

Light neutrino mass spectrum with one or two right-handed singlet fermions added

Darius Jurčiukonis^a, Thomas Gajdosik^b, Andrius Juodagalvis^a

^aVilnius University, Institute of Theoretical Physics and Astronomy, A. Goštauto st. 12, LT-01108 Vilnius, Lithuania

^bVilnius University, Physics Faculty, Saulėtekio al. 9, LT-10222 Vilnius, Lithuania

Abstract

We analyse two cases of the minimal extension of the Standard Model when one or two right-handed fields are added to the three left-handed fields. A second Higgs doublet (two Higgs doublet model – 2HDM) is included in our model. We calculate one-loop radiative corrections to the mass parameters which produce mass terms for the neutral leptons. In both cases we numerically analyse light neutrino masses as functions of the heavy neutrino masses. Parameters of the model are varied to find light neutrino masses that are compatible with experimental data of solar Δm_{\odot}^2 and atmospheric Δm_{atm}^2 neutrino mass differences for normal hierarchy. We choose values for the parameters of the tree-level by numerical scans, where we look for the best agreement between computed and experimental neutrino oscillation angles.

Keywords:

Neutrino, seesaw mechanism, Higgs doublet

1. Description of the model

The mass terms for the neutrinos can be written in a compact form as a mass term with a $(n_L + n_R) \times (n_L + n_R)$ symmetric mass matrix

$$M_\nu = \begin{pmatrix} 0 & M_D^T \\ M_D & \hat{M}_R \end{pmatrix}, \quad (1)$$

where M_D is a $n_L \times n_R$ Dirac neutrino mass matrix, while the hat indicates that \hat{M}_R is a diagonal matrix. M_ν can be diagonalized as

$$U^T M_\nu U = \hat{m} = \text{diag}(m_1, m_2, \dots, m_{n_L+n_R}), \quad (2)$$

where the m_i are real and non-negative. In order to implement the seesaw mechanism [1] we assume that the elements of M_D are of order m_D and those of M_R are of order m_R , with $m_D \ll m_R$. Then, the neutrino masses m_i with $i = 1, 2, \dots, n_L$ are of order m_D^2/m_R , while those with $i = n_L + 1, \dots, n_L + n_R$ are of order m_R .

In the standard seesaw, one-loop corrections to the mass matrix, i.e. the self energies, are determined by

the neutrino interactions with the Z boson, the neutral Goldstone boson G^0 , and the Higgs boson h^0 . Each diagram contains a divergent piece but the sum of these three contributions turns out to be finite.

Once the one-loop corrections are taken into account the neutral fermion mass matrix is given by

$$M_\nu^{(1)} = \begin{pmatrix} \delta M_L & M_D^T + \delta M_D^T \\ M_D + \delta M_D & \hat{M}_R + \delta M_R \end{pmatrix} \approx \begin{pmatrix} \delta M_L & M_D^T \\ M_D & \hat{M}_R \end{pmatrix} \quad (3)$$

where the $0_{3 \times 3}$ matrix appearing at tree level (1) is replaced by the contribution δM_L . This correction is a symmetric matrix, it has the largest influence as compared to other corrections.

The expression for one-loop corrections is given by [2]

$$\delta M_L = \sum_b \frac{1}{32\pi^2} \Delta_b^T U_R^* \hat{m} \left(\frac{\hat{m}^2}{m_{H_b^0}^2} - \mathbb{1} \right)^{-1} \ln \left(\frac{\hat{m}^2}{m_{H_b^0}^2} \right) U_R^\dagger \Delta_b + \frac{3g^2}{64\pi^2 m_W^2} M_D^T U_R^* \hat{m} \left(\frac{\hat{m}^2}{m_Z^2} - \mathbb{1} \right)^{-1} \ln \left(\frac{\hat{m}^2}{m_Z^2} \right) U_R^\dagger M_D, \quad (4)$$

with $\Delta_b = \sum_k b_k \Delta_k$, where b_k are two-dimensional complex unit vectors, Δ_k are the Yukawa matrices, and the sum \sum_b runs over all neutral physical Higgses H_b^0 .

Neutral Higgses are characterized by the unit-length vectors b which hold special orthogonality conditions. We use the set of values

$$b_Z = \begin{pmatrix} i \\ 0 \end{pmatrix}, b_1 = \begin{pmatrix} 1 \\ 0 \end{pmatrix}, b_2 = \begin{pmatrix} 0 \\ 1 \end{pmatrix}, b_3 = \begin{pmatrix} 0 \\ i \end{pmatrix}. \quad (5)$$

It coincides with a special case of the invariant mixing angles that define the diagonalization matrix of the neutral Higgs squared-mass matrix [3]. The Higgs potential of the 2HDM becomes CP-conserving in this case.

2. Case $n_R = 1$

First we consider the minimal extension of the standard model by adding only one right-handed field ν_R to the three left-handed fields contained in ν_L .

We define $\Delta_i = (\sqrt{2} m_D / v) \vec{a}_i^T$, where $\vec{a}_1^T = (0, 0, 1)$ and $\vec{a}_2^T = (0, n, e^{i\phi} \sqrt{1-n^2})$. Diagonalization of the symmetric mass matrix M_ν (1) in block form is

$$U^T M_\nu U = U^T \begin{pmatrix} 0_{3 \times 3} & m_D \vec{a}_1 \\ m_D \vec{a}_1^T & \hat{M}_R \end{pmatrix} U = \begin{pmatrix} \hat{M}_l & 0 \\ 0 & \hat{M}_h \end{pmatrix}. \quad (6)$$

The non zero masses in \hat{M}_l and \hat{M}_h are determined analytically by finding eigenvalues of the hermitian matrix $M_\nu M_\nu^\dagger$. These eigenvalues are the squares of the masses of the neutrinos $\hat{M}_l = \text{diag}(0, 0, m_l)$ and $\hat{M}_h = m_h$. Solutions $m_D^2 = m_h m_l$ and $m_R^2 = (m_h - m_l)^2 \approx m_h^2$ correspond to the seesaw mechanism.

It is possible to estimate masses of the light neutrinos from experimental data of solar and atmospheric neutrino oscillations [4] assuming that the lightest neutrino has initial mass value m_{in} . Numerically varying the value of m_{in} we choose parameters for the tree level that agree with the experimental oscillation data most accurately.

For the calculation of radiative corrections we use the orthogonal complex vectors b defined in eq. (5). Diagonalization of the mass matrix corrected by one-loop contributions is performed with a unitary matrix $U_{\text{loop}} = U_{\text{egv}} U_\varphi(\varphi_1, \varphi_2, \varphi_3)$, where U_{egv} is an eigenmatrix of $M_\nu^{(1)} M_\nu^{(1)\dagger}$, and U_φ is a phase matrix. The second light neutrino obtains its mass from radiative corrections. The third light neutrino remains massless.

The numerical analysis shows (see Fig. 1) that we can reach the allowed neutrino mass ranges for a heavy singlet with the mass close to 10^4 GeV when the angle of oscillations θ_{atm} is fixed to the experimental 3σ range [4].

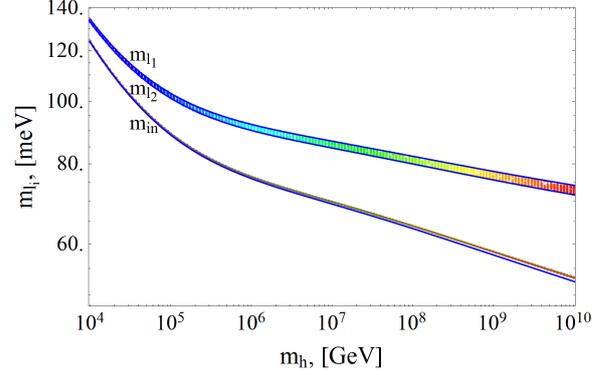


Figure 1: (Color online) Calculated masses of two light neutrinos as a function of the heavy neutrino mass m_h . Values of m_{l1} (m_{l2}) are shown as a band (brown line). The blue solid line represents the value of m_{in} used for the calculations of the tree level masses.

The free parameters n , ϕ , $m_{H_2^0}$ and $m_{H_3^0}$ are restricted by the parametrization used and by the oscillation data. Figure 2 illustrates the allowed values of Higgs masses for different values of the heavy singlet. The values of Higgs masses spread to two separated sets.

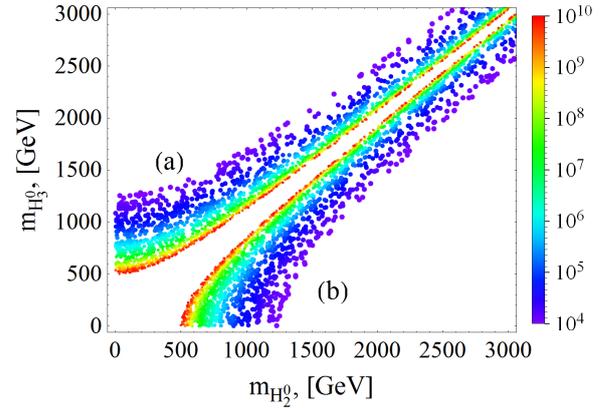


Figure 2: (Color online) The values of the free parameters $m_{H_2^0}$ and $m_{H_3^0}$ as functions of the heaviest right-handed neutrino mass m_h . The scale of the m_h values is shown on the right. The mass of the SM Higgs boson is fixed to $m_{H_1^0} = 125$ GeV.

3. Case $n_R = 2$

When we add two singlet fields ν_R to the three left-handed fields ν_L , the radiative corrections give masses to all three light neutrinos.

We parametrize $\Delta_i = \frac{\sqrt{2}}{v} (m_{D_2} \vec{a}_i^T, m_{D_1} \vec{b}_i^T)^T$ with $|\vec{a}| = 1$, and $|\vec{b}| = 1$. Diagonalizing the symmetric mass matrix

M_ν (1) in block form we write:

$$U^T M_\nu U = U^T \begin{pmatrix} 0_{3 \times 3} & m_{D_2} \vec{a} & m_{D_1} \vec{b} \\ m_{D_2} \vec{a}^T & \hat{M}_l & 0 \\ m_{D_1} \vec{b}^T & 0 & \hat{M}_h \end{pmatrix} U = \begin{pmatrix} \hat{M}_l & 0 \\ 0 & \hat{M}_h \end{pmatrix}. \quad (7)$$

The non zero masses in \hat{M}_l and \hat{M}_h are determined by the seesaw mechanism: $m_{D_i}^2 \approx m_{h_i} m_{l_i}$ and $m_{R_i}^2 \approx m_{h_i}^2$, $i = 1, 2$. We use $m_1 > m_2 > m_3$ ordering of masses. At tree level the third light neutrino is massless.

In numerical calculations the model parameters as well as the derived masses of the light neutrinos are obtained in several steps. First, the diagonal mass matrix for the tree level is constructed. Having chosen the mass for the third neutrino m_{in} , the masses of the other two light neutrinos are estimated from experimental data on solar and atmospheric neutrino oscillations [4]. The masses of the heavy right-handed neutrinos are input parameters. This diagonal tree level matrix is used to constrain the parameters that enter the tree-level mass matrix M_ν and its diagonalization matrix. Then the diagonalization matrix is used to evaluate one-loop corrections to the mass matrix. Diagonalization of the corrected mass matrix yields masses for the three light neutrinos. If the calculated mass difference is compatible with the experimental oscillations data, the parameter set is kept. Otherwise, another set of parameters is generated. Figure 3 illustrates the obtained results for normal neutrino mass ordering.

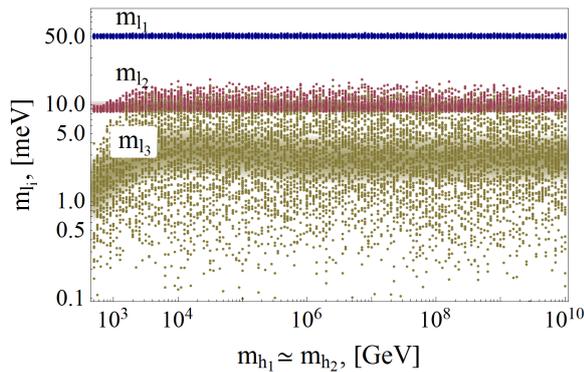


Figure 3: (Color online) The masses m_i of the light neutrinos as functions of the heaviest right-handed neutrino mass $m_{h_{1,2}}$ for the normal hierarchy of the light neutrinos. The wide solid lines indicate the area of the most frequent values of the scatter data.

Figure 4 illustrates the allowed values of Higgs masses for different values of the heavy singlet. The values of Higgs masses spread to two separated sets for the $n_R = 1$ case.

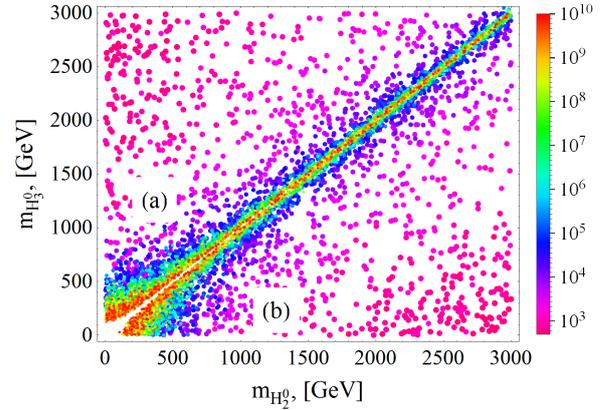


Figure 4: (Color online) The values of the free parameters $m_{H_2^0}$ and $m_{H_3^0}$ as functions of the heaviest right-handed neutrino mass $m_{h_{1,2}}$ for the normal hierarchy of the light neutrinos.

In the case of $n_R = 2$ the inverted hierarchy is also allowed [5].

4. Conclusions

1. For the case $n_R = 1$ we can match the differences of the calculated light neutrino masses to Δm_\odot^2 and Δm_{atm}^2 with the mass of a heavy singlet larger than 10^4 GeV. The parametrization used for this case and restrictions from the neutrino oscillation data limit the values of free parameters. One light neutrino remains massless.
2. In the case $n_R = 2$ we obtain three non vanishing masses. The numerical analysis shows that the values of parameters and Higgs masses depend on the choice of the set of b vectors and the heavy neutrino masses. The radiative corrections generate the lightest neutrino mass and have a big impact on the second lightest neutrino mass.

Acknowledgement

The authors thank the Lithuanian Academy of Sciences for the support (the project DaFi2014, No. CERN-VU-2014-1).

References

- [1] M. Gell-Mann, P. Ramond, and R. Slansky, in *Supergravity, Proceedings of the Workshop, Stony Brook, New York, 1979*, edited by F. van Nieuwenhuizen and D. Freedman (North Holland, Amsterdam, 1979).
- [2] W. Grimus and L. Lavoura, *Phys. Lett. B* **546** 86 (2002).
- [3] H.E. Haber and D. O’Neil, *Phys. Rev. D* **83** 055017 (2011).
- [4] D. V. Forero, M. Tortola and J. W. Valle, arXiv:1405.7540.
- [5] D. Jurčiukonis, T. Gajdosik and A. Juodagalvis, paper in progress.