

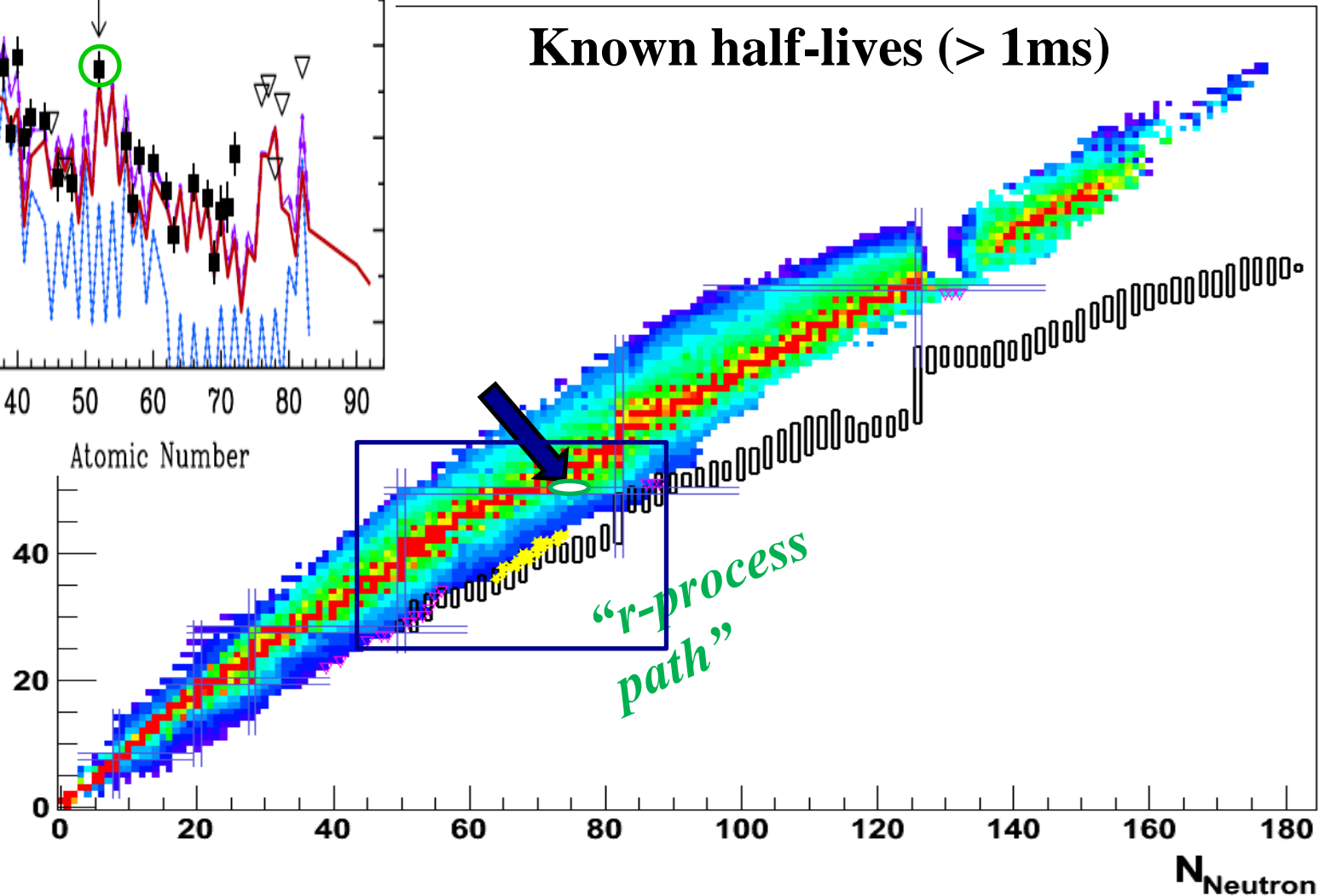
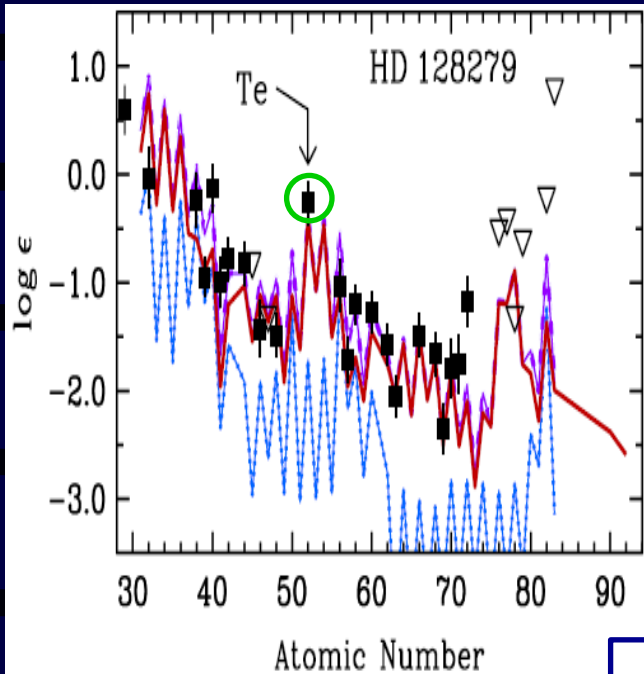
# **Proposal RIKEN (I)**

## **Mass region $A=110-125$**

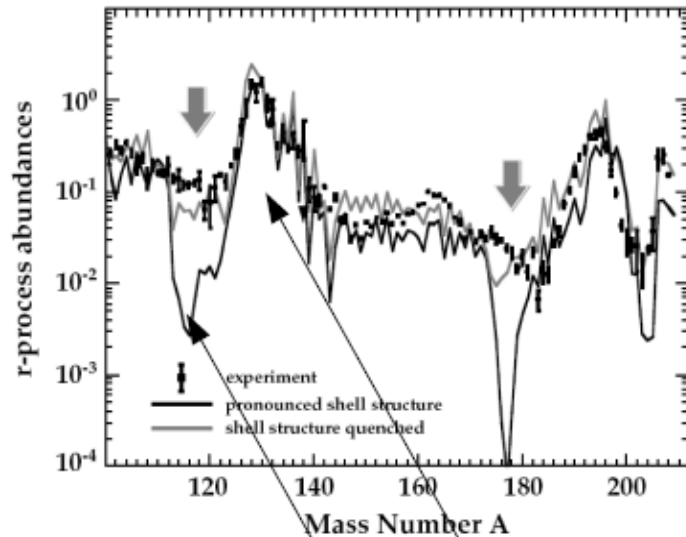
**Shunji NISHIMURA**  
**( RIKEN Nishina Center )**

# Motivation

I.U.Roederer, ApJ 747, L8 (2012)



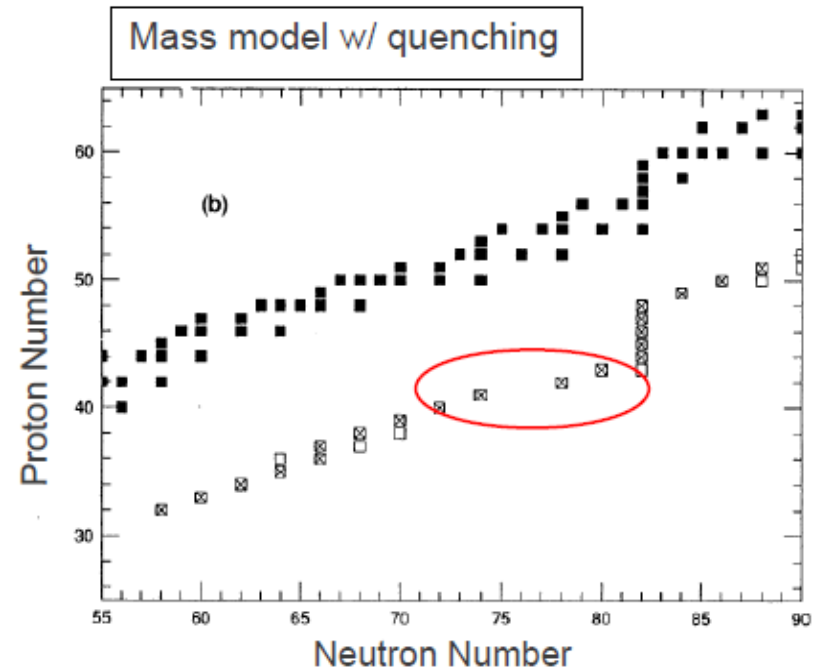
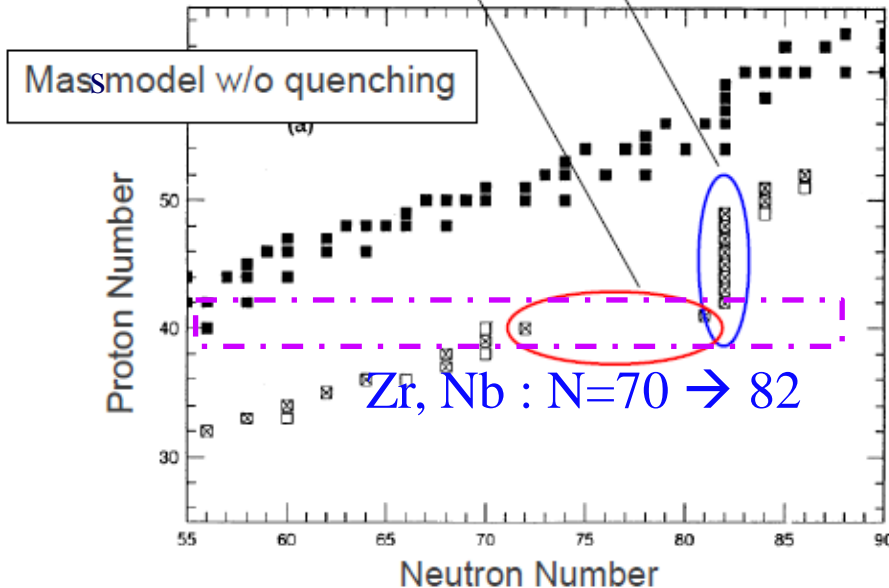
# 2<sup>nd</sup> r-process peak and path around $Z = 38 \sim 46$



Profound astrophysical impact of:

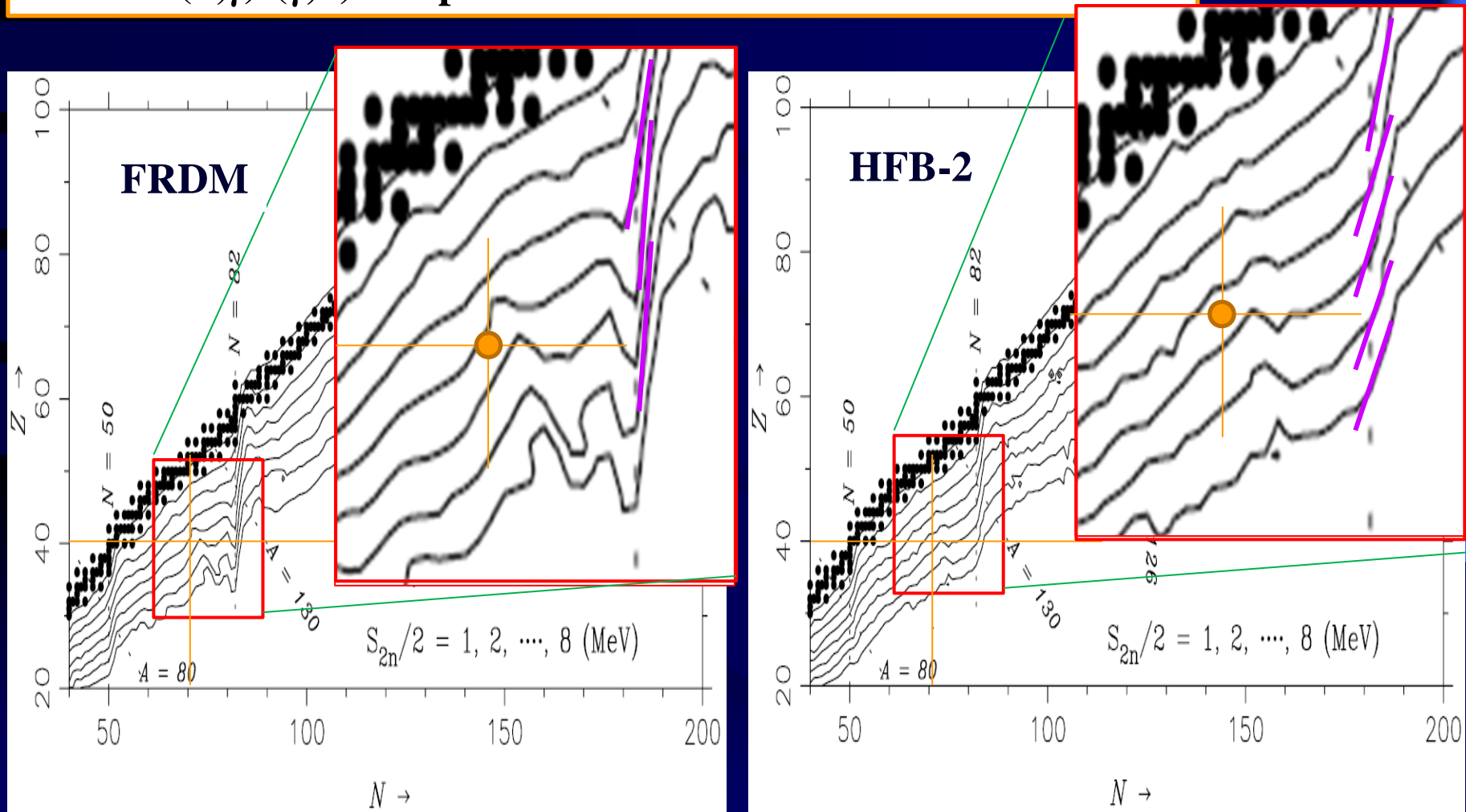
- Quenching of the  $N=82$  shell gap  
(need to study region around  $^{128}\text{Pd}_{82}$ )
- appearance of the a  $N=70$  sub-shell closure  
(need to study region around  $^{110}\text{Zr}_{70}$ )

K.-L.Kratz (1993)



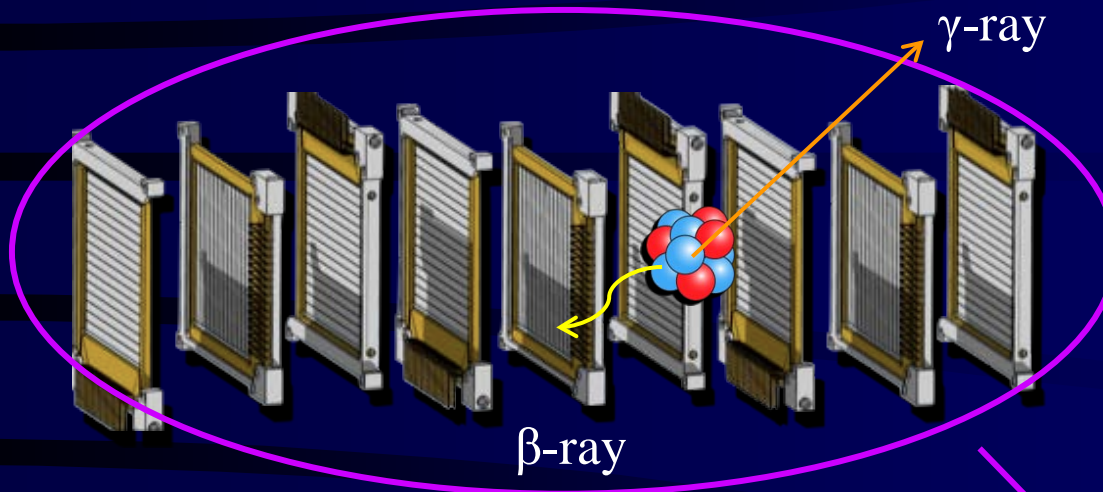
# Neutron separation energies around Mass A=110-125

Location of r-process path depends on  $S_n$  (2 – 3 MeV)  
(n, $\gamma$ )-( $\gamma$ ,n) competition



S.Wanajo, S.Goriely, M.Samyn, N.Itoh, Astro.J 606 (2004).

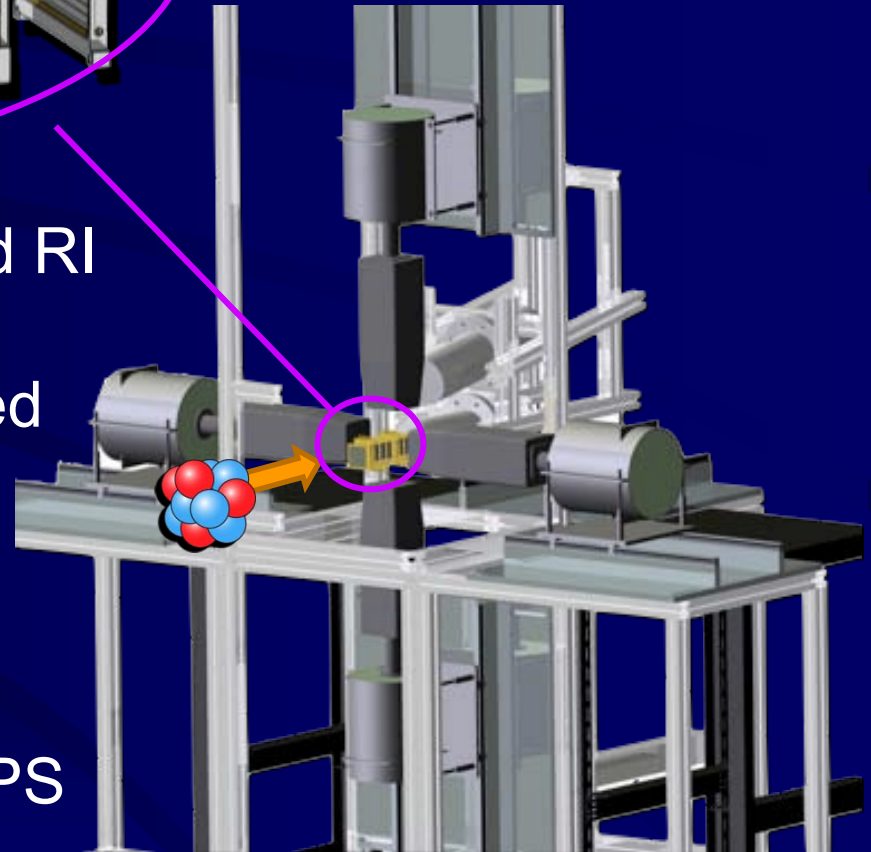
# Experimental Setup in 2009



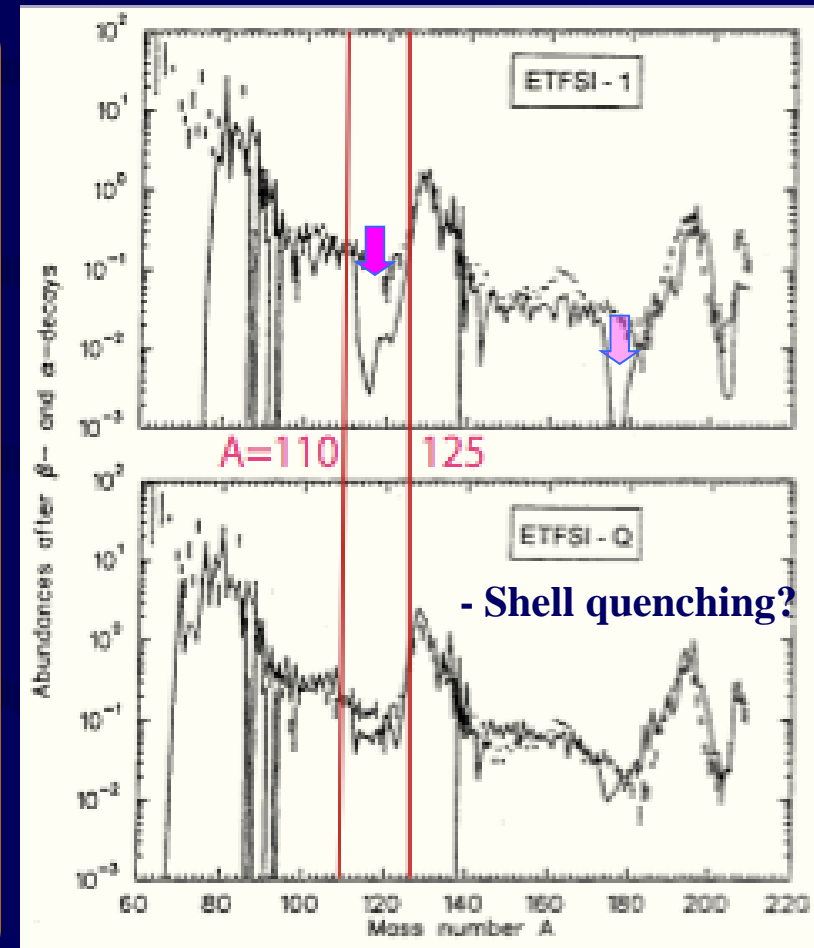
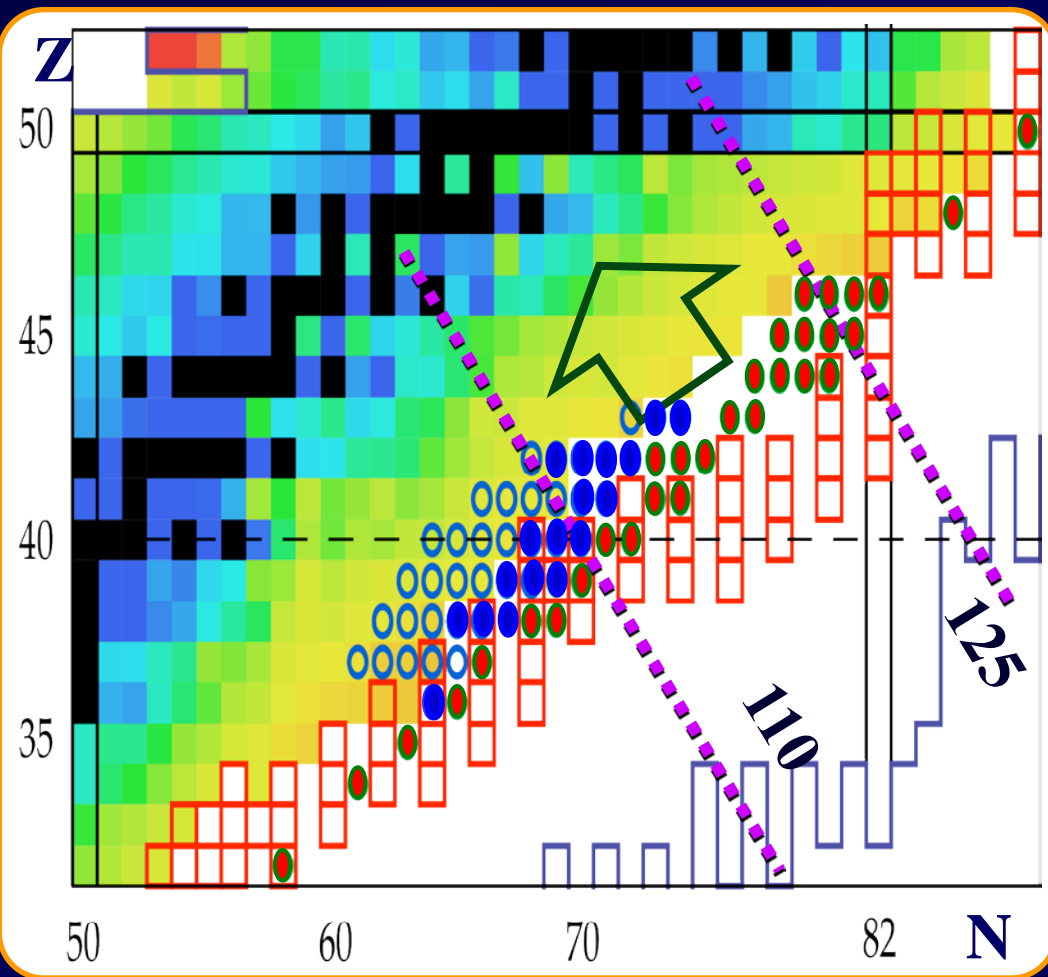
- **RI &  $\beta$ -ray detection**
  - 9 DSSDs ( $50 \times 50 \times 1 \text{ mm}^3$ )
  - $16 \times 16$  strips
  - $\sim 2000$  pixels in total

➤ The implantation of an identified RI is associated with the following  $\beta$ -decay events that are detected in the same DSSSD pixel

➤  $\Delta E$ -TOF-Bp method using the focal plane detectors in BigRIPS



# R-process Abundance around 2<sup>nd</sup> peak

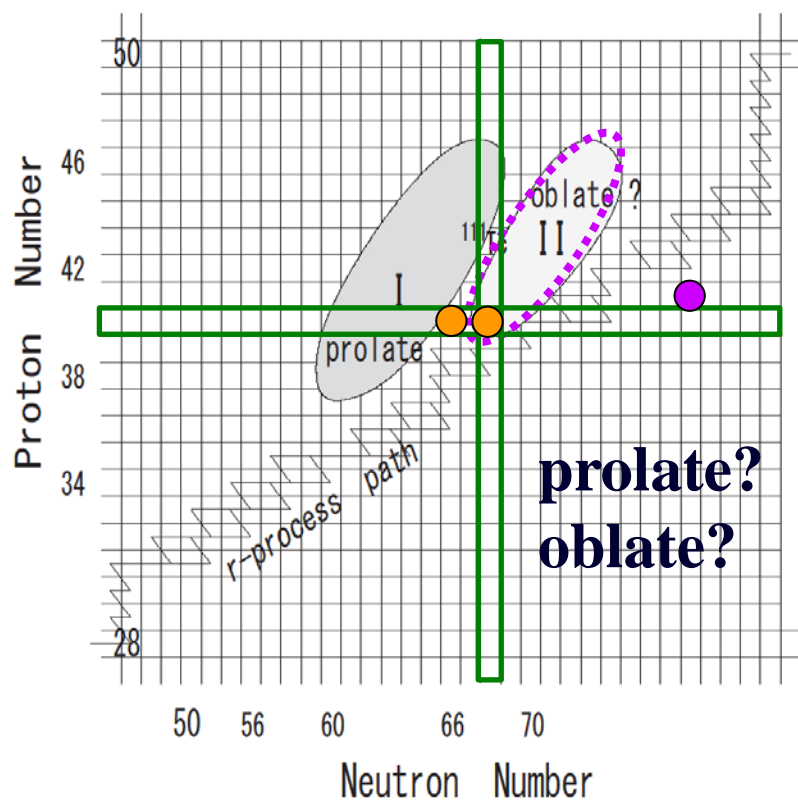


$$1/T_{1/2} = \sum_{E_i \geq 0}^{E_i \leq Q_\beta} S_\beta(E_i) \times f(Z, Q_\beta - E_i);$$

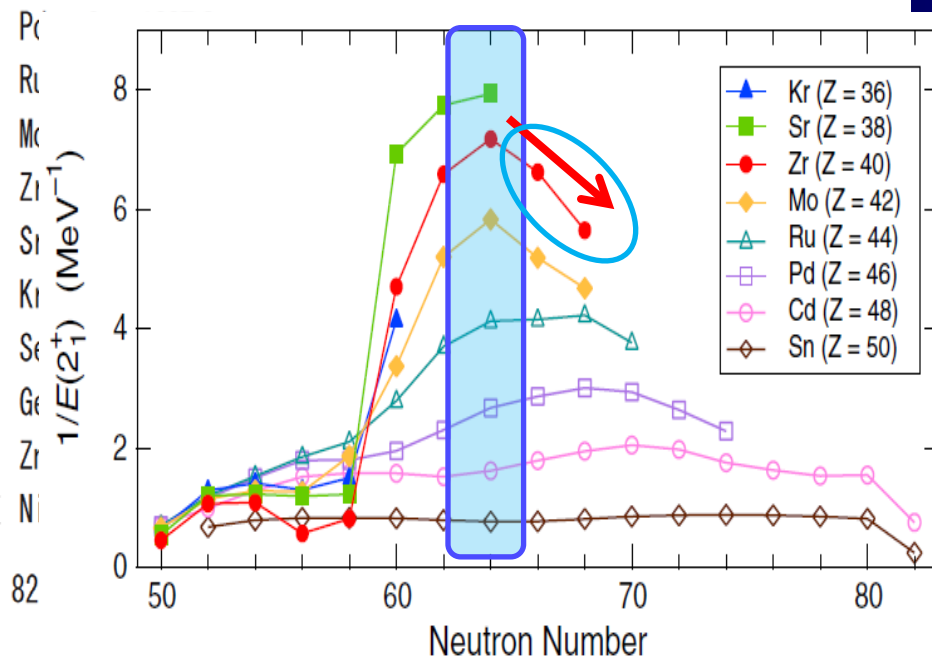
$$f \sim (Q_\beta - E_i)^5$$

# Deformation

- Phys. Lett. B 696 (2011) H.Watanabe  
Oblate shape isomer for  $^{109}\text{Nb}$ ?



## Structural Evolution in $^{106}\text{Zr}$ and $^{108}\text{Zr}$ T.Sumikama, et al., PRL 106 (2011)

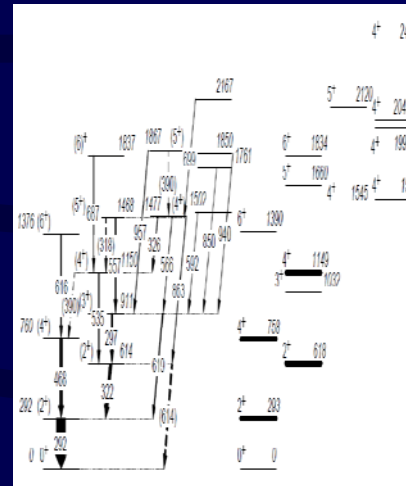
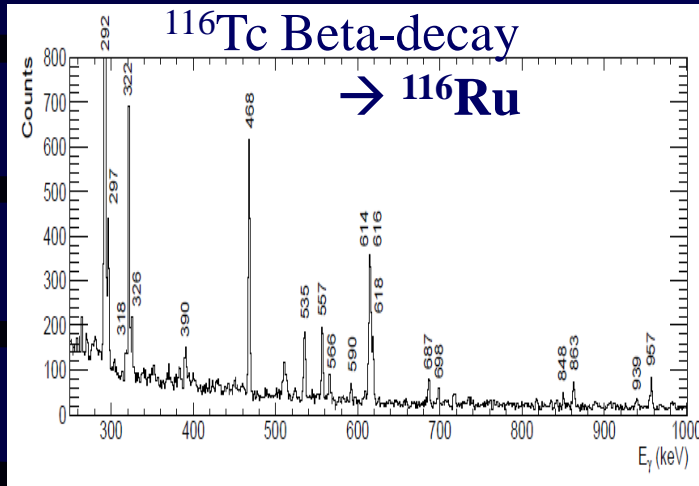


- Beta-delayed gamma :  $^{106}\text{Y} \rightarrow ^{106}\text{Zr}$
- Isomeric states :  $^{108}\text{Zr}$

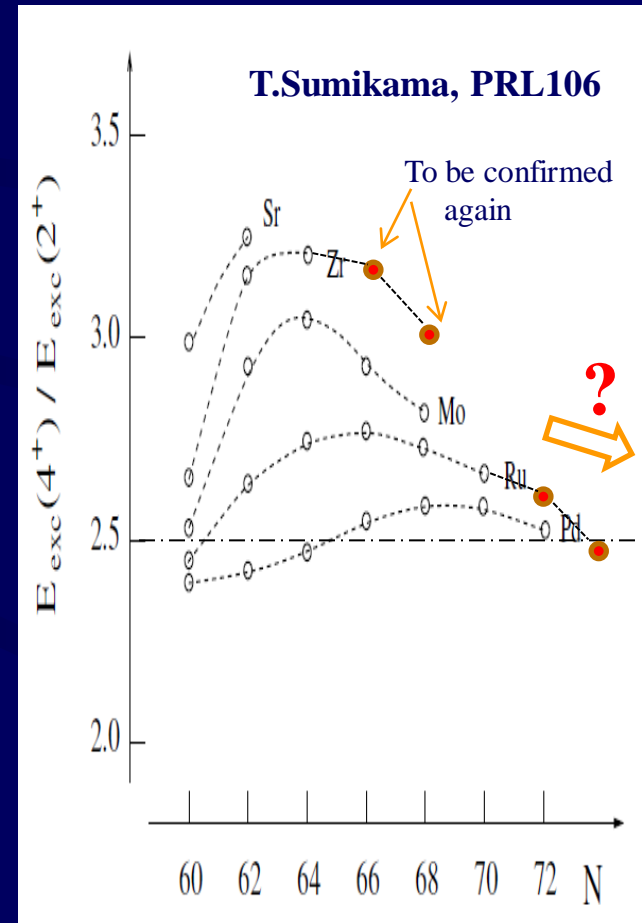
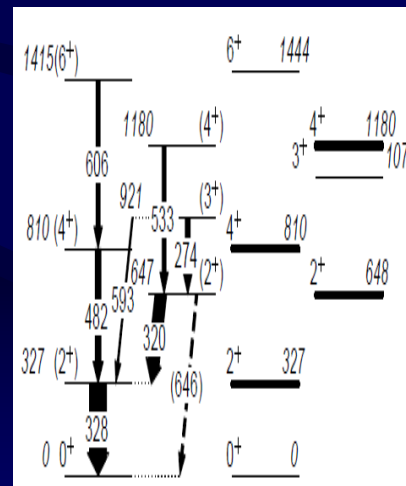
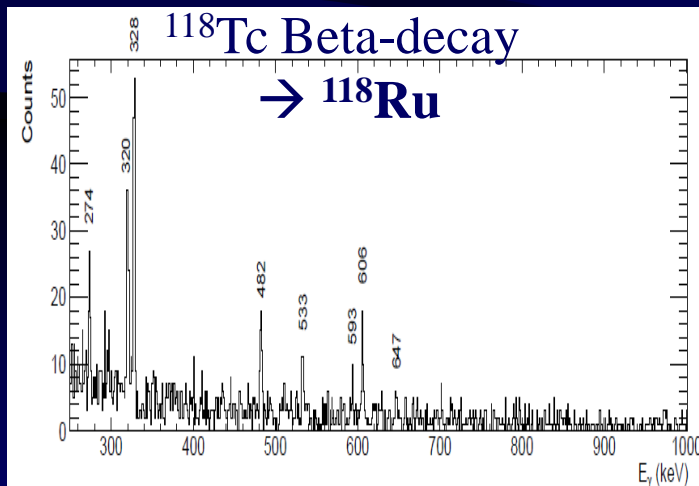


# Shape evolution in $^{116,118}\text{Ru}$ (EURICA)

Spokesperson: G.Lorusso, H. Watanabe



P.-A. Söderström, et al.  
Phys. Rev. C, accepted.

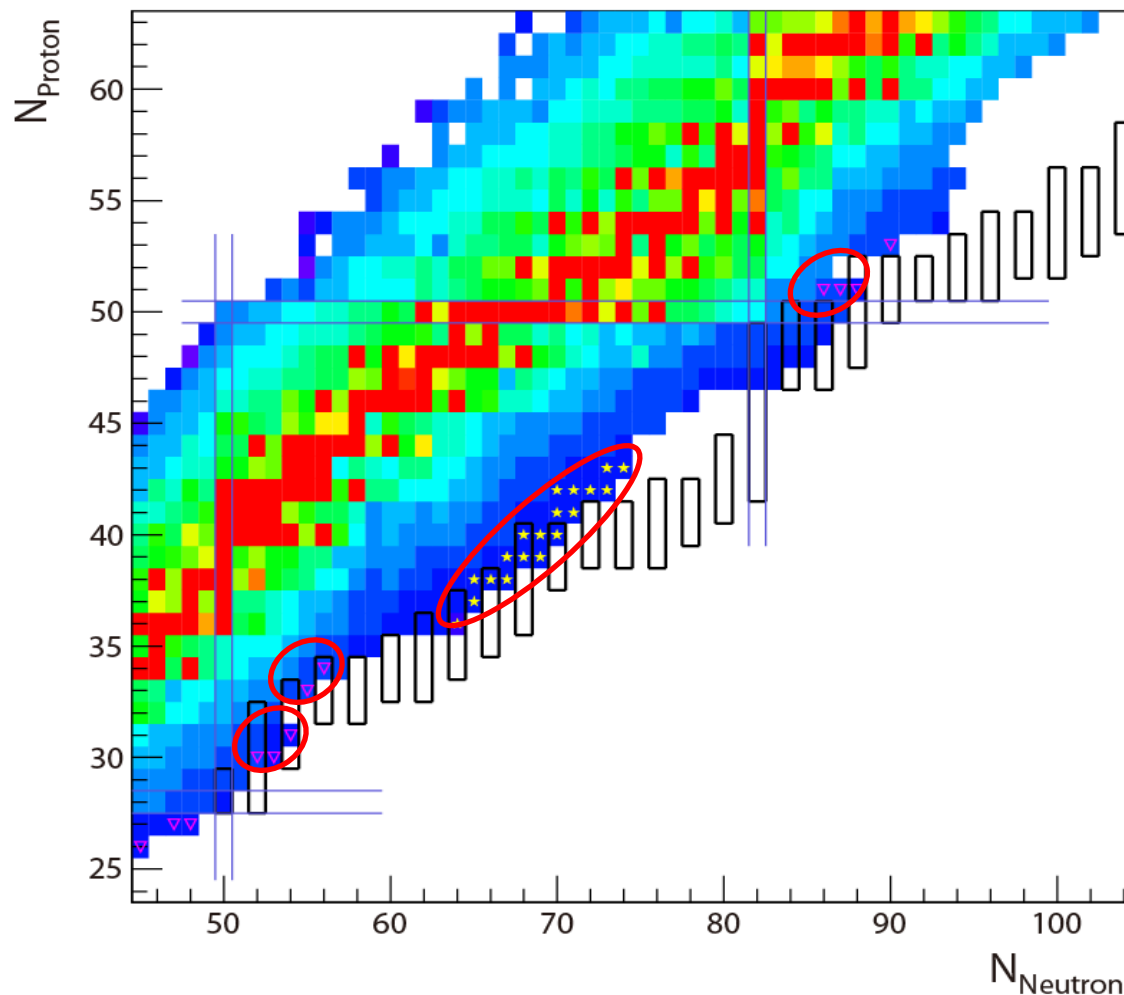


Urban, EPJ A 20, 381 (2004)

Analysis is in progress in other isotopes around  $^{110}\text{Zr}$  region



# Beta-decay half-lives since 2007



2011 .. O.Arndt,  
PRC 84,  
**3**  $T_{1/2}$  (ISOLDE)

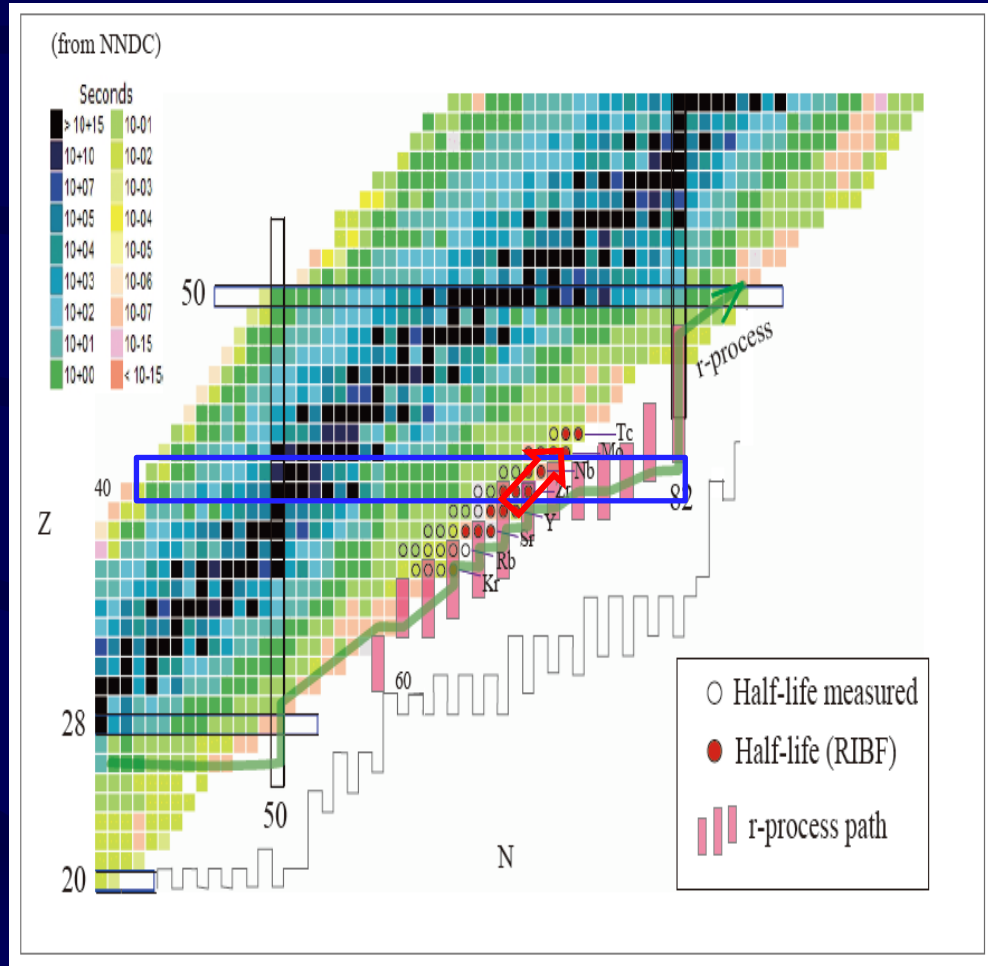
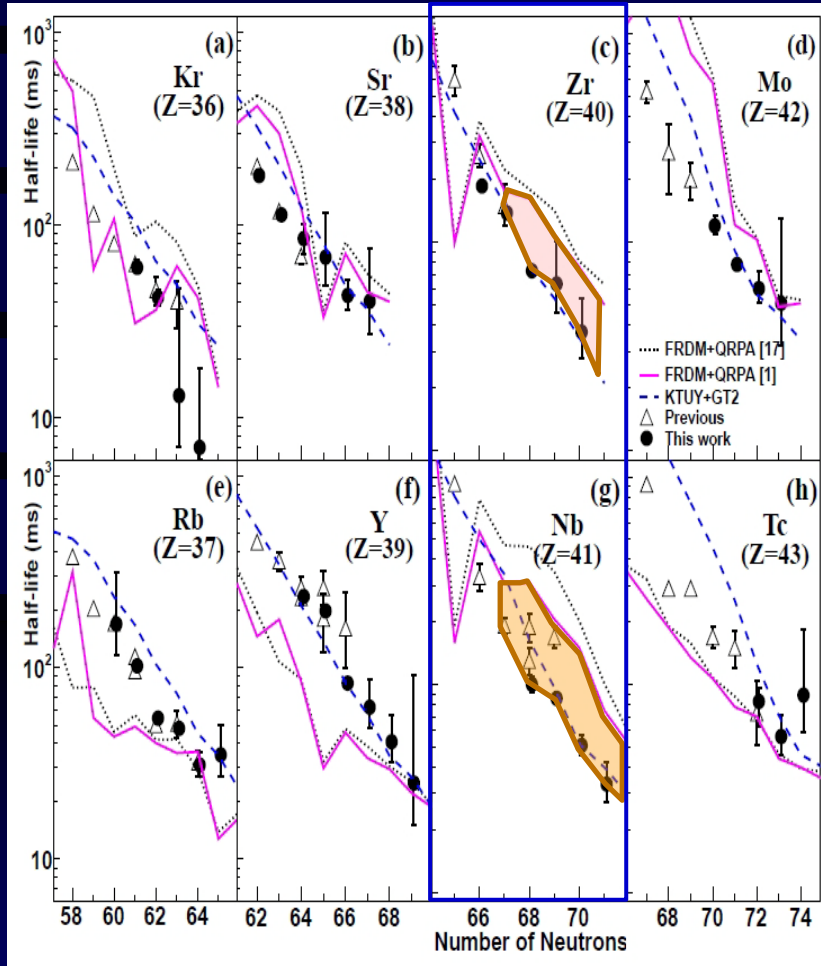
2011 .. SN,  
PRL 106,  
**18**  $T_{1/2}$  (RIBF)

2012 .. M.Quinn,  
PRC 85,  
**2**  $T_{1/2}$  (MSU)

2012 .. M.Madurga,  
PRL 109 ,  
**3**  $T_{1/2}$  (ORNL)

2013 ... C.Mazzocchi  
**1**  $T_{1/2}$  PRC (ORNL)

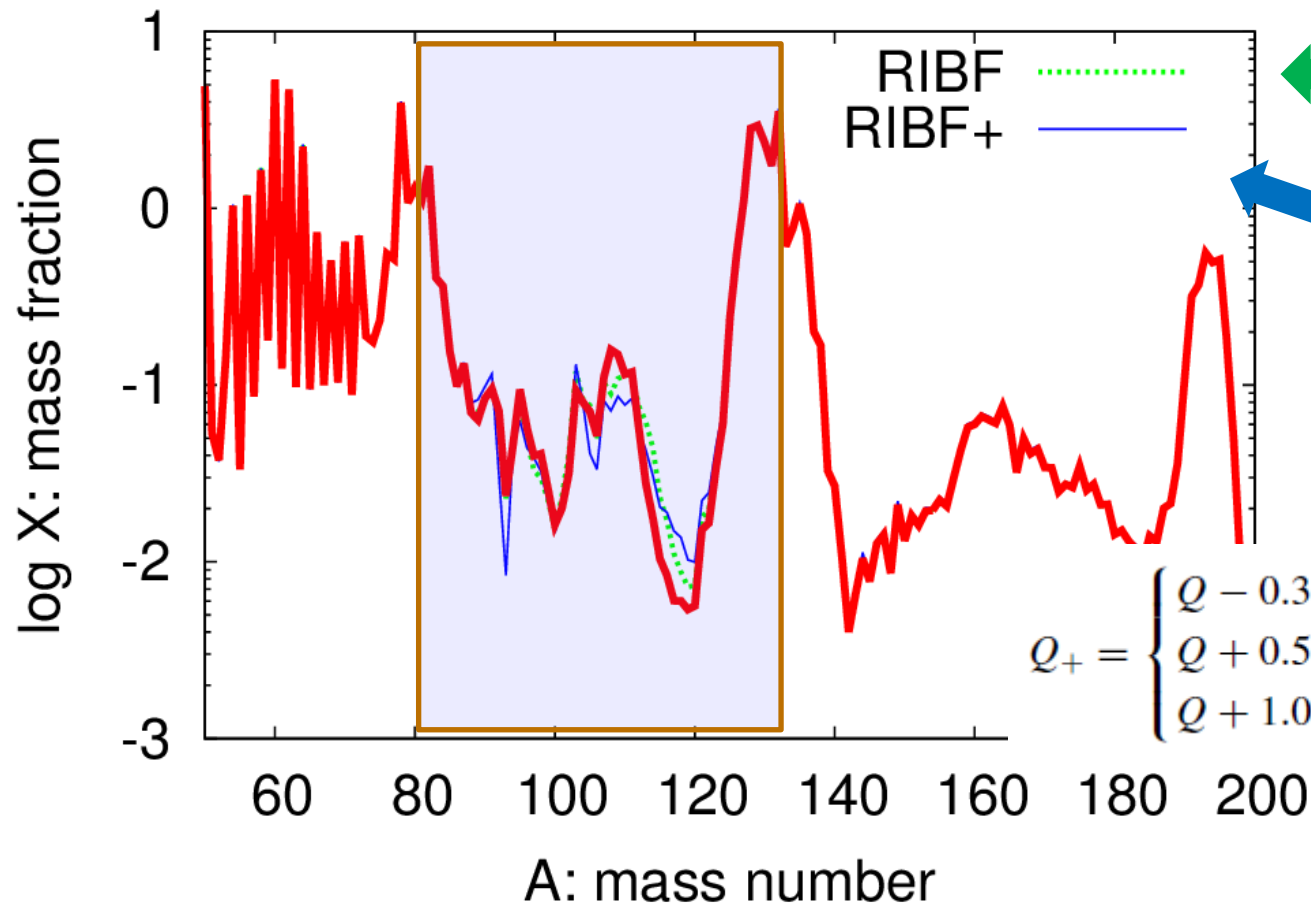
# Decay Experiment in 2009



Zr and Nb decay faster than expected by FRDM+QRPA ( $T_{1/2} : 1/2 \sim 1/3$ )

# RIBF data → Impact to r-process abundance

Nobuya Nishimura, T.Kajino, G.Mathew, SN, T.Suzuki, PRC 85 (2012)



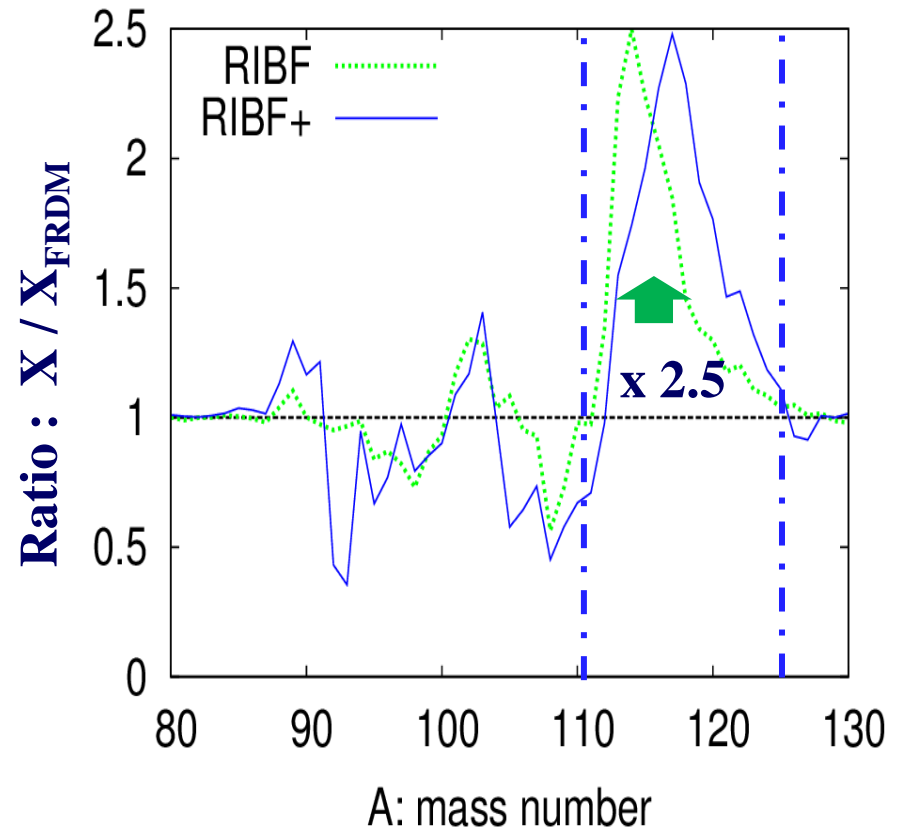
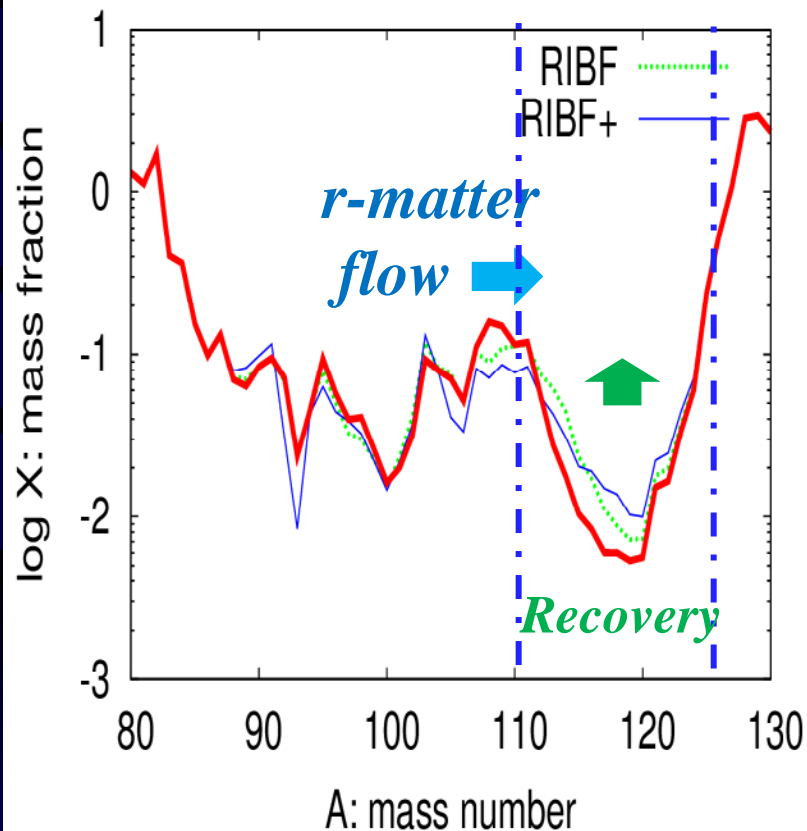
**38 Half-lives  
from RIBF**

**38 Half-lives  
+  $\Delta Q_{\beta}$   
from RIBF**

MHD supernova  
explosion model

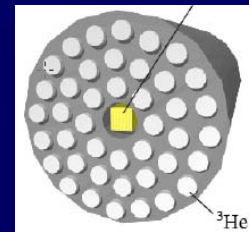
# RIBF data $\rightarrow$ Impact to r-process abundance

Nobuya Nishimura, T.Kajino, G.Mathew, SN, T.Suzuki, PRC 85 (2012)

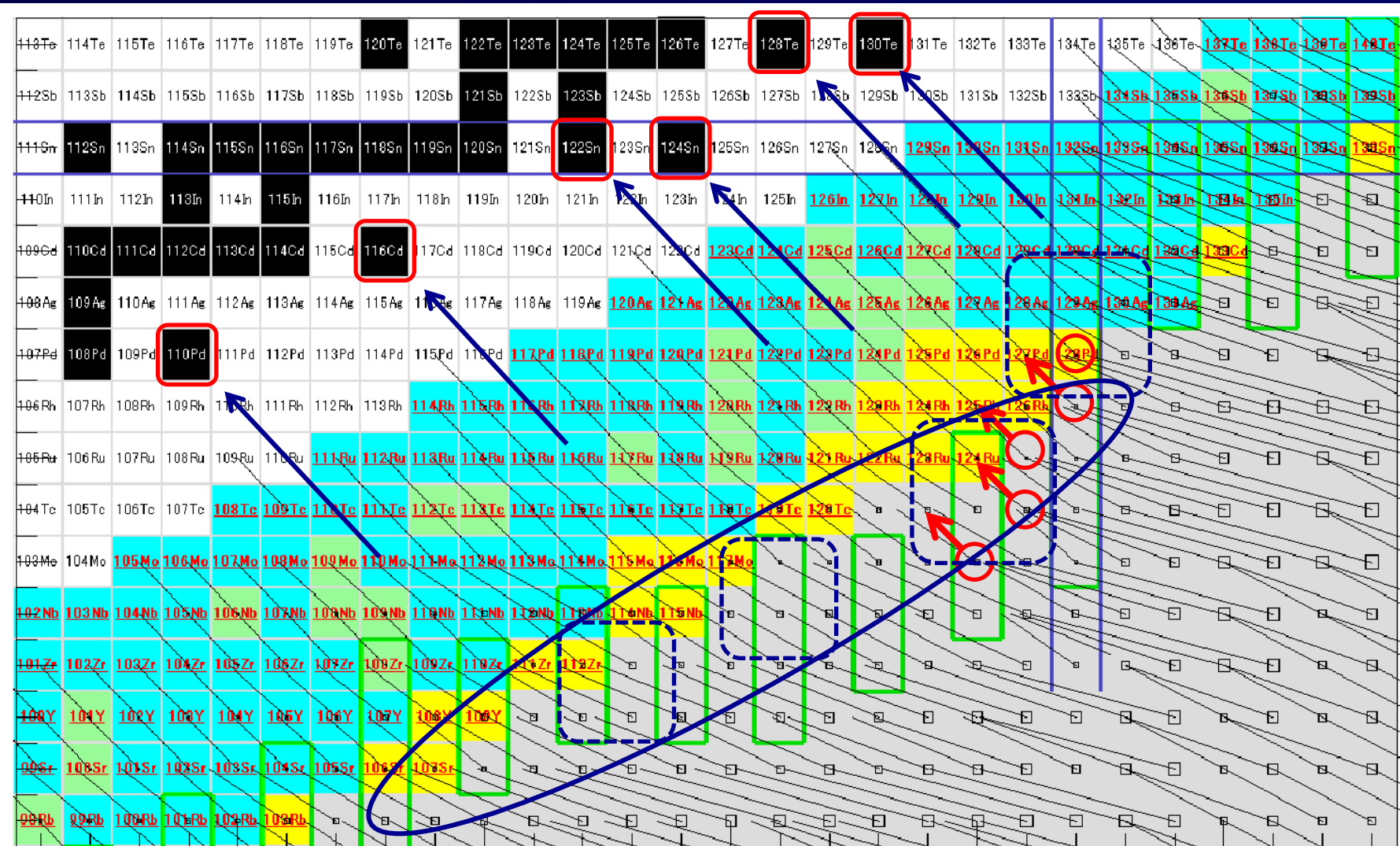


**The calculated r-process abundance is improved by factor of  $\times 2.5$ . But, there is still issue remaining in mass  $A=110 - 125$ !**

# Expected Beta-decay Flow (Beta-delayed Neutron Emission)

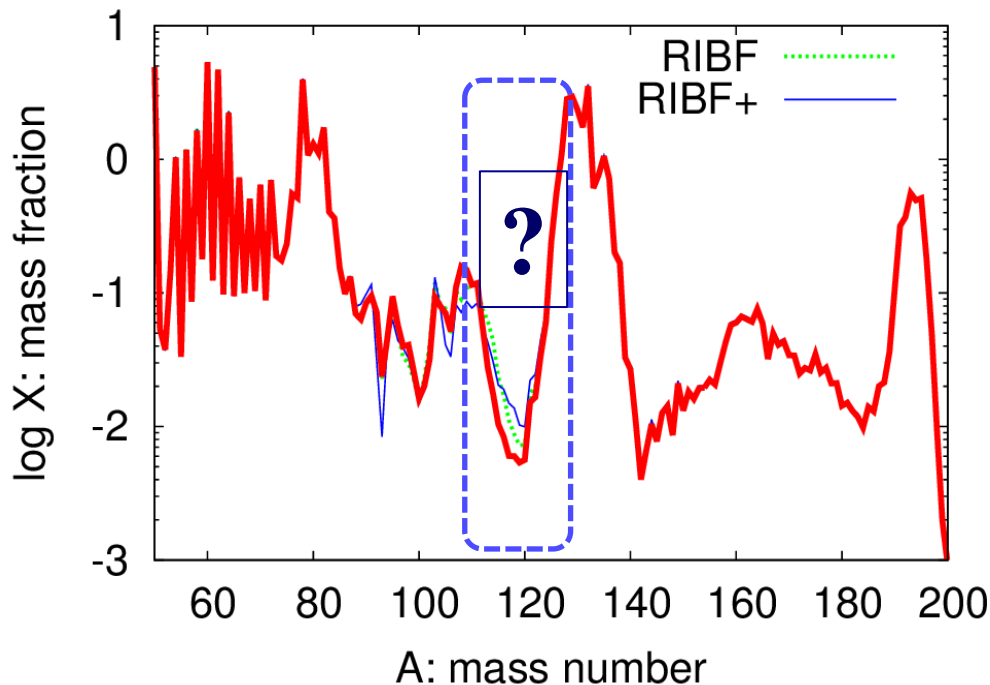


For example, FRDM+QRPA (P.Moller)



# Summary

- Mysteries of the r-process abundance around 2<sup>nd</sup> peak.
  - Mass  $A = 110-125$ 
    - Influence of Larger / smaller neutron emission ?
    - Multi neutron emission  $P_n, P_{2n}, \dots$  ?
    - Mass ? and Half-lives?



Do we have problems in

Nuclear theory?

or

Nuclear astrophysics

( or both... )