

Mass Measurements with Ion Traps

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Outline

1. Confinement of charged particles
2. Mass measurements
 - Importance of the mass
 - Physics goals & mass uncertainty
 - Existing facilities
3. Penning-Trap Techniques
4. The TRAPSENSOR laboratory in Granada
 - Physics motivation
 - The QS concept
 - The QS project

CONFINEMENT OF CHARGED PARTICLES

Introduction

Confinement of charged particles

- Required:
3-dimensional potential minimum.
Force acting towards a center.

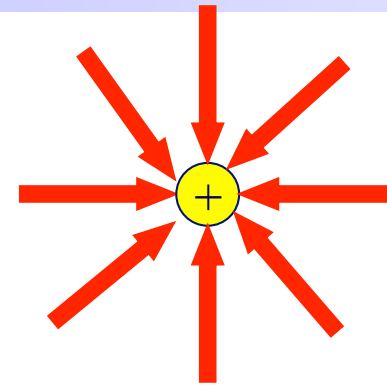
- Convenience:
a) $F \sim -r$ (Harmonic oscillator)
b) Rotational symmetry

- Consequences:
From a) $\vec{F} = -e\nabla\Phi \propto -\vec{r}$
 $\Rightarrow \Phi = ax^2 + by^2 + cz^2$

Laplace equation

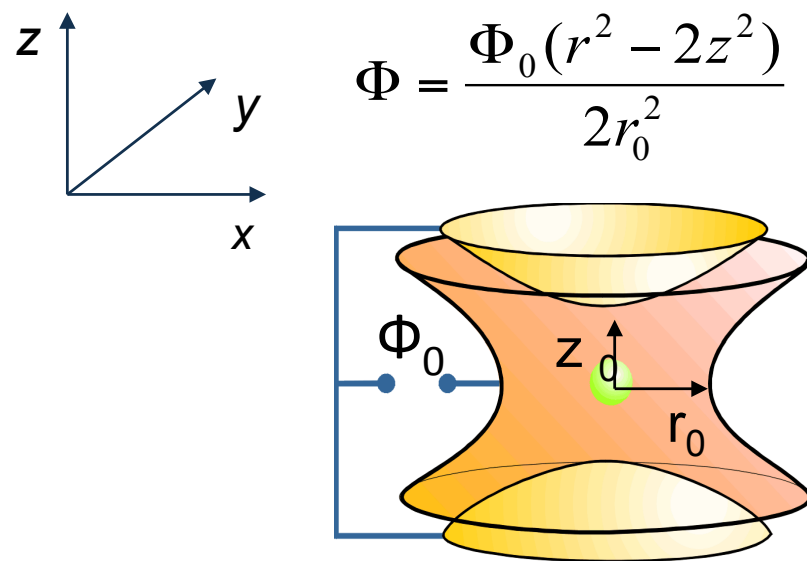
Rotational symmetry around z-axis

- THERE IS NOT 3-DIMENSIONAL CONFINEMENT WITH ELECTROSTATIC FIELDS



Introduction

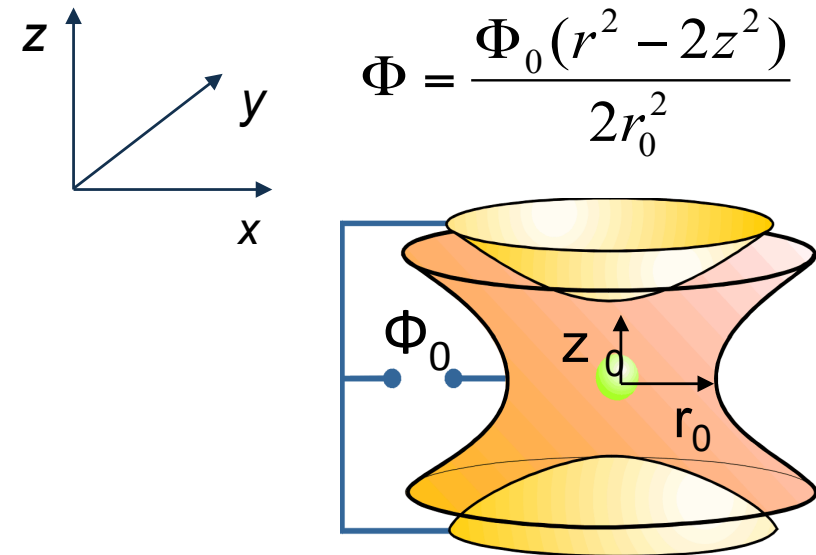
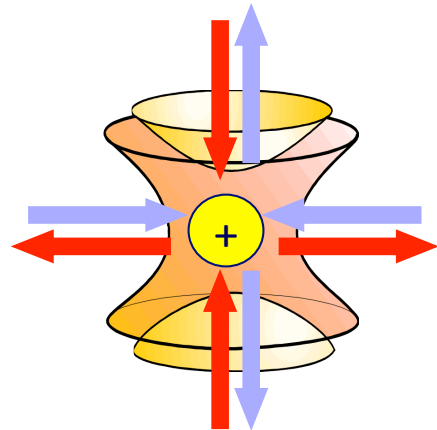
Confinement of charged particles



$\Phi = \Phi(t)$

↓

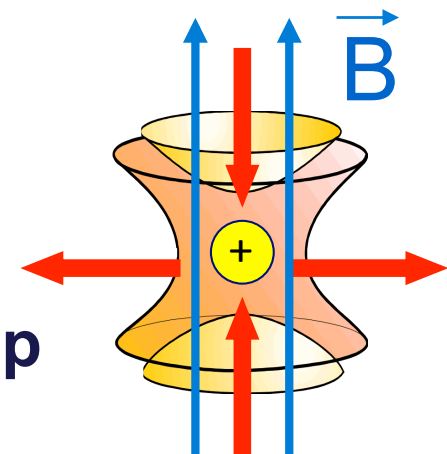
Paul trap



$\vec{E}(\Phi) + \vec{B}$

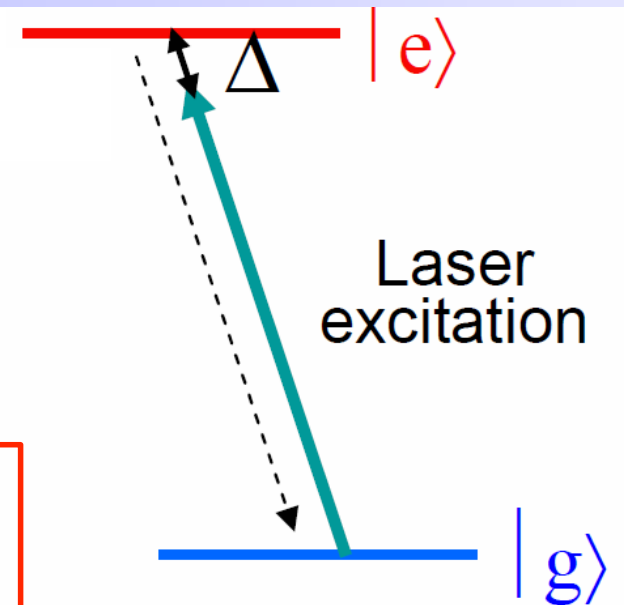
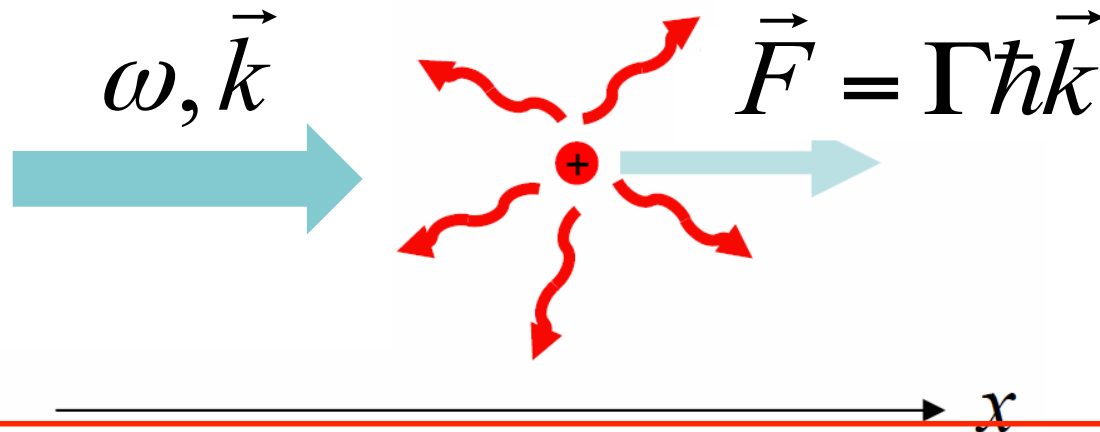
↓

Penning trap

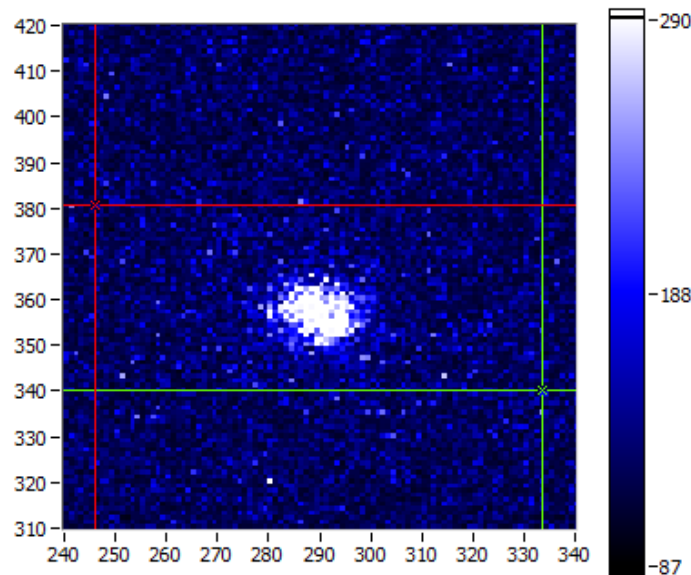


Cooling in Penning traps

Doppler cooling

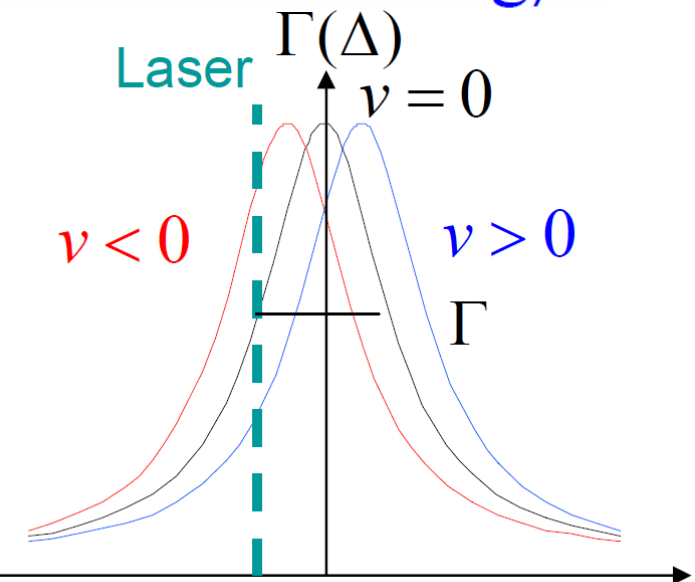


- Cooling until the Doppler limit is reached



$$T_{\text{limit}} = \frac{\hbar \Gamma}{2k_B}$$

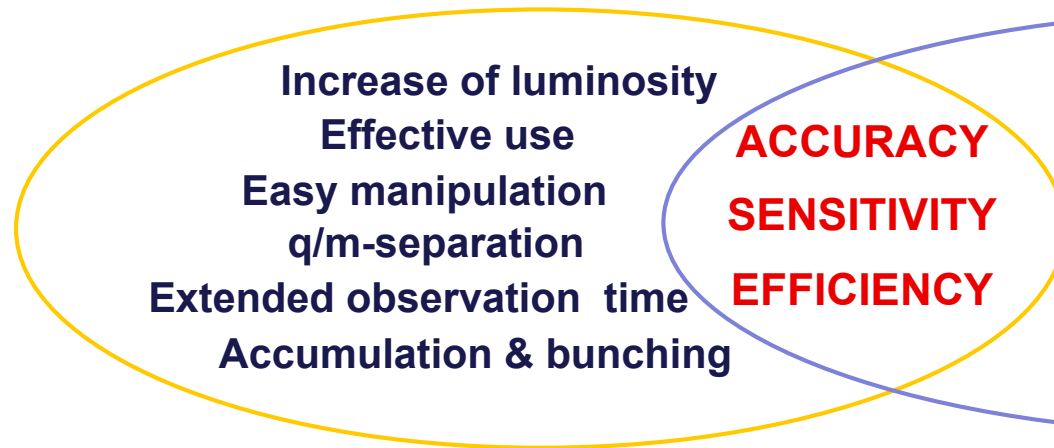
$$T_{\text{limit}} \approx 1 \text{ mK}$$



Introduction

Confinement of charged particles

Trapping



Cooling

Cooling (Blue Oval):

- Increase of luminosity
- Small amplitudes
- q/m-separation
- Emittance improvement
- Doppler width reduction
- Doppler effect elimination

Examples

- Experiments with rare species (radioactive isotopes, HCl, antiprotons, anti-H,....)
- Ultra-high accuracy (mass, g -factor, α , optical clocks, relativity)
- Single-particle detection and observation (quantum logic spectroscopy)
- Quantum simulations and quantum computing
- Fundamental interactions (EDM, PV, search for non V-A interaction,.....)
- Accelerator device (accumulator, charge breeder, isomer separation, buncher,....)

Penning traps

Different geometries (heavy-ion masses)



CERN



MSU



GSI-Darmstadt



TRIGA
Mainz



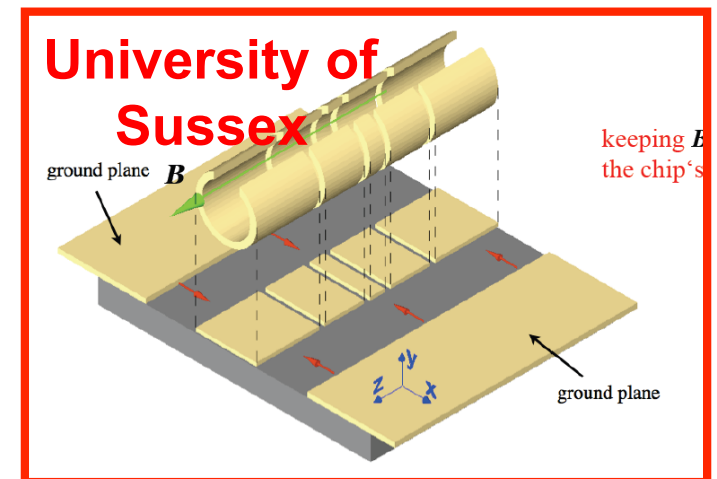
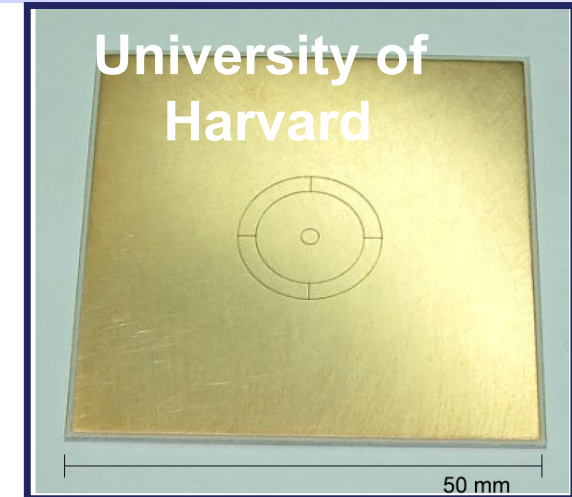
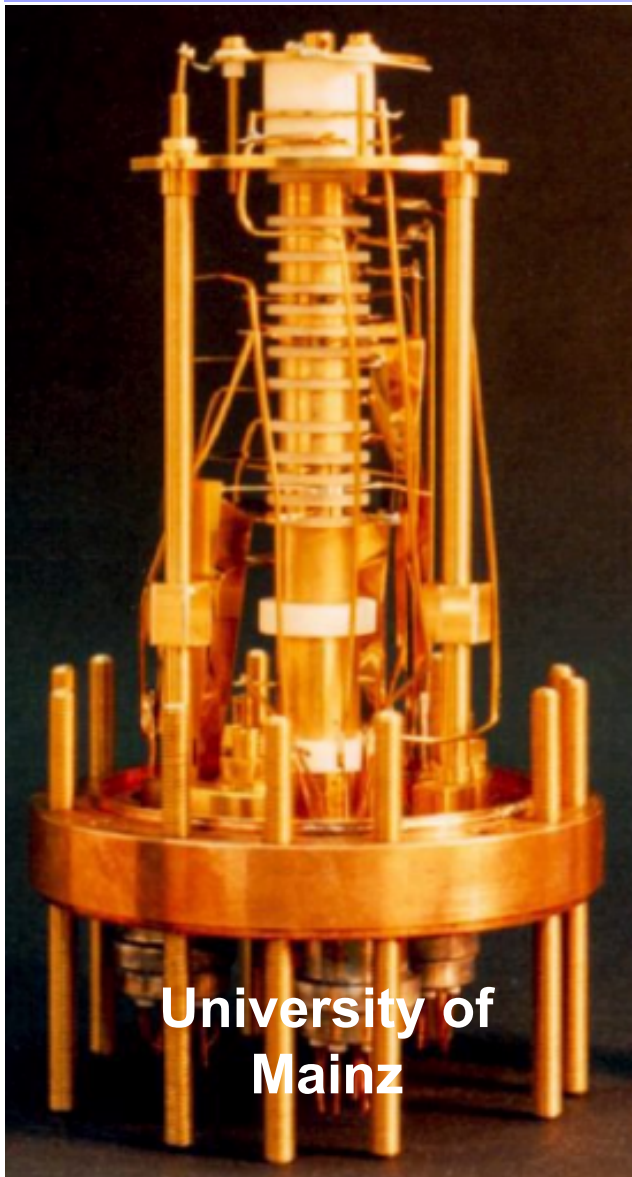
MPIK



MPIK

Penning traps

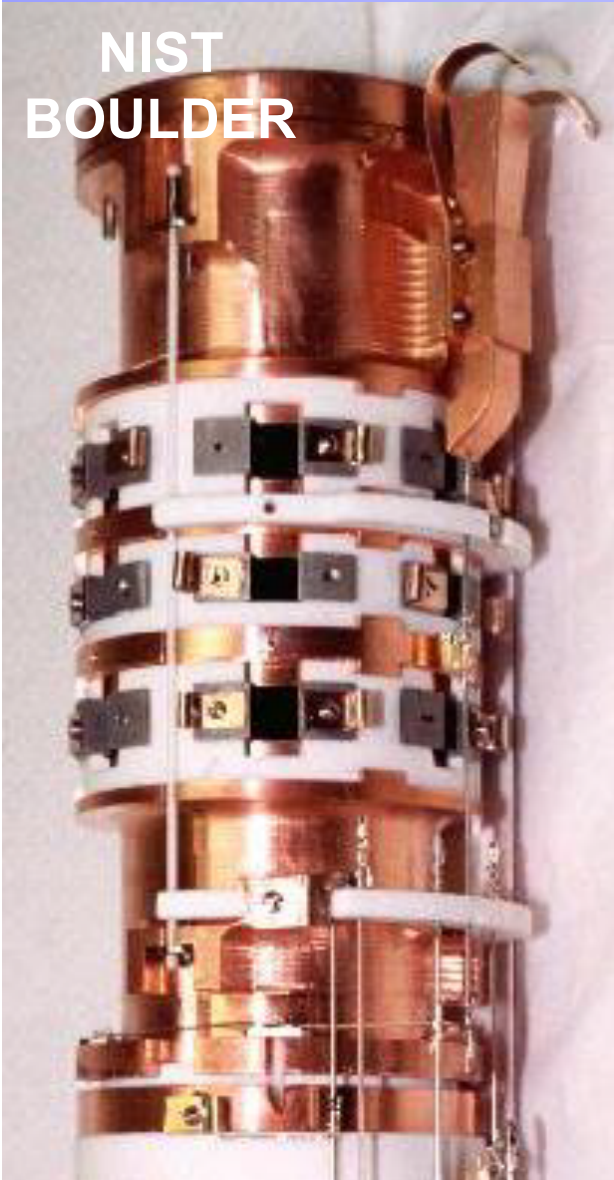
Different geometries (HCl, electrons)



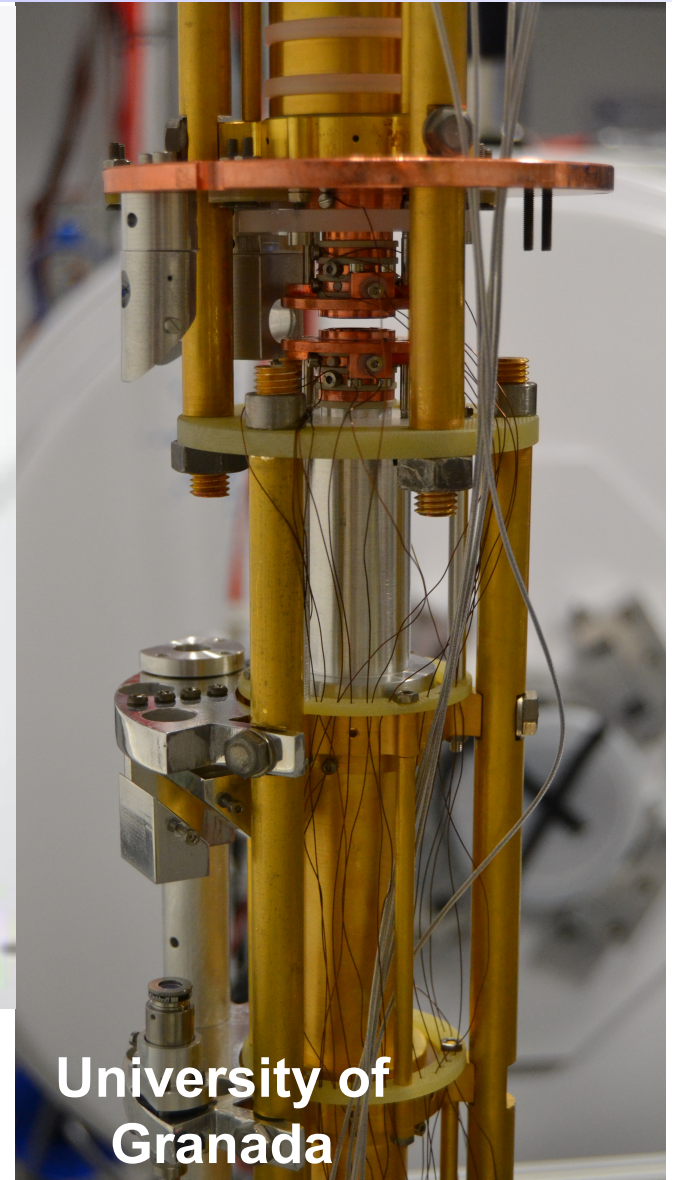
Penning traps

Different geometries (fluorescence collection)

NIST
BOULDER



Imperial
College
(& GSI-Darmstadt)

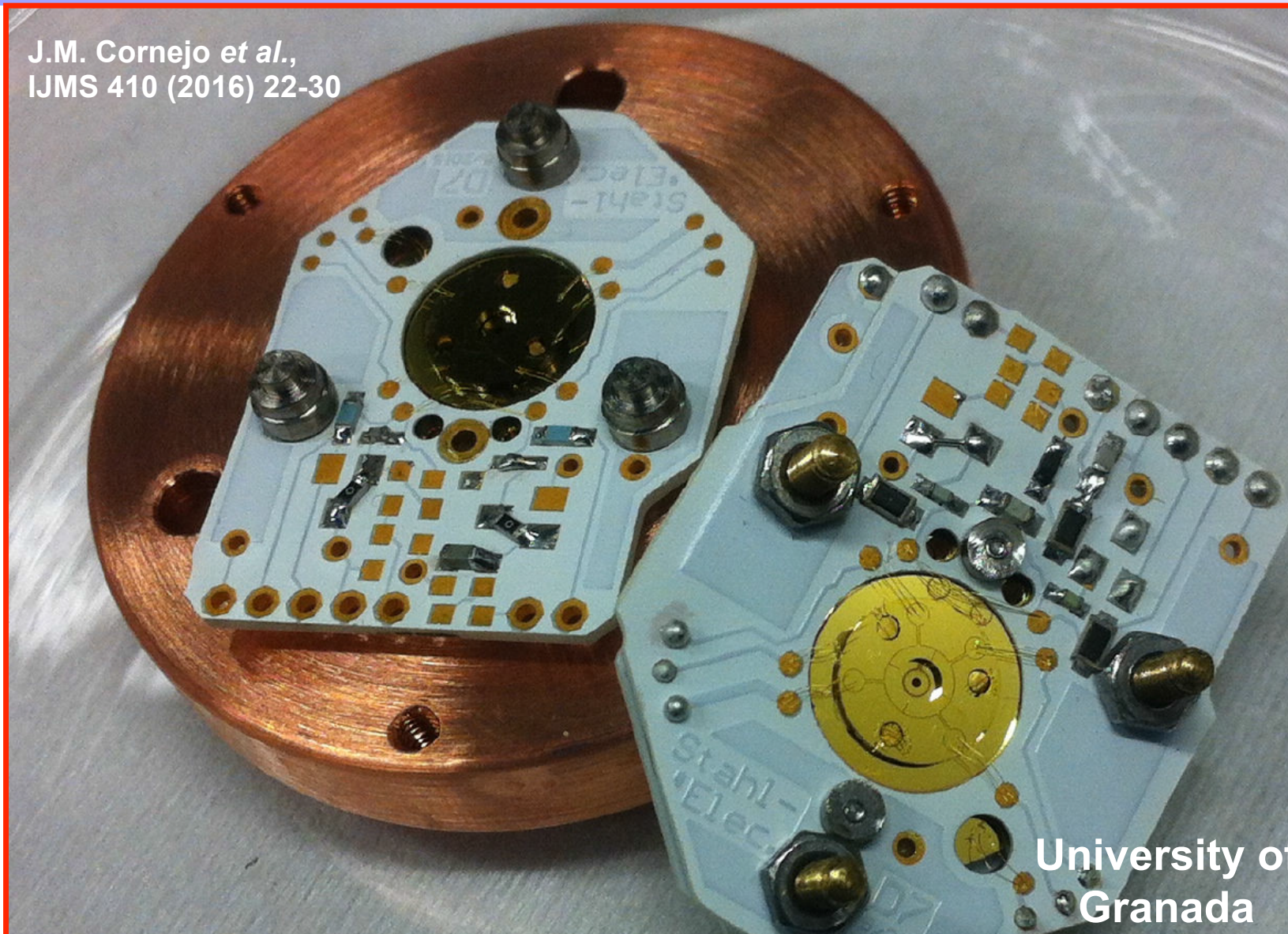


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Granada

Penning traps

Different geometries (miniaturization)

J.M. Cornejo *et al.*,
IJMS 410 (2016) 22-30



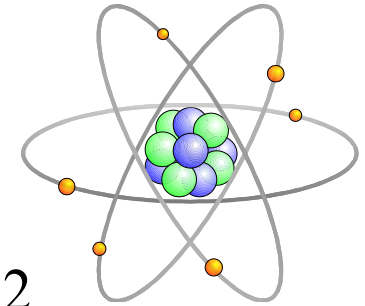
University of
Granada

MASS MEASUREMENTS

Importance of the mass

The nuclear binding energy (B)

The nuclear binding energy reflects all forces acting in a nucleus



$$B(N, Z) = \left\{ Zm_p + Nm_n - \underbrace{M(N, Z)}_{\text{Nuclear mass}} \right\} \cdot c^2$$

$$B(N, Z) = \left\{ Zm_p + Nm_n - \left[\underbrace{m(^A X)}_{\text{Atomic mass}} - Zm_e - \sum_i B_i^e \right] \right\} \cdot c^2$$

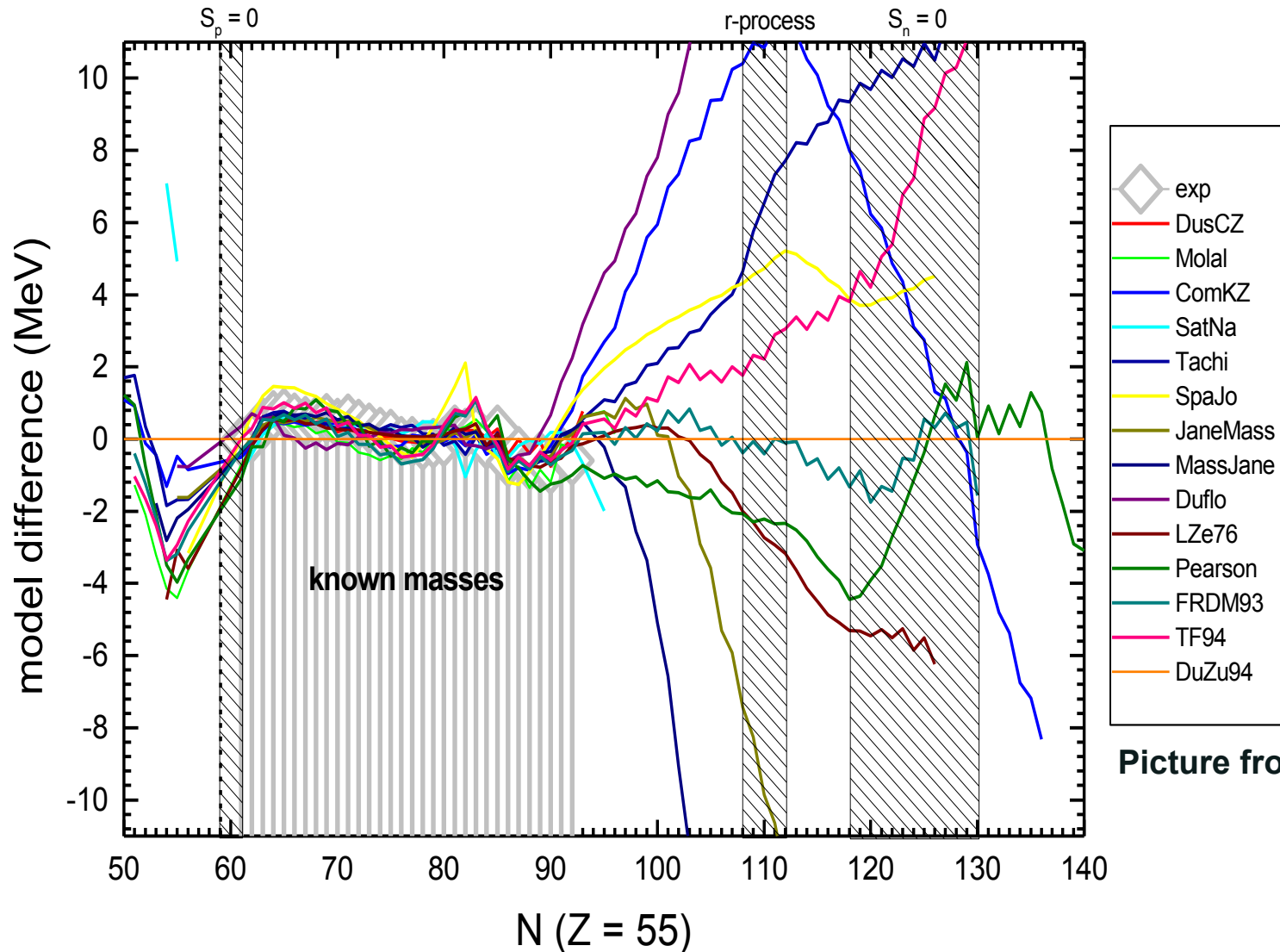
$$ME(N, Z) = m(^A X) - A \cdot u$$

Mass excess (keV)

$$1 \text{ keV} \rightarrow \frac{\delta m}{m} \approx 10^{-8}$$

Importance of the mass

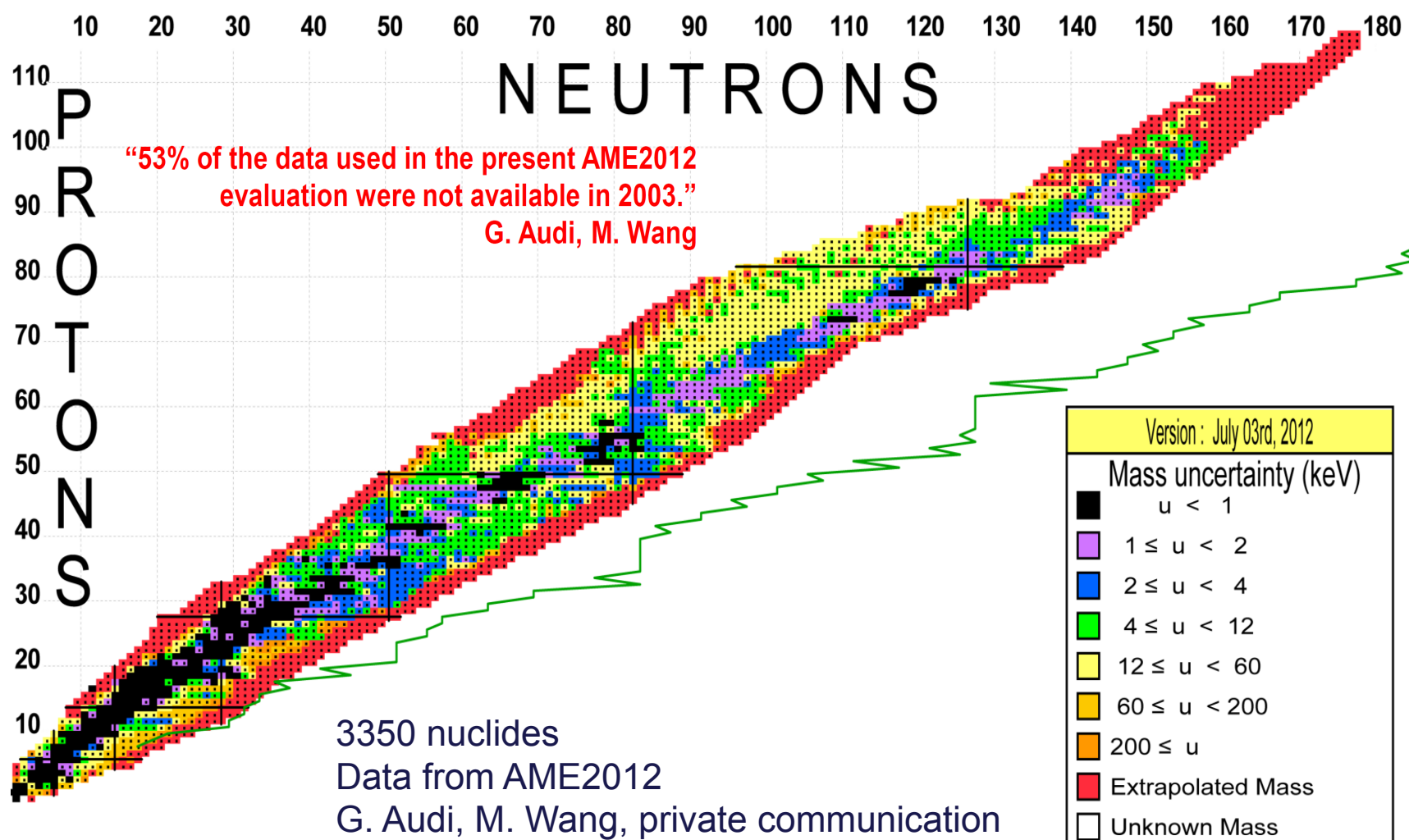
Discrepancies between the models



Picture from D. Lunney

Mass measurements in the last decade

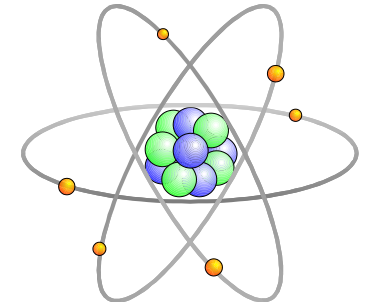
The 2012 Atomic Mass Evaluation



Importance of the mass

Nucleon separation energies (S_n , S_p , S_{2n} , S_{2p})

Nuclear Structure & astrophysics



$$S_{2n} = B(N, Z) - B(N - 2, Z)$$

$$S_{2p} = B(N, Z) - B(N, Z - 2)$$

$$S_n = B(N, Z) - B(N - 1, Z)$$

$$S_p = B(N, Z) - B(N, Z - 1)$$

Other quantities are also important for nuclear structure,
all of them obtained from nuclear binding energies

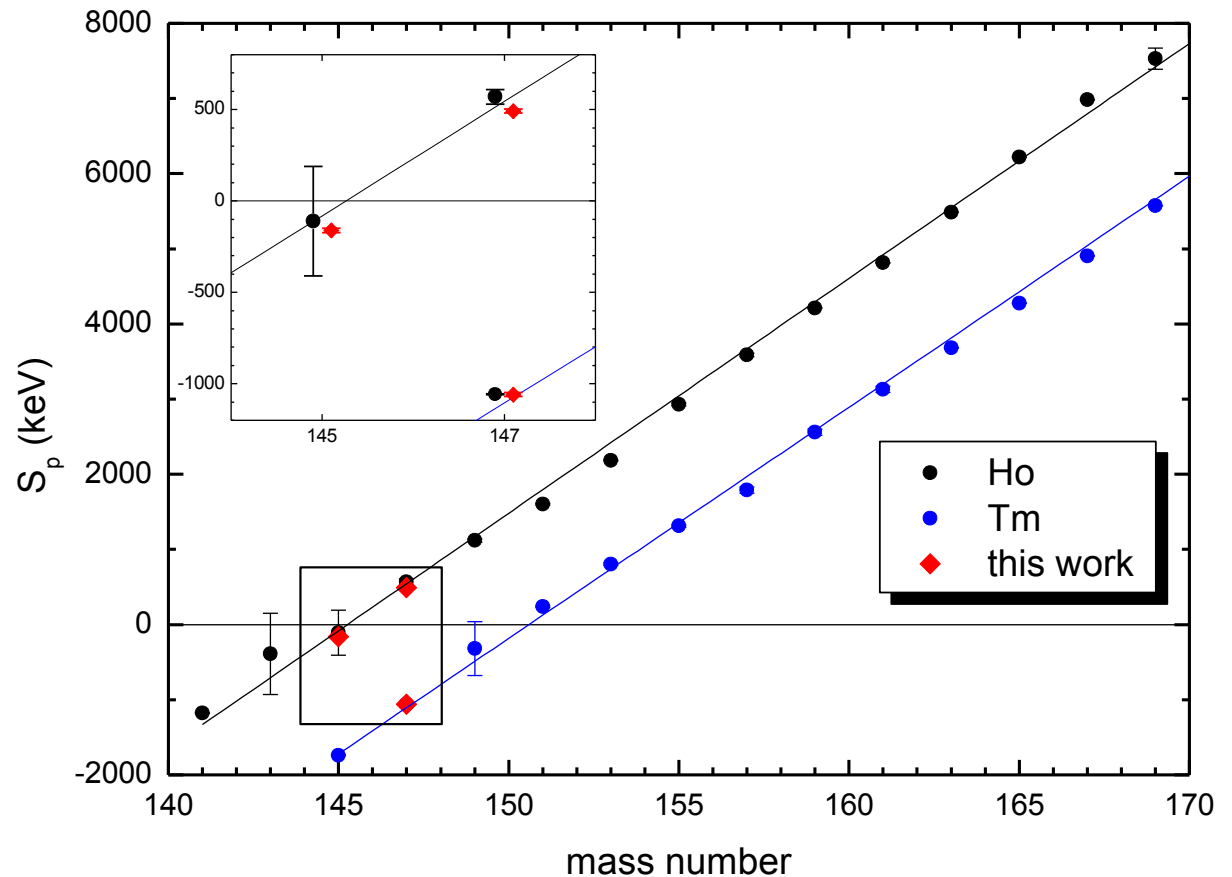
Importance of the mass

Nucleon separation energies (S_n , S_p , S_{2n} , S_{2p})

- Location of the proton-drip line

146 69 Tm 77	147 69 Tm 78	148 69 Tm 79	149 69 Tm 80	150 69 Tm 81	151 69 Tm 82
72 ms (10 ⁺) E _{ex} 71 (6) p ⁺ =2 p ⁺ =16%	360 μs 30 ⁺ E _{ex} 63 (5) p ⁺ =100% p ⁺ =100%	900 ns 110 ⁺ M ⁺ 332704 (4004) p ⁺ =100% p ⁺ =100%	700 ms (10 ⁺) M ⁺ 332704 (4004) p ⁺ =100% p ⁺ =100%	900 ms (11/2 ⁺) M ⁺ 440404 (3004) p ⁺ =100% p ⁺ =100%	52 ms (10 ⁺) E _{ex} 41 (4) p ⁺ =100% p ⁺ =100%
145 68 Er 77	146 68 Er 78	147 68 Er 79	148 68 Er 80	149 68 Er 81	150 68 Er 82
900 ms 1/2 ⁺ M ⁺ 396904 (4004) p ⁺ =100% p ⁺ =100%	1.7 s 0 ⁺ M ⁺ 447104 (3004) p ⁺ =100% p ⁺ =100%	25 s (11/2 ⁺) E _{ex} 104 (6) p ⁺ =100% p ⁺ =100%	4.6 s 0 ⁺ M ⁺ 516504 (2004) p ⁺ =100% p ⁺ =100%	8.9 s (11/2 ⁺) E _{ex} 741.8 (2) p ⁺ =100% p ⁺ =100%	4 s (10 ⁺) M ⁺ 53742 (20) p ⁺ =100% p ⁺ =100%
144 67 Ho 77	145 67 Ho 78	146 67 Ho 79	147 67 Ho 80	148 67 Ho 81	149 67 Ho 82
700 ms M ⁺ 452004 (3004) p ⁺ =100% p ⁺ =100%	100 ms 50 ⁺ E _{ex} 104 (6) p ⁺ =100% p ⁺ =100%	24 s (11/2 ⁺) M ⁺ 515704 (2004) p ⁺ =100% p ⁺ =100%	5.8 s (11/2 ⁺) E _{ex} 5837 (28) p ⁺ =100% p ⁺ =100%	22 ms (10 ⁺) E _{ex} 104 (6) p ⁺ =100% p ⁺ =100%	56 s (11/2 ⁺) E _{ex} 46.80 (20) p ⁺ =100% p ⁺ =100%
143 66 Dy 77	144 66 Dy 78	145 66 Dy 79	146 66 Dy 80	147 66 Dy 81	148 66 Dy 82
58 s (11/2 ⁺) E _{ex} 357.7 (3) p ⁺ =100% p ⁺ =100%	9.1 s 0 ⁺ M ⁺ 56500 (30) p ⁺ =100% p ⁺ =100%	14 s (11/2 ⁺) E _{ex} 112.2 (2) p ⁺ =100% p ⁺ =100%	150 ms 10 ⁺ E _{ex} 258.7 (3) p ⁺ =100% p ⁺ =100%	55 s (11/2 ⁺) E _{ex} 758.5 (4) p ⁺ =100% p ⁺ =100%	3.3 m 0 ⁺ M ⁺ 67859 (11) p ⁺ =100% p ⁺ =100%
142 65 Tb 77	143 65 Tb 78	144 65 Tb 79	145 65 Tb 80	146 65 Tb 81	147 65 Tb 82
300 ms (5 ⁺) E _{ex} 282.2 (14) p ⁺ =100% p ⁺ =100%	597 ms 1 ⁺ E _{ex} 282.2 (14) p ⁺ =100% p ⁺ =100%	<21 s 5/2 ⁺ E _{ex} 94 (104) p ⁺ =100% p ⁺ =100%	28 s (6 ⁺) E _{ex} 112.2 (2) p ⁺ =100% p ⁺ =100%	30.9 s (11/2 ⁺) E _{ex} 112.2 (2) p ⁺ =100% p ⁺ =100%	1.87 m (11/2 ⁺) E _{ex} 92.6 (3) p ⁺ =100% p ⁺ =100%

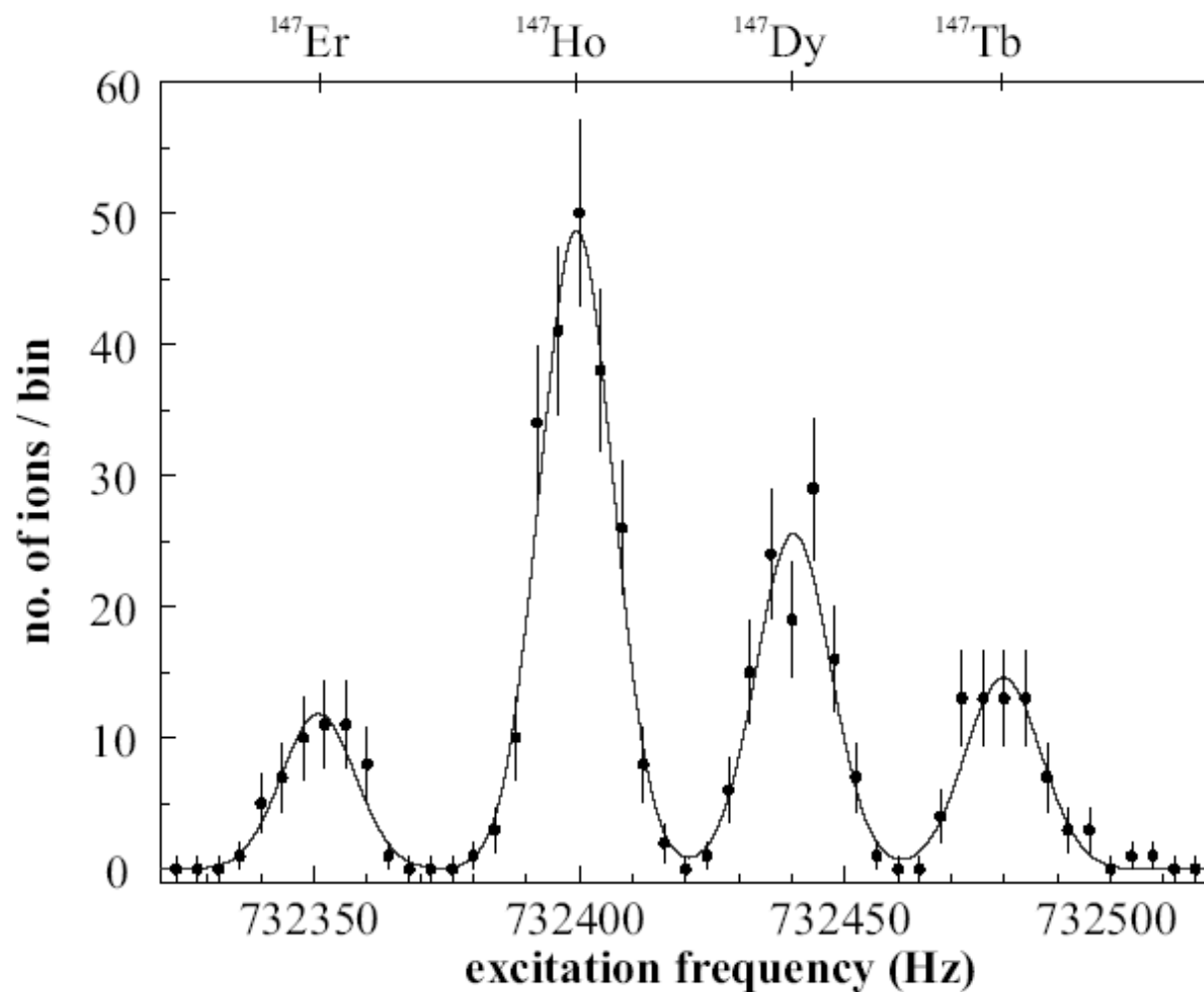
- One-proton separation energy for odd-A (even-N) isotopes of Ho (Z=67) and Tm (Z=69):



C. Rauth *et al.*, PRL 100 (2008) 012501

Importance of the mass

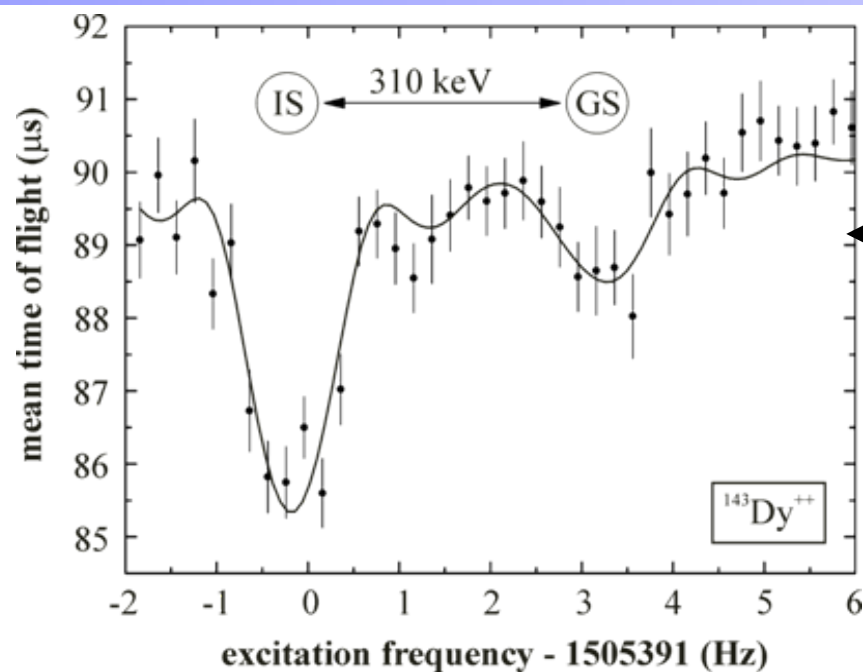
A remark



^{146}Tm 68 77 72 ms (10%) Ex 71 (6) p=100% p=100%	^{147}Tm 69 78 340 ms (6%) Ex 71 (6) p=100% p=100%	^{148}Tm 69 79 300 ms (10%) Ex 71 (6) p=100% p=100%	^{149}Tm 69 80 900 ms (11%) Ex 71 (6) p=100% p=100%	^{150}Tm 69 81 320 ms (10%) Ex 71 (6) p=100% p=100%	^{151}Tm 69 82 450 ms (10%) Ex 71 (6) p=100% p=100%
^{145}Er 68 77 900 ms (12%) Ex 71 (6) p=100% p=100%	^{146}Er 68 78 1.7 s (1%) Ex 71 (6) p=100% p=100%	^{147}Er 68 79 1.5 s (1%) Ex 71 (6) p=100% p=100%	^{148}Er 68 80 4.6 s (1%) Ex 71 (6) p=100% p=100%	^{149}Er 68 81 4 s (1%) Ex 71 (6) p=100% p=100%	^{150}Er 68 82 18.5 s (1%) Ex 71 (6) p=100% p=100%
^{144}Ho 67 77 700 ms Ex 71 (6) p=100% p=100%	^{145}Ho 67 78 1.4 s (1%) Ex 71 (6) p=100% p=100%	^{146}Ho 67 79 3.6 s (1%) Ex 71 (6) p=100% p=100%	^{147}Ho 67 80 5.8 s (1%) Ex 71 (6) p=100% p=100%	^{148}Ho 67 81 25 ms (10%) Ex 71 (6) p=100% p=100%	^{149}Ho 67 82 31 s (1%) Ex 71 (6) p=100% p=100%
^{143}Dy 66 77 3.0 s (1%) Ex 71 (6) p=100% p=100%	^{144}Dy 66 78 8.1 s (1%) Ex 71 (6) p=100% p=100%	^{145}Dy 66 79 14 s (1%) Ex 71 (6) p=100% p=100%	^{146}Dy 66 80 150 ms (10%) Ex 71 (6) p=100% p=100%	^{147}Dy 66 81 65 s (1%) Ex 71 (6) p=100% p=100%	^{148}Dy 66 82 3.3 m (1%) Ex 71 (6) p=100% p=100%
^{142}Tb 65 77 300 ms (10%) Ex 71 (6) p=100% p=100%	^{143}Tb 65 78 42 s (1%) Ex 71 (6) p=100% p=100%	^{144}Tb 65 79 330 ms (10%) Ex 71 (6) p=100% p=100%	^{145}Tb 65 80 33 s (10%) Ex 71 (6) p=100% p=100%	^{146}Tb 65 81 18 ms (10%) Ex 71 (6) p=100% p=100%	^{147}Tb 65 82 1.07 m (10%) Ex 71 (6) p=100% p=100%

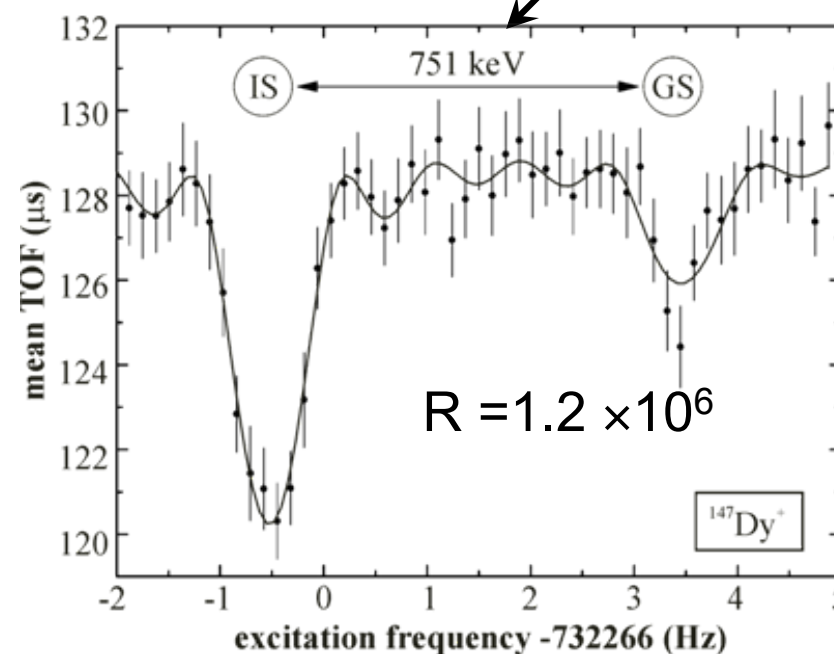
Importance of the mass

A remark



C. Rauth *et al.*, Eur. Phys. J ST 150 (2007) 329

145 68 Er 77	146 68 Er 78	147 68 Er 79	148 68 Er 80	149 68 Er 81	150 68 Er 82
800 ms (1/2) ⁺ M = 145.9064 (30) B = 100% P = 100%	1.7 s (1) M = 146.9050 (30) B = 100% P = 100%	254 (11/2) ⁺ B = 100% P = 100%	4.8 s (1) M = 147.9036 (30) B = 100% P = 100%	8.5 (1/2) ⁺ B = 100% P = 100%	18.5 s (1) M = 148.9022 (30) B = 100% P = 100%
144 67 Ho 77	145 67 Ho 78	146 67 Ho 79	147 67 Ho 80	148 67 Ho 81	149 67 Ho 82
700 ms M = 144.9126 (30) B = 100% P = 100%	144 (1/2) ⁺ B = 100% P = 100%	3.6 s (1/2) ⁺ M = 145.9112 (30) B = 100% P = 100%	5.8 s (1/2) ⁺ M = 146.9098 (30) B = 100% P = 100%	14.0 (1/2) ⁺ B = 100% P = 100%	21.1 (1/2) ⁺ B = 100% P = 100%
143 66 Dy 77	144 66 Dy 78	145 66 Dy 79	146 66 Dy 80	147 66 Dy 81	148 66 Dy 82
50 (1/2) ⁺ B = 100% P = 100%	3.3 s (1) M = 143.9096 (30) B = 100% P = 100%	4.1 (1/2) ⁺ B = 100% P = 100%	10 ms (1/2) ⁺ B = 100% P = 100%	33.2 (1/2) ⁺ B = 100% P = 100%	48.1 (1/2) ⁺ B = 100% P = 100%
143 65 Tb 77	144 65 Tb 78	145 65 Tb 79	146 65 Tb 80	147 65 Tb 81	148 65 Tb 82
200 ms (1/2) ⁺ B = 100% P = 100%	21 s (1/2) ⁺ M = 143.9082 (30) B = 100% P = 100%	3.1 (1/2) ⁺ B = 100% P = 100%	3.1 (1/2) ⁺ B = 100% P = 100%	1.8 (1/2) ⁺ B = 100% P = 100%	1.8 (1/2) ⁺ B = 100% P = 100%



Importance of the mass

Astrophysics

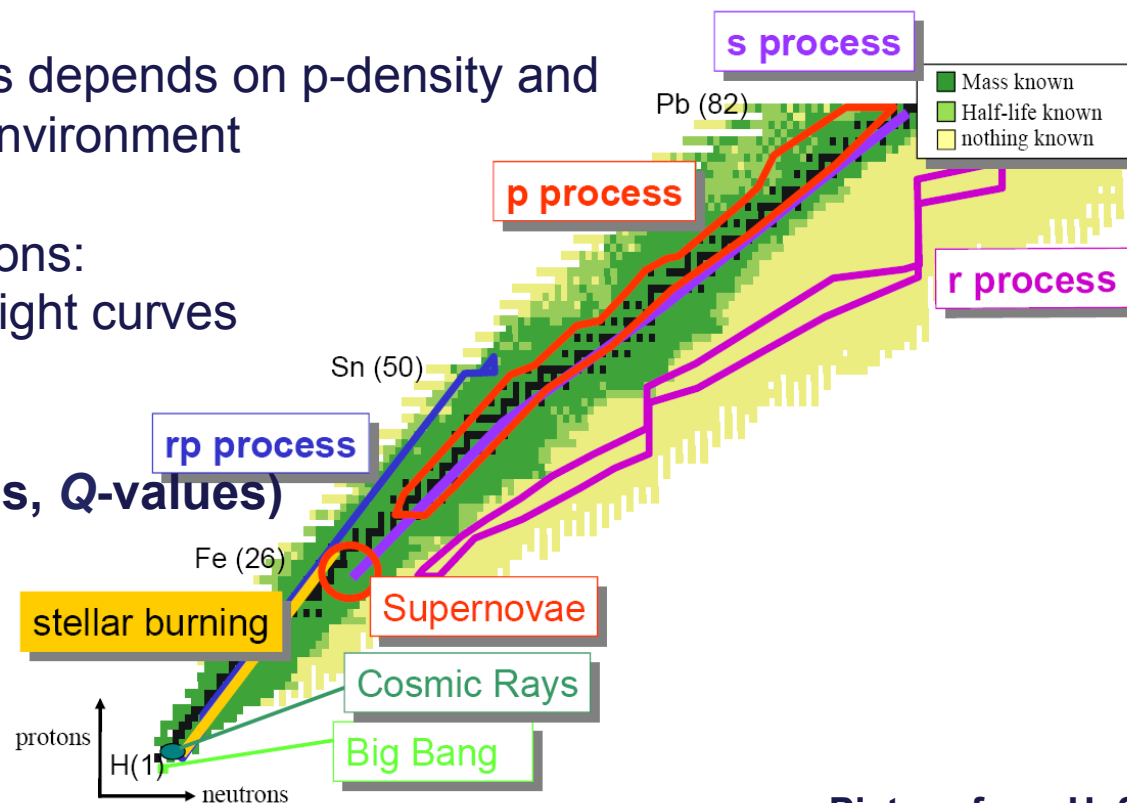
- Nucleosynthesis e.g. on accreting neutron stars
- Explosive hydrogen burning (X-ray bursts)
- Steady-state burning

Evolution of the process depends on p-density and temperature in stellar environment

Astrophysical observations:
elemental abundance, light curves

Required nuclear data:

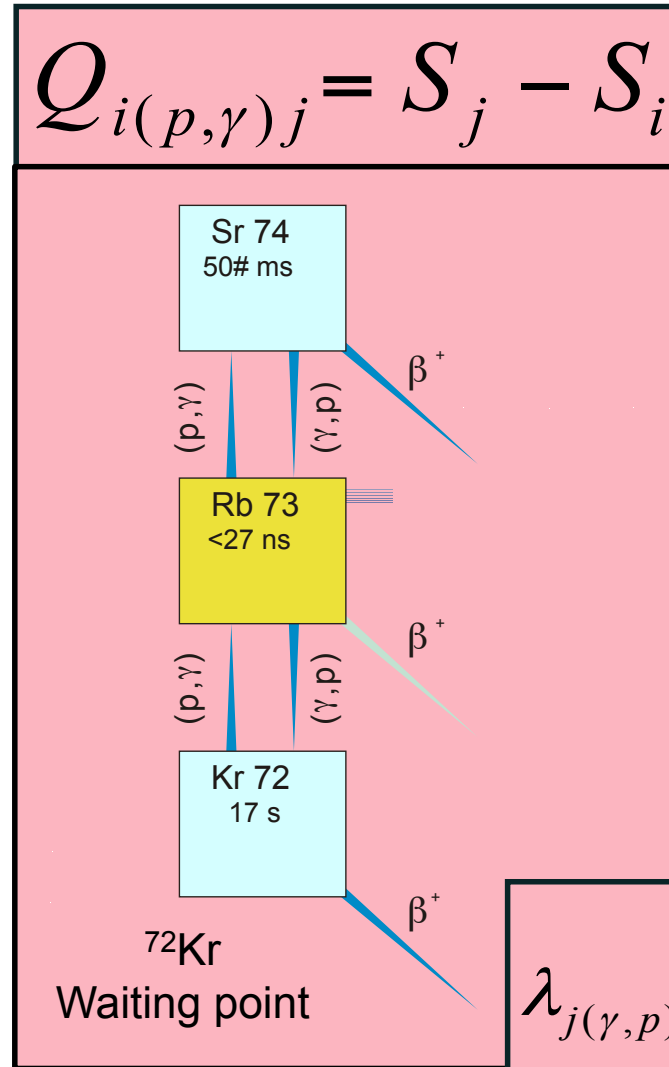
- **Masses (sep. energies, Q-values)**
- β decay half-lives
- Reaction rates



Picture from H. Schatz

Importance of the mass

Nucleon separation energies (S_n , S_p , S_{2n} , S_{2p})



structure & astrophysics

$$B(N, Z) - B(N - 2, Z)$$

$$B(N, Z) - B(N, Z - 2)$$

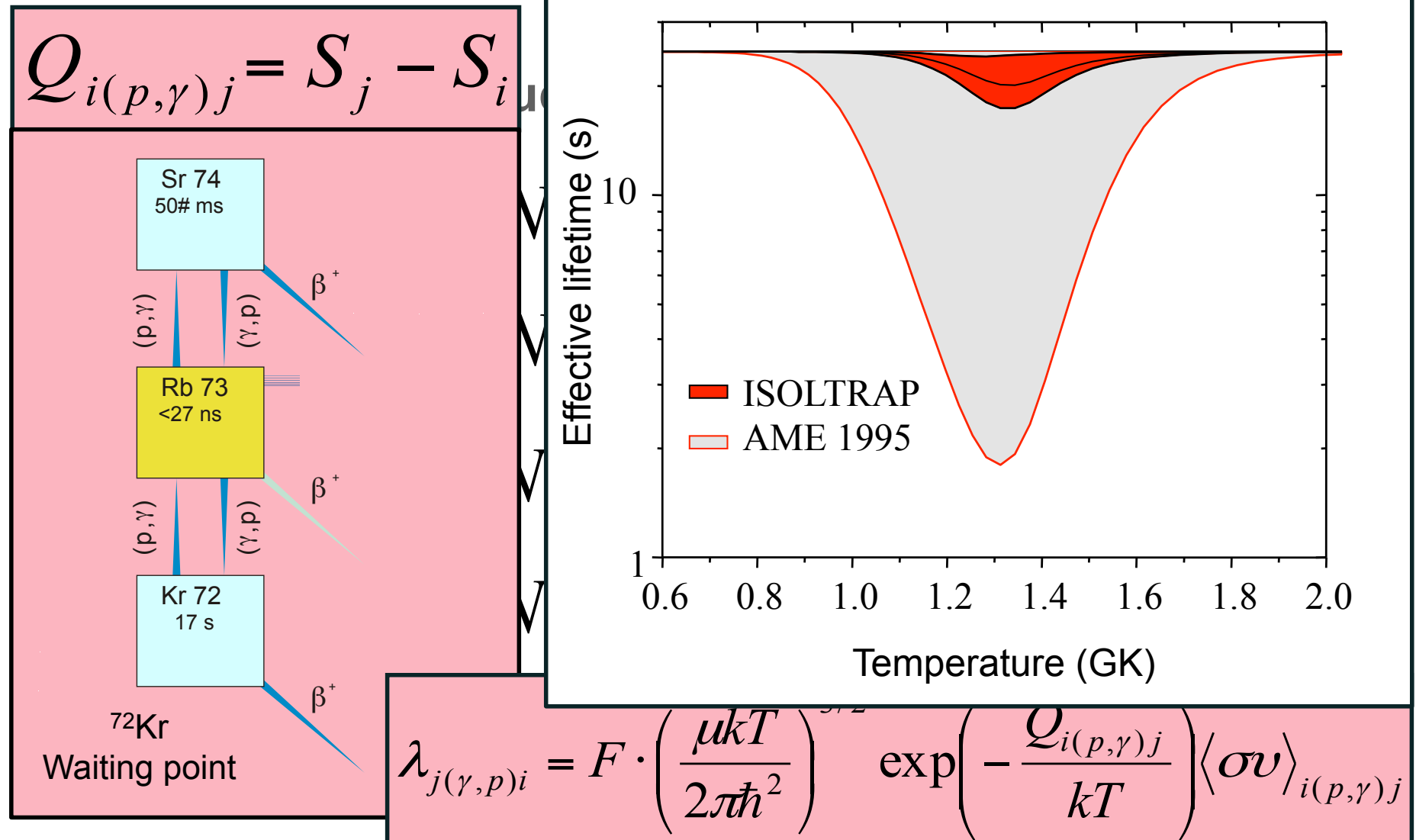
$$B(N, Z) - B(N - 1, Z)$$

$$B(N, Z) - B(N, Z - 1)$$

$$\lambda_{j(\gamma, p)i} = F \cdot \left(\frac{\mu k T}{2\pi \hbar^2} \right)^{3/2} \exp\left(-\frac{Q_{i(p, \gamma)j}}{kT} \right) \langle \sigma v \rangle_{i(p, \gamma)j}$$

Importance of the mass

Nucleon separation energies (S_n , S_p , S_{2n} , S_{2p})



Importance of the mass

Q-values from β -decay

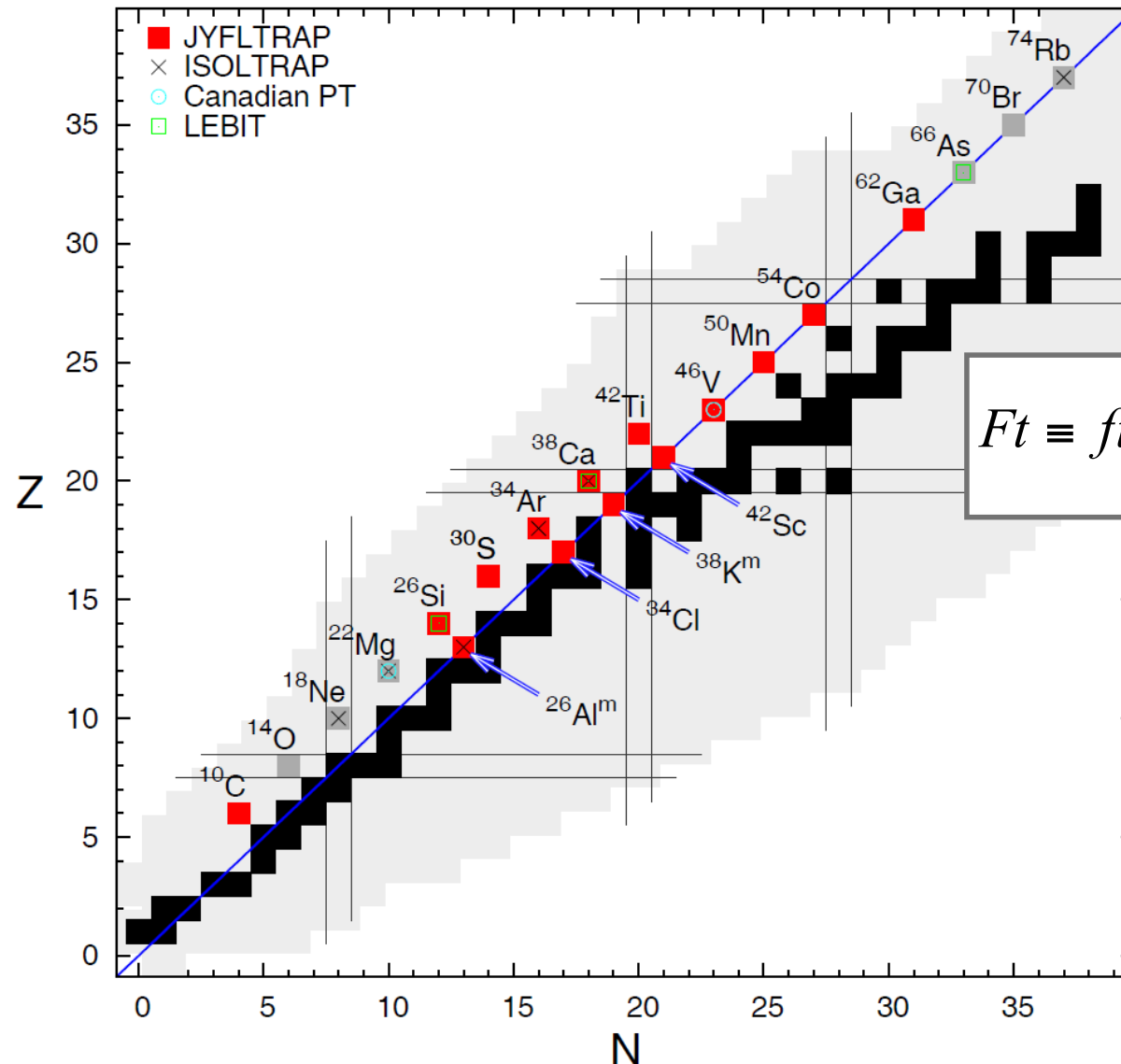
Fundamental symmetries & Neutrino-related physics

$$Q_{EC} = ME(Z, N) - ME(Z - 1, N + 1)$$

$$Q_{\beta^+} = ME(Z, N) - ME(Z - 1, N + 1) - 2m_e$$

$$Q_{\beta^-} = ME(Z, N) - ME(Z + 1, N - 1)$$

Importance of the mass Q-values from β -decay



CVC hypothesis
Superallowed
 β emitters
 $0^+ \rightarrow 0^+$

$$Ft \equiv ft(1 + \delta_R)(1 - \delta_C) = \frac{K}{2G_V^2(1 + \Delta_R^V)}$$

$f(Q^5)$ statistical rate function
 t ($t_{1/2}$, $1/R$) partial half-life

$$V_{ud}^2 = G_V^2 / G_\mu^2$$

$$|V_{ud}^2| + |V_{us}^2| + |V_{ub}^2| = 1 - \Delta$$

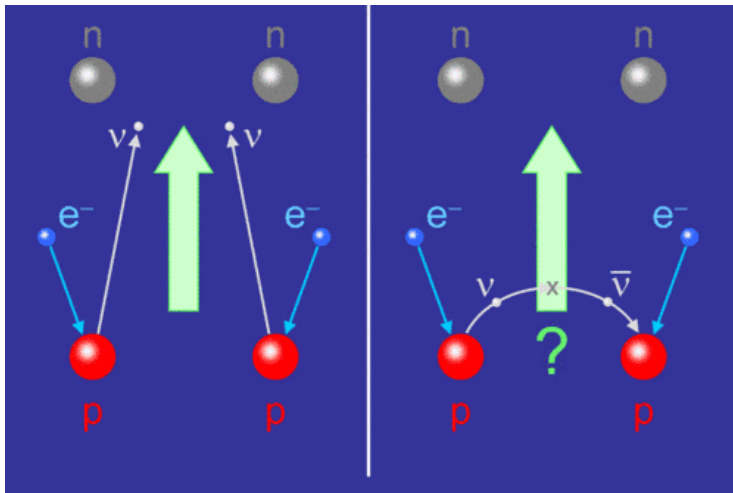
Importance of the mass Q-values from EC-decay

Is the neutrino a Majorana or Dirac particle?

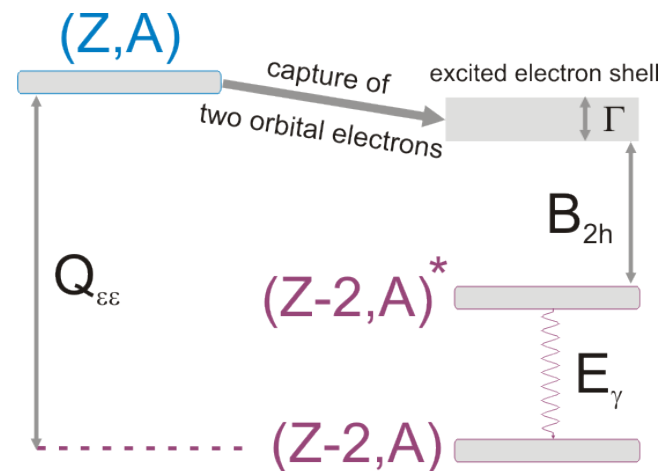
$2\nu 2EC$ ($T_{1/2} > 10^{24}$ y)

$0\nu 2EC$ ($T_{1/2} > 10^{30}$ y)

$$\frac{1}{T_{1/2}} = C \times m_\nu^2 \times |M|^2 \times |\Psi_{1e}|^2 \times |\Psi_{2e}|^2 \times \frac{\Gamma}{(Q - B_{2h} - E_\gamma^2)^2 + \frac{1}{4}\Gamma^2}$$



$0\nu 2EC$ might be resonantly enhanced ($T_{1/2} \sim 10^{25}$ y)



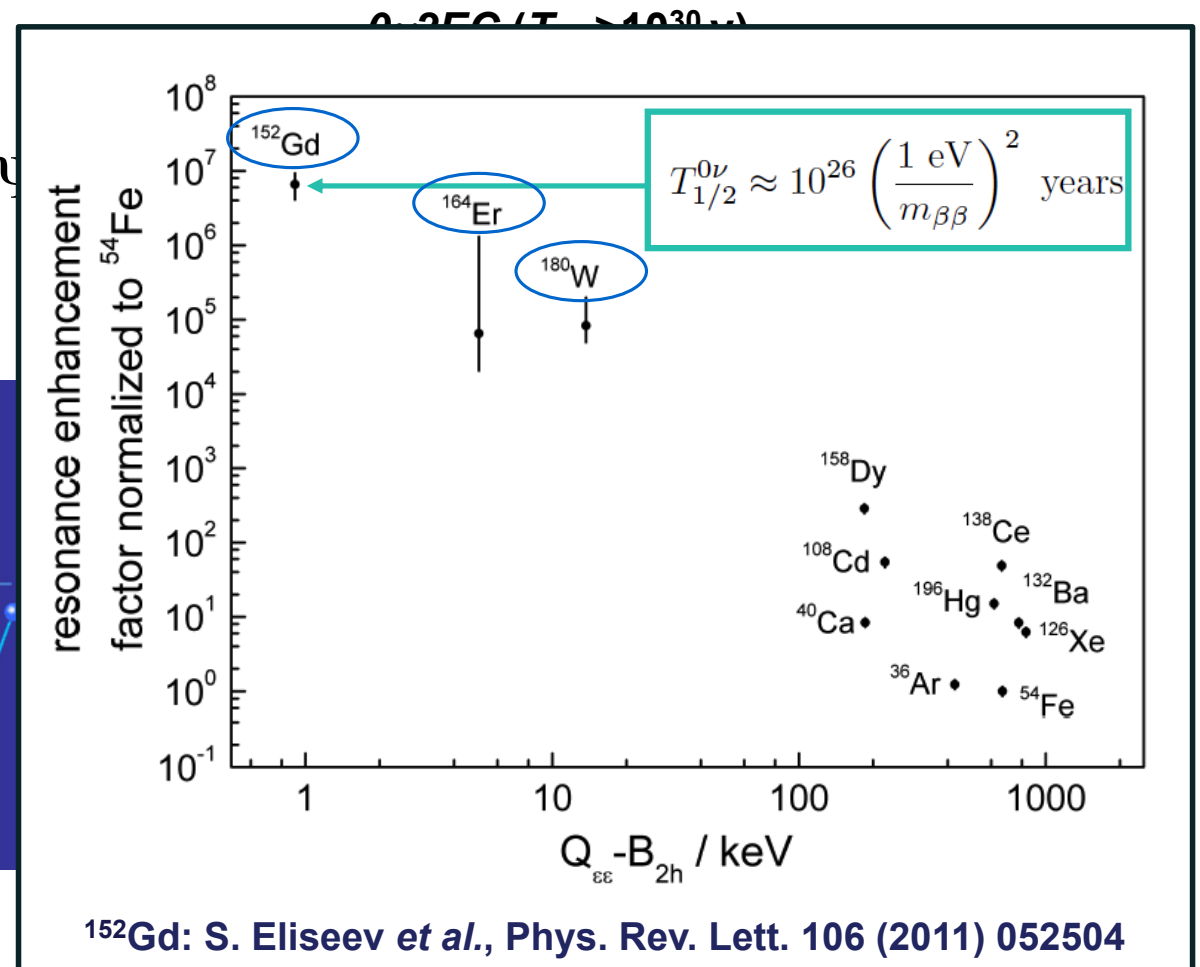
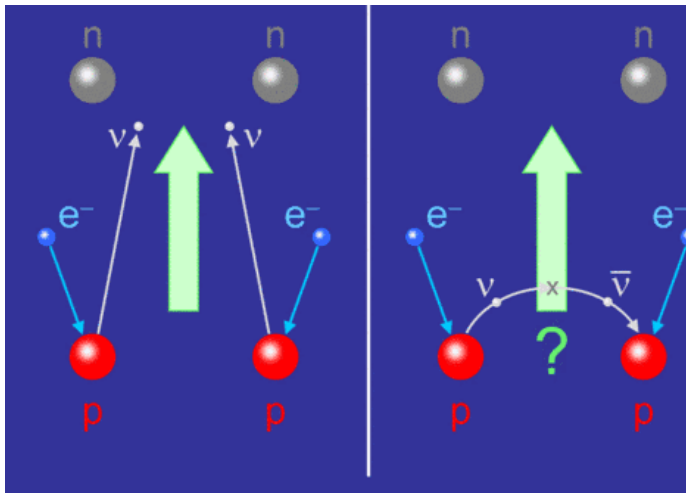
Search for nuclides with $D = (Q_{ee} - B_{2h} - E_\gamma) < 1$ keV

Importance of the mass Q-values from EC-decay

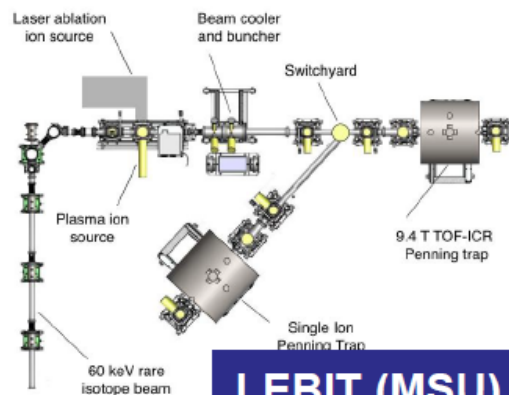
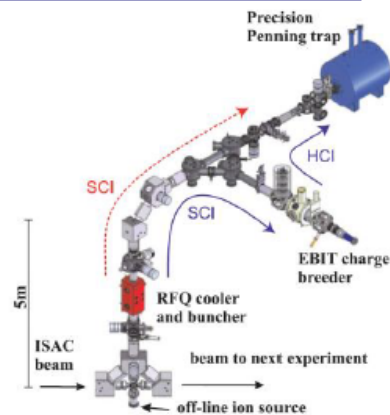
Is the neutrino a Majorana or Dirac particle?

$2\nu 2EC$ ($T_{1/2} > 10^{24}$ y)

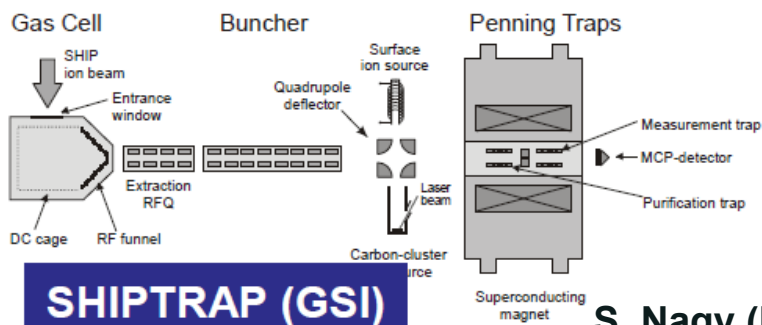
$$\frac{1}{T_{1/2}} = C \times m_\nu^2 \times |M|^2 \times |\psi|^2$$



TITAN (TRIUMF)

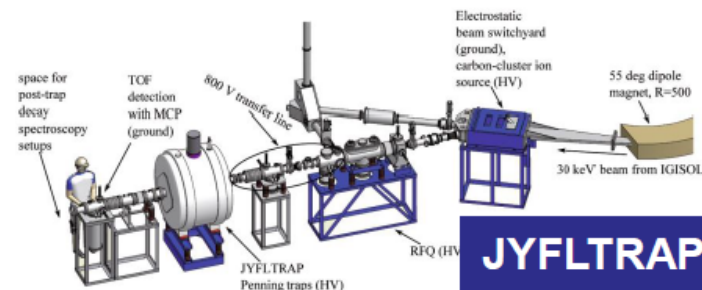


LEBIT (MSU)

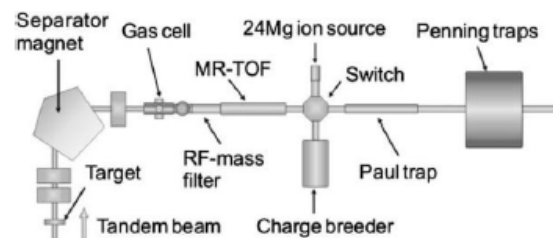


SHIPTRAP (GSI)

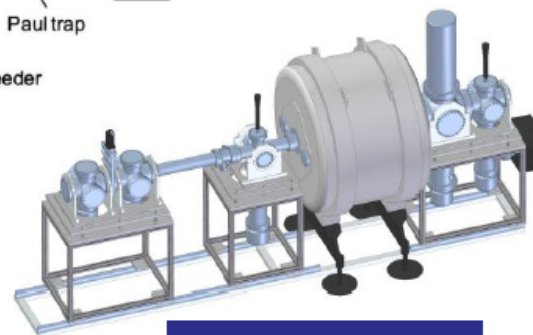
Growing facilities



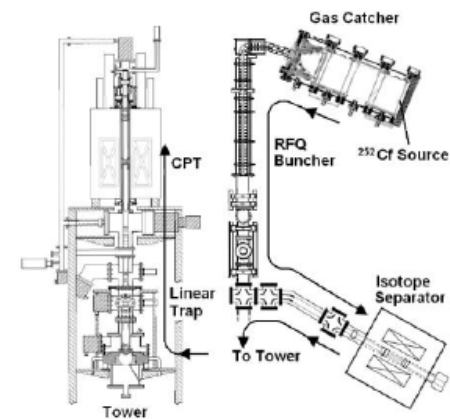
JYFLTRAP



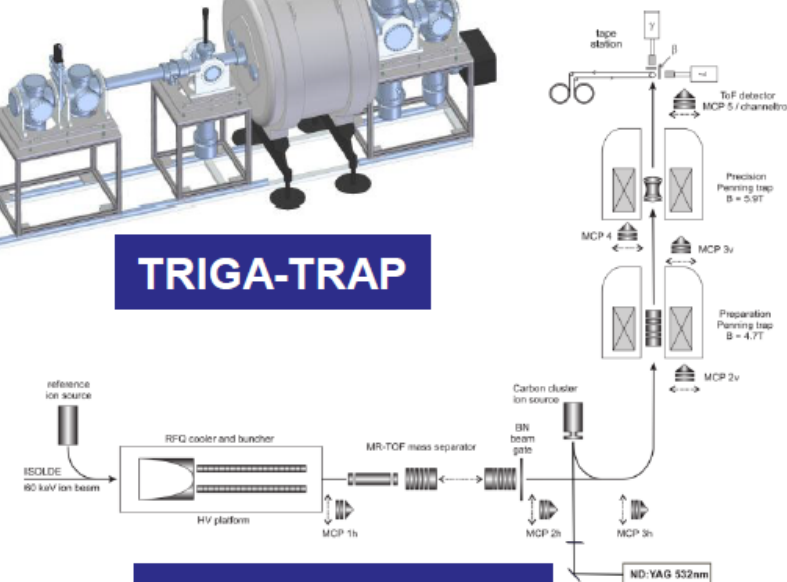
MLLTRAP



TRIGA-TRAP



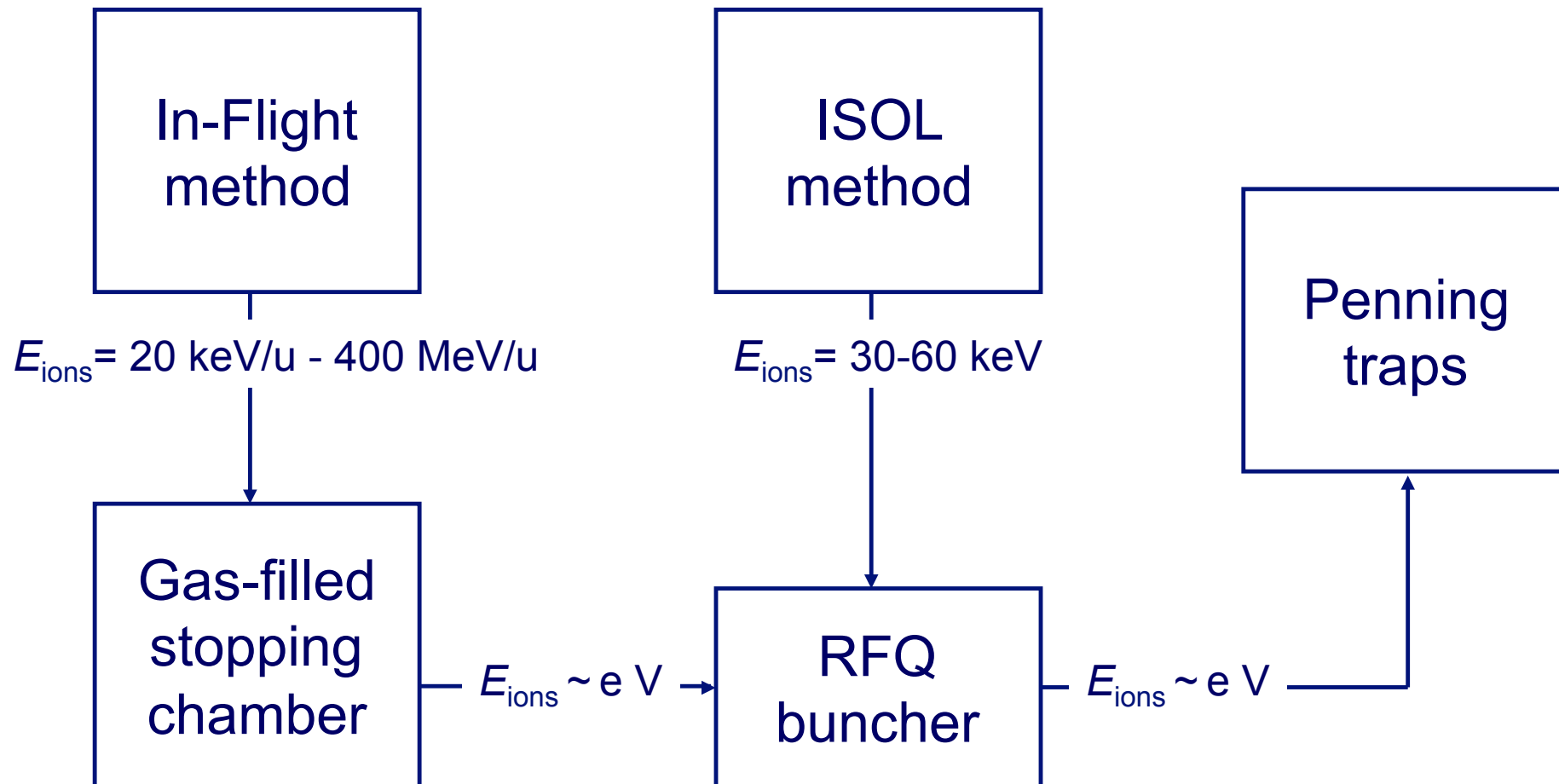
CPT (ARGONNE)



ISOLTRAP(CERN)

S. Nagy (EMIS2012)

Coupling Penning traps to RIB facilities



Coupling Penning traps to RIB facilities



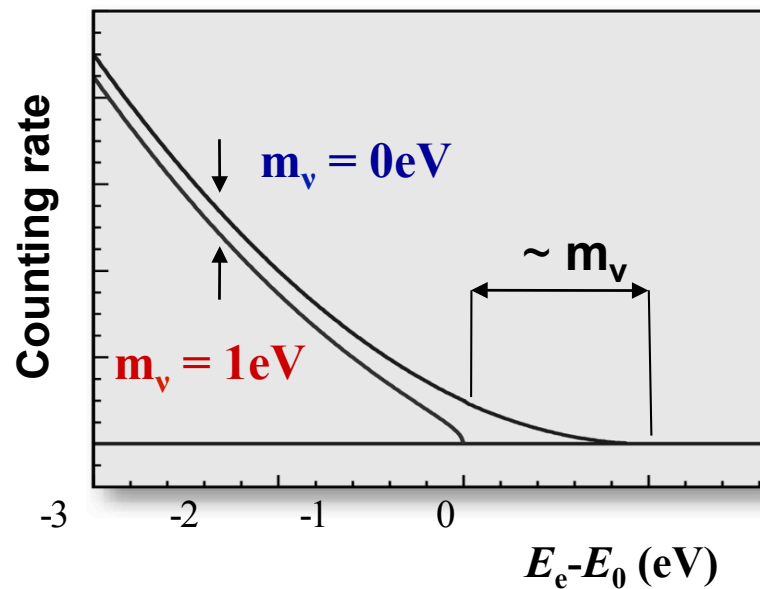
Importance of the mass

Q-values from β -decay

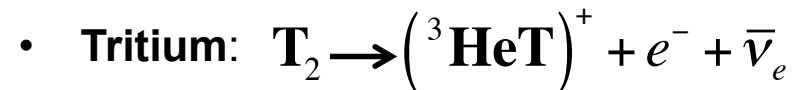
What is the mass of the electron (anti)neutrino?

$$\frac{dN}{dE} = k \times F(E, Z) \times p \times E_{total} \times (E_0 - E_e)^2 \times \left[(E_0 - E_e)^2 - m_\nu^2 \right]^{1/2}$$

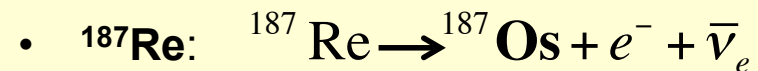
$$F(\delta E) = \int_{-\delta E}^0 N(E_\beta) dE \approx 2 \left(\frac{\delta E}{E_0} \right)^3$$



- The best candidates will be decaying nuclei with low *end-point energies*



$$Q = 18.6 \text{ keV} \quad T_{1/2} = 12.3 \text{ y}$$



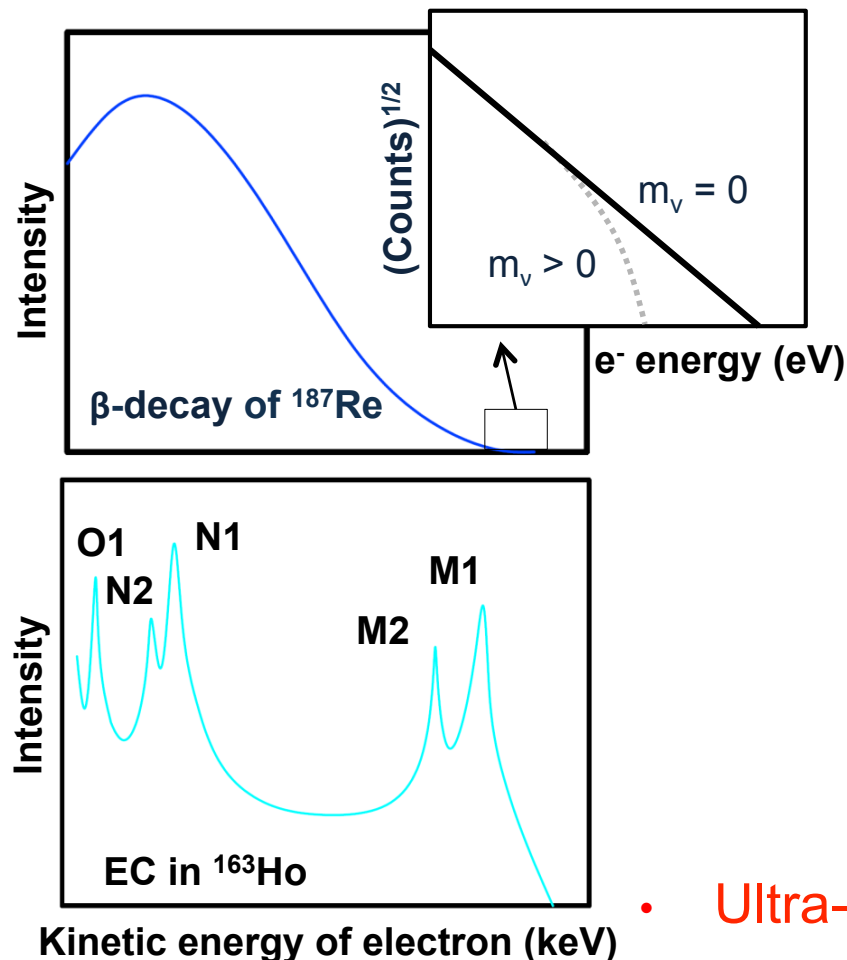
$$Q = 2.5 \text{ keV} \quad T_{1/2} = 43.2 \times 10^9 \text{ y}$$

Q has to be unambiguously determined (with $\delta m/m$ of at least 10^{-11}) from the mass difference using a Penning trap

Importance of the mass

Q-values from β -decay and EC-decay

What is the mass of the electron (anti)neutrino?



- β -decay of ^{187}Re - The MARE project



$$E_Q = 2.4653(17) \text{ keV} \quad \text{M. Sisti et al., NIMA (2004)}$$

- EC in ^{163}Ho - The ECHo, HOLMES and NuMECS projects



$$E_Q = 2.80(5) \text{ keV} \quad \text{L. Gastaldo et al., JLTP (2014)}$$

$$\Delta m = 2833(30_{\text{stat}})(15_{\text{sys}}) \text{ eV}/c^2$$

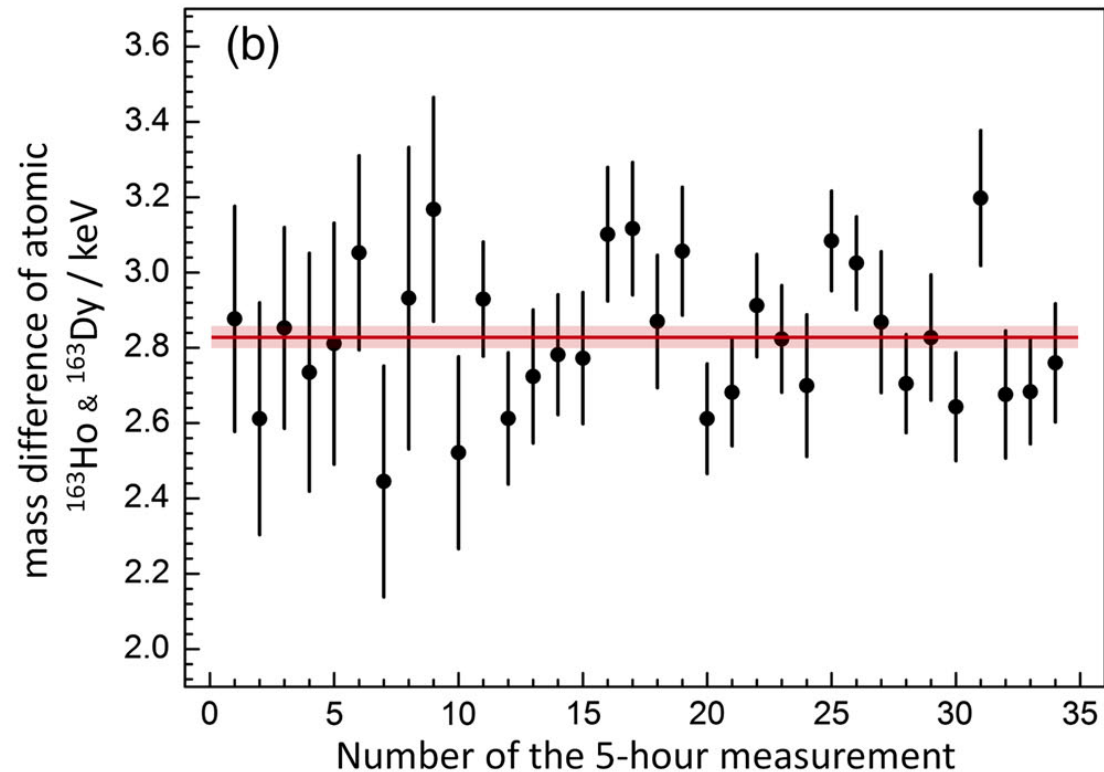
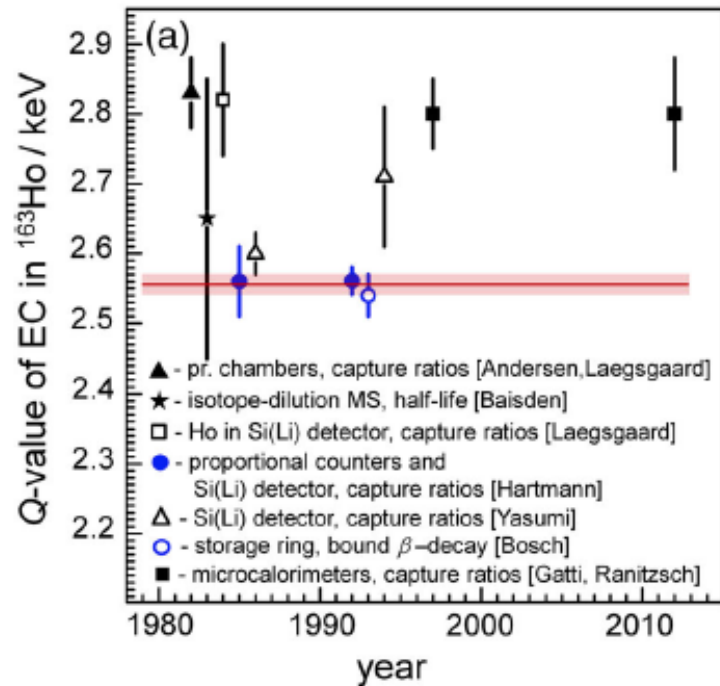
$$\text{S. Eliseev et al., PRL 115, (2015) 062501}$$

- Ultra-precise masses provide accurate Q-values

Recent highlights

Penning-trap Q-value for neutrino physics at GSI

$$\Delta m = 2833(30_{\text{stat}})(15_{\text{sys}}) \text{ eV}/c^2$$



Pictures from S. Eliseev *et al.*, PRL 115, (2015) 062501

Physics goals

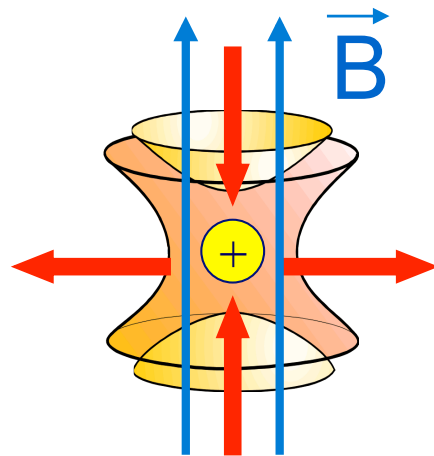
Mass uncertainty required

Field	Topic	Mass uncertainty
Nuclear structure physics	Shell and sub-shells, deformations, halos, S_{2p} , S_{2n} , SHEs	$10^{-6} - 10^{-8}$
Nuclear models and mass formulas	Nuclear models and mass formulas	$10^{-6} - 10^{-8}$
Astrophysics	(r)-process, (rp)-process, waiting point nuclides	10^{-7}
Weak interactions studies	CVC hypothesis, CKM matrix unitary	10^{-8}
Fundamental constants	Fine-structure constant	10^{-9}
Metrology	Silicon mass	$< 10^{-9}$
Neutrino physics	$0\nu\beta\beta$, $0\nu 2EC$ Neutrino mass	$10^{-8} - 10^{-9}$ $< 10^{-11}$
CPT Test	m_p and $m_{\bar{p}}$	$< 10^{-11}$

Penning traps

Fundamentals

Magnetic + Electrostatic field



Axial motion

$$\omega_z = \sqrt{\frac{q\Phi_0}{md^2}}$$

Reduced cyclotron motion

Magnetron motion

$$\omega_+ = \frac{1}{2} \left[\omega_c + \sqrt{(\omega_c^2 - 2\omega_z^2)} \right]$$

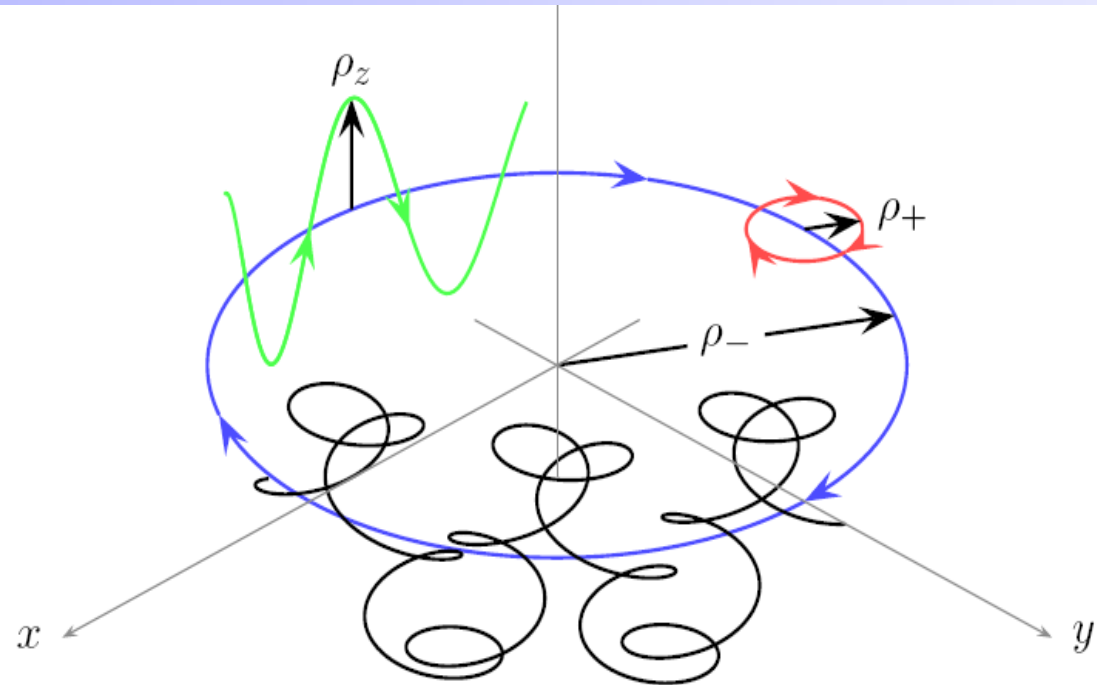
$$\omega_- = \frac{1}{2} \left[\omega_c - \sqrt{(\omega_c^2 - 2\omega_z^2)} \right]$$

$$\omega_c = \frac{q}{m} B$$

$$\omega_c = \omega_+ + \omega_-$$

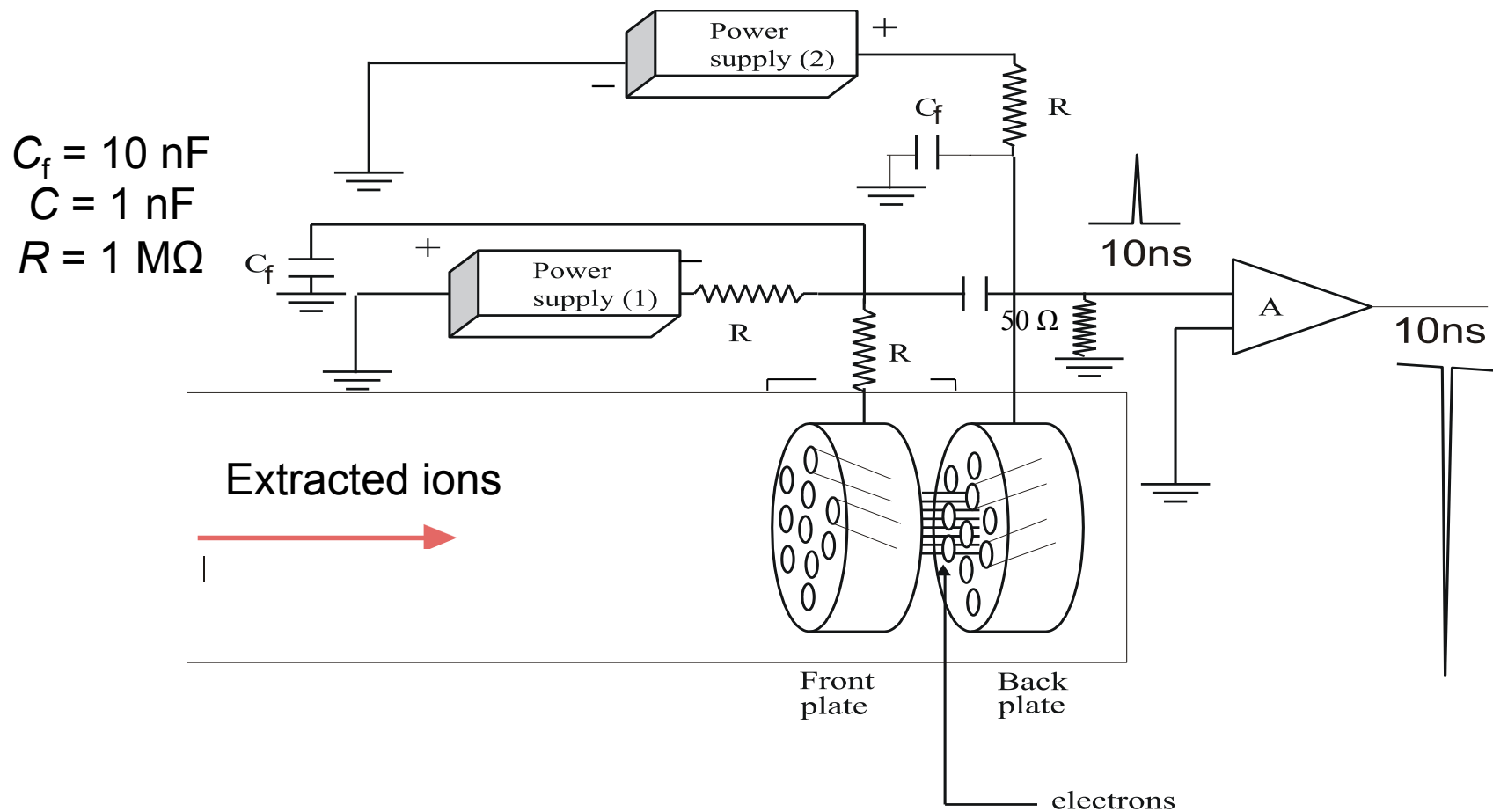
$$\omega_+ > \omega_z > \omega_-$$

$$\omega_+^2 + \omega_z^2 + \omega_-^2 = \omega_c^2$$



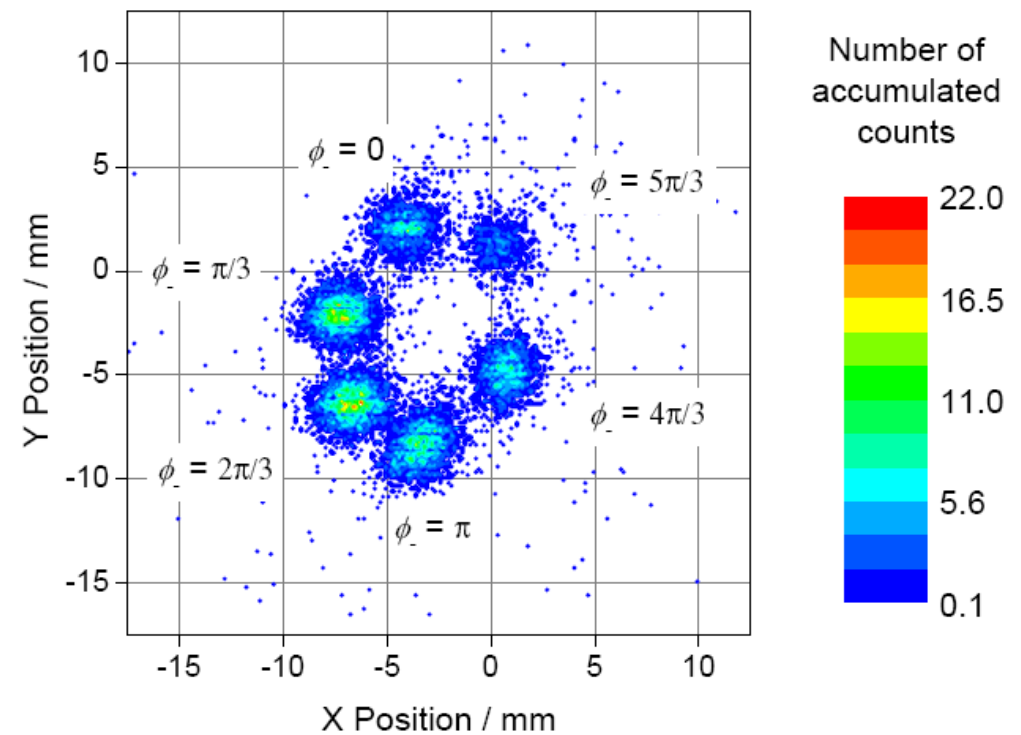
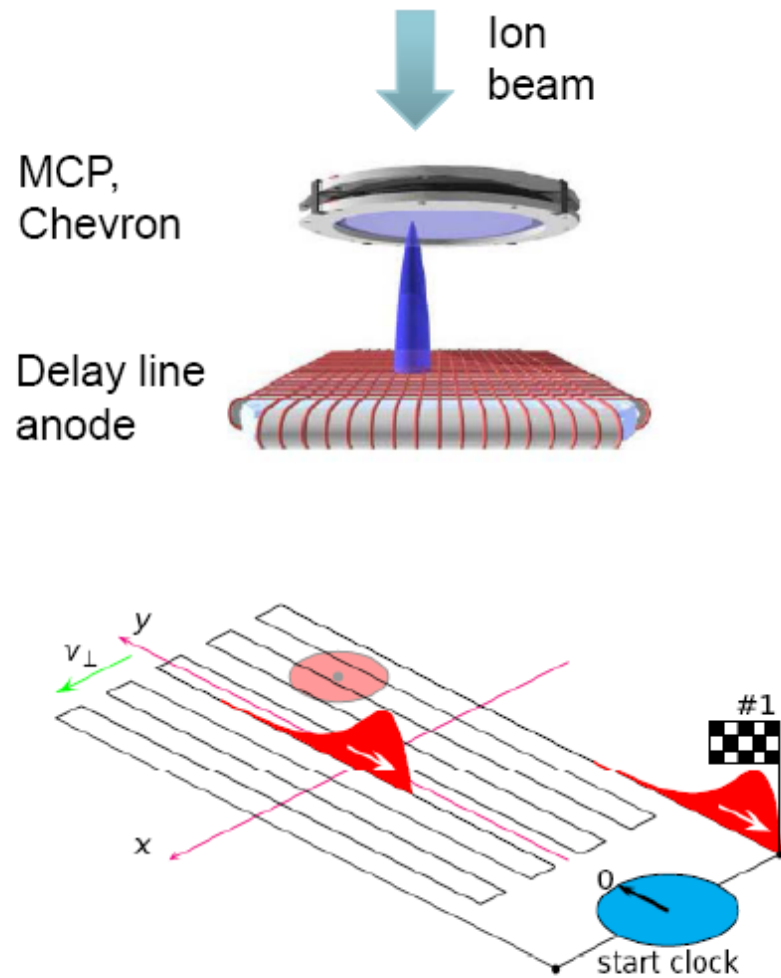
Penning traps

Detection



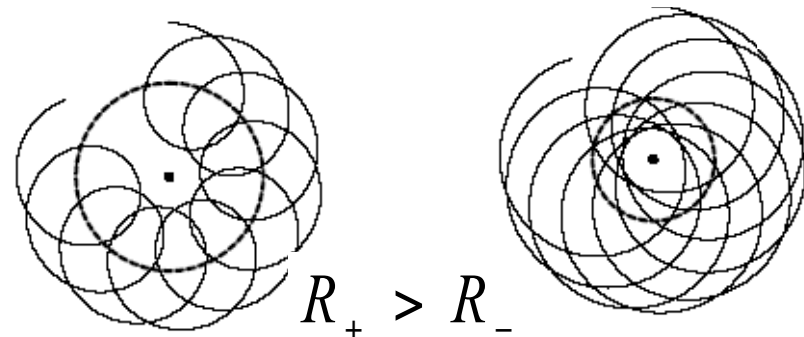
Penning traps

Detection



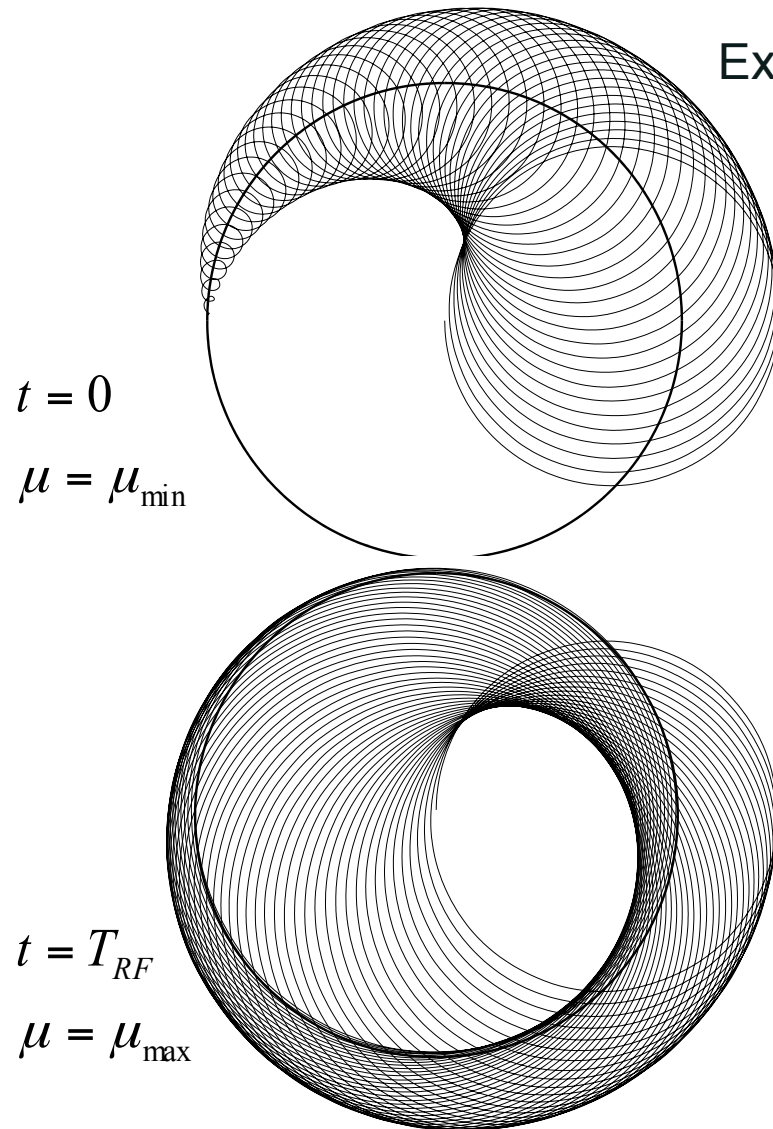
G. Eitel *et al.*, Nucl. Instrum. Methods A, 606, (2009) 475

$$R_+ < R_-$$

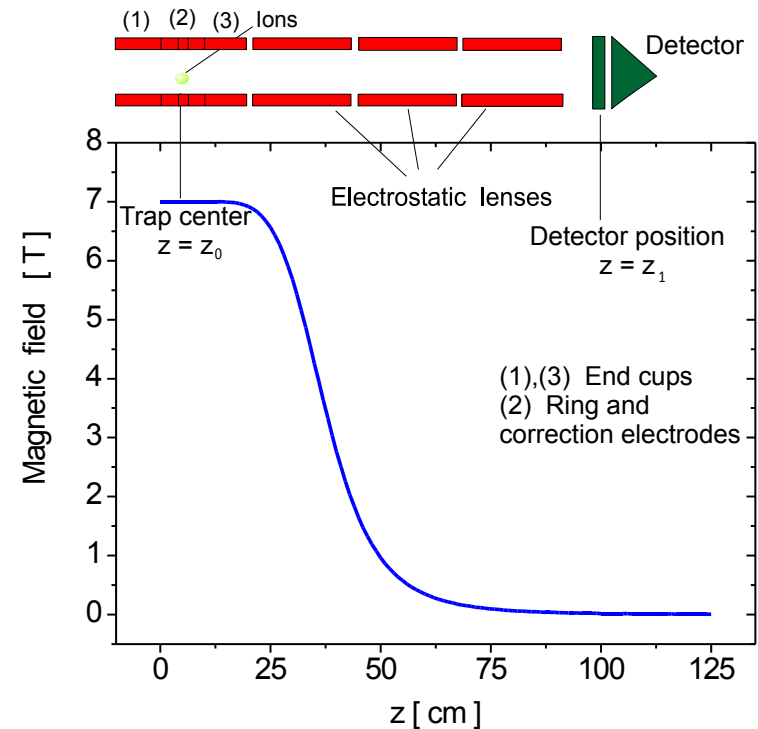


Penning-trap techniques

Time-of-Flight Ion Cyclotron Resonance (TOF-ICR)



External RF quadrupole field with $\omega_{RF} = \omega_c (= \omega_+ + \omega_-)$



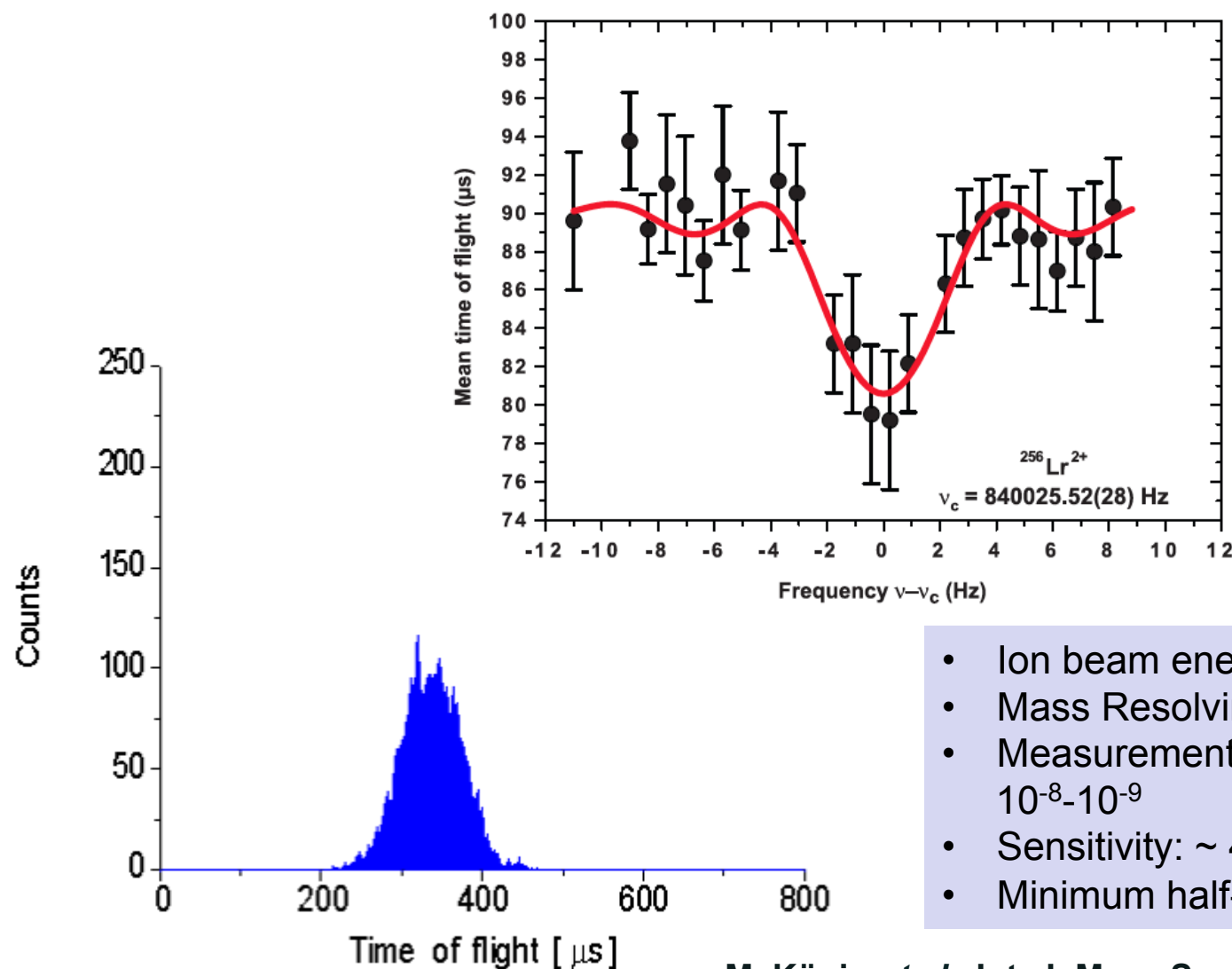
$$F_z = \mu \frac{\partial B}{\partial z}$$

$$TOF_{\min} \Leftrightarrow \mu_{\max} \Leftrightarrow \omega_{RF} = \omega_c$$

G. Gräff et al., Z. Phys. A 297 (1980) 35

Penning-trap techniques

Time-of-Flight Ion Cyclotron Resonance (TOF-ICR)

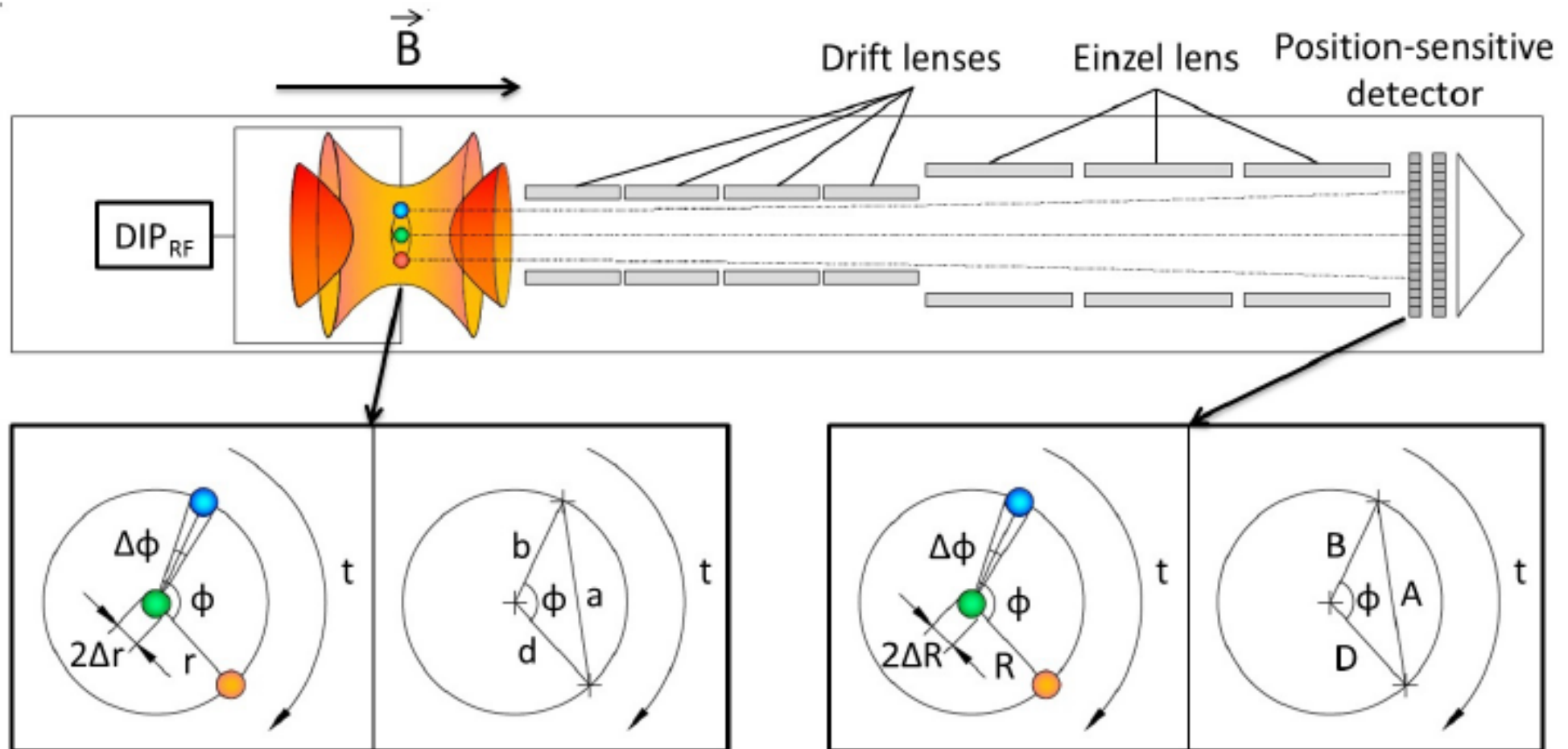


- Ion beam energy: ~ 1 eV
- Mass Resolving Power: 1×10^6
- Measurement Uncertainty: 10^{-8} - 10^{-9}
- Sensitivity: ~ 40 ions
- Minimum half-life: 8 ms

M. König *et al.*, Int. J. Mass Spectrom. 142 (1995) 95

Penning-trap techniques

Phase-Imaging Ion Cyclotron Resonance (PI-ICR)



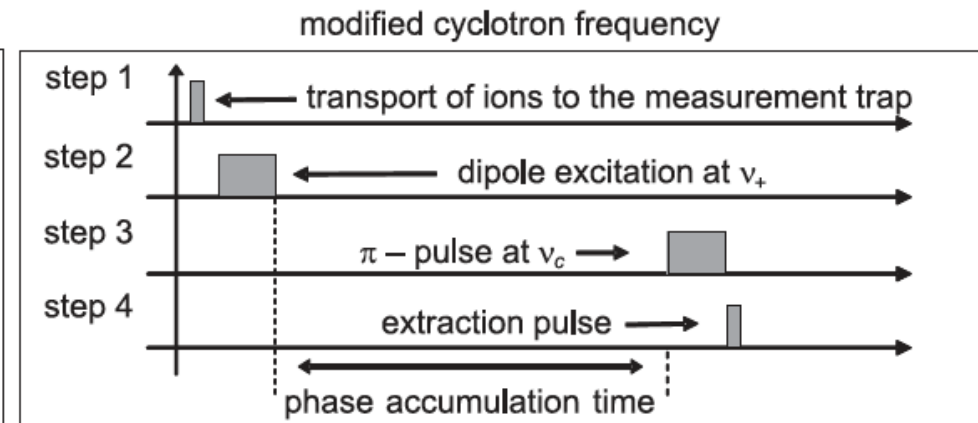
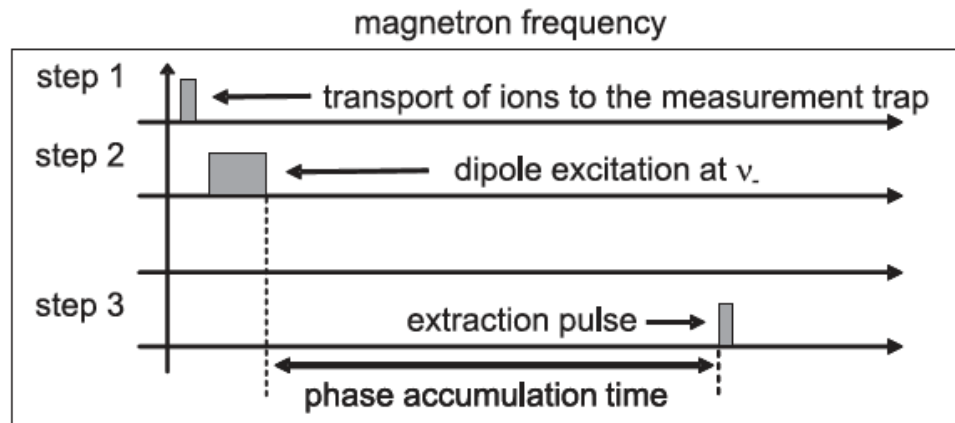
- Ion beam energy: ~ 1 eV
- Mass Resolving Power: 1×10^7
- Measurement Uncertainty: 10^{-9} - 10^{-10}

- Sensitivity: ~ 10 -20 ions
- Minimum half-life: ~ 1 ms

S. Eliseev *et al.*, PRL 110, (2013) 082501

Penning-trap techniques

Phase-Imaging Ion Cyclotron Resonance (PI-ICR)



$$\nu = \frac{\phi + 2\pi n}{2\pi t}$$

$$\delta\nu = \frac{\delta k}{2\pi t \sqrt{1 - k^2}}$$

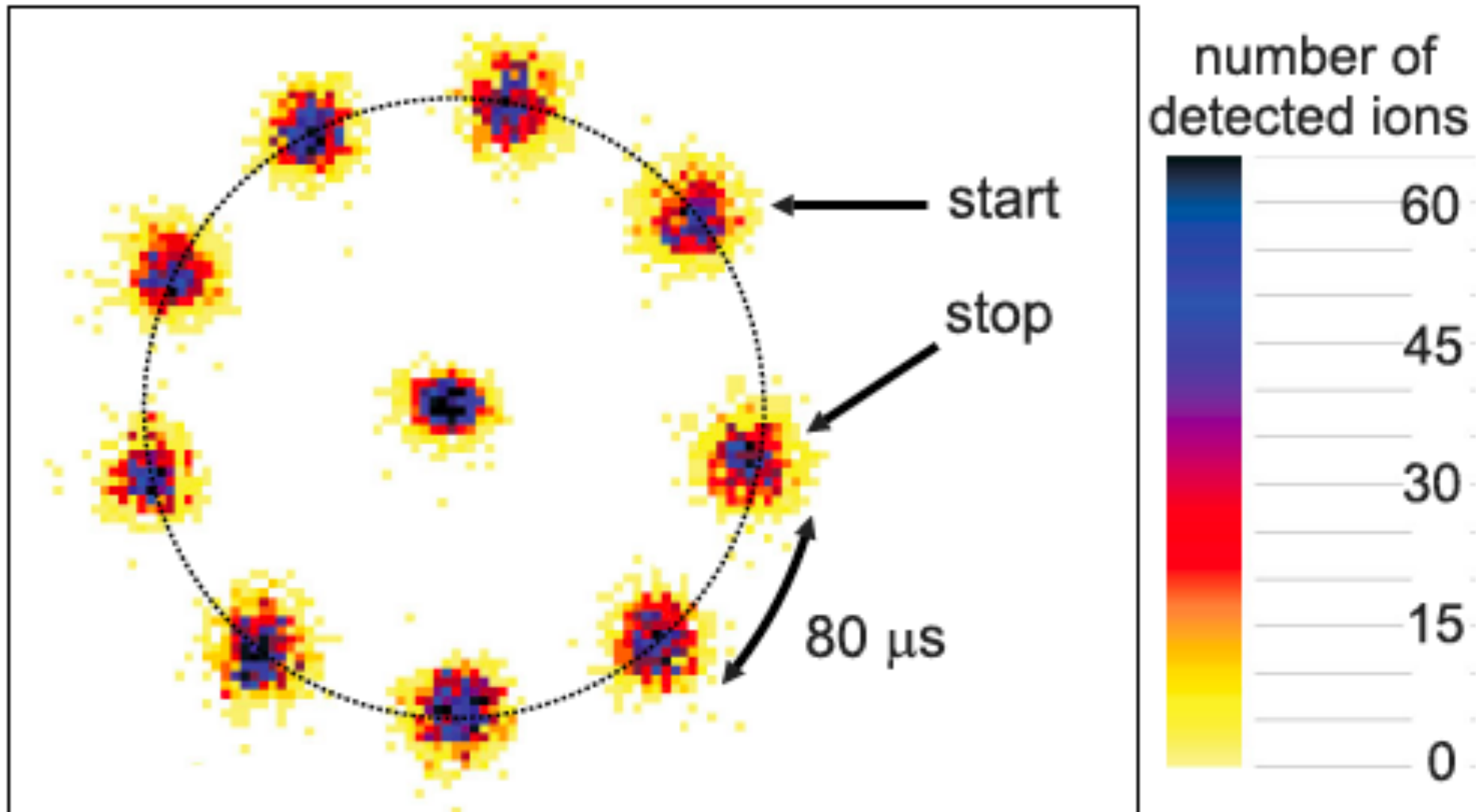
$$\frac{\nu_c}{\Delta\nu_c} \approx \frac{\nu_+}{\Delta\nu_+} = \frac{\phi + 2\pi n}{\Delta\phi} = \frac{\pi\nu_+ t R_+}{\Delta R_+}.$$

$$\delta\nu_c \approx \frac{1}{\pi t \sqrt{N}} \times \frac{\Delta R}{R}$$

S. Eliseev *et al.*, PRL 110, (2013) 082501

Mass measurements in nuclear physics

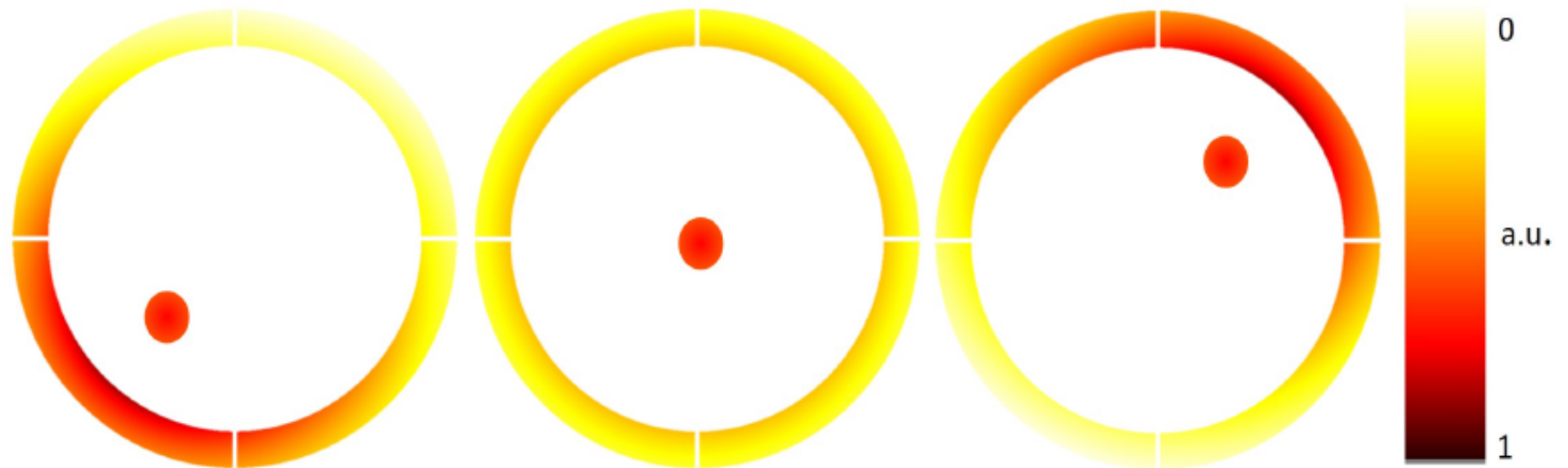
Mass measurements techniques: PI-ICR



S. Eliseev *et al.*, PRL 110, (2013) 082501

Penning-trap techniques

Induced-image current detection



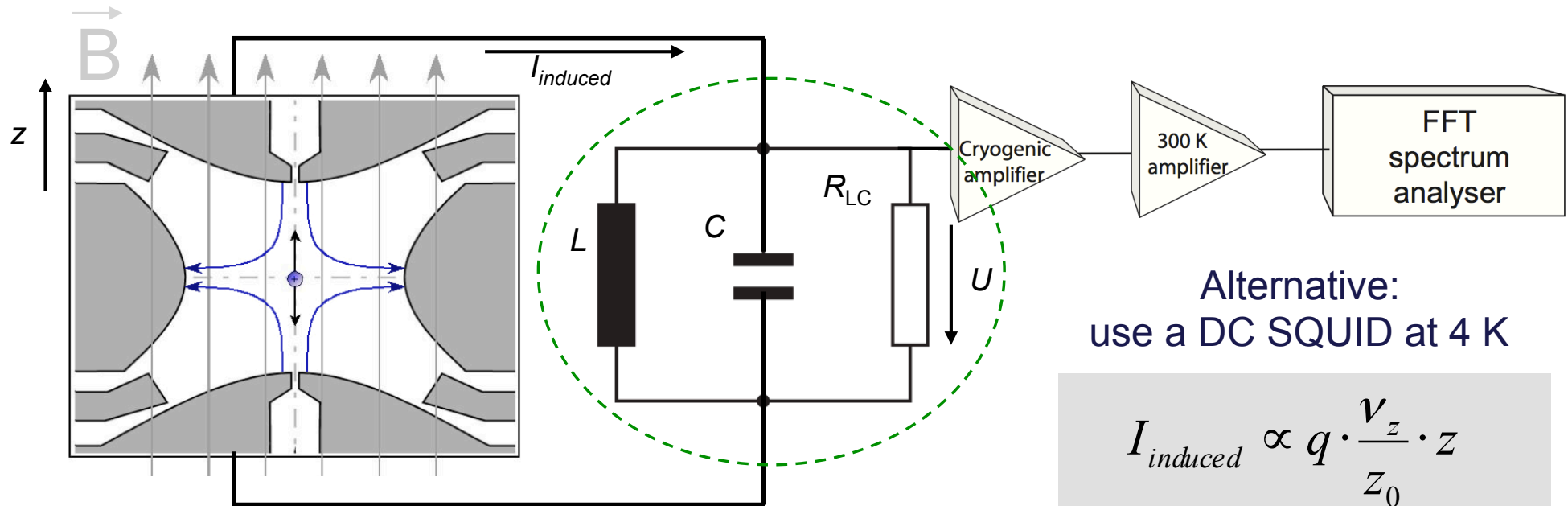
$$Q_i(x) = -q \frac{\Delta\psi(x)}{\Delta U}$$

$$Q_i(x) = -q \frac{x}{d}$$

$$i = \frac{dQ_i(x, t)}{dt} = -\frac{q}{d} \frac{dx}{dt} = 2\pi q \frac{x(t)\nu}{d}$$

Penning-trap techniques

Induced-image current detection

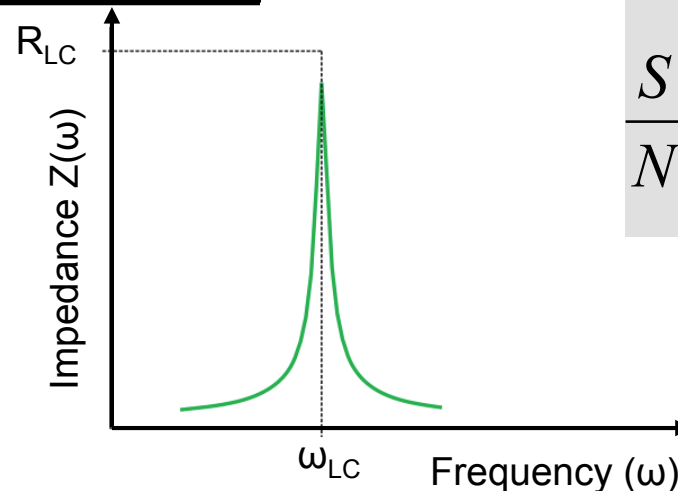


Alternative:
use a DC SQUID at 4 K

$$I_{induced} \propto q \cdot \frac{v_z}{z_0} \cdot z$$

$$\frac{S}{N} \propto q \cdot \frac{z}{z_0} \cdot \sqrt{\frac{v_z}{\Delta \nu}} \cdot \sqrt{\frac{Q}{TC}}$$

$$Z = \frac{Q}{2\pi\nu C}$$



This technique was originally developed mainly at:

- Univ. of Harvard ($\bar{p}$, p)
- Univ. of Washington (low m/q) (now at MPIK in Heidelberg)
- MIT (m/q up to ~ 50) (now at Florida State Univ.)
- University of Mainz (now also at MPIK)

THE TRAPSENSOR LABORATORY IN GRANADA



Physics Motivation

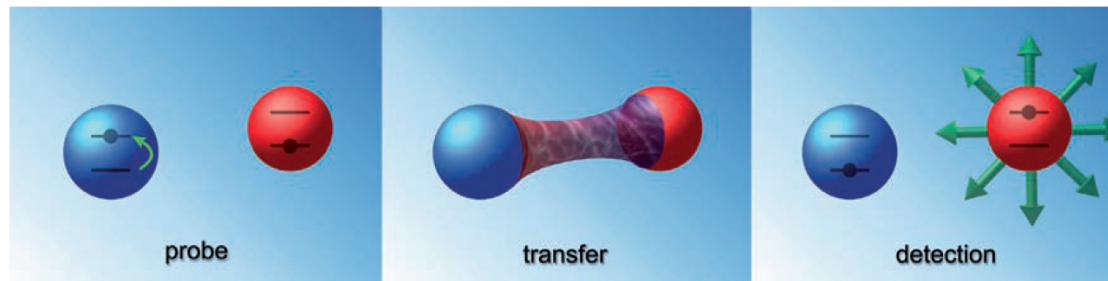
Precision sensing and metrology with a single ion

- **Two techniques in the framework of quantum-optics experiments**

→ Ions stored in the same trap

- Quantum-logic spectroscopy
(Same potential well)

P. O. Schmidt *et al.*, Science 309 (2005) 749-752
(Wineland's group)

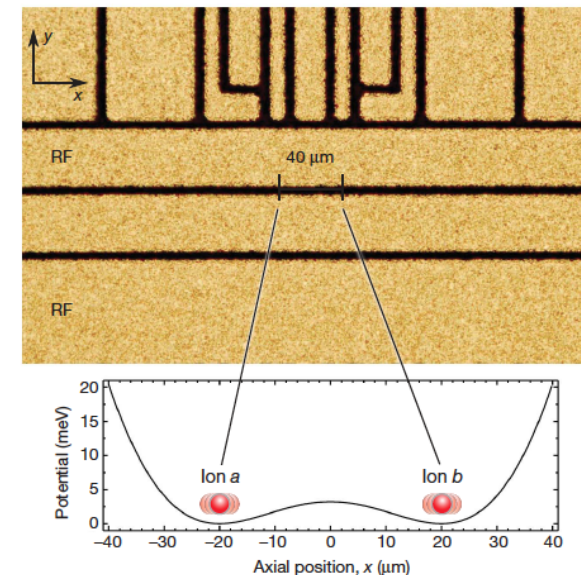


Picture from P.O. Schmidt *et al.*, Special Issue / PTB-Mitteilungen 119 (2009), No. 2

- Extension to molecular ions

F. Wolf *et al.*, Nature 530 (2016) 457
(Schmidt's group)

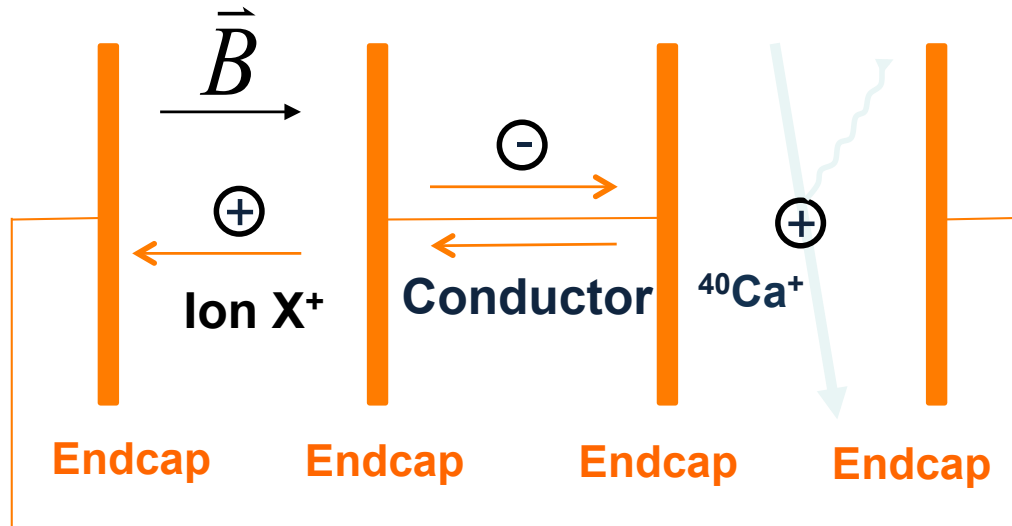
- Coupled quantized mechanical oscillators
(Different potential well)



K.E. Brown *et al.*, Nature 471 (2016) 196-199
(Wineland's group)

The QS technique

The principle of the stored ion calorimeter



D. J. Heinzen & D. J. Wineland, Phys. Rev. A 42(5) (1990) 2977-2994

1. COOLING

Conductor ON

2. PROBING THE ION X^+

Conductor OFF

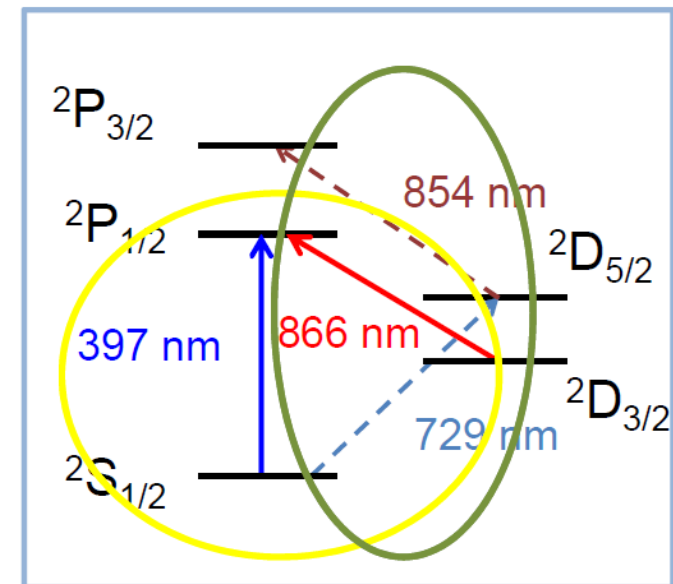
3. ION-ION INTERACTION

Conductor ON/ Lasers OFF

4. DETECTION

Conductor ON

$$T(^{40}\text{Ca}^+) \approx 1 \text{ mK}$$
$$T(^{40}\text{Ca}^+) \ll; \quad \langle n \rangle \sim 0$$

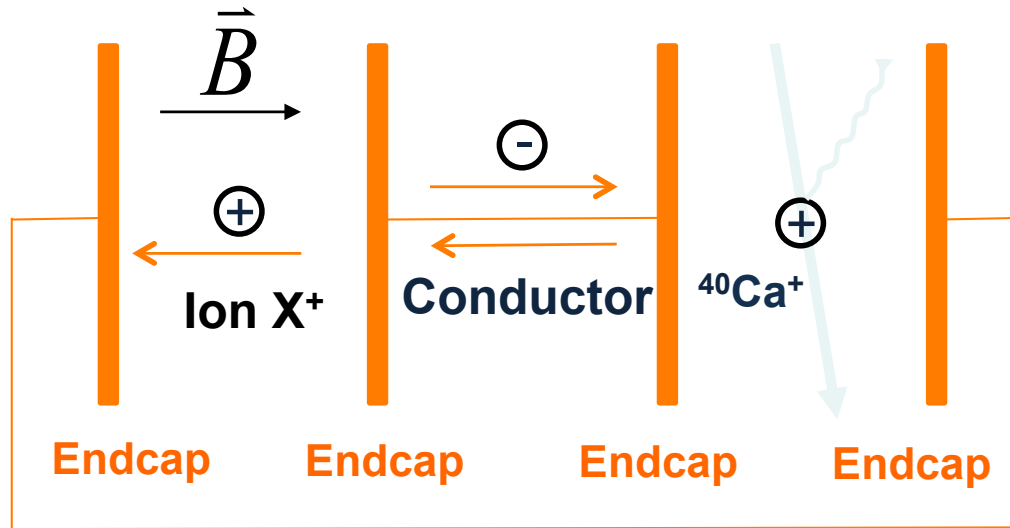


Doppler cooling

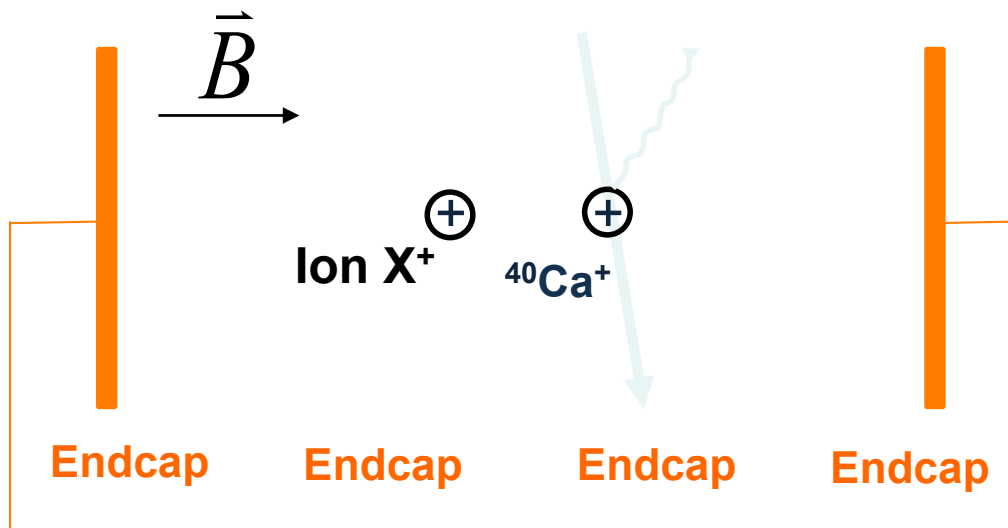
Ground-state cooling

The QS technique

The principle of the stored ion calorimeter

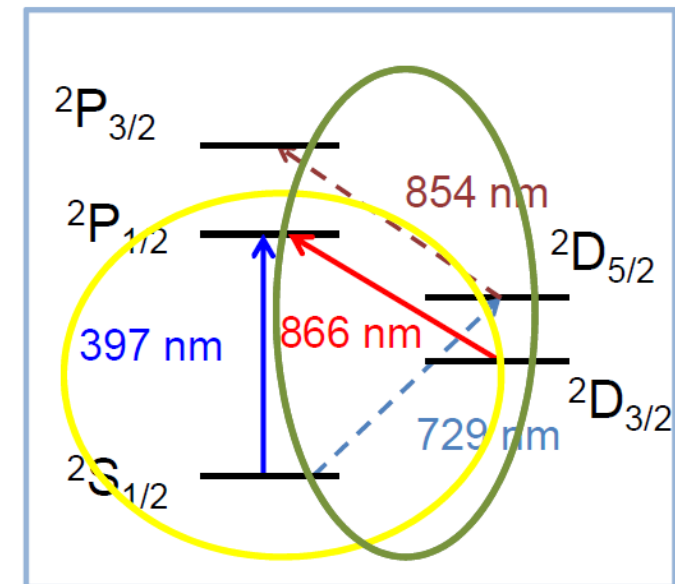


D. J. Heinzen & D. J. Wineland, Phys. Rev. A 42(5) (1990) 2977-2994



$$T(^{40}\text{Ca}^+) \approx 1 \text{ mK}$$

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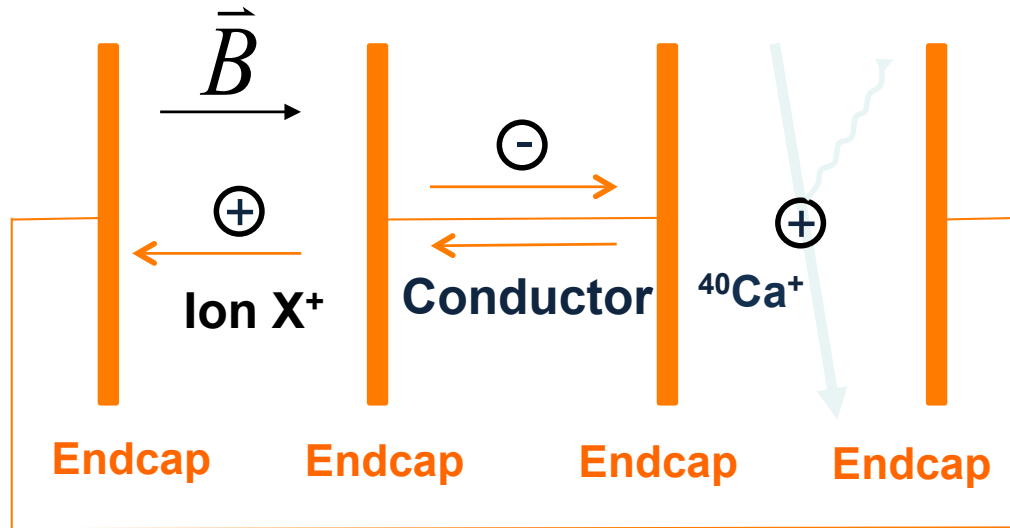


Doppler cooling

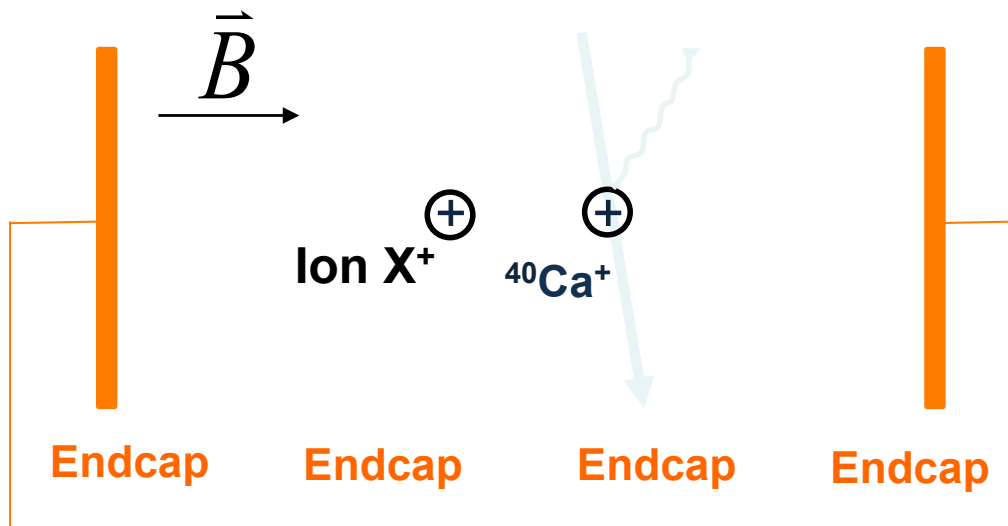
Ground-state cooling

The QS technique

The principle of the stored ion calorimeter



D. J. Heinzen & D. J. Wineland, Phys. Rev. A 42(5) (1990) 2977-2994



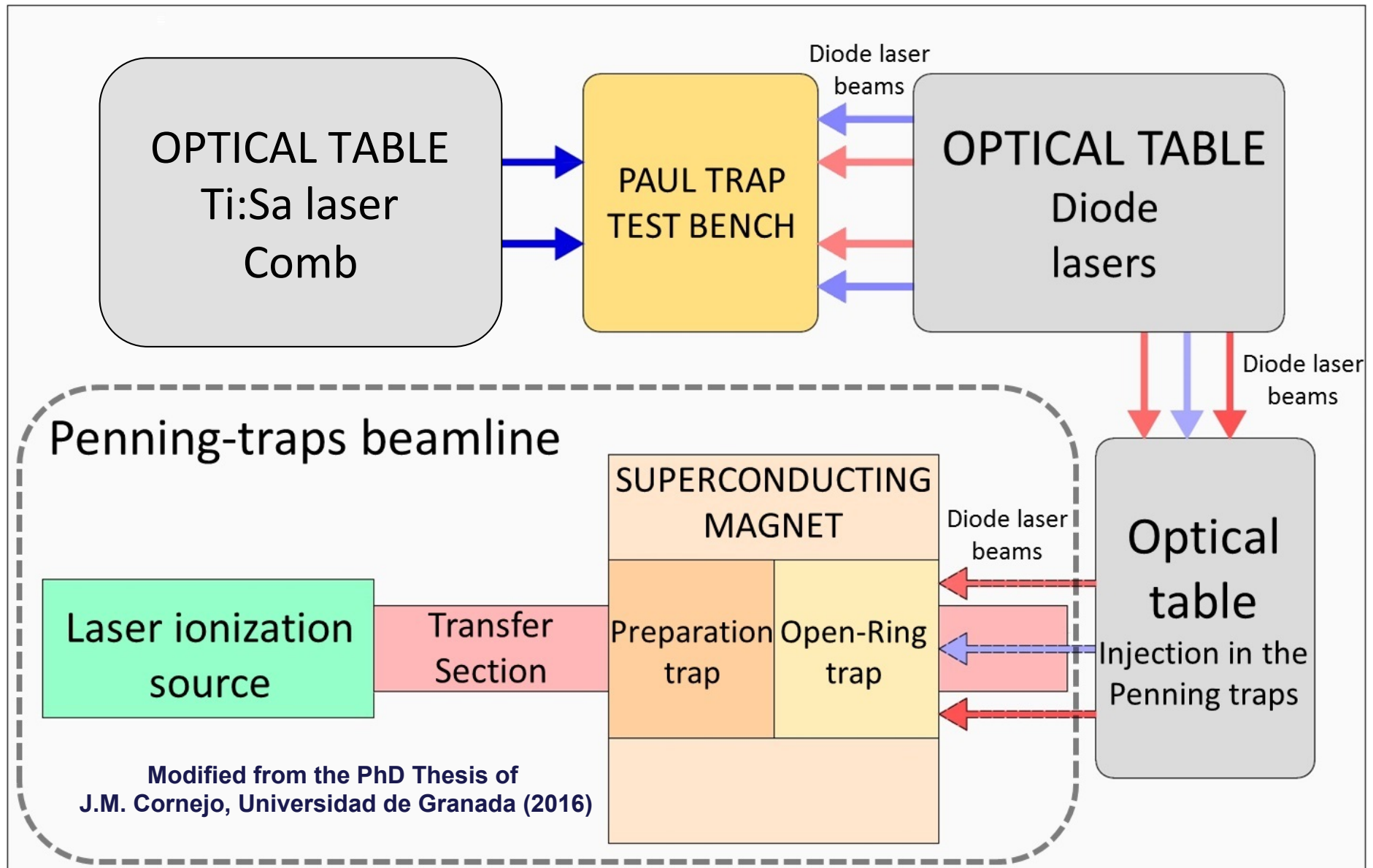
$$T(^{40}\text{Ca}^+) \approx 1 \text{ mK}$$

- Ion X^+ can be any ion regardless of its mass, charge and polarity.
- Very small oscillation amplitudes can be detected → **highest sensitivity as sensor & highest precision.**

FOR SUPERHEAVY ELEMENTS AND/
OR NEUTRINO MASSES

The QS project

Layout (TRAPSENSOR laboratory)



The QS Project

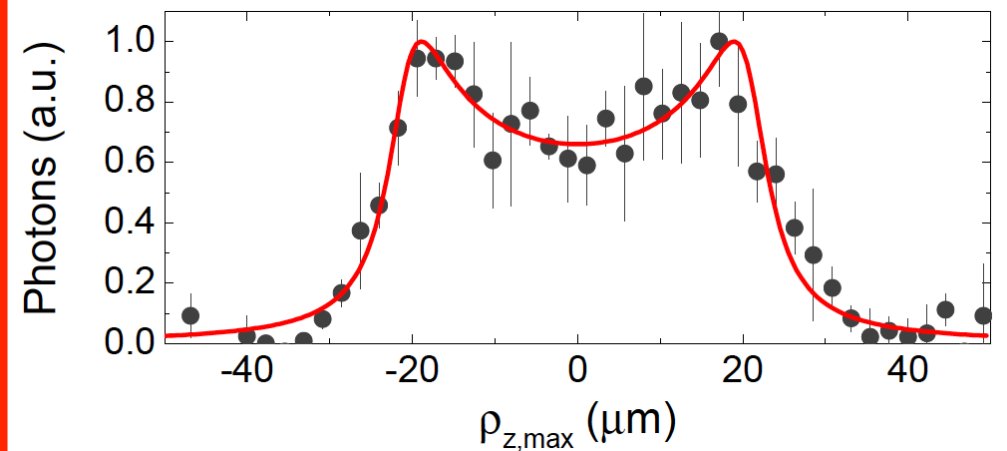
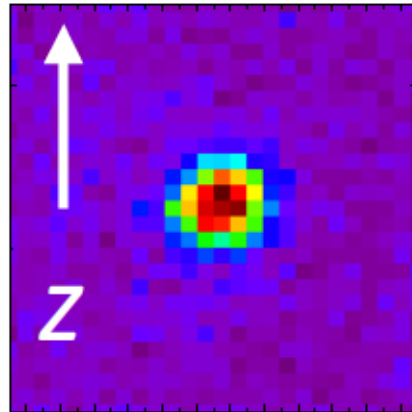
Single ion

F. Domínguez, *et al.*, Scientific Reports 7, 8336 (2017)

**Thermal
equilibrium**

$$\nu_z \approx 100 \text{ kHz}$$

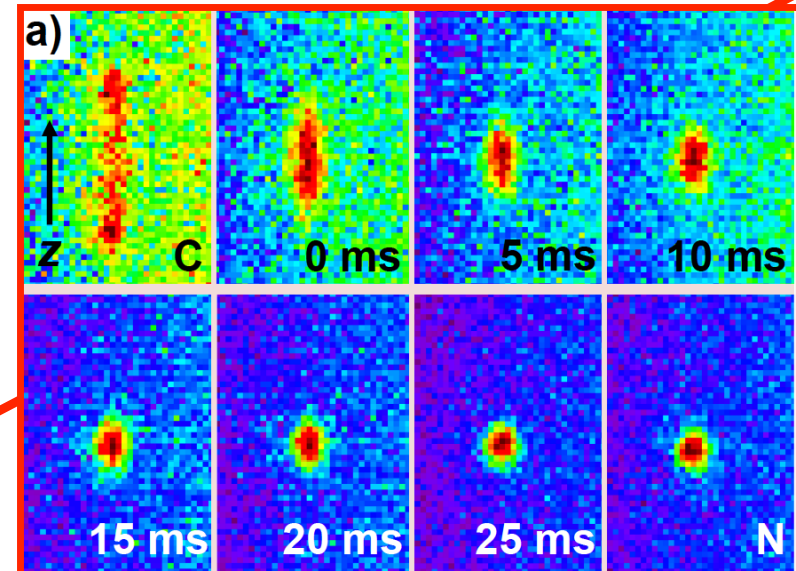
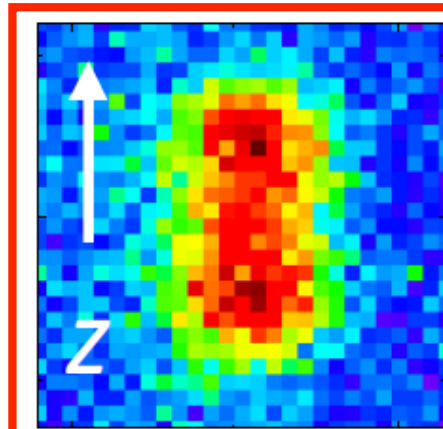
$$\rho_z^{\text{max}} \approx 2.6 \text{ } \mu\text{m}$$



**External
RF field**

$$\nu_z \approx 100 \text{ kHz}$$

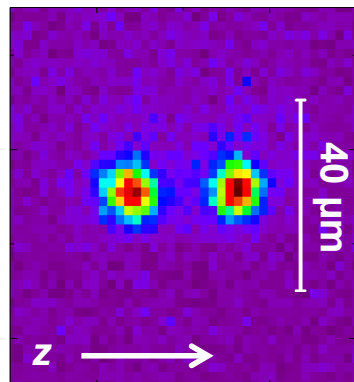
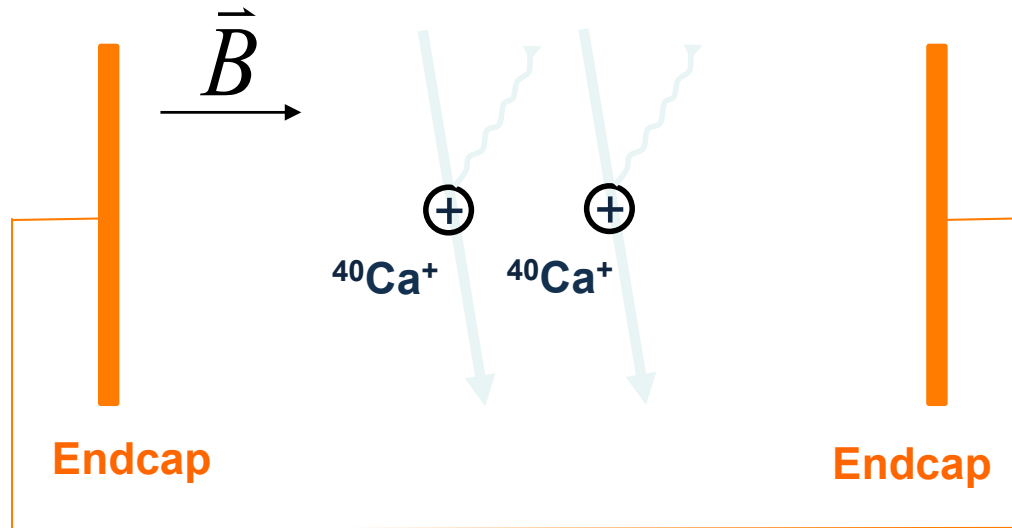
$$\rho_z^{\text{max}} \approx 21 \text{ } \mu\text{m}$$



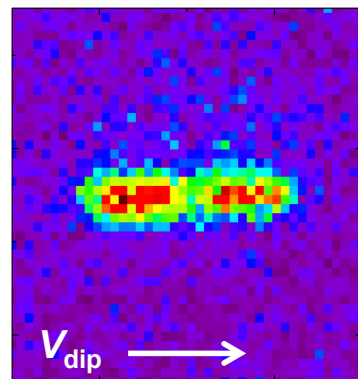
In a trap with characteristics
dimensions of about 1 cm.

The QS Project

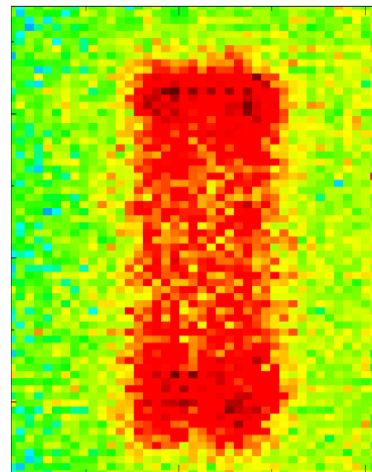
Two-ion crystal



No excitation



Axial excitation
Axial orientation



Axial excitation
Radial orientation

- Use different ions (e.g. $m/q = 187$) and monitor the ion of interest through the $^{40}\text{Ca}^+$ ion.

$M, M = \mu m$ → Axial frequency shift
 $\mu = 4.675$

$$\Omega_{\pm}^2 = \omega_{\text{trap}}^2 \left(1 + \frac{1}{\mu} \pm \sqrt{1 + \frac{1}{\mu^2} - \frac{1}{\mu}} \right)$$

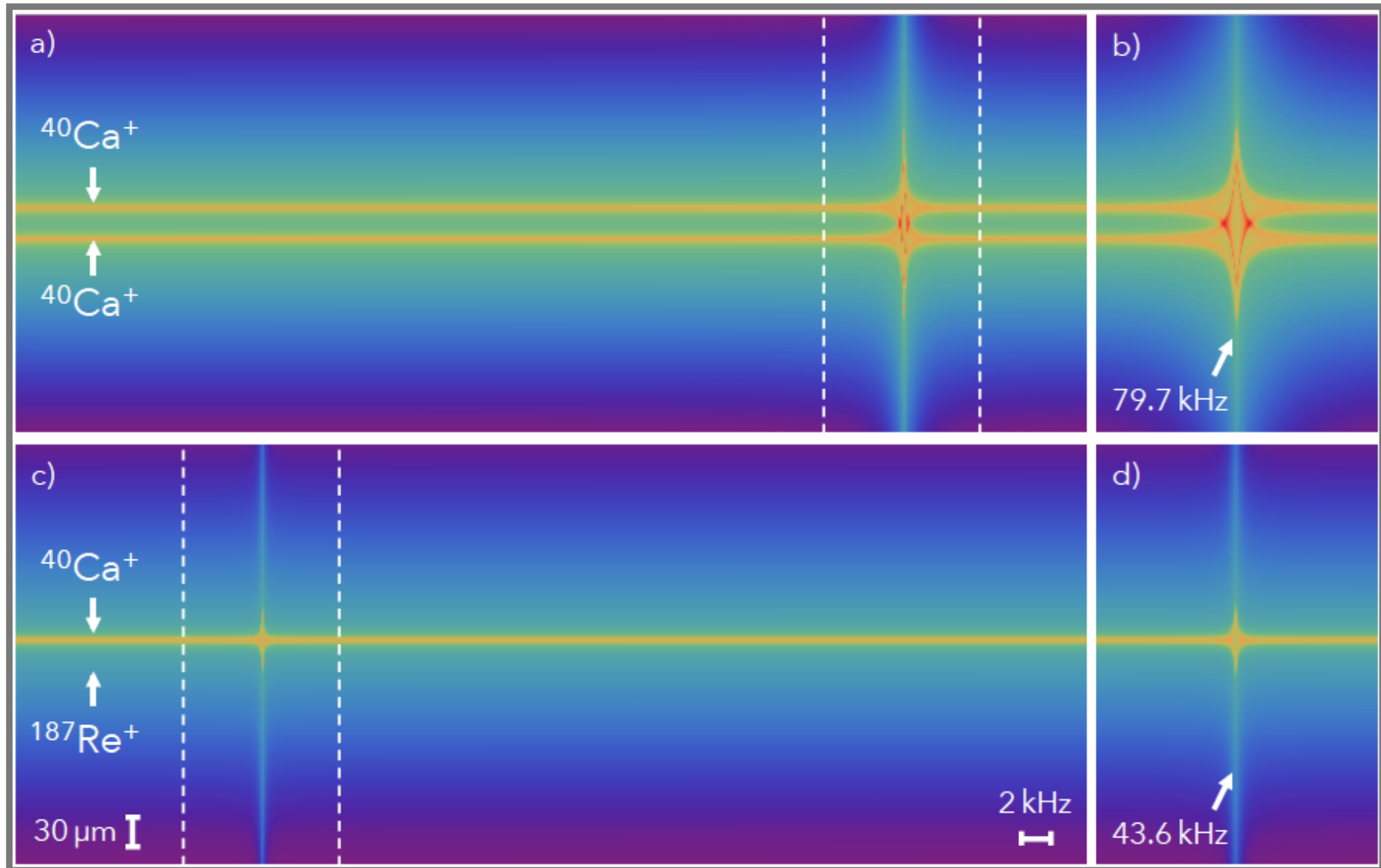
$$\rightarrow \Omega_- = 0.55 \times \omega_{\text{trap}}$$

G. Morigi and H. Walter, *Eur. Phys. J. D* 13(2) 261-269 (2001)

M. Drewsen *et al*, *Phys. Rev. Lett.* 93 243201 (2004)

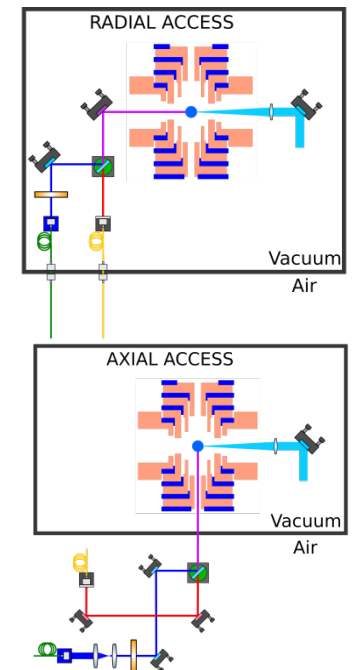
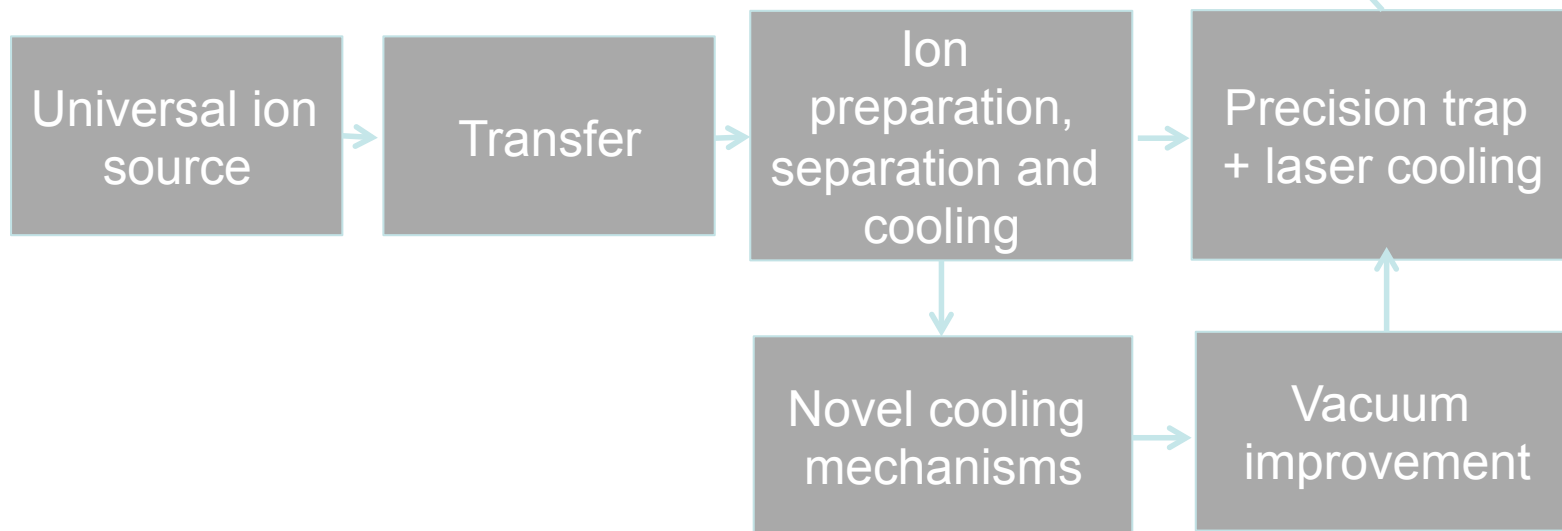
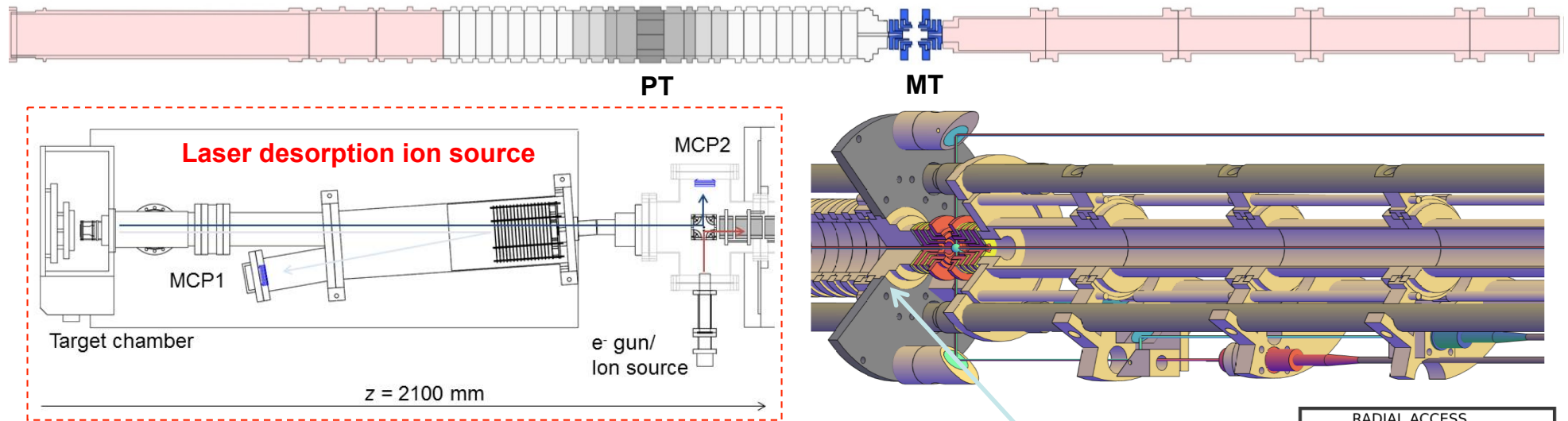
The QS Project

Two-ion crystal



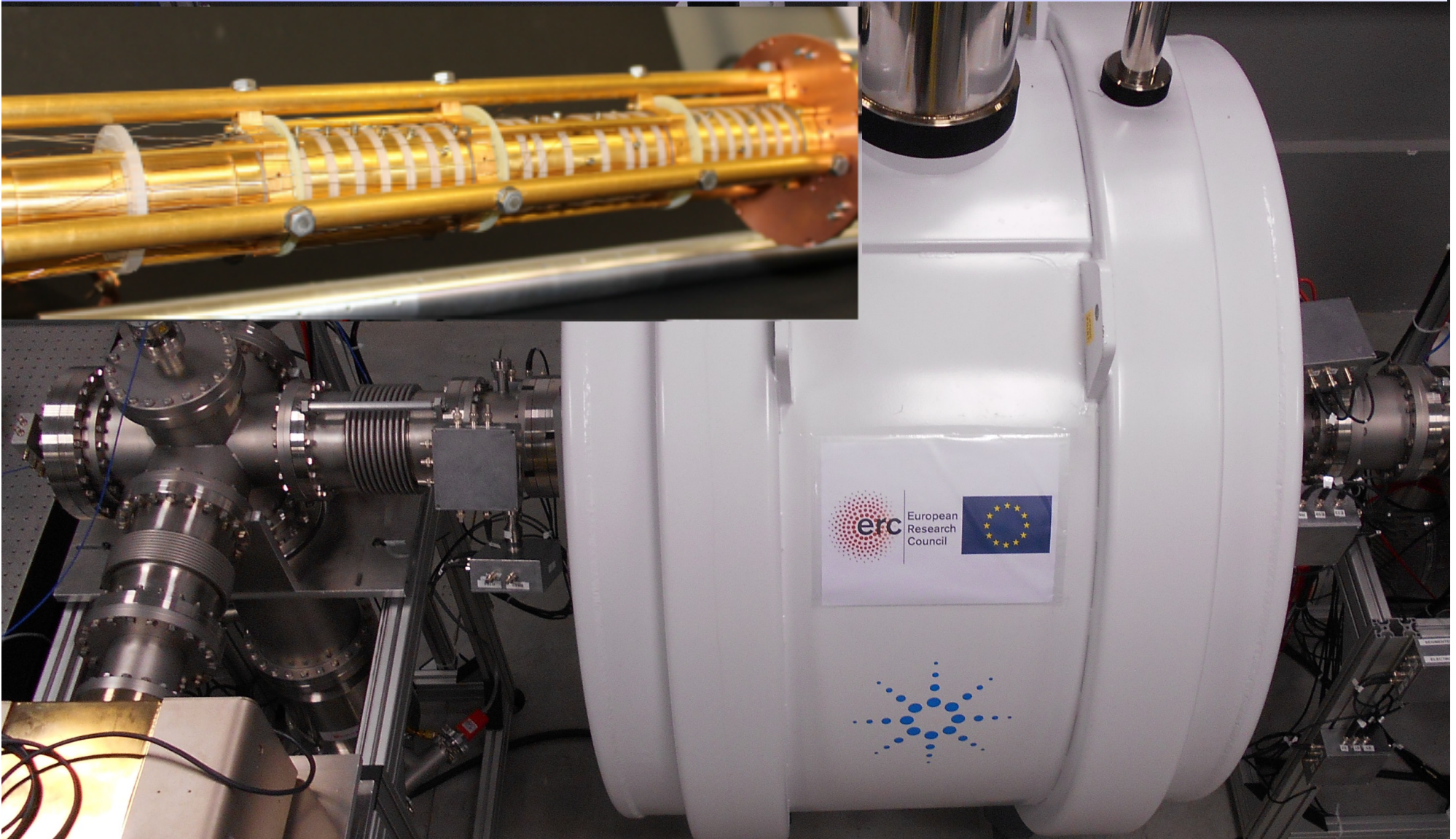
The QS Project

The Penning-traps beamline



The QS Project

The Penning-traps beamline



The QS Project

The Penning-traps beamline

