |     |     |    |    |    |    |    |   |   |    |    |  |  |
|    |    |    | Ce | Pr | Nd | Pm | Sm | Eu | Gd | Tb | Dy  | Ho  | Er  | Tm  | Yb | Lu |    |    |    |   |   |    |    |  |  |
|    |    |    | Th | Pa | U  | Np | Pu | Am | Cm | Bk | Cf  | Es  | Fm  | Md  | No | Lr |    |    |    |   |   |    |    |  |  |

ION SOURCE

|       |         |        |
|-------|---------|--------|
| +     | SURFACE | -      |
| hot   | PLASMA  | cooled |
| LASER |         |        |

Many experiments have been carried out  
During these 50 years with beams right after  
the separation  
(non accelerated beams or slow beams)  
The activity of the IFIC gamma Group focuss  
On beta decay activities here  
(see talk by A. Algora)



**High Resolution Separator**

**P beam**

**General Purpose Separator**

## ISOLDE

The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

17 DECEMBER 1964

CERN approves the online separator project

View 



## ISOLDE

The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

26 JUNE 1992

First experiment at the ISOLDE Proton-Synchrotron Booster

View 



## ISOLDE

The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

31 OCTOBER 2001

New accelerator, REX-ISOLDE, goes live

View 



## ISOLDE

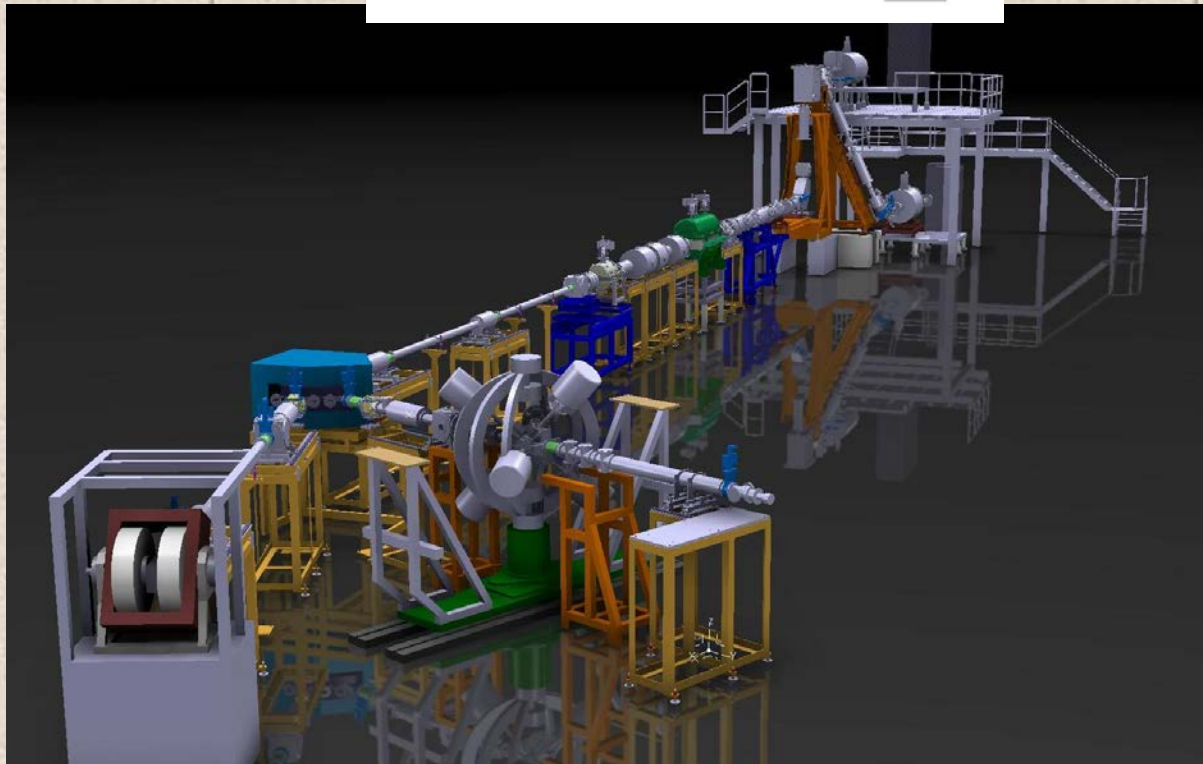
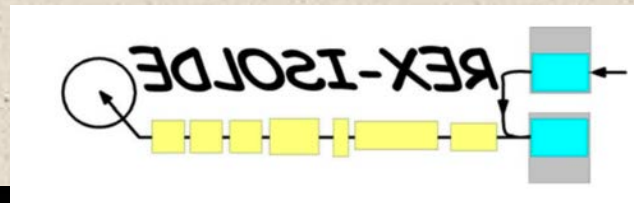
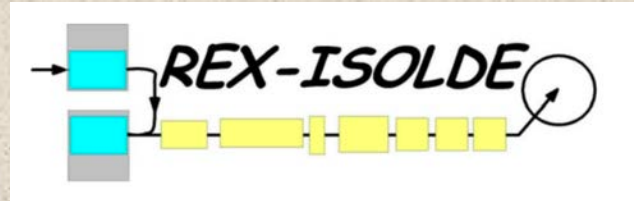
The On-Line Isotope Mass Separator ISOLDE is a facility dedicated to the production of a large variety of radioactive ion beams

22 OCTOBER 2015

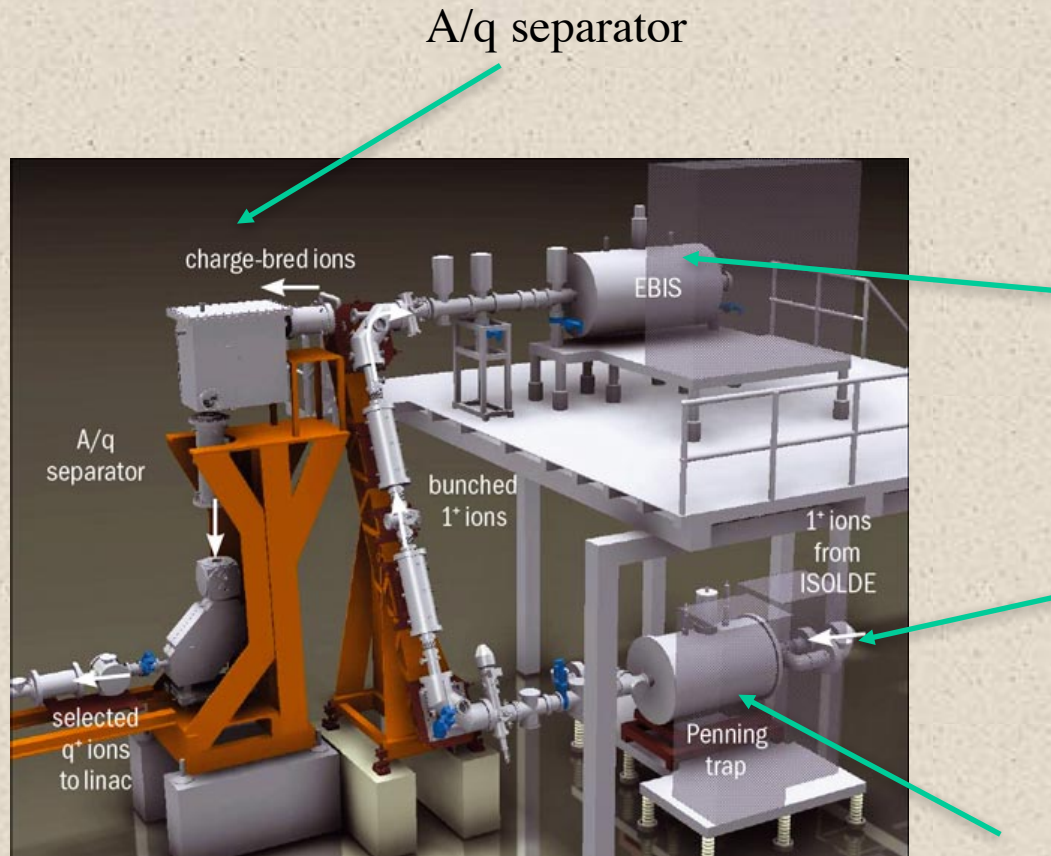
First radioactive isotope beam accelerated in HIE ISOLDE

View 





# EBIS Ion Source and beam handling: 5-15% efficiency

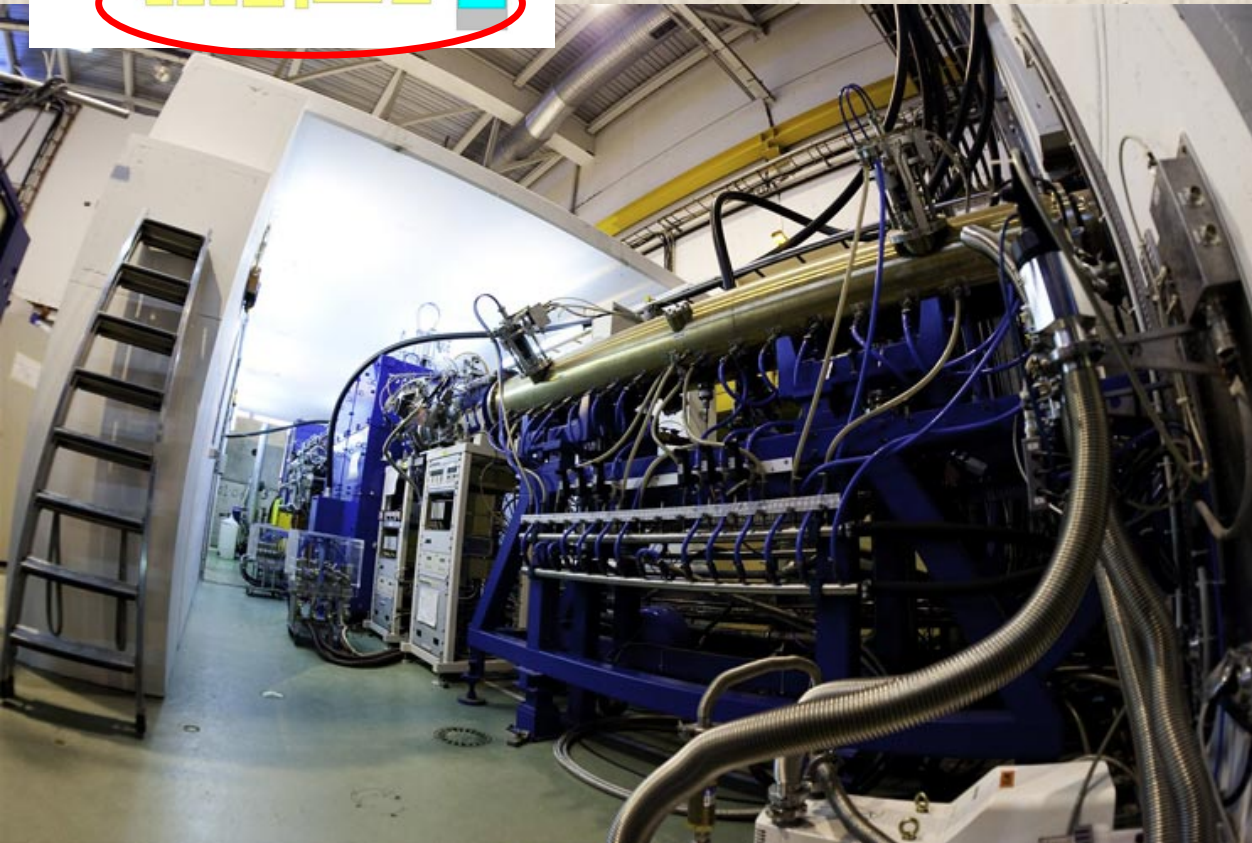
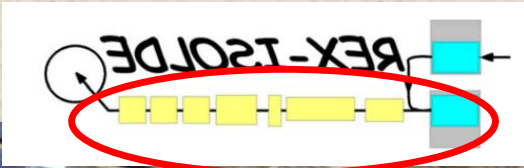


In the EBIS the  $1^+$  to  $n^+$  ion conversion takes place

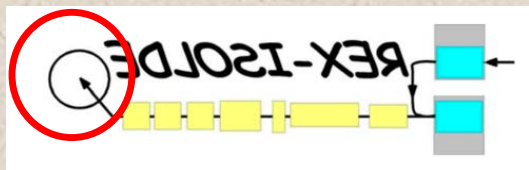
Beams from ISOLDE with  $1^+$  charge state

Preparation trap

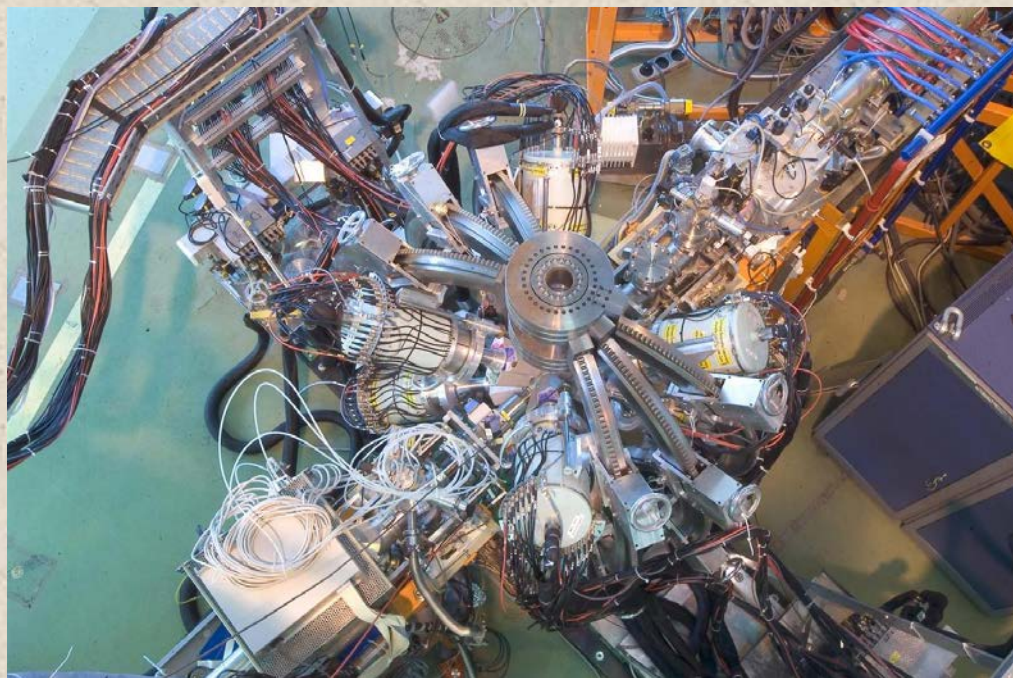
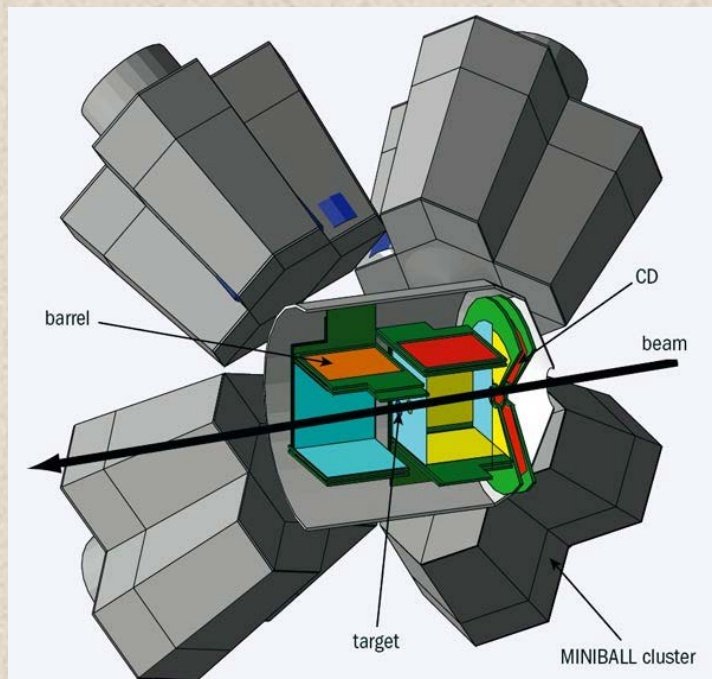
# Post accelerator: Normal Conducting LINAC



However, energies only 0.8 to 3 MeV/u which too low for Nuclear reactions



## Miniball gamma array

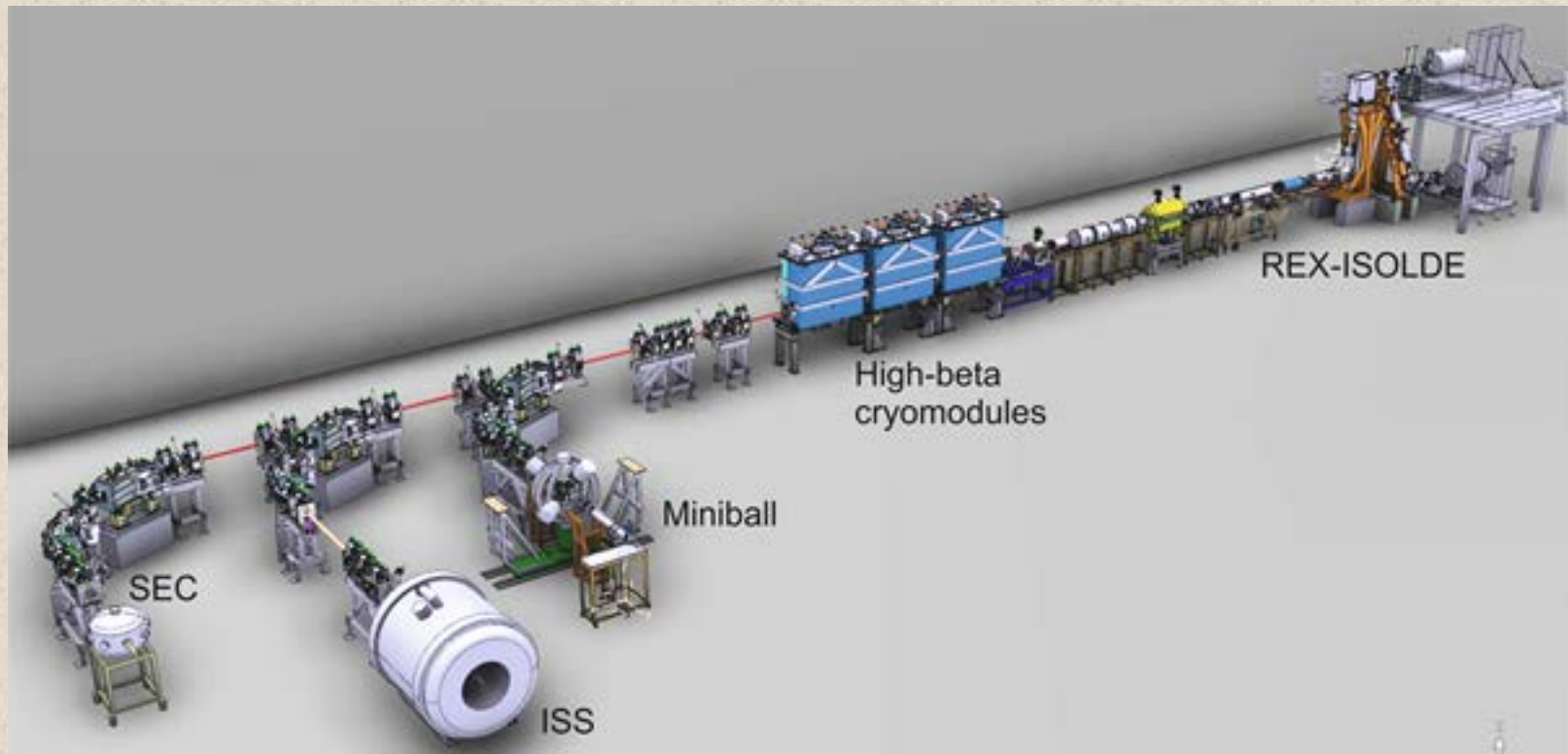


24 six-fold segmented, tapered, encapsulated high-purity germanium crystals

# HIE ISOLDE

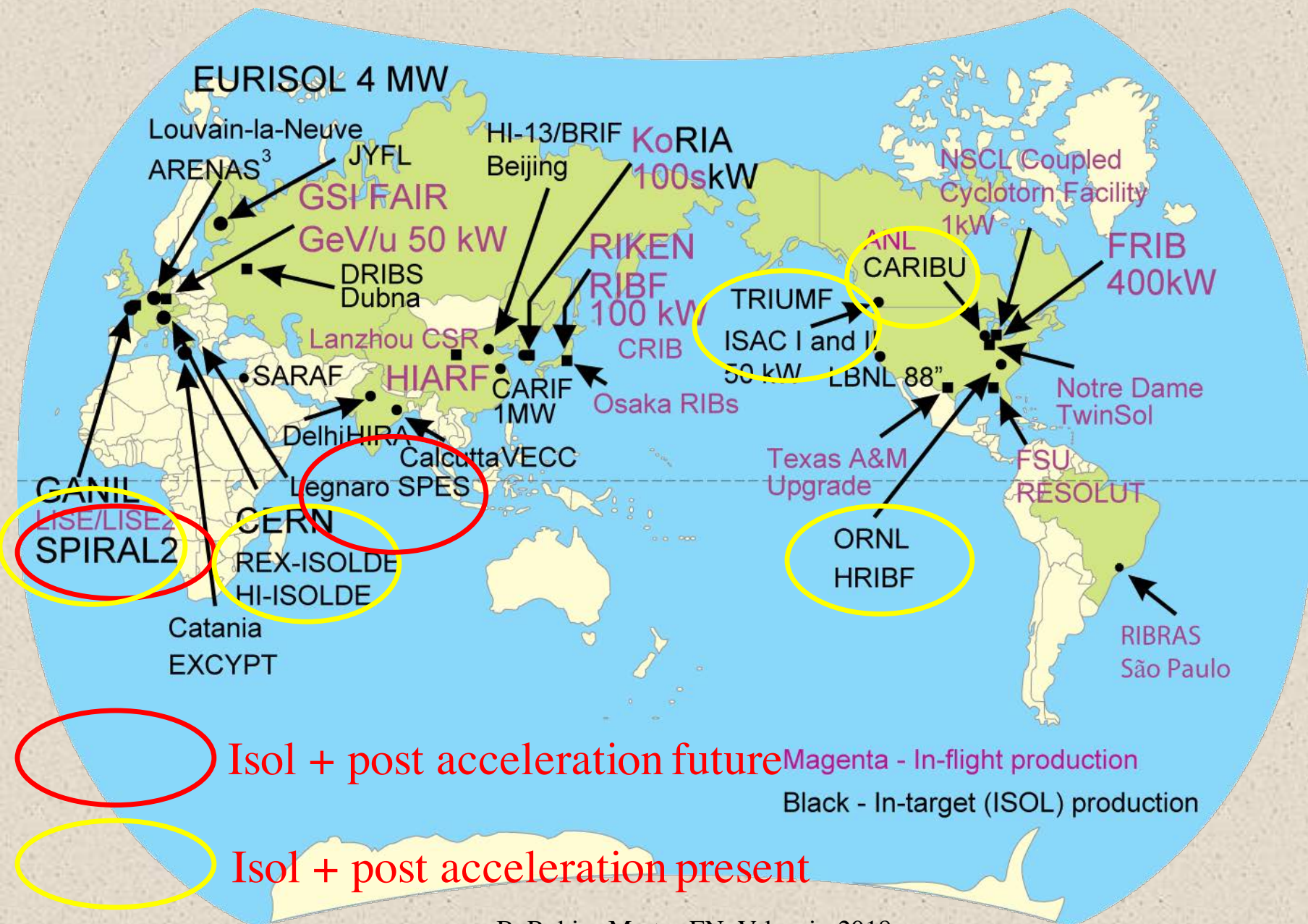
High Intensity and Energy ISOLDE (up to 10 MeV per nucleon)

4.3 MeV per nucleon



First experiments in 2015

B. Rubio. Master FN, Valencia 2018



# Future ISOL facilities

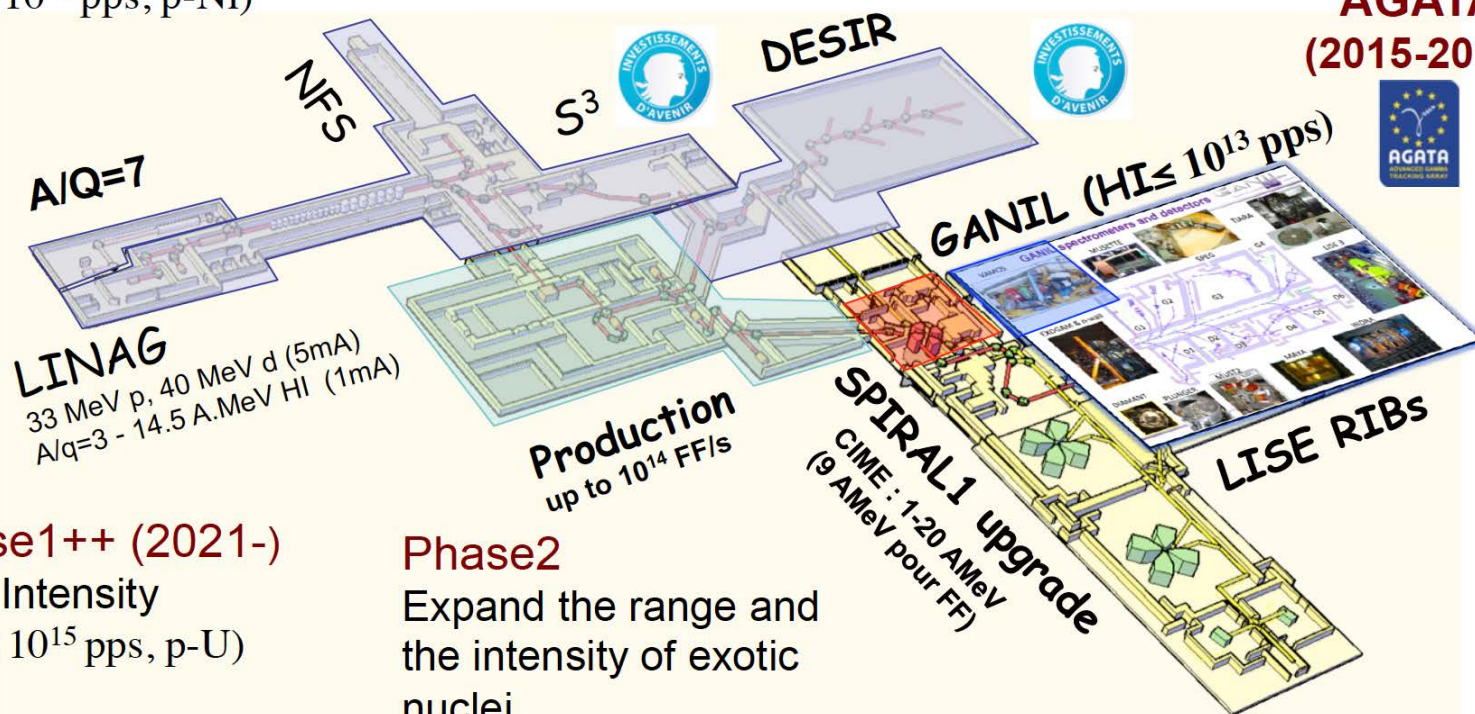
## Phase1 (2016-)

Increase the intensity of stable beams  
High intense neutron source  
( $HI \leq 10^{15}$  pps, p-Ni)

## DESIR Phase1+ (2020-)

Low energy facility

## AGATA (2015-2018)



## Phase1++ (2021-)

High Intensity  
( $HI \leq 10^{15}$  pps, p-U)

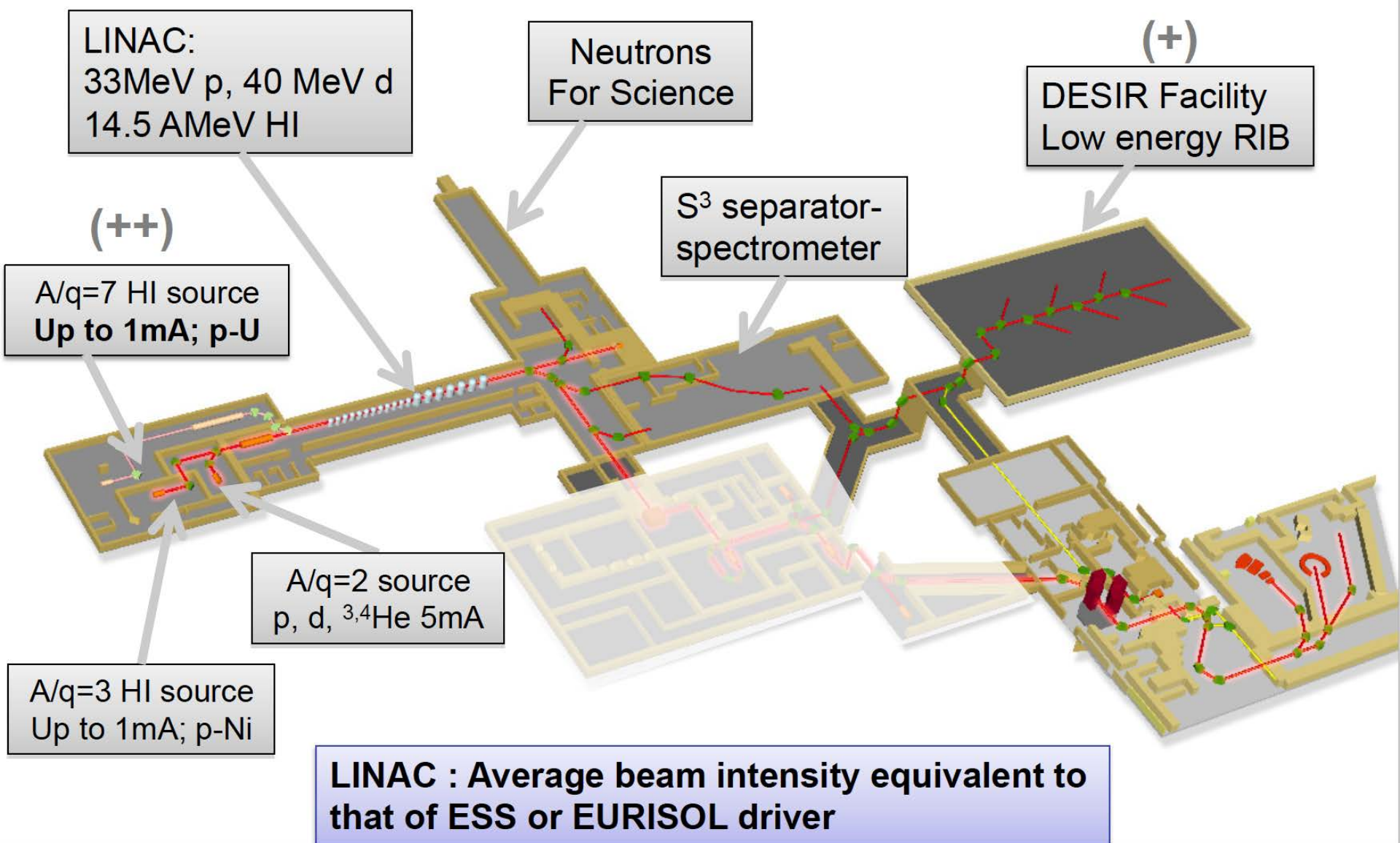
## Phase2

Expand the range and  
the intensity of exotic  
nuclei

## SPIRAL1 Upgrade (2017-)

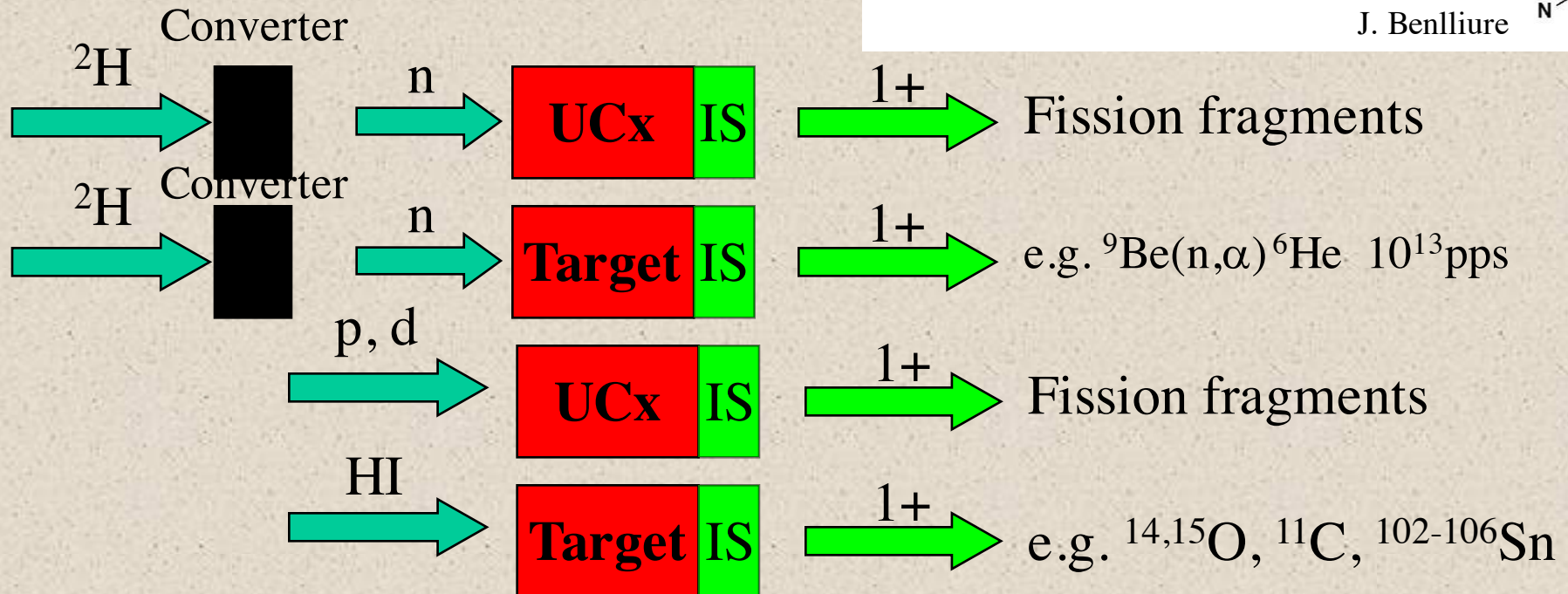
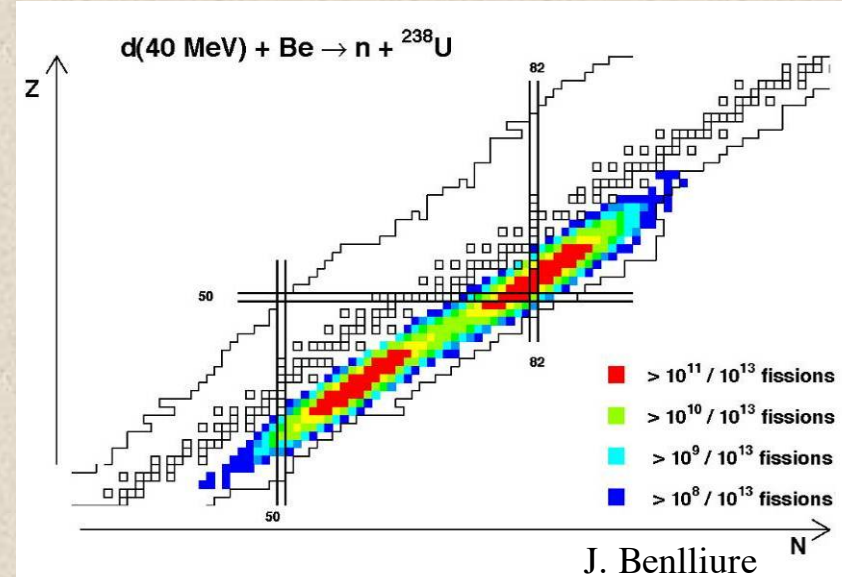
New light RIBs from  
beam/target fragmentation

# Near Future : SPIRAL2 Phase 1



# ISOL Rare Isotope Beams at SPIRAL 2

Unfortunately the neutron rich side of The Spiral 2 Facility is not financed at The moment



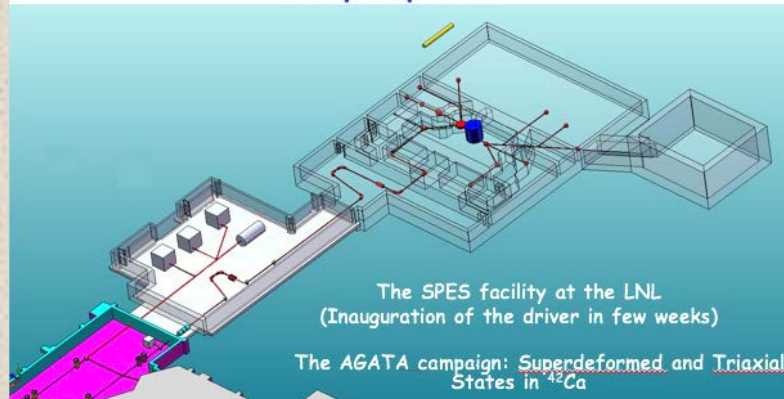
Fortunately and strategically a new project for neutron rich (fission), has been approved in Legnaro (Italy), and is already under construction

## SPES Selective Production of Exotic Species

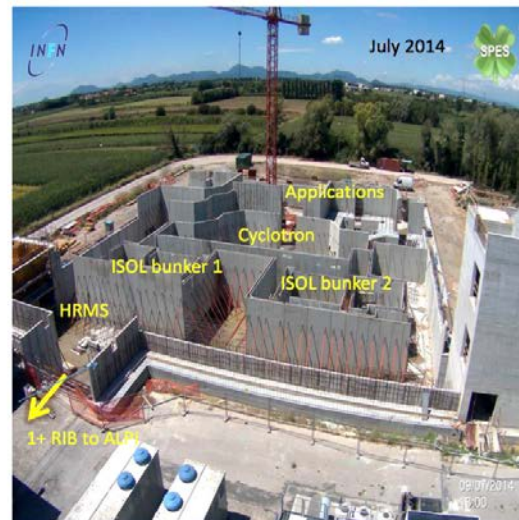


### SPES Project

#### The SPES radioactive ion beam facility: status and perspectives



Giacomo de Angelis  
INFN Laboratori Nazionali di Legnaro  
for The SPES collaboration



July 2014



SPES building

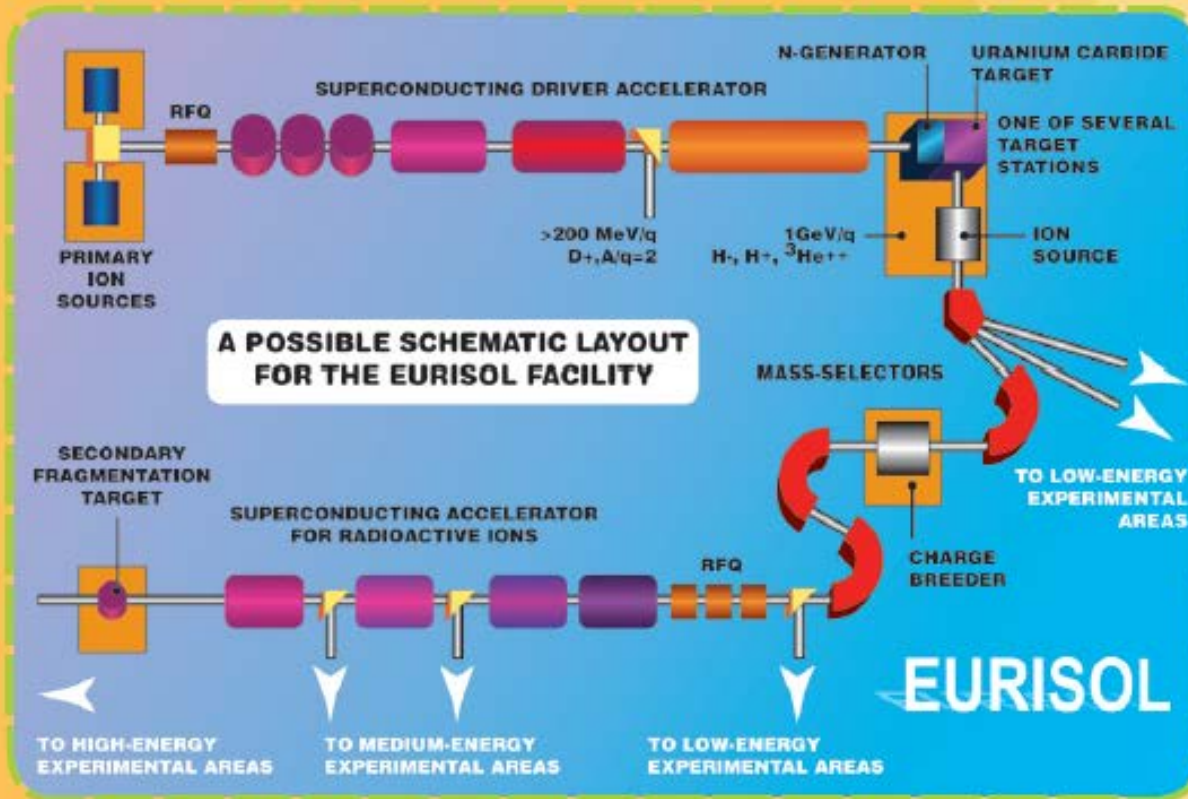
50x60 m<sup>2</sup>  
-3 to +11 m height  
24.000 m<sup>3</sup> of concrete  
1.150 tons iron  
3-4 m shielding wall thick



Oct 2013

## Far in the Future ISOL facility

# Furthering target and accelerator technology



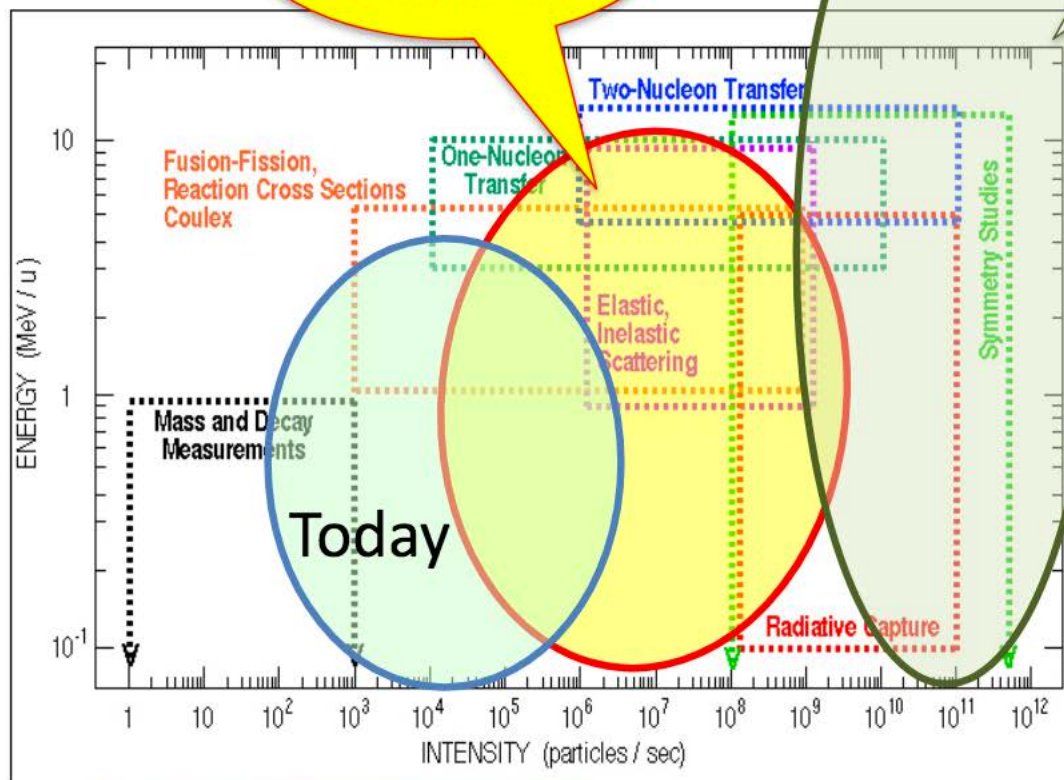
A 4-year Design Study began in 2005 in order to work on the technologically challenging aspects of the project, the instrumentation and the radiation safety issue. Synergy with other projects is being examined, including a feasibility study for the new 'beta-beam' neutrino proposal, forming an integral part of the Design Study. After this, possible sites will be evaluated, and the community will be ready for a full Engineering Design, to be followed by construction of the facility.

Researchers and engineers of several European laboratories are collaborating in twelve tasks to further the EURISOL design.

The EURISOL layout consists of a superconducting linear accelerator providing protons of energy 1 GeV and an impressive power of 5 MW, but also capable of accelerating deuterons,  $^3\text{He}$  and ions up to mass 40. The beams will impinge simultaneously on two types of targets, either directly or after conversion of the protons into neutrons through a loop containing 1 ton of mercury surrounded by kilograms of fissile material. The unstable nuclei produced diffuse out of the target, are ionized and selected, and can be used directly at low energy or reaccelerated by another linear accelerator to energies up to 150 MeV per nucleon in order to induce nuclear reactions.

# Physics with ISOL RIB Intensity & Energy domains

# Precision nuclear structure physics & applications



HIE-ISOLDE,  
SPES, SPIRAL2,  
ISOL@MYRRHA  
**EURISOL-DF**

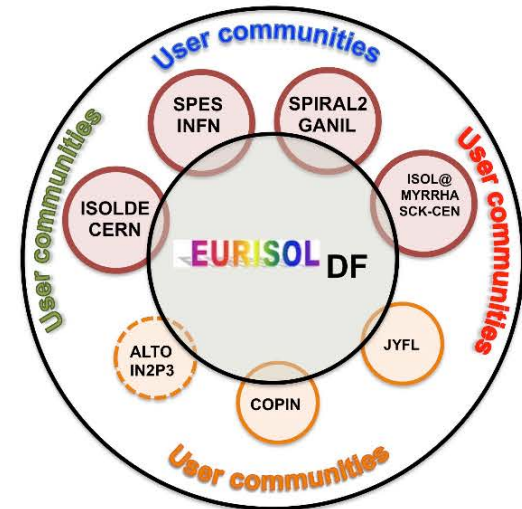
**EURISOL**



-> EURISOL-DF (Distributed Facility) Initiative from 2014 as an intermediate step towards EURISOL

**EURISOL DF**

# EURISOL – Distributed Facility (DF)

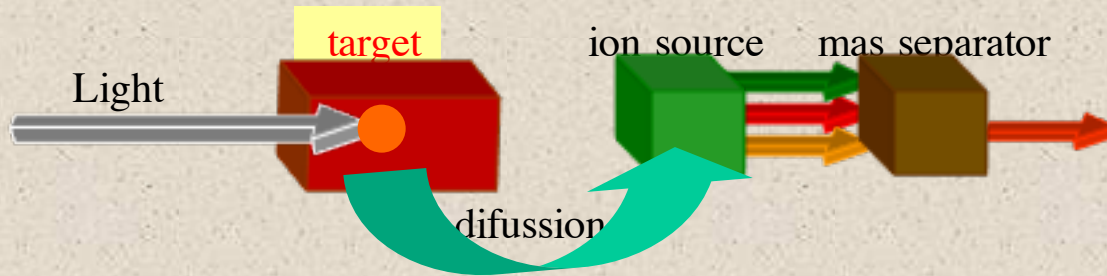


**Core members :**  
HIE-ISOLDE/CERN  
SPES-INFN  
SPIRAL2-GANIL  
ISOL@MYRRHA-SCK\*CEN

**Associated Members**  
JYFL, Finland  
COPIN Consortium, Poland  
(ALTO, Orsay)

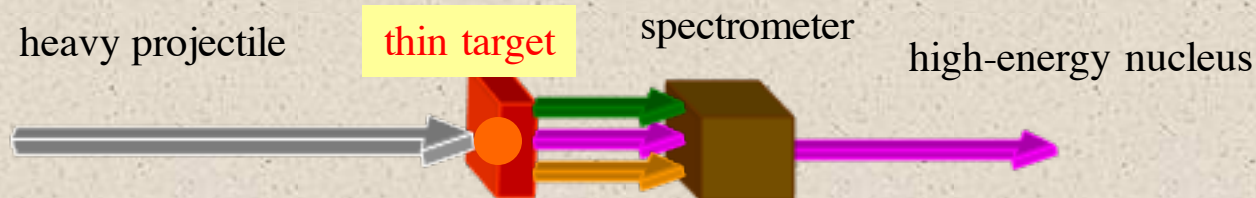
# Radioactive nuclei production techniques

## Isotopic separation on-line (ISOL)



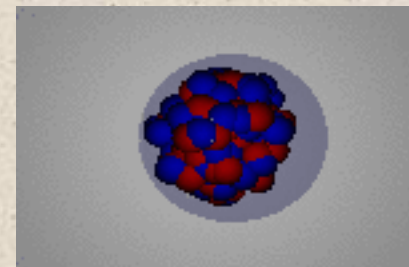
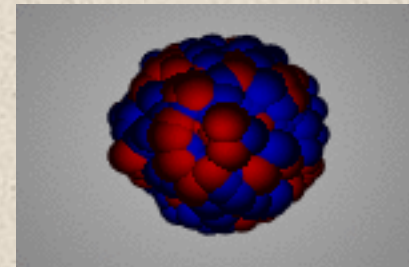
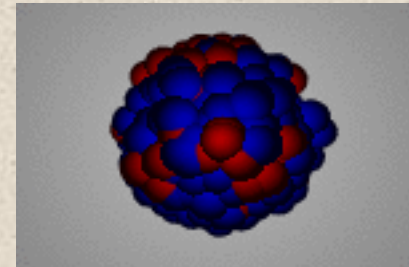
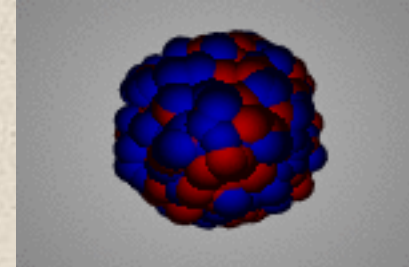
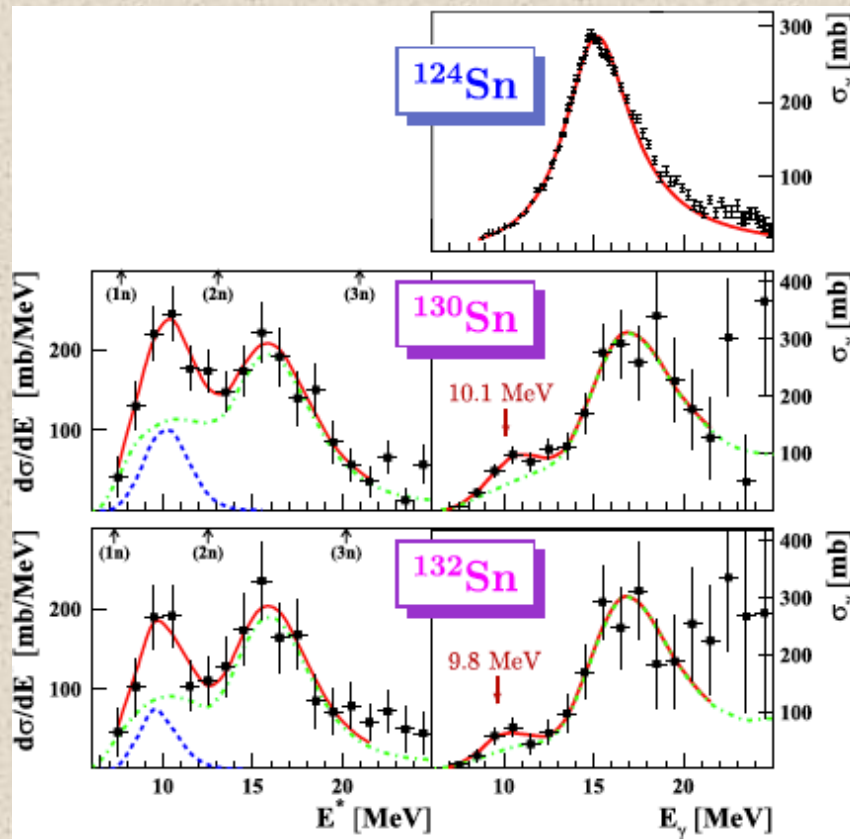
- 😊 We can use thick production targets (100% of energy range)  
=> and high injector beam current (upto  $10^{16} \text{ s}^{-1}$ )
- 😞 long extraction and ionization time (ms)
- 😞 chemistry dependent (disadvantages and some advantages)

## In-flight fragmentation

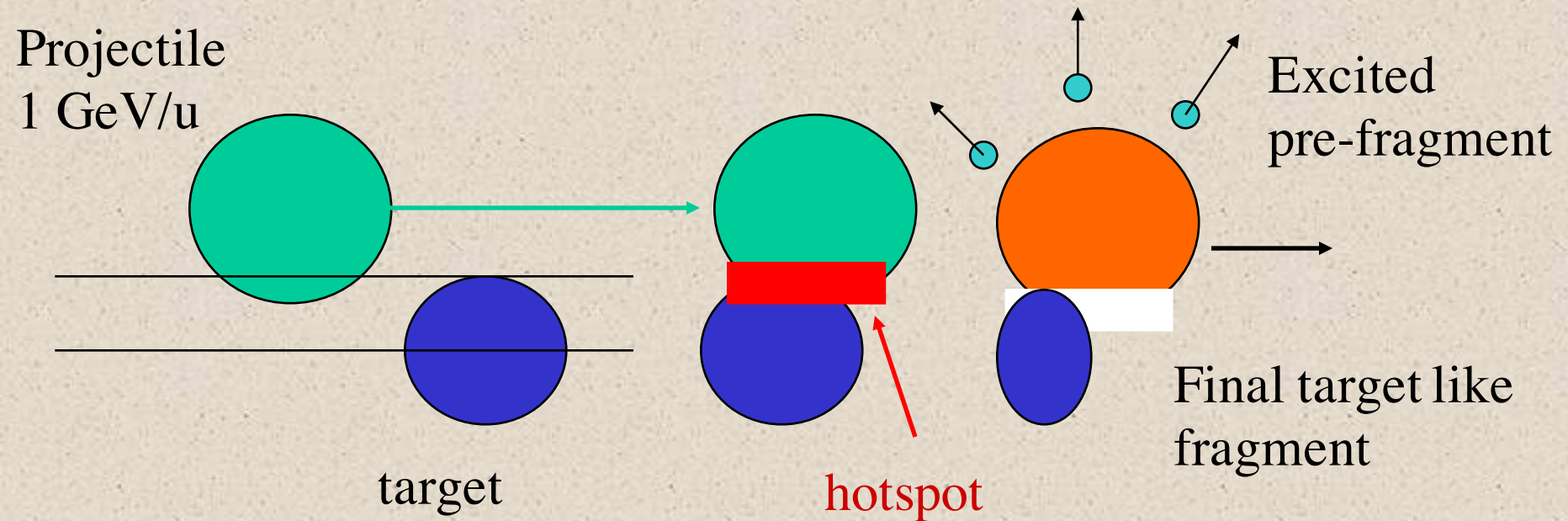


- 😊 short separation+identification time (100 ns)
- 😞 thinner targets (10% of range) => lower beam currents (upto  $10^{12} \text{ s}^{-1}$ )
- 😊 chemistry independent

# Los haces radioactivos

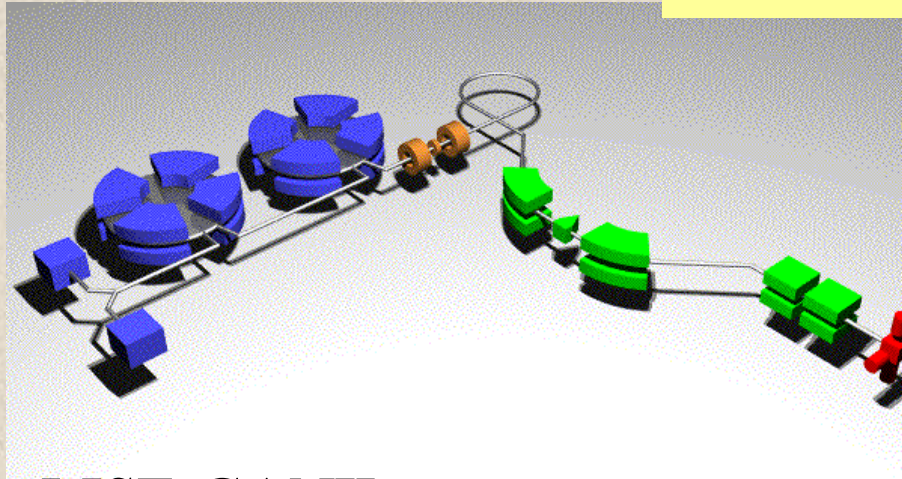


# Projectile Fragmentation Reactions

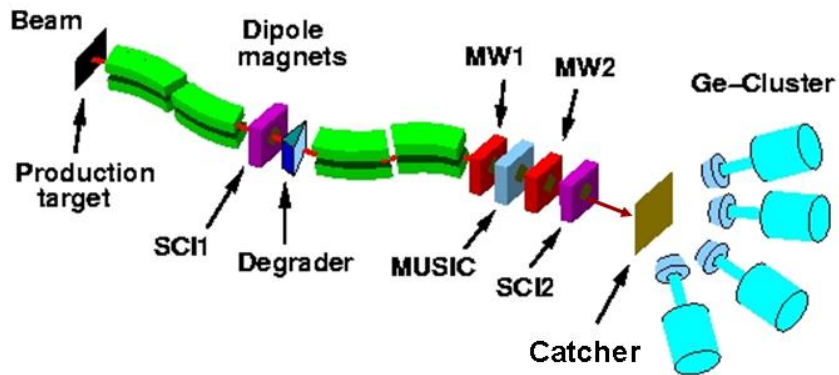


**Energy (velocity) of beam > Fermi velocity** inside nucleus  $\sim 30$  MeV/u  
Can 'shear off' different combinations of protons and neutrons.  
Large variety of exotic nuclear species created, all at forward angles with  $\sim$ beam velocity.

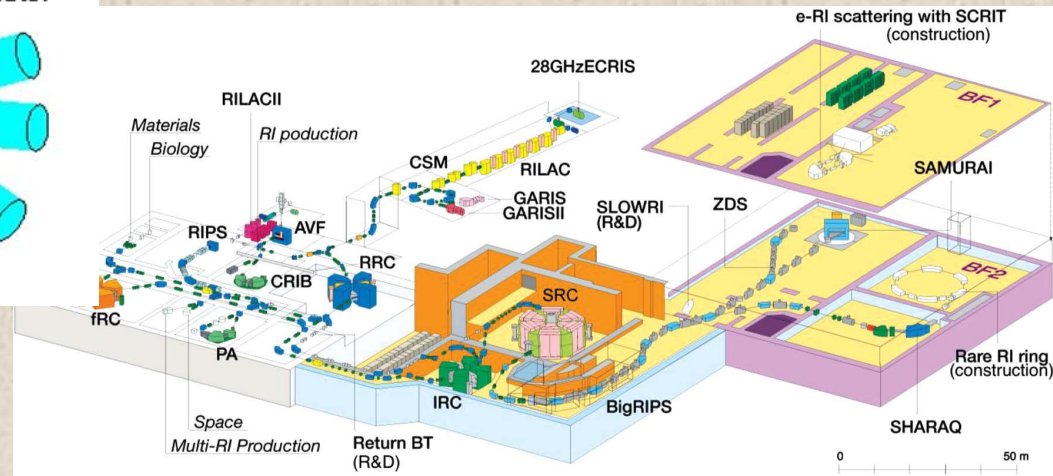
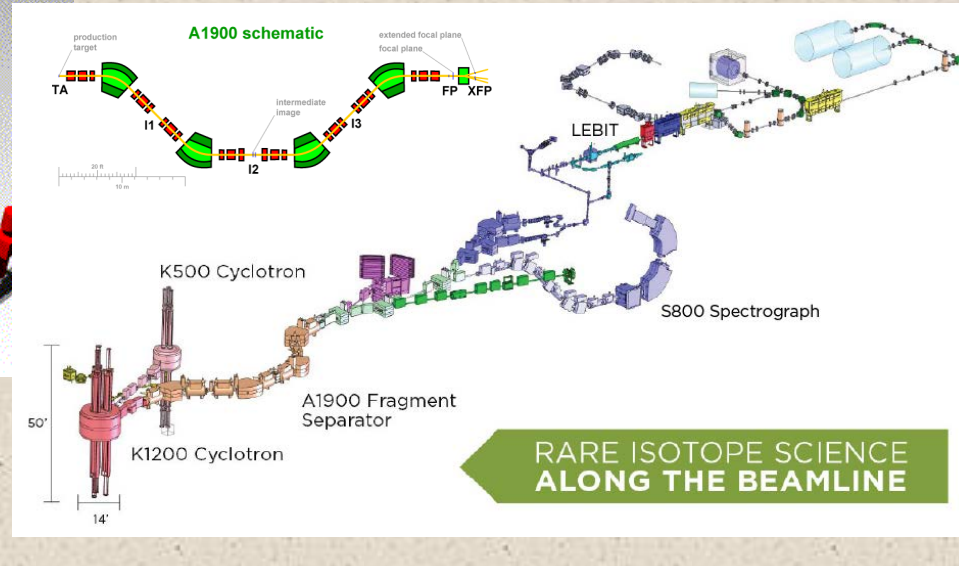
# Actual Fragment Separators

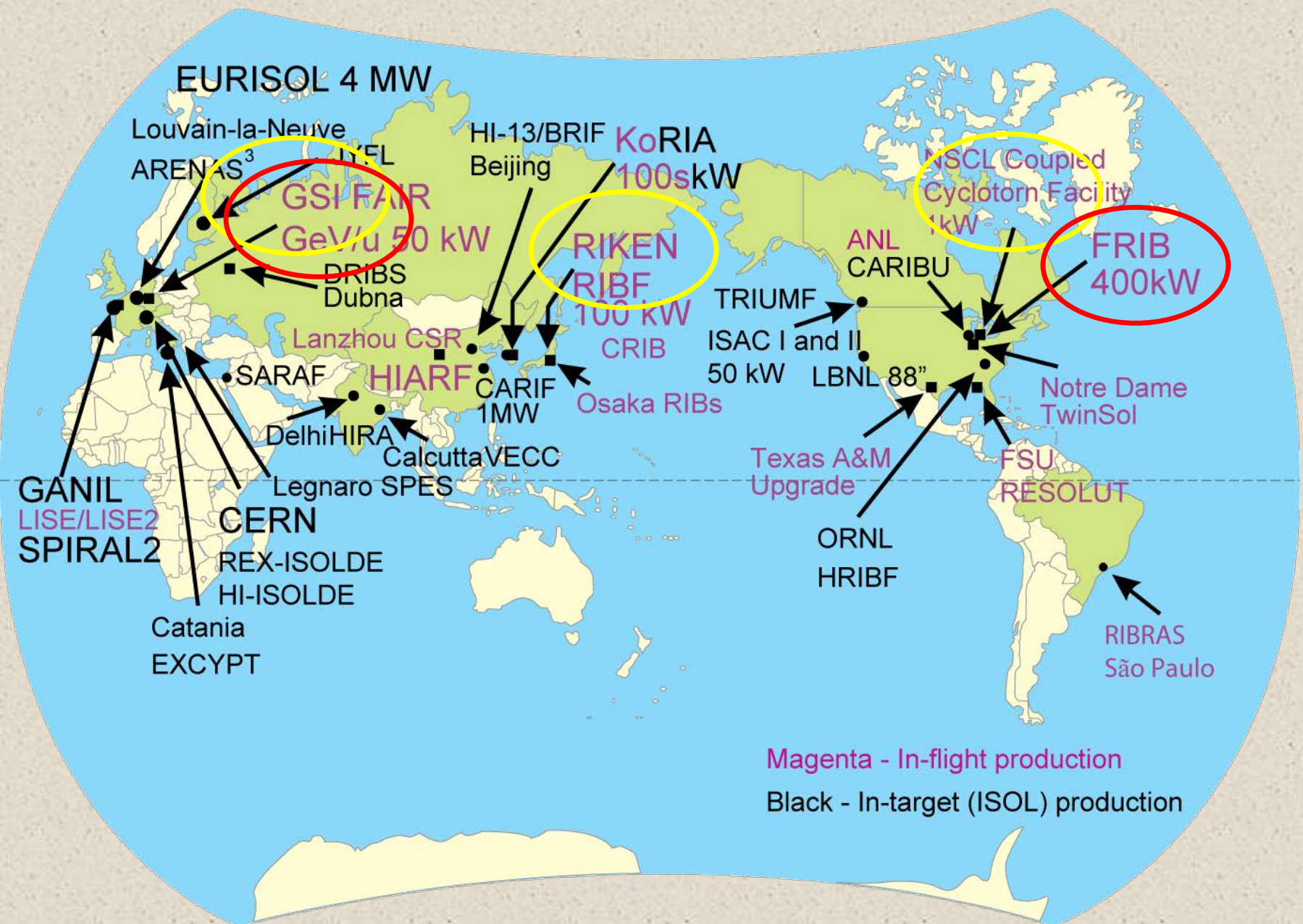


LISE-GANIL



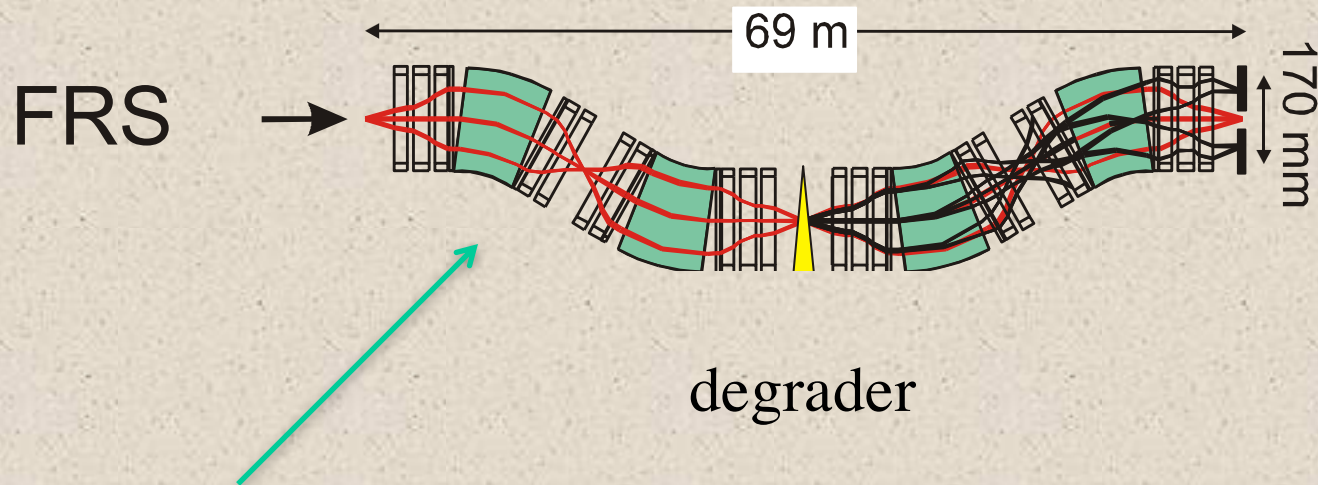
FRS GSI





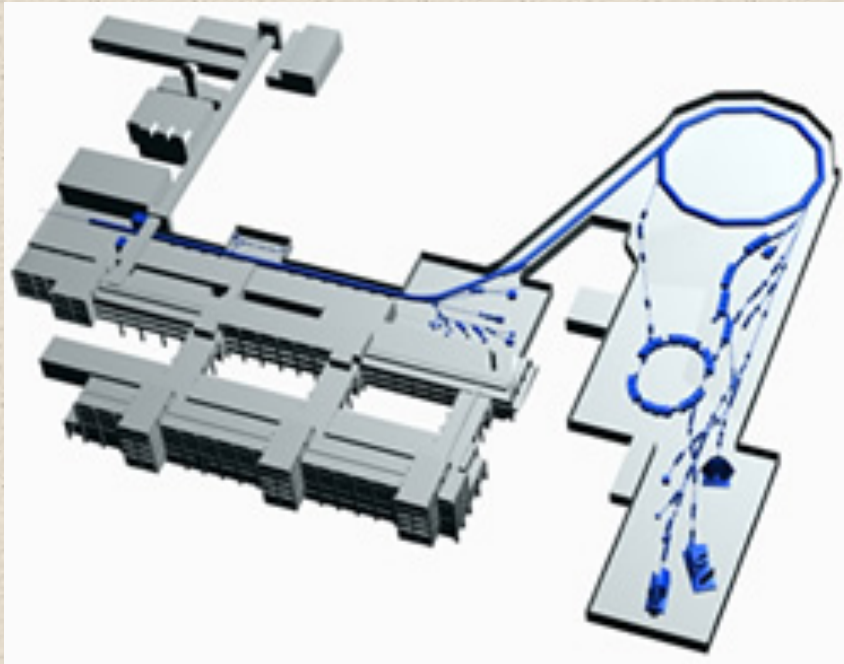
### $B\rho$ - $\Delta E$ - $B\rho$ Separation Method

Optimised to separate very fast fragment or fission products



Ions are selected according to the momentum to charge ratio  
 $Mv/q = Mv/Z = B\rho$ . Primary beam is rejected.

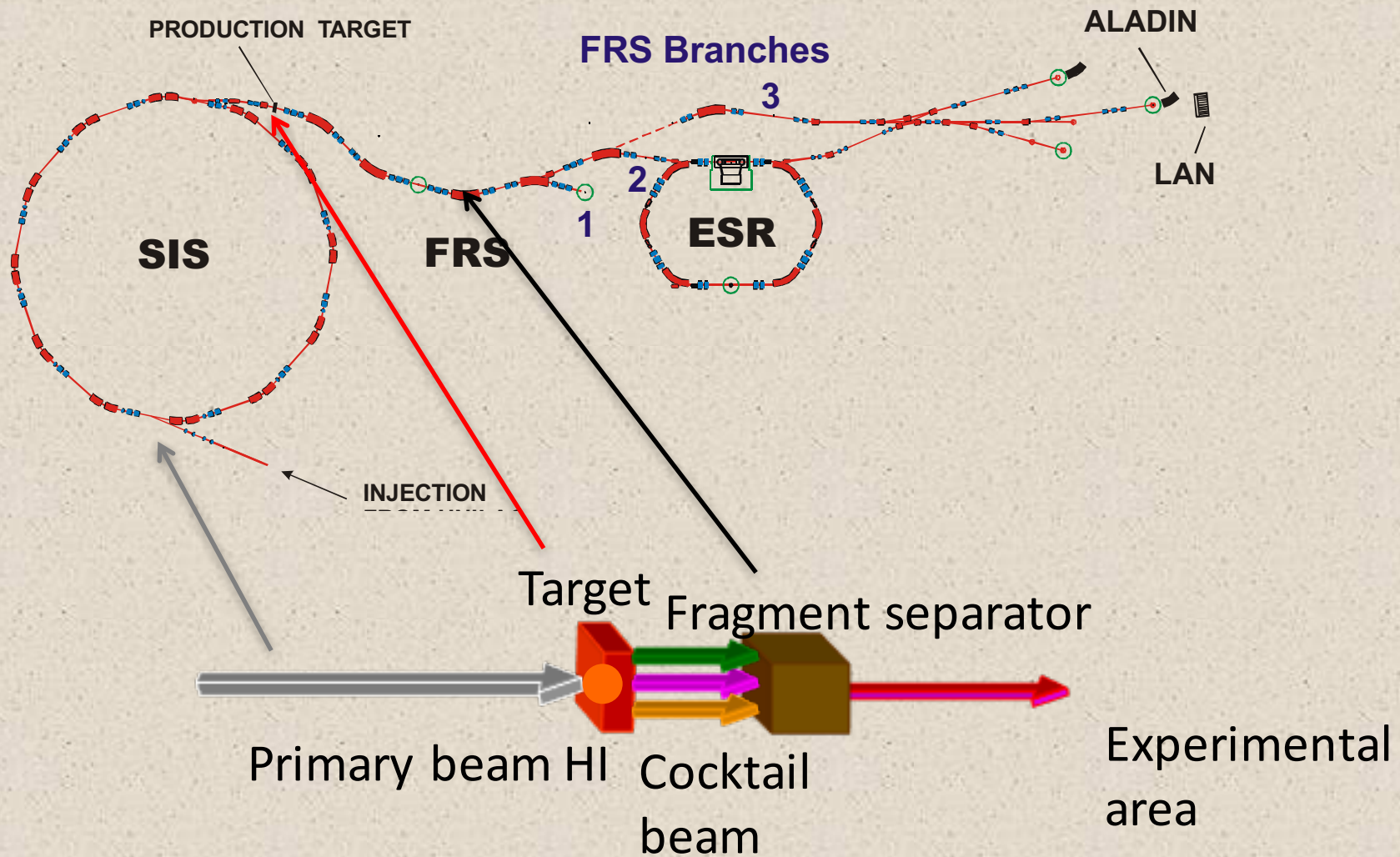
# The GSI fragmentation facility

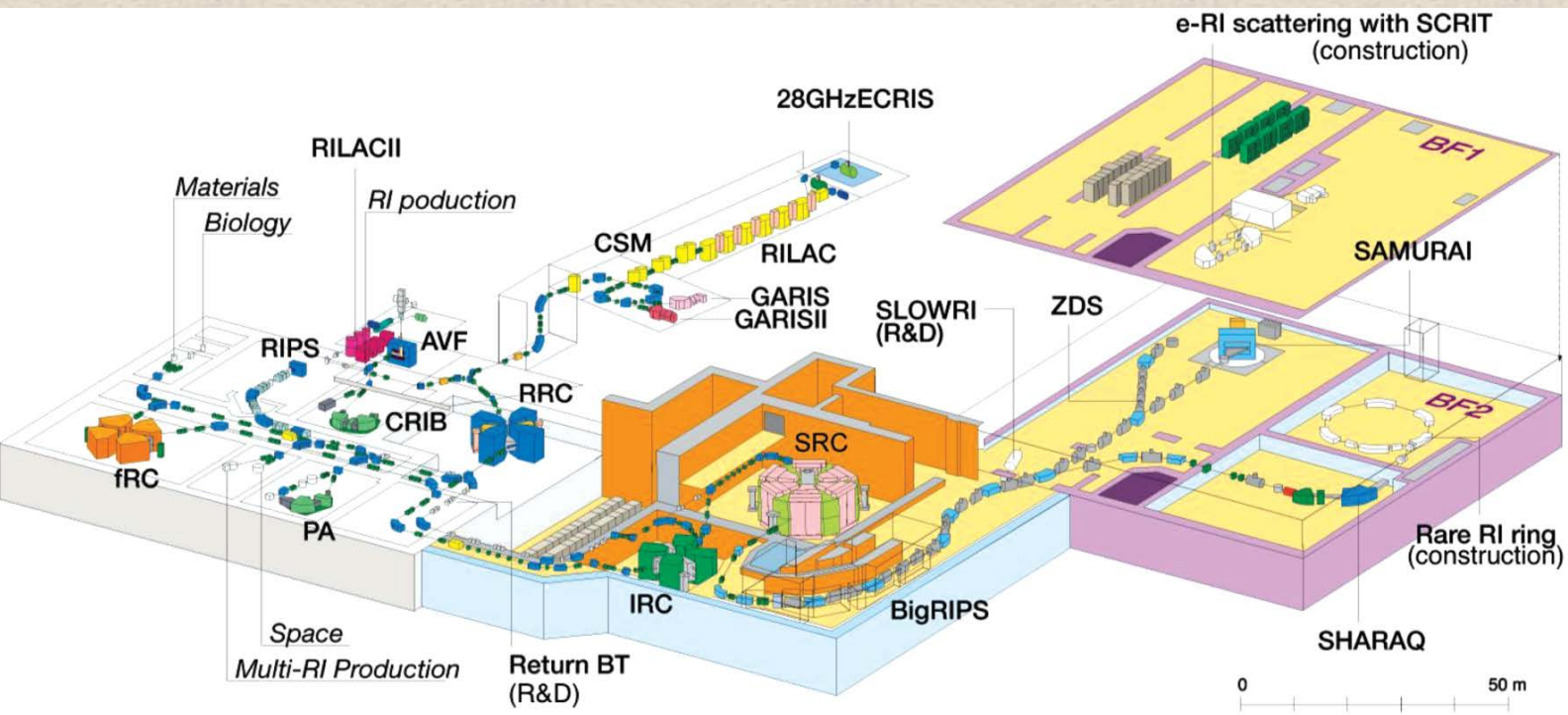


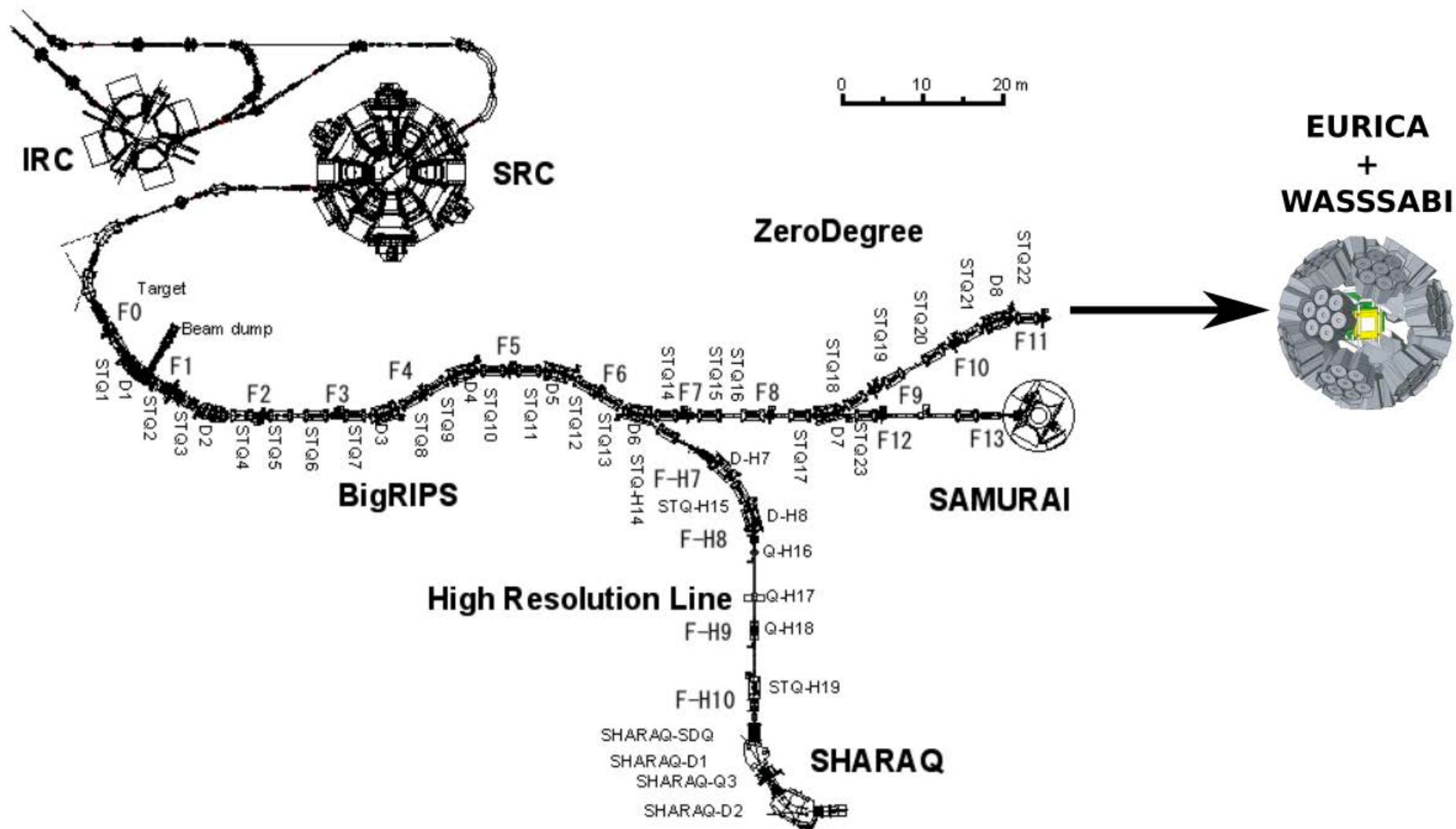
**It is possible to accelerate stable Nuclei from H ( $Z=1$ ) to U ( $Z=92$ ), in the UNILAC + Synchrotron.**

**Fragmentation of the primary ion beam:  
“cocktail” of ions**

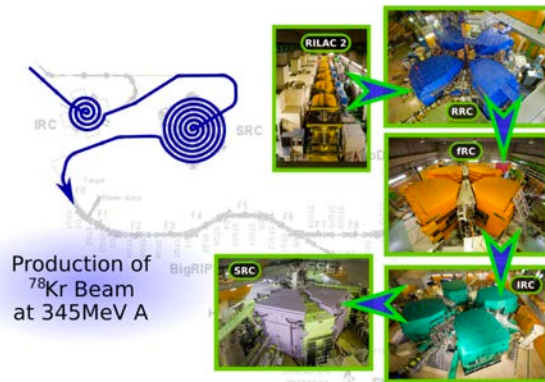




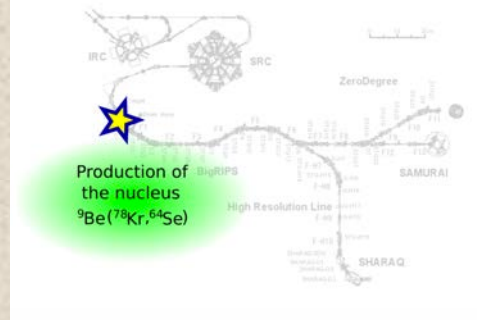




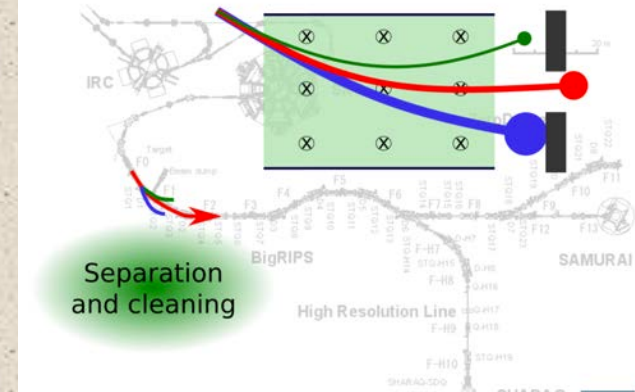
# Experimental Setup BigRIPs In-Flight Separator



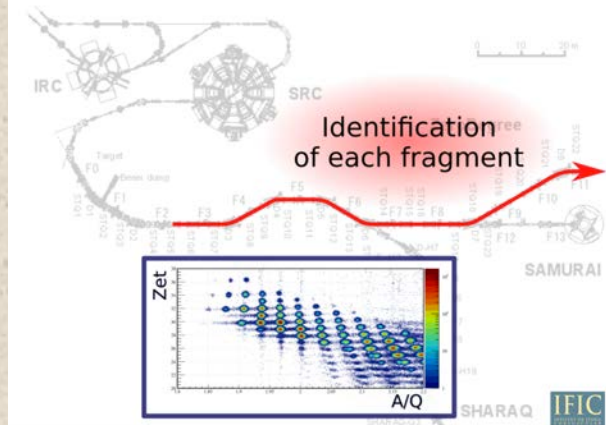
# Experimental Setup BigRIPs In-Flight Separator



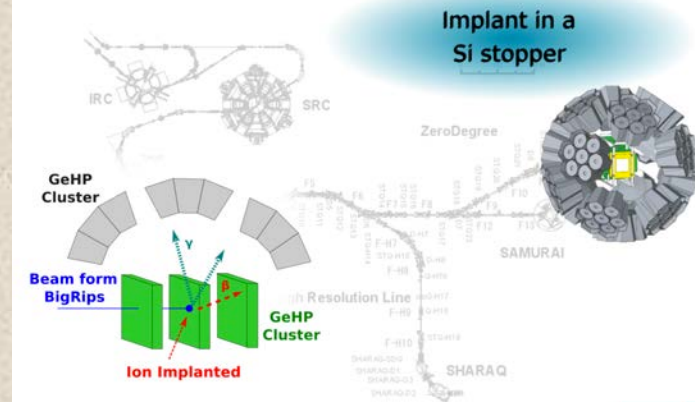
# Experimental Setup BigRIPs In-Flight Separator



# Experimental Setup BigRIPs In-Flight Separator



# Experimental Setup BigRIPs In-Flight Separator





# Estudios de núcleos exóticos en RIKEN (Japón)

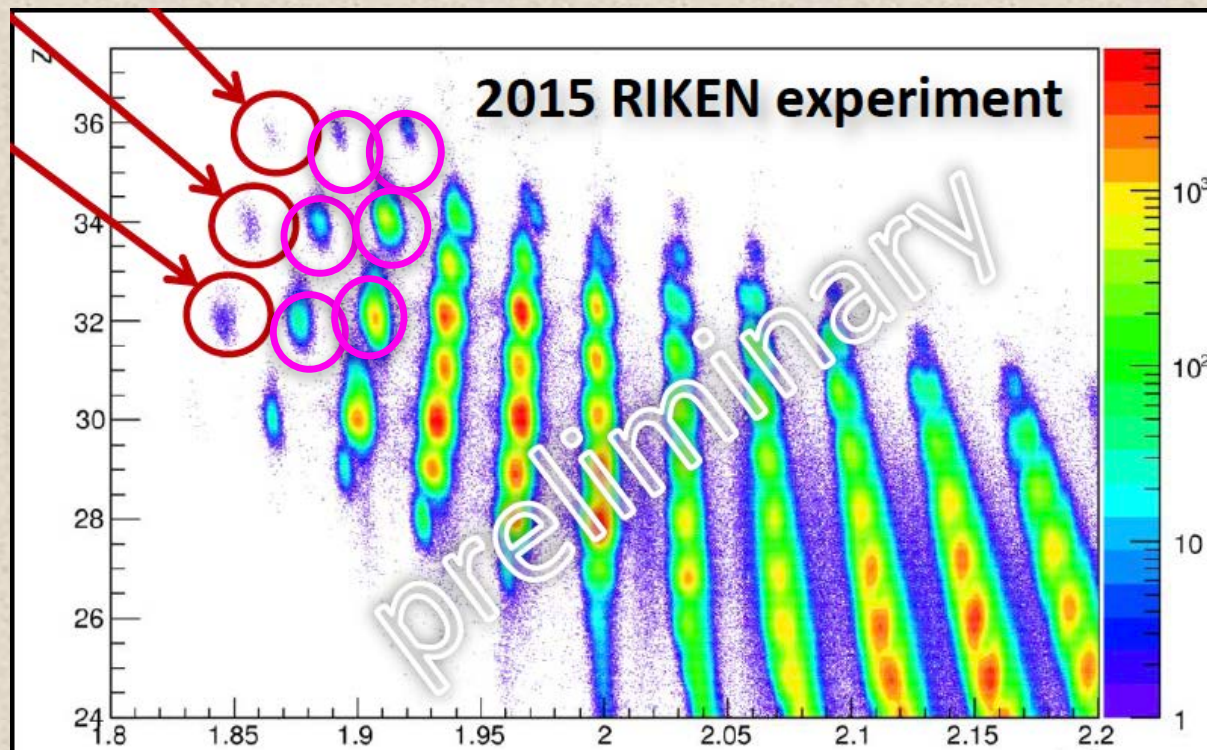
A. Algora et al.

B. Blank et al.

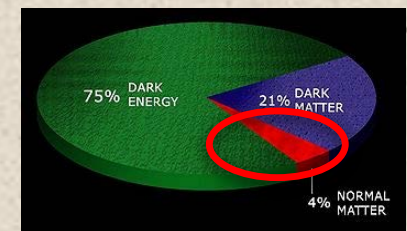
B. Rubio et al.

Descubrimiento de nuevos isótopos

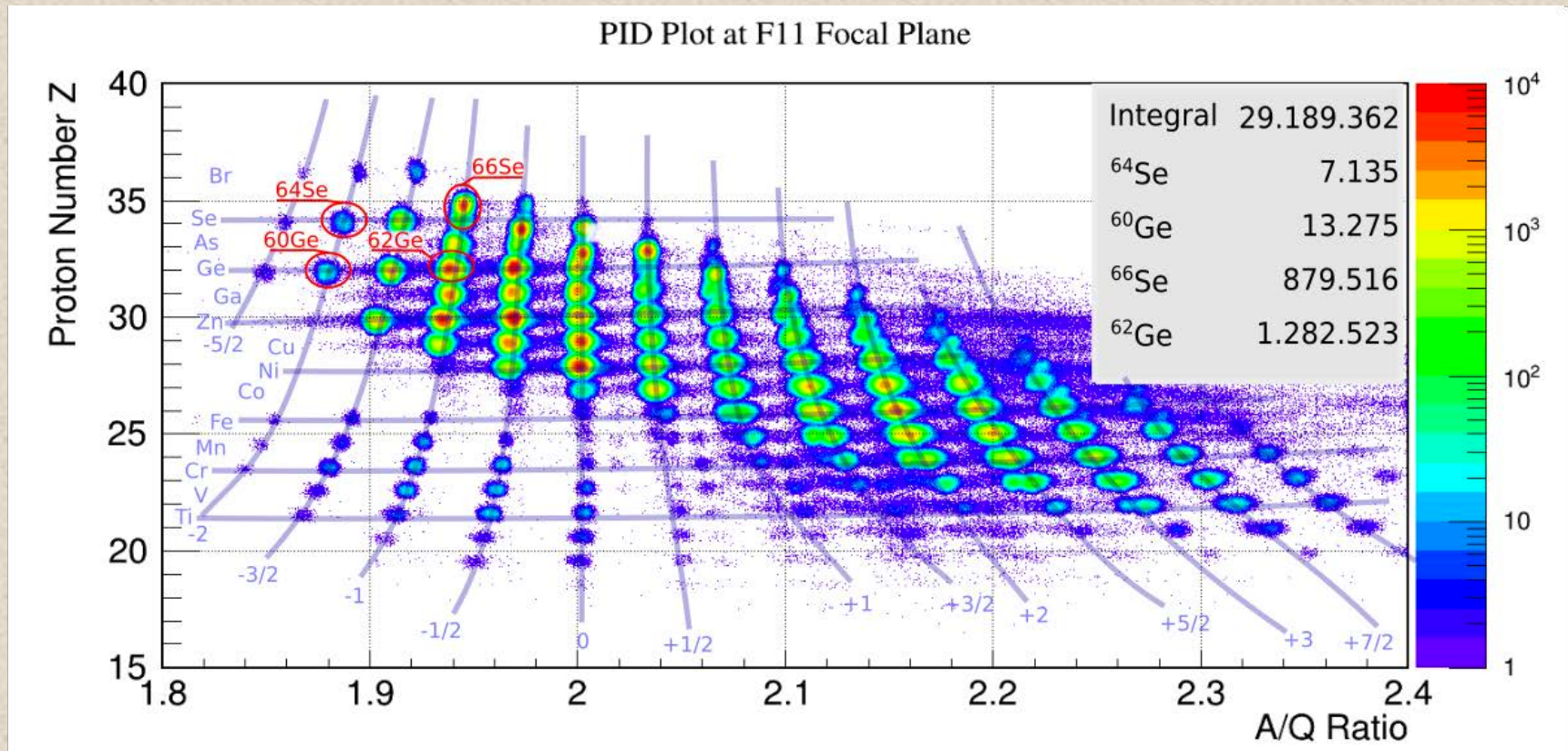
Estudio de los núcleos de Ge, Se, y Kr más exóticos



Alcanzado el límite de Existencia de los núcleos



At a fragmentation Facility, normally, many nuclei arrive  
**AT THE SAME TIME** at the experimental set up area  
 (implantation detector, secondary target....),.  
 But in an event by event data taking mode, they can be  
 separated (characterised) a posteriori.





B. Rubio. Master FN, Valencia 2018

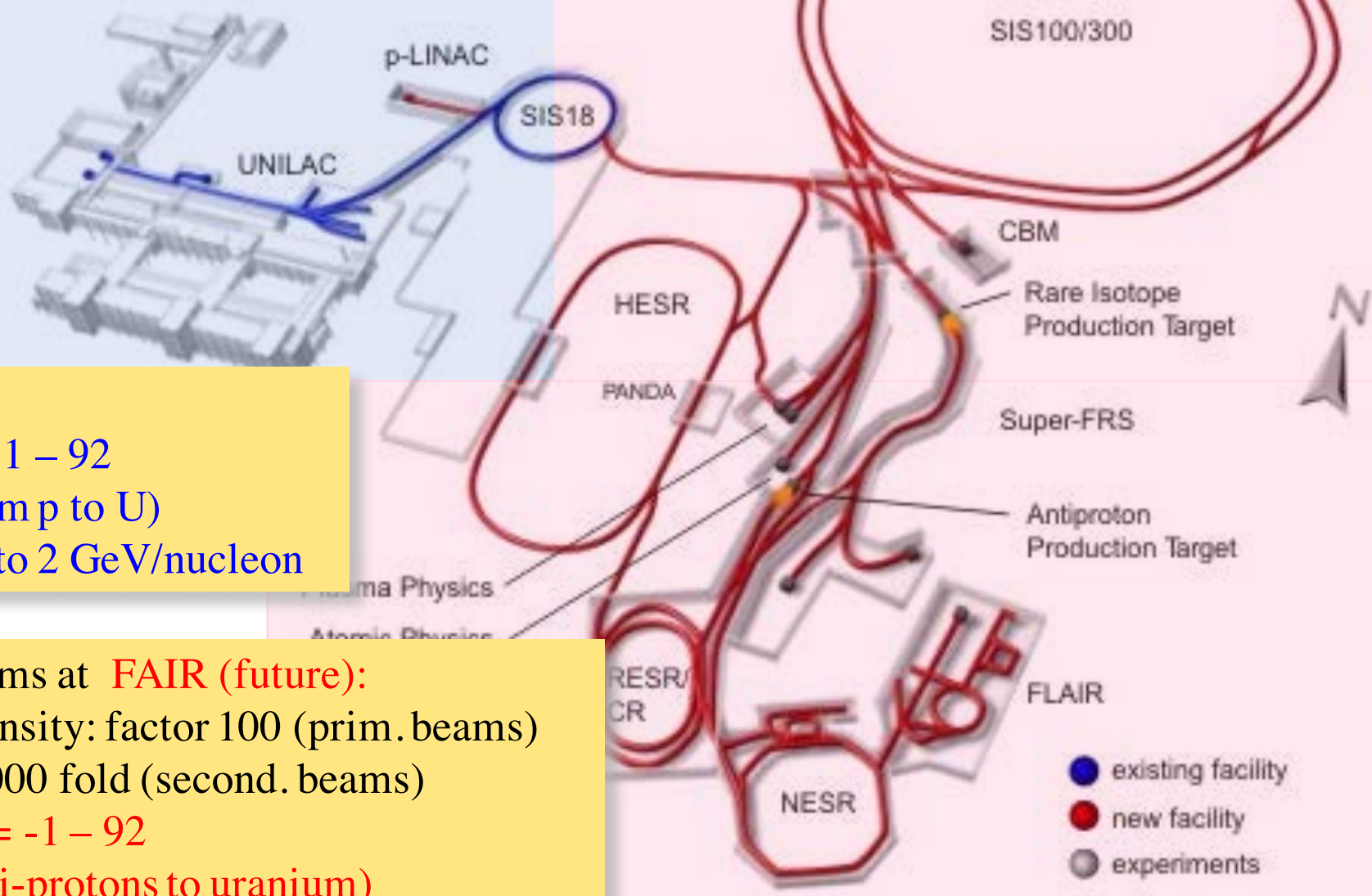
Future (in Europe)



B. Rubio. Master FN, Valencia 2018

## Existing GSI facility

## New facility



GSI

$Z = 1 - 92$

(from p to U)

Up to 2 GeV/nucleon

Beams at **FAIR (future):**

Intensity: factor 100 (prim. beams)

10 000 fold (second. beams)

$Z = -1 - 92$

(anti-protons to uranium)

Up to 35 - 45 GeV/u

„full beam cooling“

## FRS to Super-FRS

**B $\rho$ - $\Delta$ E-B $\rho$  Separation Method**

FRS

69 m

170 mm

Degradar

Super-FRS

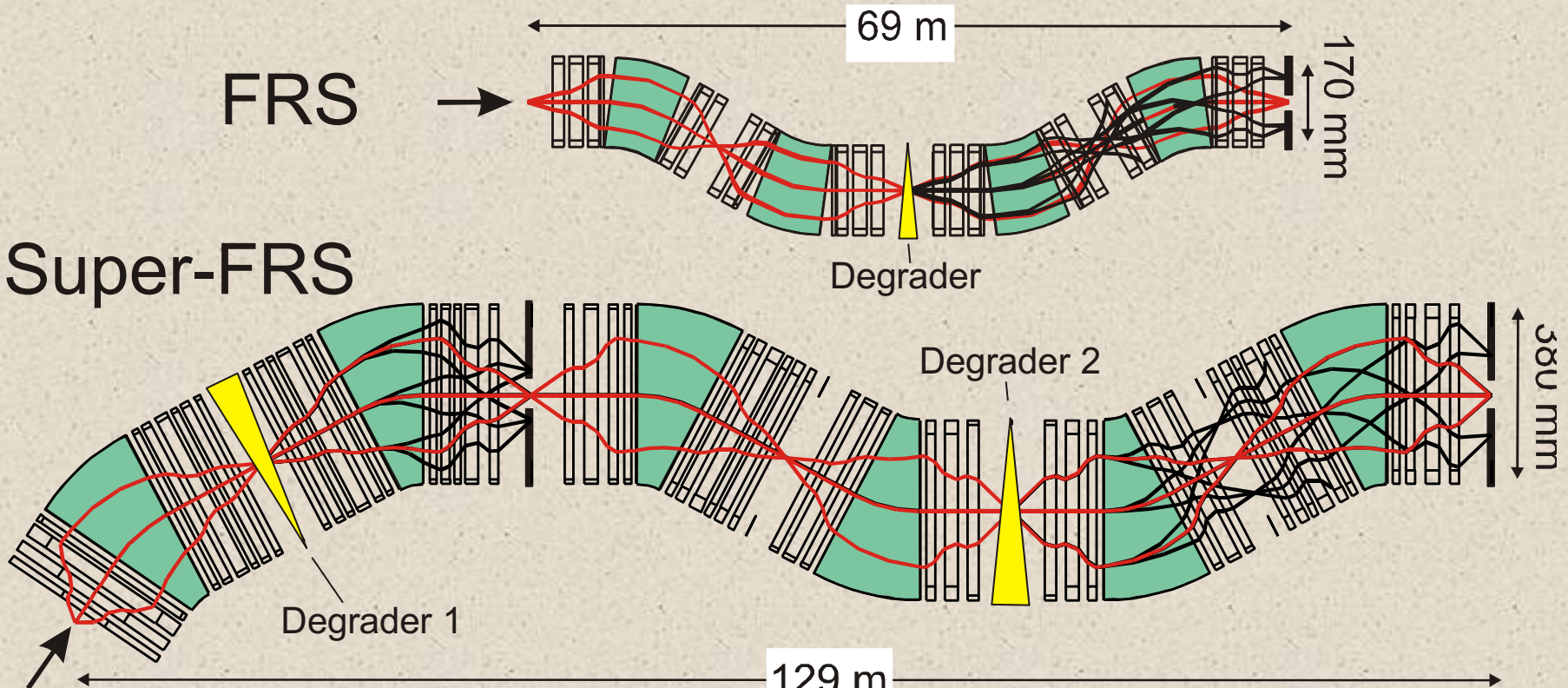
Degradar 2

380 mm

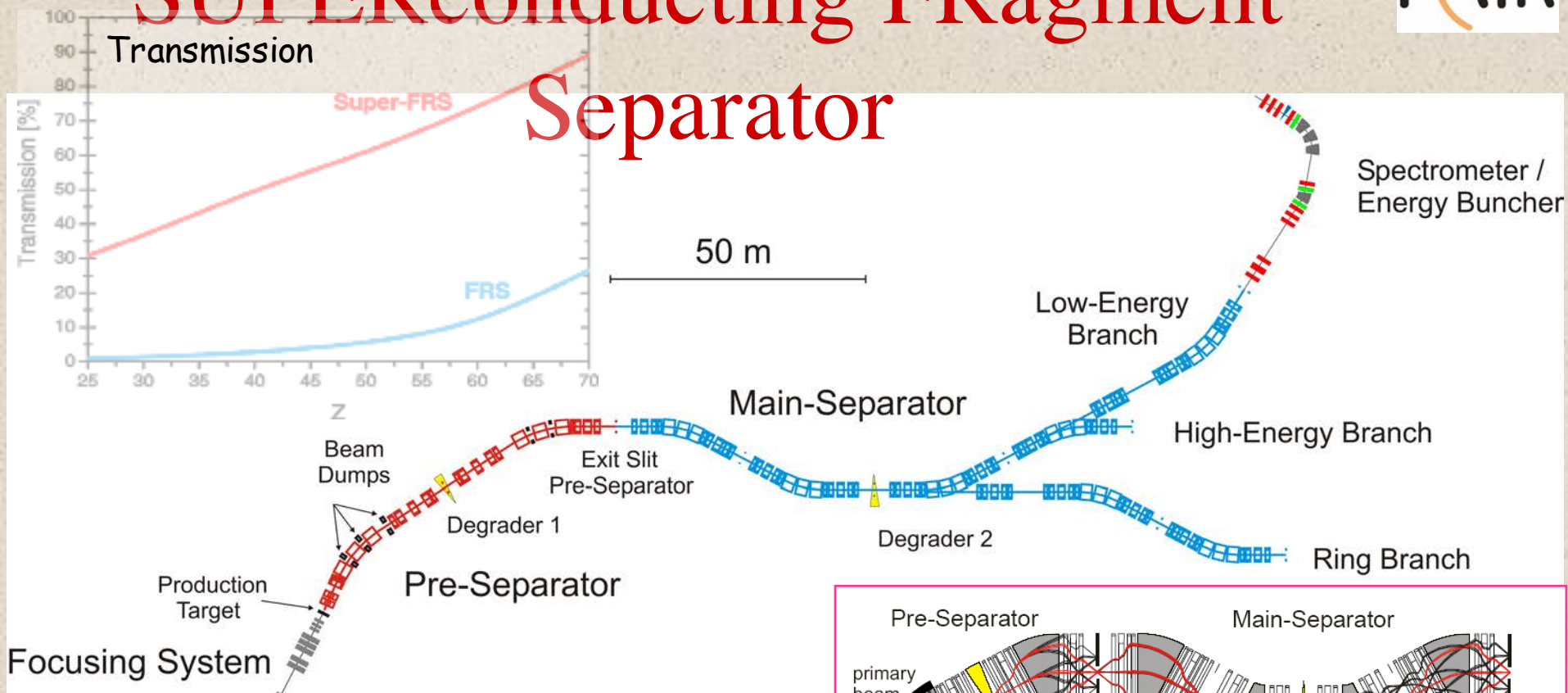
Degradar 1

129 m

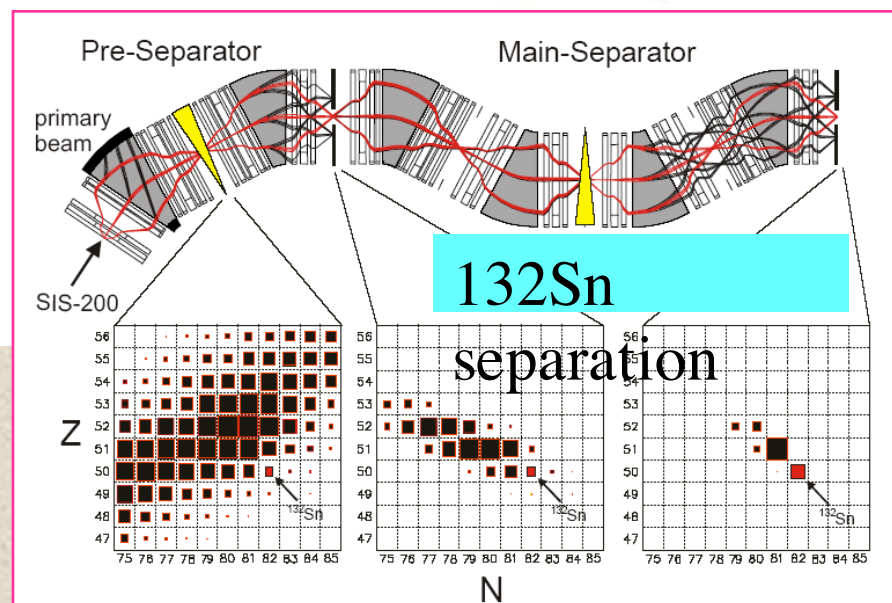
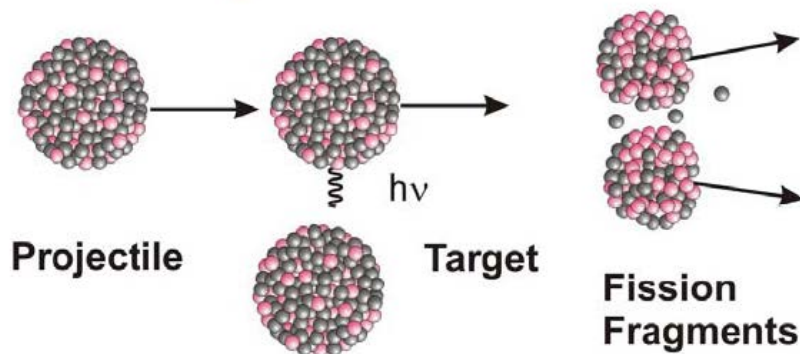
Optimised to separate  
very fast fragment or  
fission products



# SUPERconducting FRagment Separator



## Projectile Fission



# FAIR Project Progress – Civil Construction



SIS 18 in direction of pLinac :  
Preparation of western transfer  
channel (WTK)

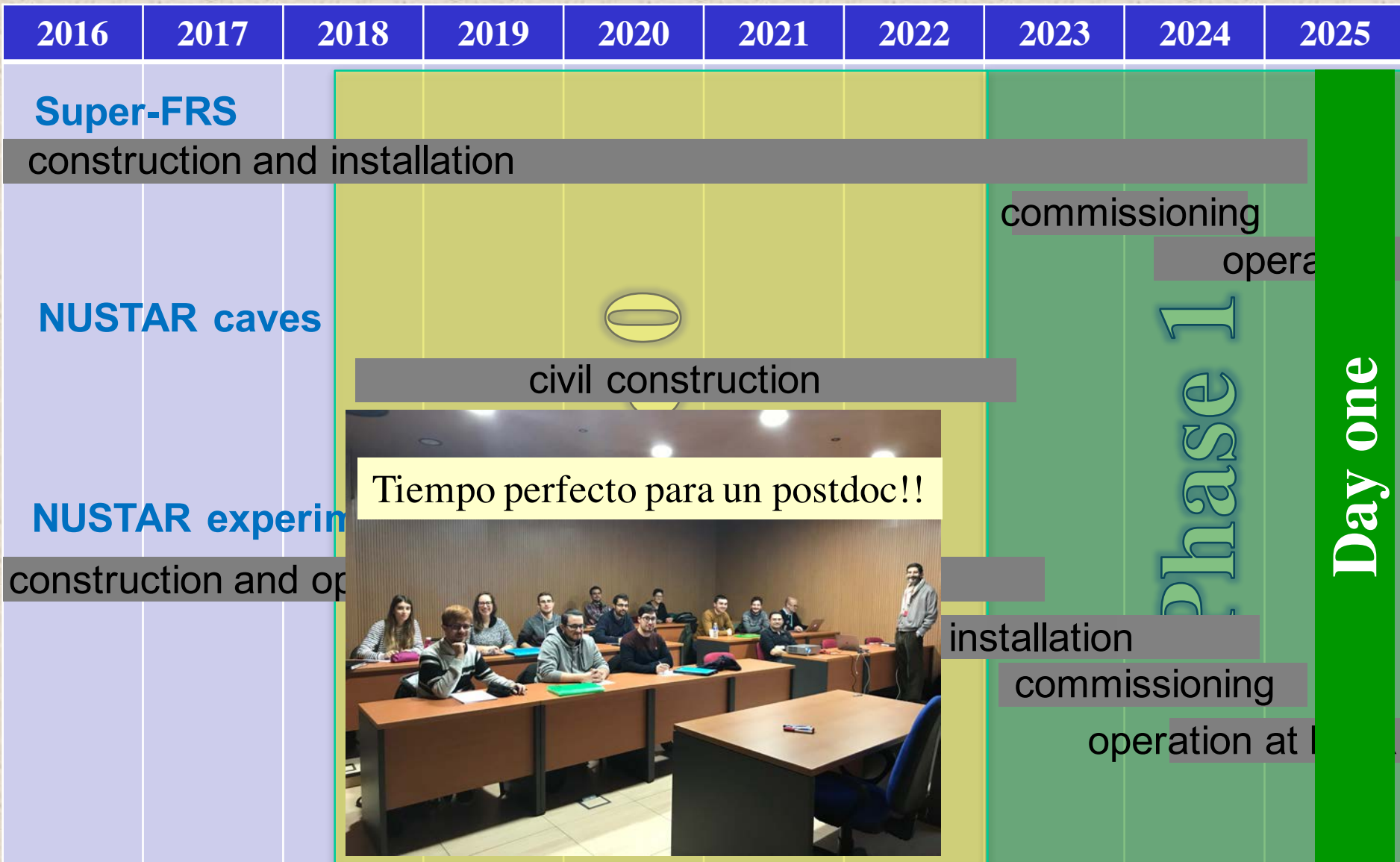
SIS 100:  
trench sheeting  
(first 8 meters  
sheet piling)



# FAIR Construction Field



# Scenario for NUSTAR: Day one



# ISOL and In-Flight facilities-Partners

It is probably true to say that if we worked at it virtually all experiments could be done with both types of facility but they are complementary.

## In-Flight

- Relativistic beams
- Universal in Z
- Down to very short  $T_{1/2}$
- Easily injected into storage rings
- More exotic beams

## ISOL

- High intensity beams with ion optics comparable to stable beams
- Easy to manipulate beam energies from keV to 10s of MeV

• Cleaner