



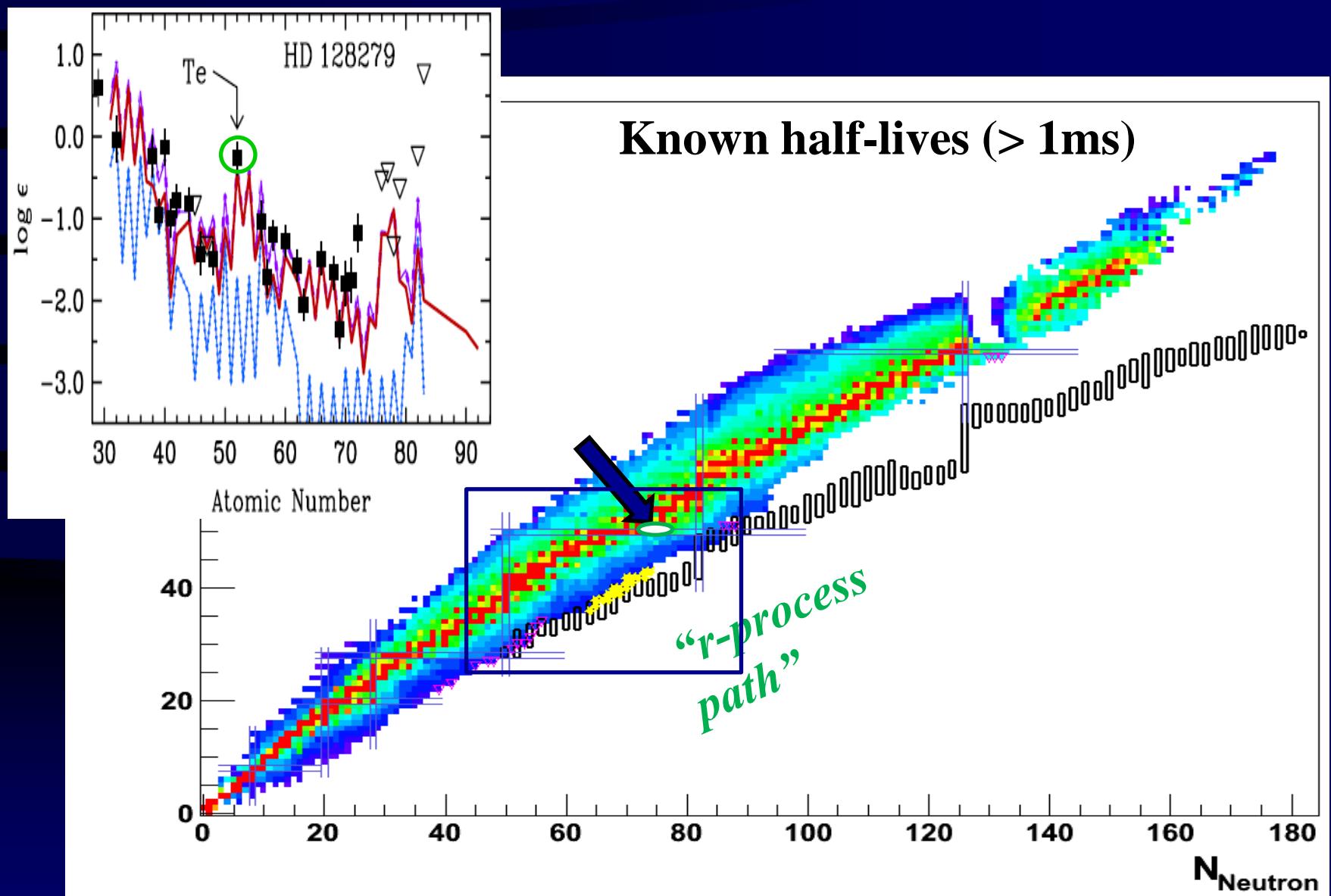
# Proposal RIKEN (I)

## Mass region A=110-125

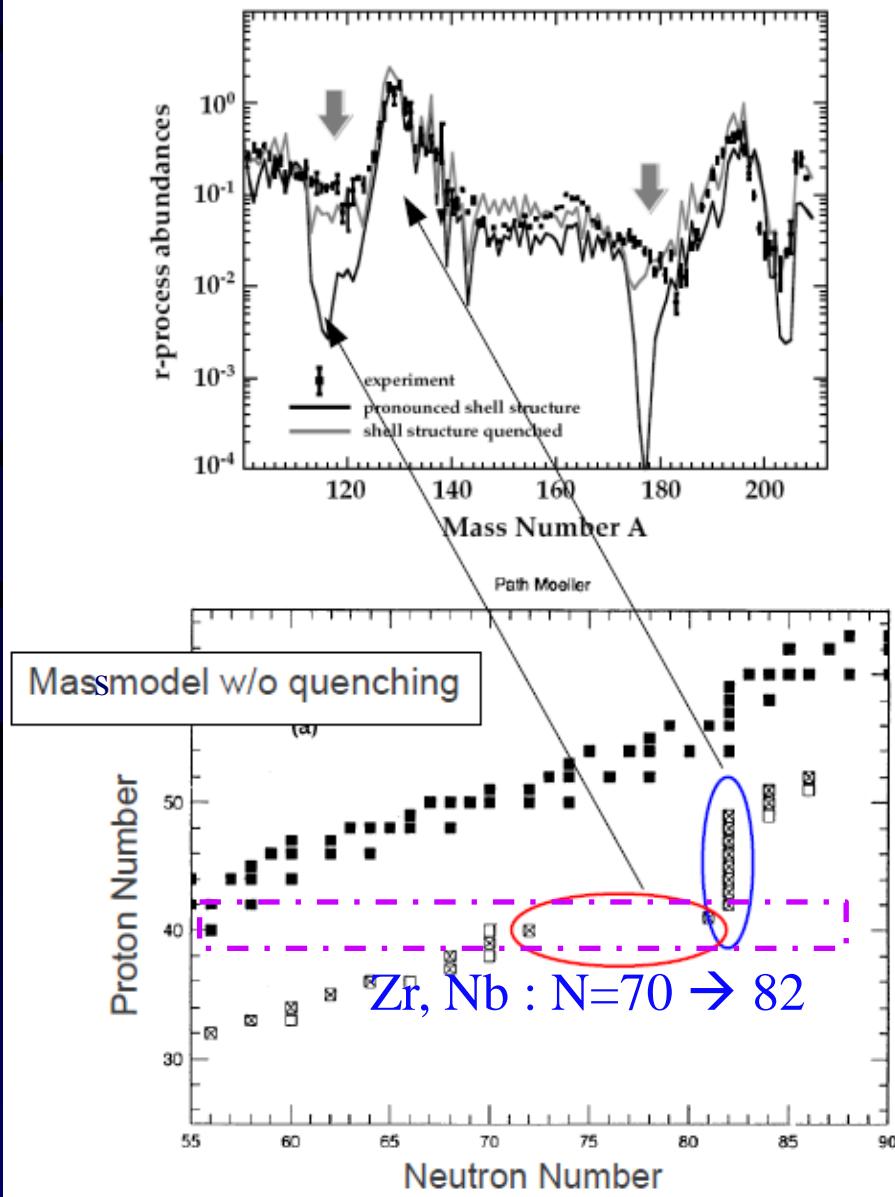
Shunji NISHIMURA

( RIKEN Nishina Center )

# Motivation



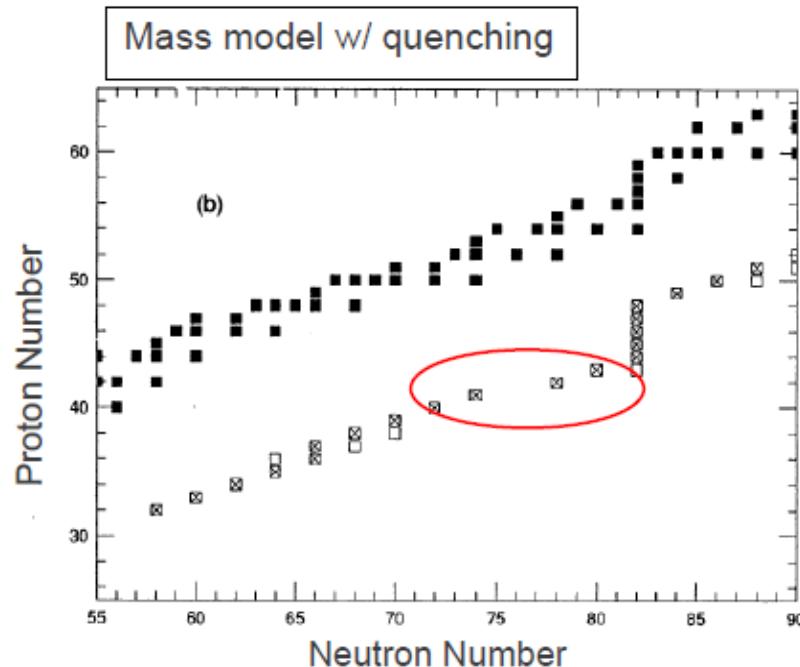
# 2<sup>nd</sup> r-process peak and path around Z = 38~46



Profound astrophysical impact of:

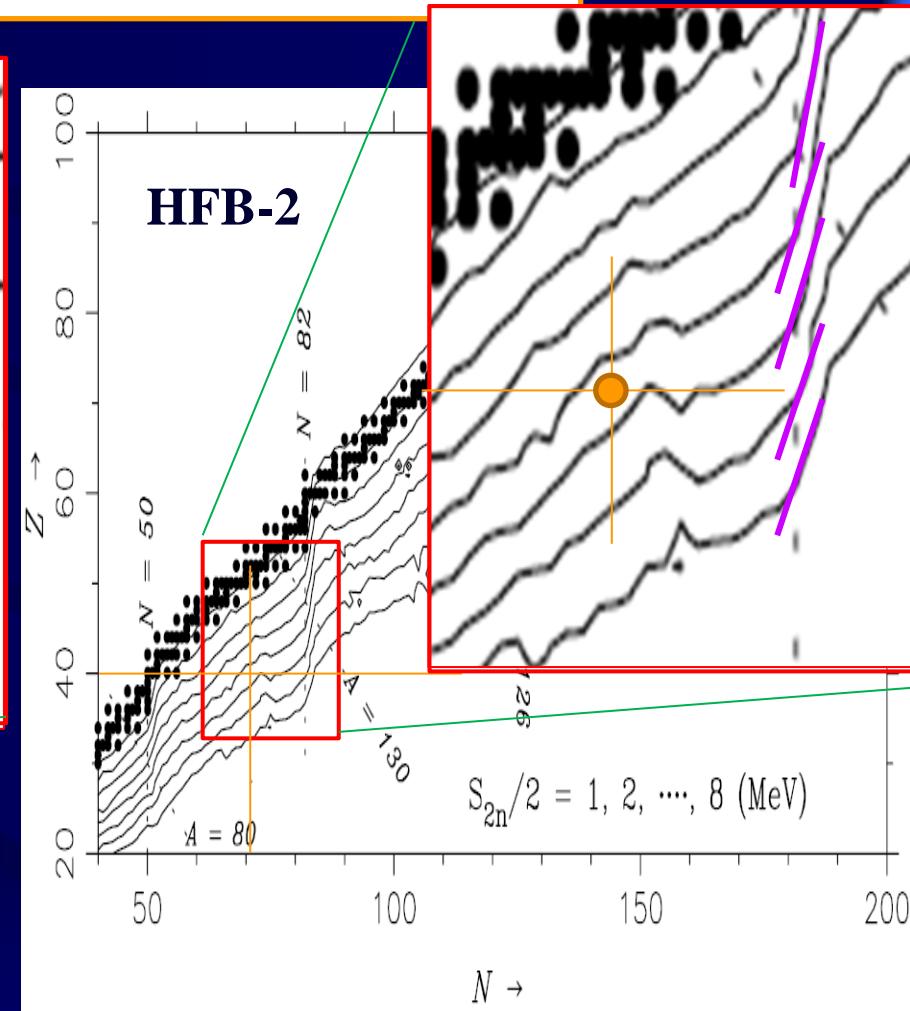
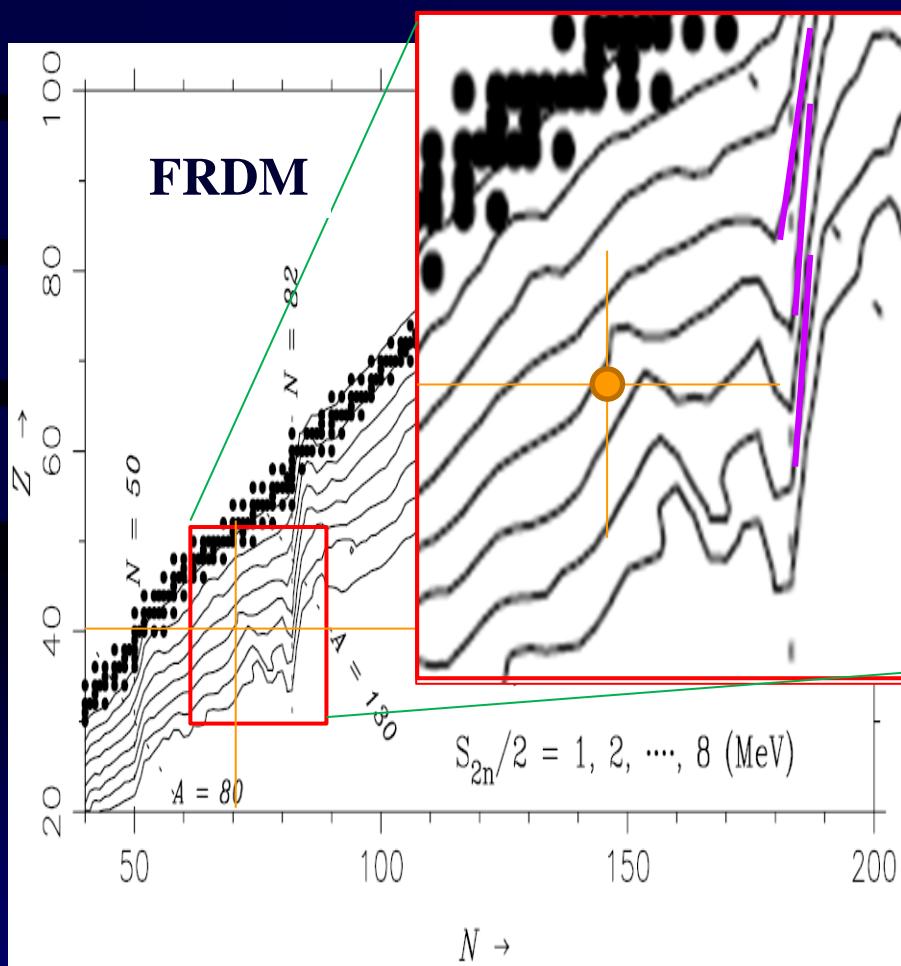
- Quenching of the N=82 shell gap  
(need to study region around  $^{128}\text{Pd}_{82}$ )
- appearance of the a N=70 sub-shell closure  
(need to study region around  $^{110}\text{Zr}_{70}$ )

K.-L.Kratz (1993)

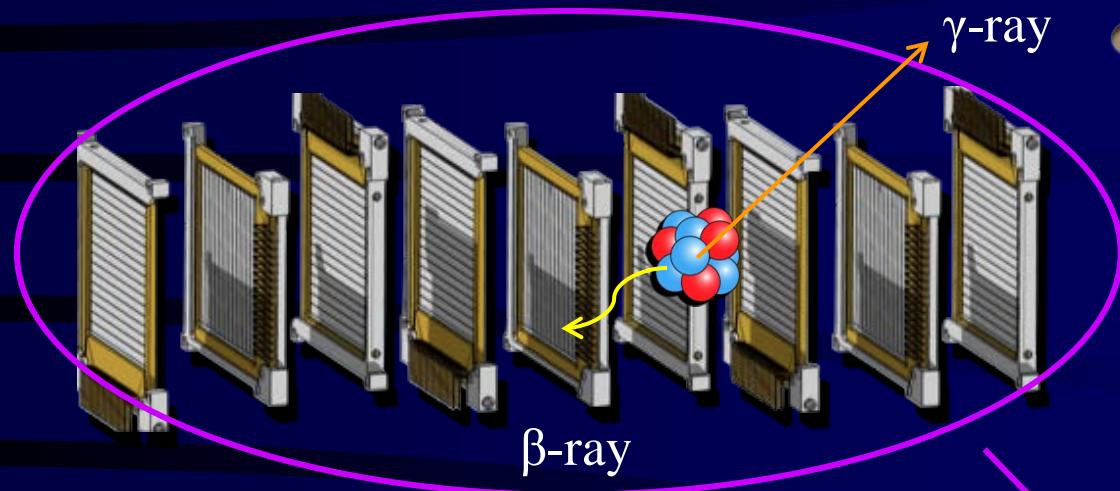


# Neutron separation energies around Mass A=110-125

**Location of r-process path depends on  $S_n$  (2 – 3 MeV)  
( $n,\gamma$ )-( $\gamma,n$ ) competition**

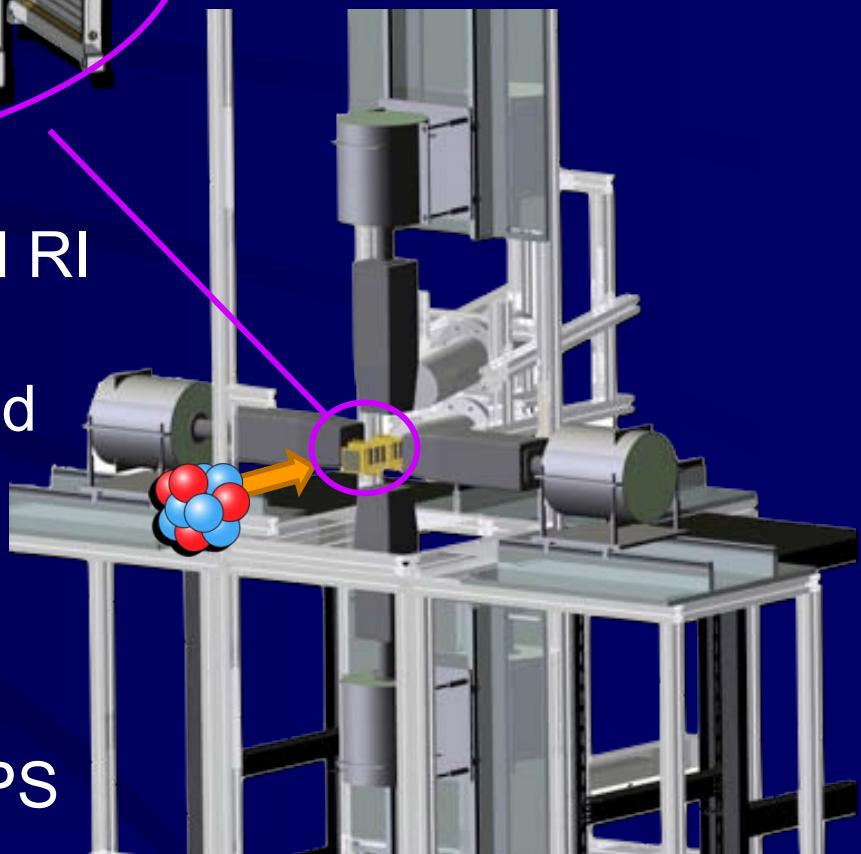


# Experimental Setup in 2009



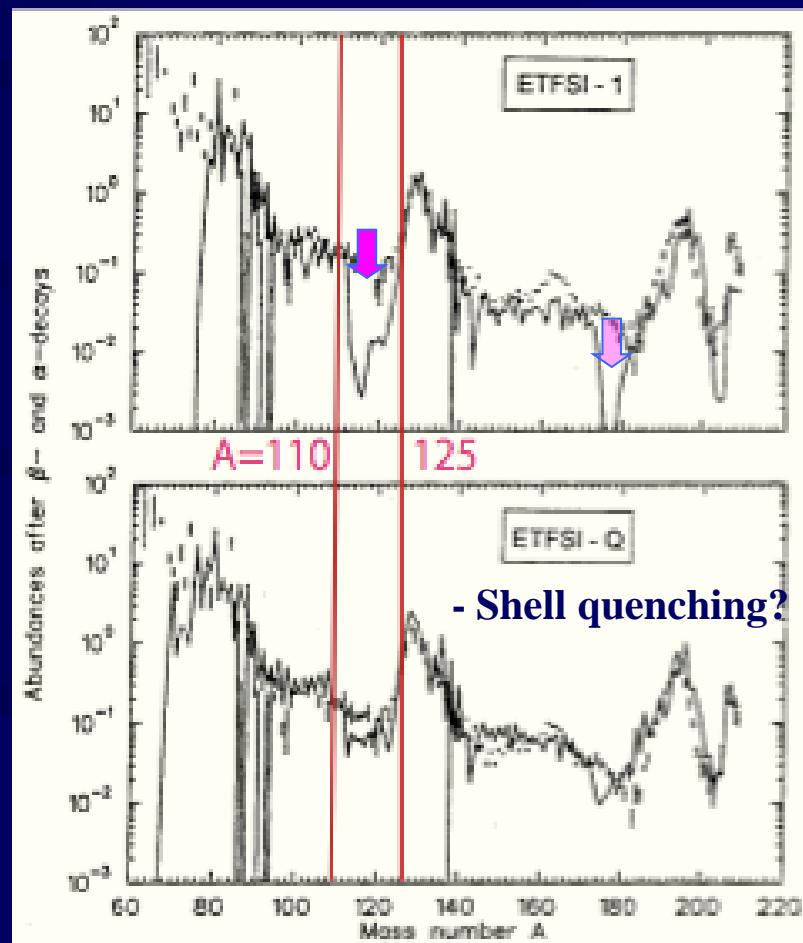
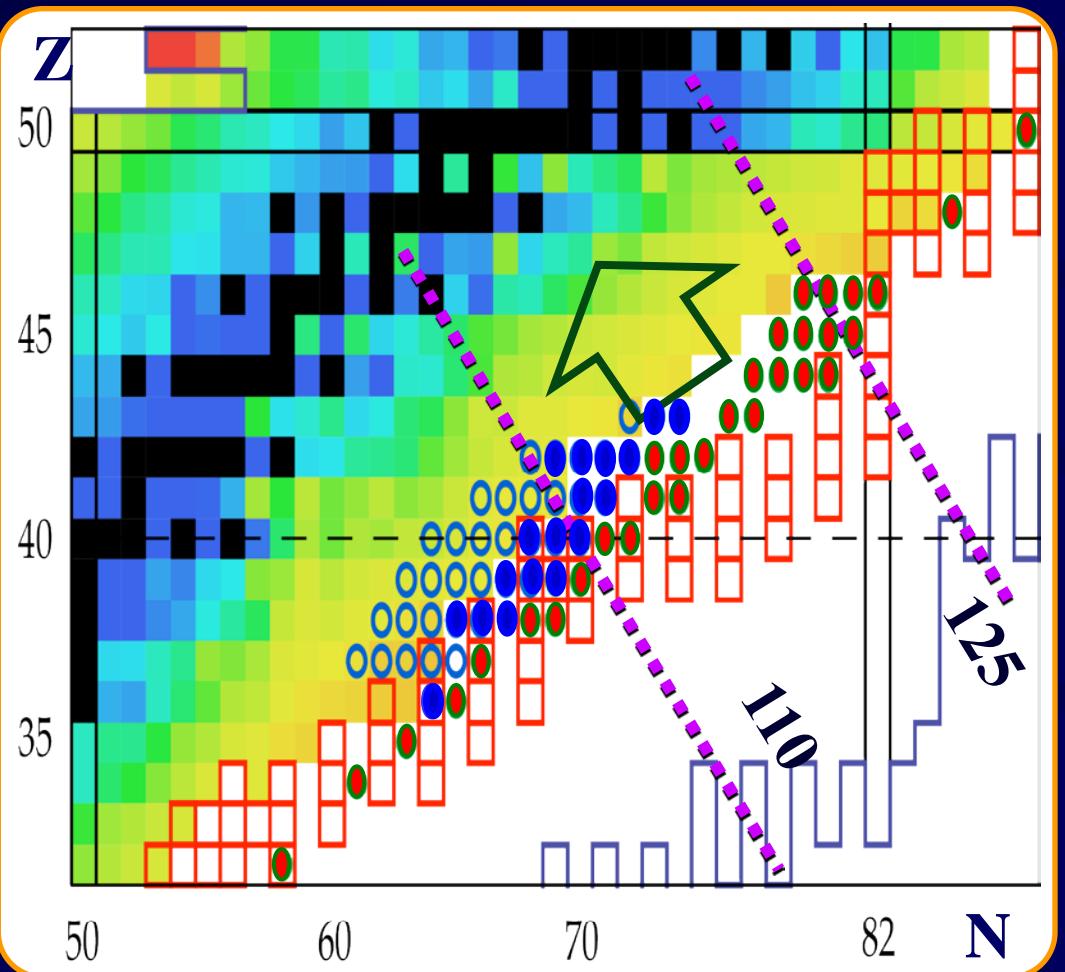
- RI &  $\beta$ -ray detection
  - 9 DSSDs ( $50 \times 50 \times 1 \text{ mm}^3$ )
  - $16 \times 16$  strips
  - $\sim 2000$  pixels in total

➤ The implantation of an identified RI is associated with the following  $\beta$ -decay events that are detected in the same DSSD pixel



➤  $\Delta E$ -TOF- $B_p$  method using the focal plane detectors in BigRIPS

# R-process Abundance around 2<sup>nd</sup> peak

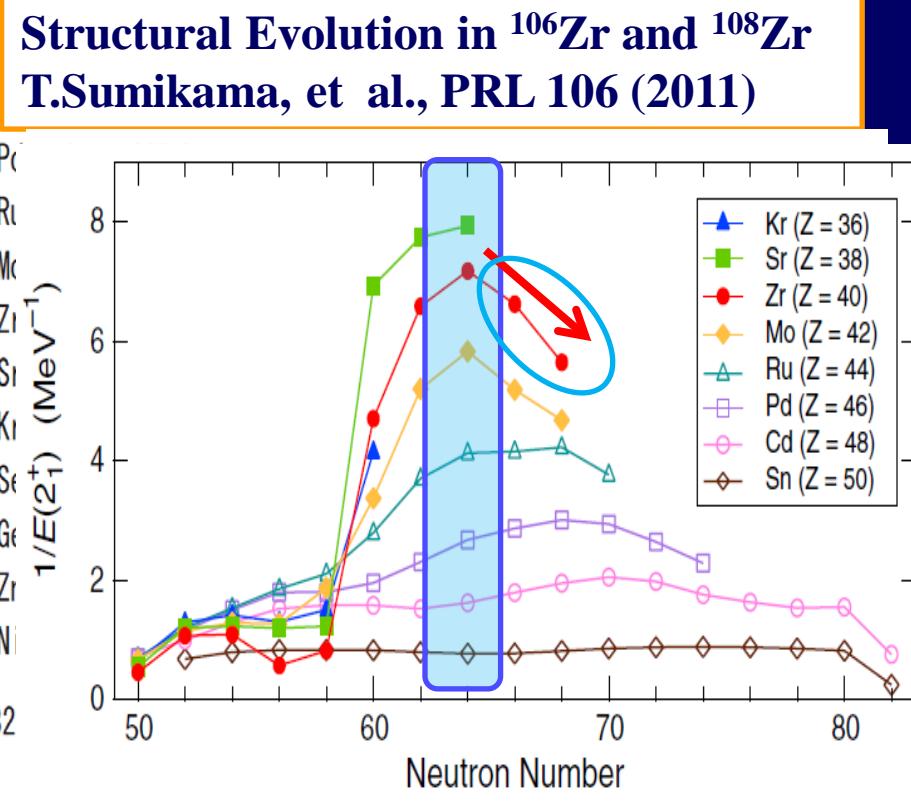
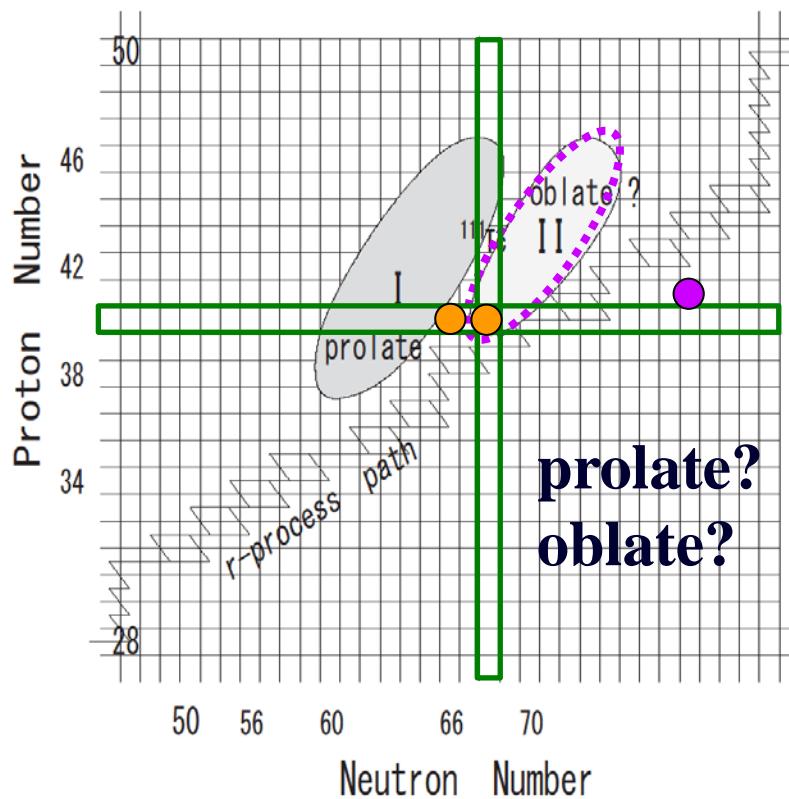


$$1/T_{1/2} = \sum_{\substack{E_i \leq Q_\beta \\ E_i \geq 0}} S_\beta(E_i) \times f(Z, Q_\beta - E_i);$$

$$f \sim (Q_\beta - E_i)^5$$

# Deformation

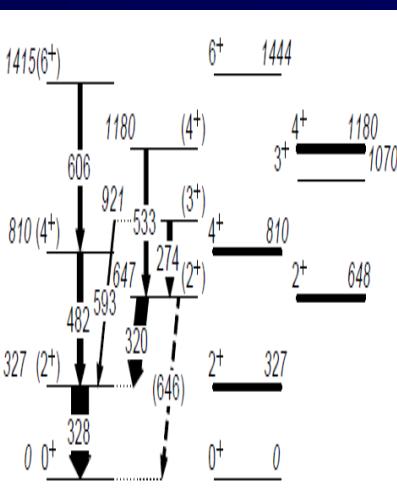
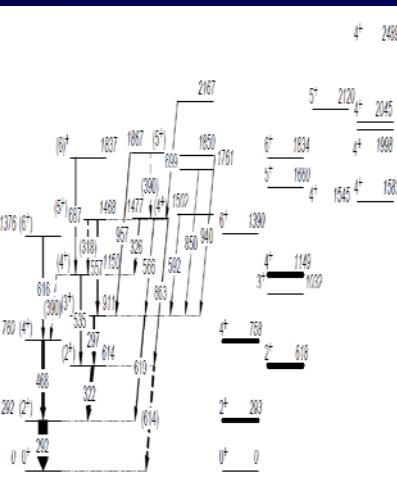
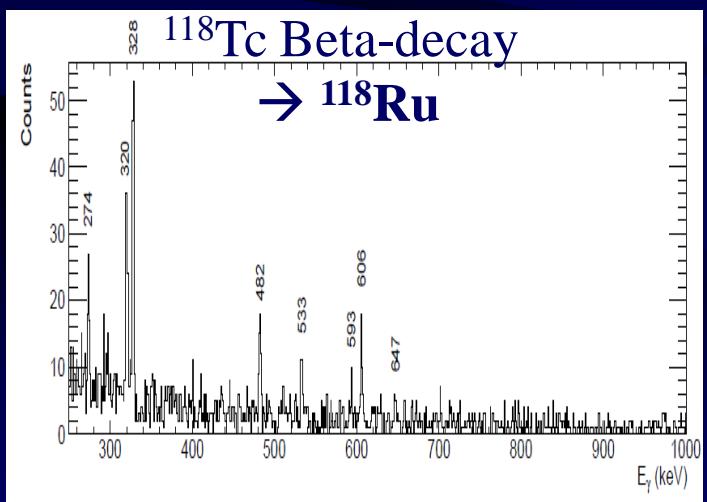
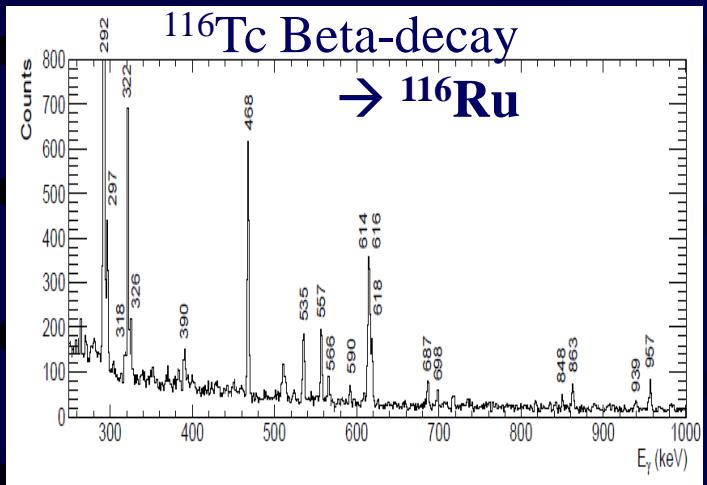
- Phys. Lett. B 696 (2011) H.Watanabe  
Oblate shape isomer for  $^{109}\text{Nb}$ ?



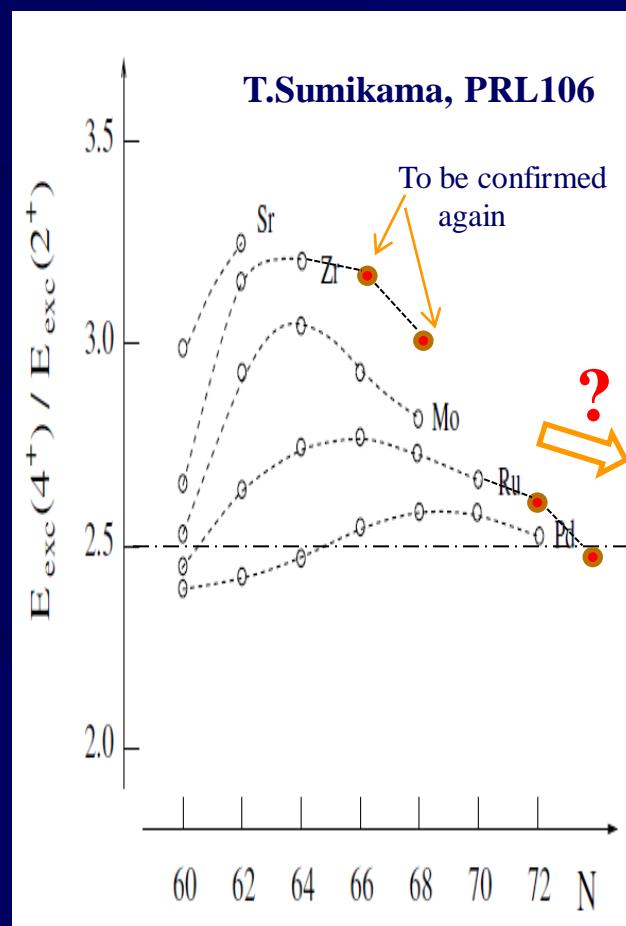
- Beta-delayed gamma :  $^{106}\text{Y} \rightarrow ^{106}\text{Zr}$
- Isomeric states :  $^{108}\text{Zr}$

# Shape evolution in $^{116, 118}\text{Ru}$ (EURICA)

Spokesperson: G.Lorusso,H,Watanabe

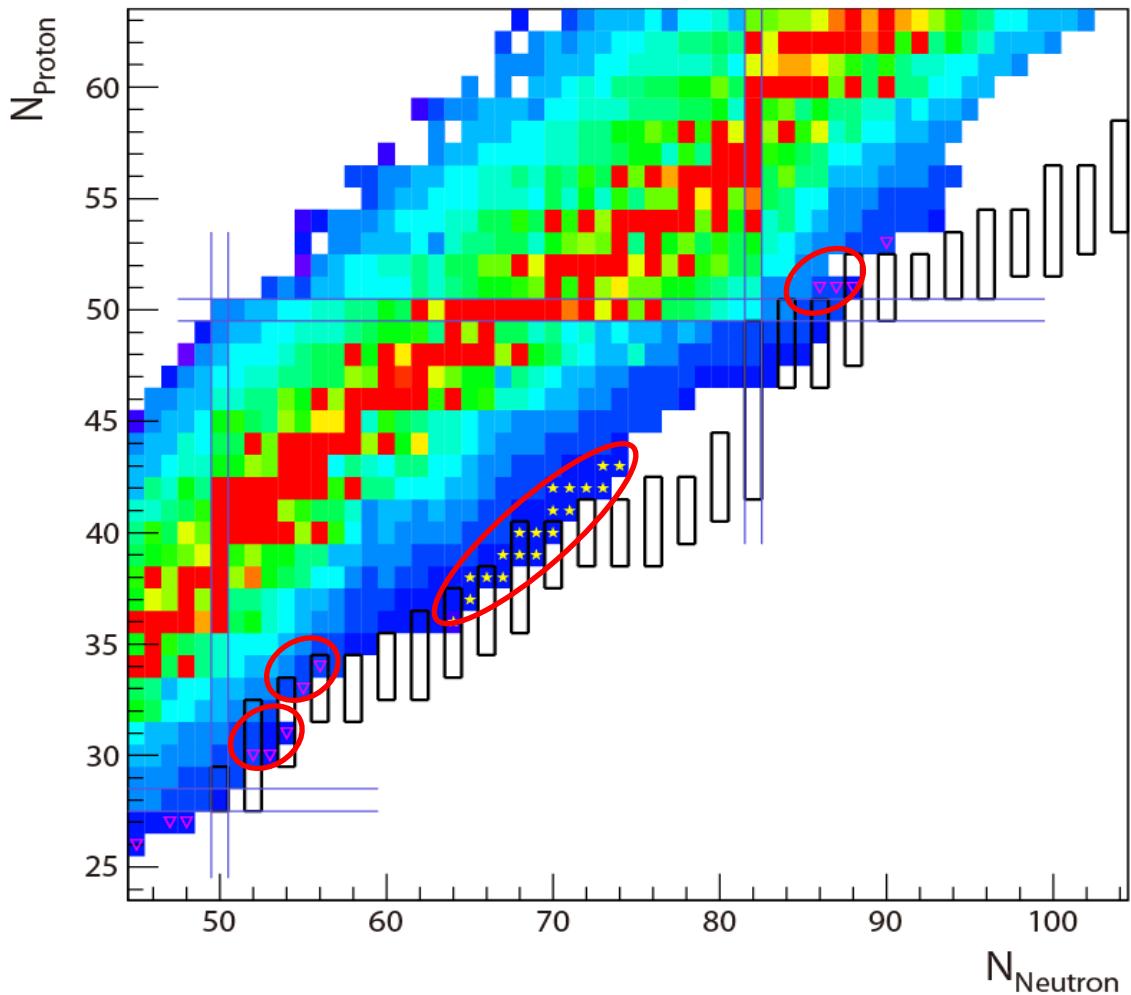


P.-A. Söderström, et al.  
Phys. Rev. C, accepted.



Analysis is in progress in other isotopes around  $^{110}\text{Zr}$  region

# Beta-decay half-lives since 2007



2011 .. O.Arndt,  
PRC 84,  
**3**  $T_{1/2}$  (ISOLDE)

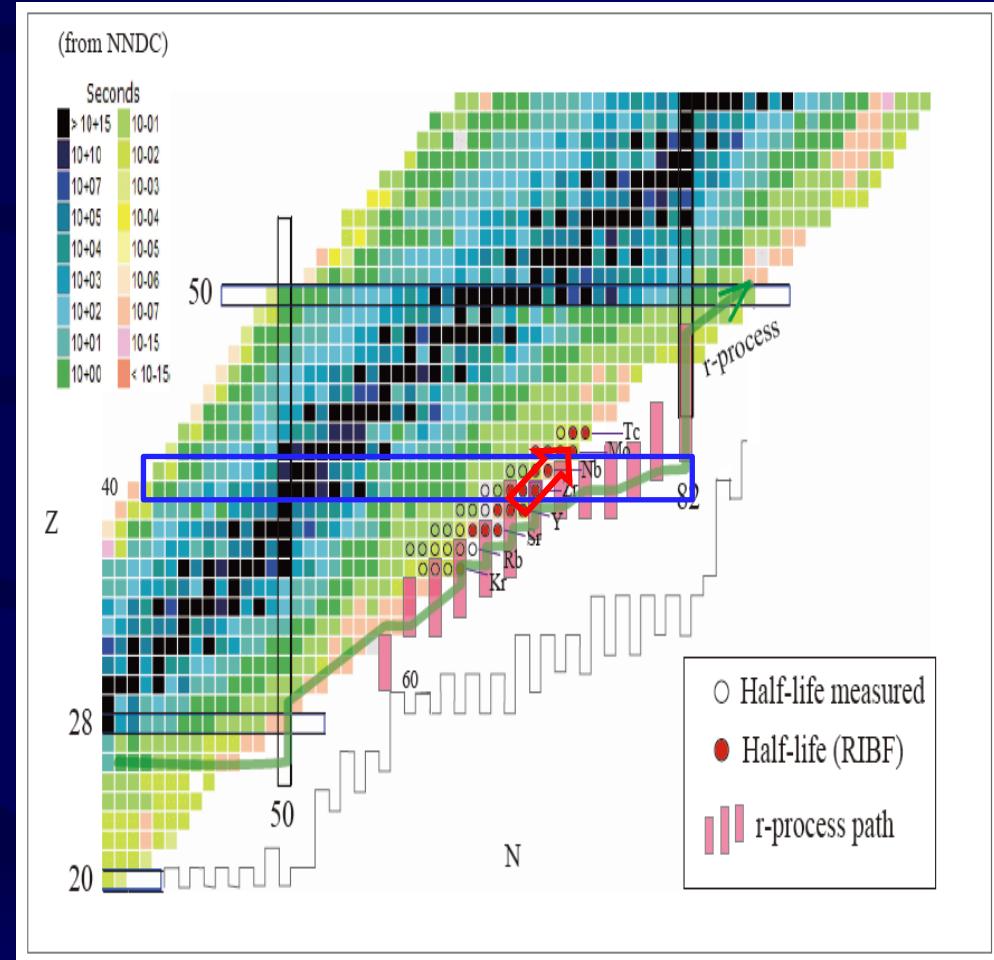
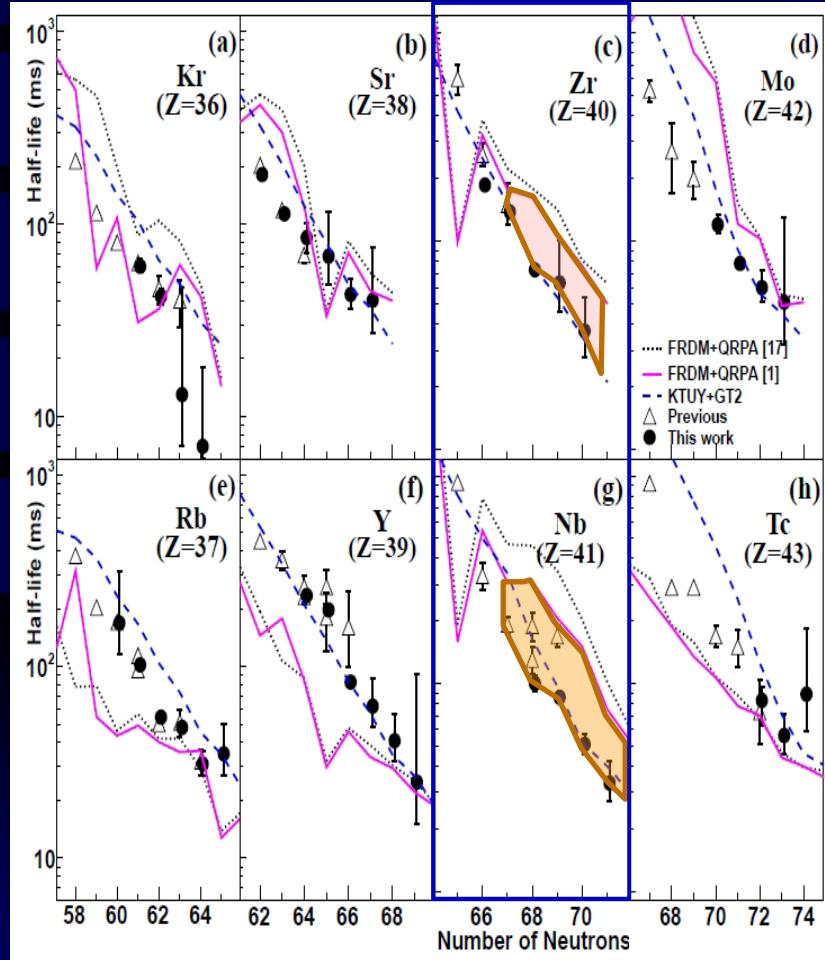
2011 .. SN,  
PRL 106,  
**18**  $T_{1/2}$  (RIBF)

2012 .. M.Quinn,  
PRC 85,  
**2**  $T_{1/2}$  (MSU)

2012 .. M.Madurga,  
PRL 109 ,  
**3**  $T_{1/2}$  (ORNL)

2013 ... C.Mazzocchi  
**1**  $T_{1/2}$  PRC (ORNL)

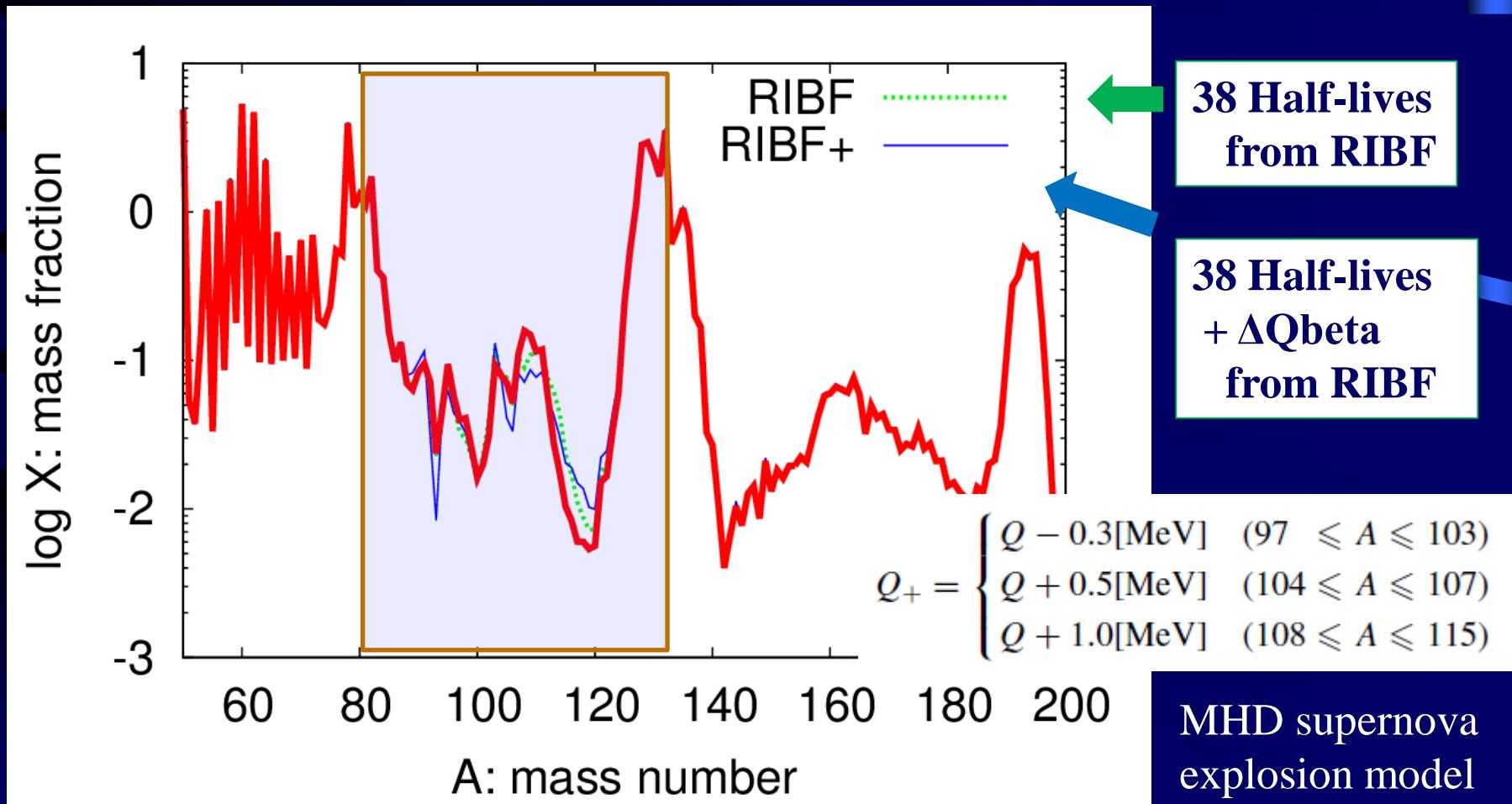
# Decay Experiment in 2009



Zr and Nb decay faster than expected by FRDM+QRPA ( $T_{1/2} : 1/2 \sim 1/3$ )

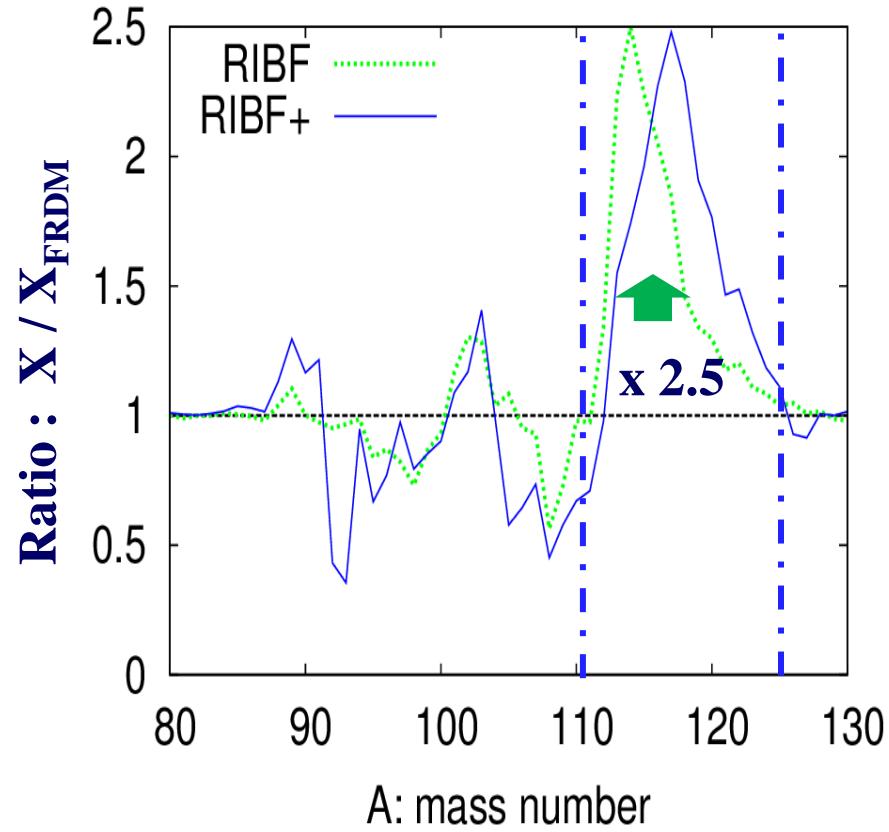
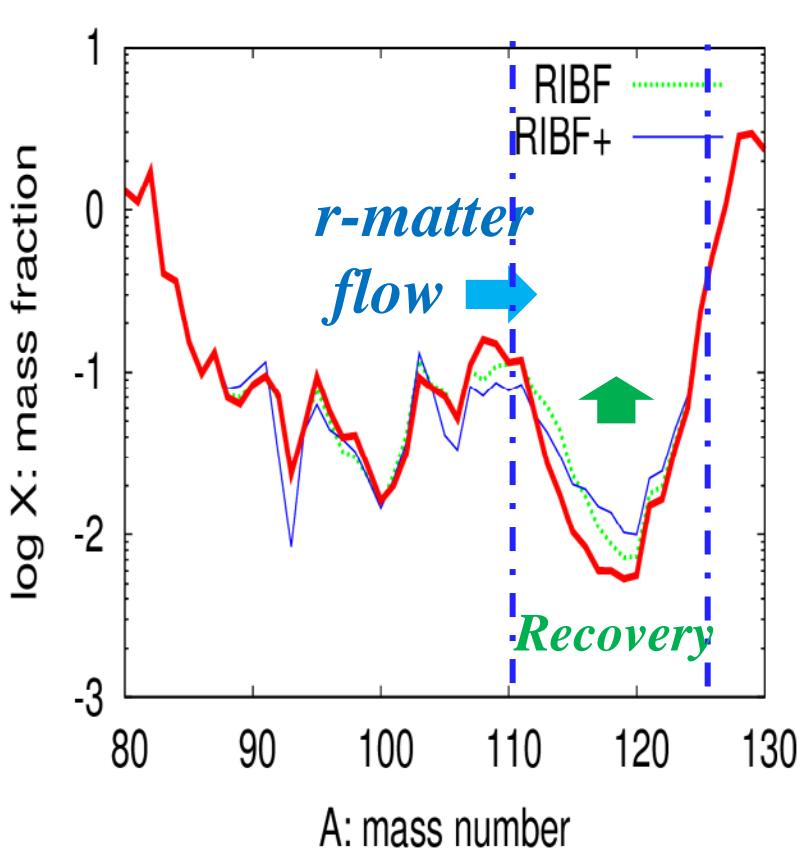
# RIBF data → Impact to r-process abundance

Nobuya Nishimura, T.Kajino, G.Mathew, SN, T.Suzuki, PRC 85 (2012)



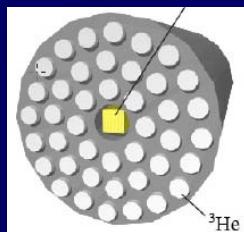
# RIBF data → Impact to r-process abundance

Nobuya Nishimura, T.Kajino, G.Mathew, SN, T.Suzuki, PRC 85 (2012)

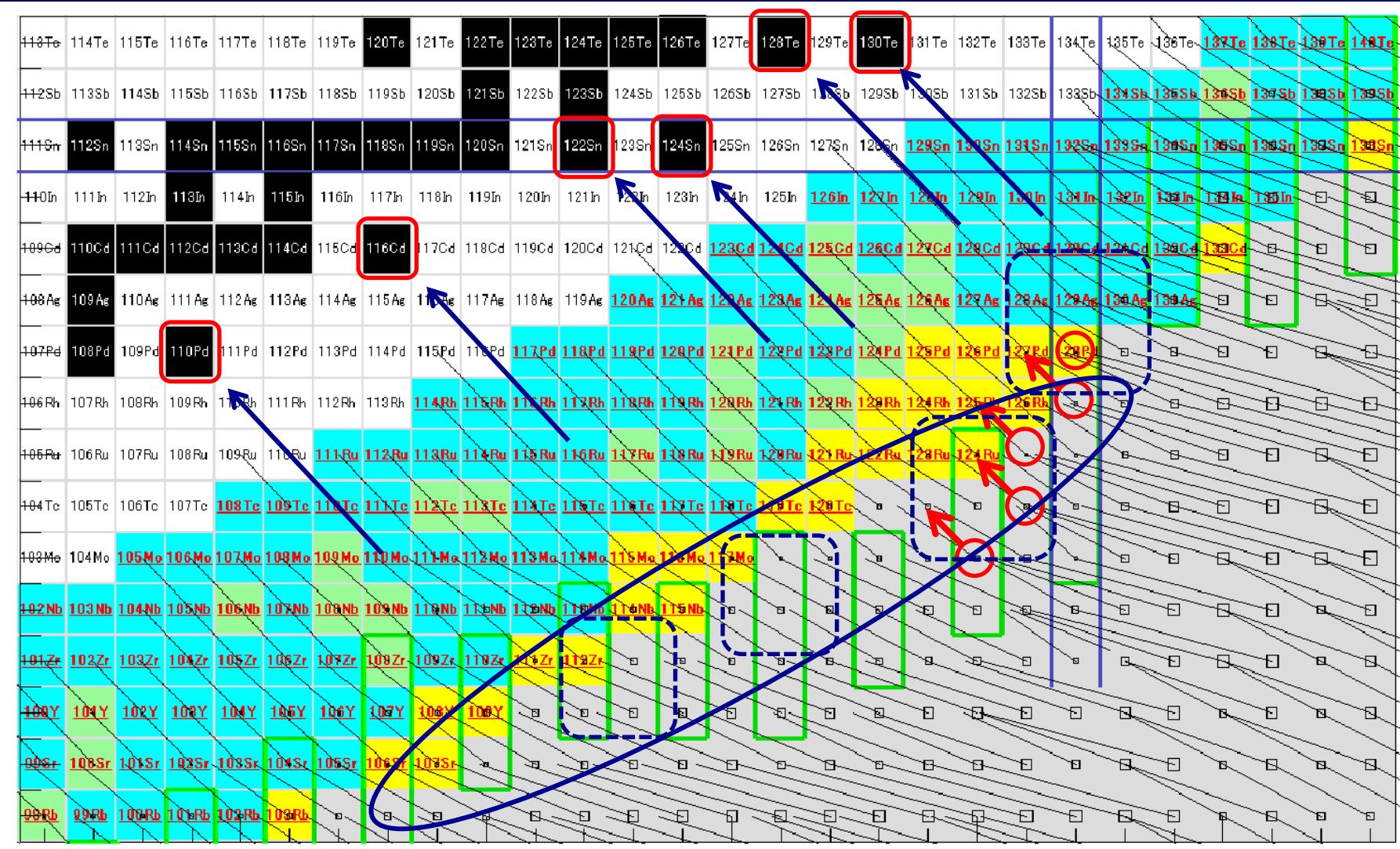


The calculated r-process abundance is improved by factor of x 2.5.  
But, there is still issue remaining in mass A=110 – 125!

# Expected Beta-decay Flow ( Beta-delayed Neutron Emission )

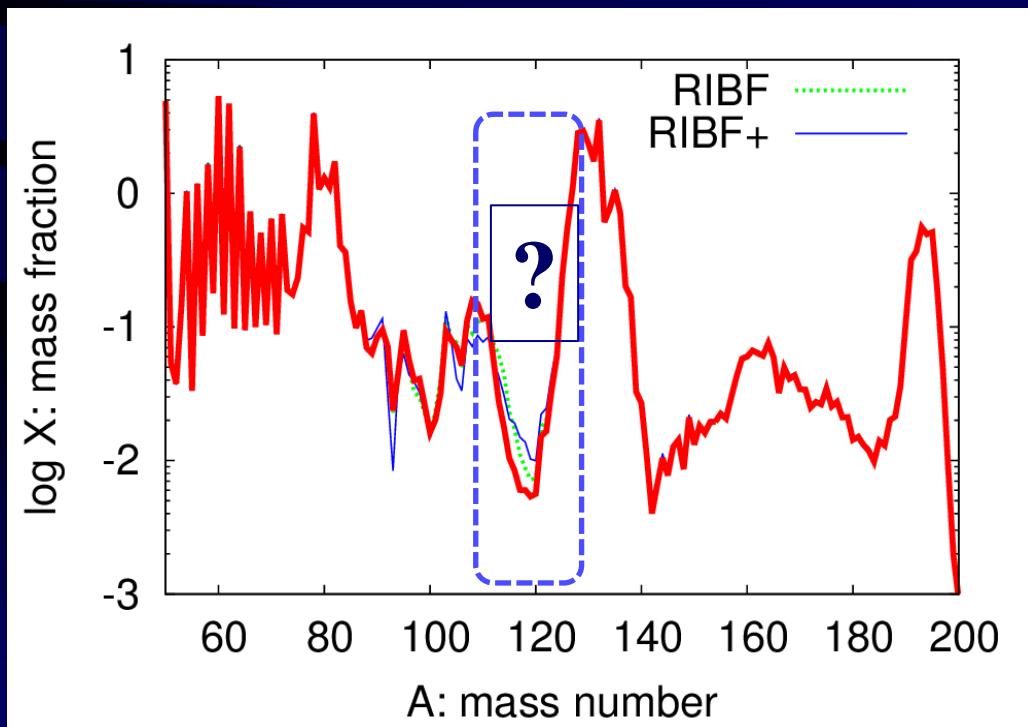


For example, FRDM+QRPA (P.Moller)



# Summary

- Mysteries of the r-process abundance around 2<sup>nd</sup> peak.
  - Mass A = 110-125
    - Influence of Larger / smaller neutron emission ?
    - Multi neutron emission Pn,P2n,.. ?
    - Mass ? and Half-lives?



Do we have problems in  
Nuclear theory?  
or  
Nuclear astrophysics  
( or both... )